

### Dielectron Production in Au+Au Collisions at BES Energies

### and its energy dependence from SPS to top RHIC energies

#### Patrick Huck for the STAR Collaboration

**RBRC** Thermal Radiation Workshop, BNL

12/06/2012

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Lawrence Berkeley National Laboratory





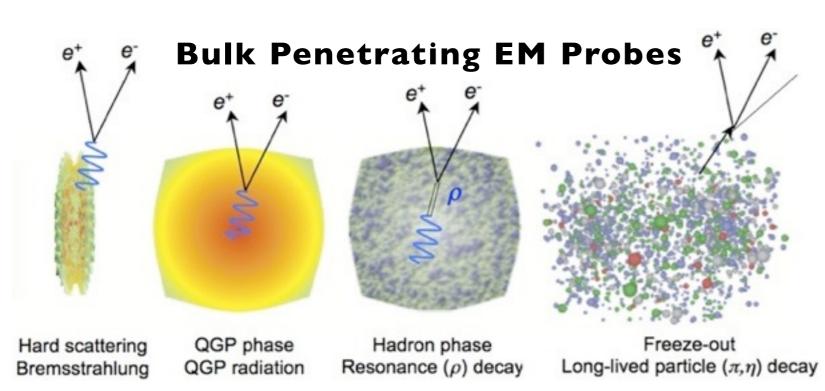


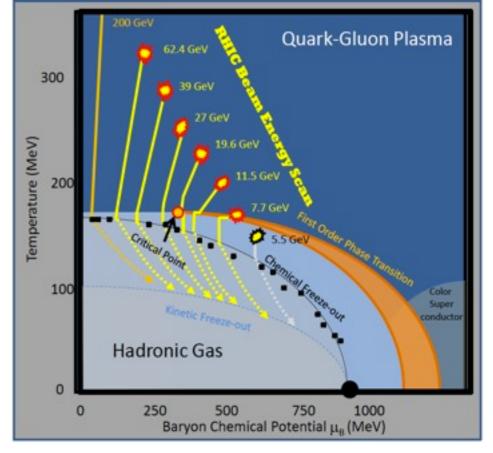
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## STAR 🛠 Study the QCD Phase Diagram

#### Fundamental Questions for HIC's

- I) thermally equilibrated matter produced through sufficient rescattering? magnitude of collective expansion indicates  $\tau < 1-2$  fm/c characterize medium by bulk thermodynamic variables
- 2) distinctive footprint of individual partons? v<sub>2</sub> NCQ scaling: collectively expanding partonic source
- 3) deconfinement? chiral restoration? spectroscopy via short-lived resonances due to inaccessible order parameter



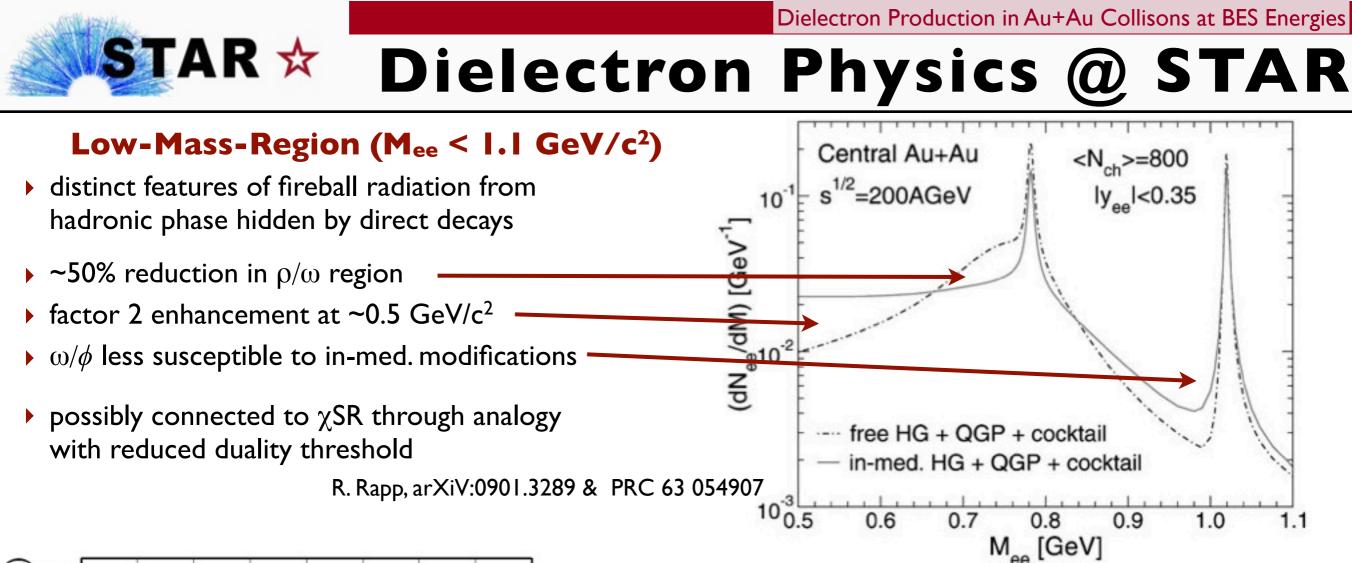


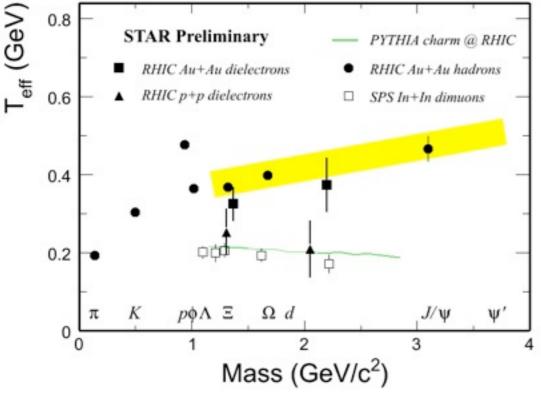
#### Beam Energy Scan Program consistently combine various signatures

over a wide range of beam energies

- access hadronic spectral functions via EM probes ( $\gamma / l^+ l^-$ ) negligible FS interactions due to  $\lambda_{mfp} >> \tau_{FB}$
- additional dynamic information about HIC stages encoded in invariant mass

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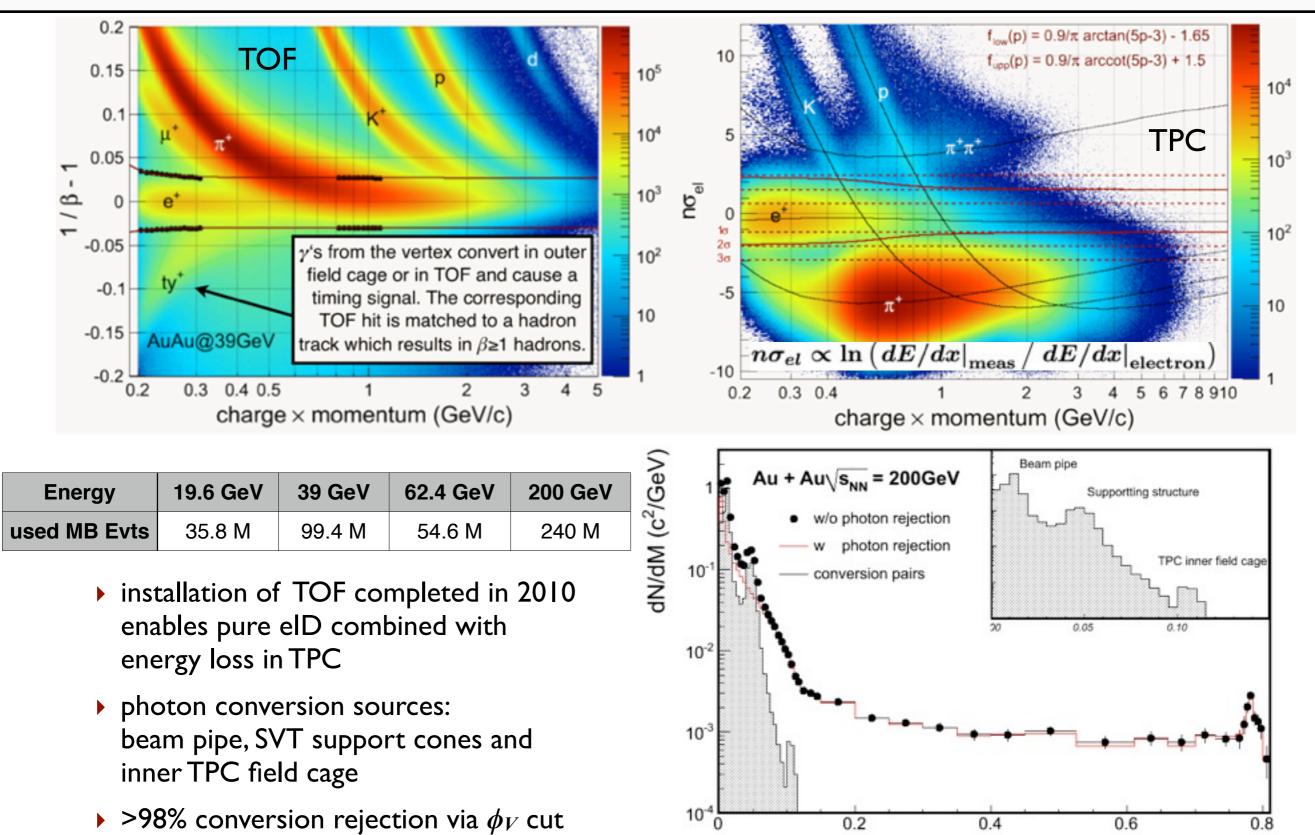
#### Intermediate-Mass-Region (I.I < $M_{ee}$ < 3 GeV/c<sup>2</sup>)

- $\blacktriangleright$  measure initial QGP temperature from IMR  $m_T$  spectra
  - however, with contributions from correlated charmed decays unknown, large systematic uncertainties arise even at 200 GeV.
  - thus, BES analyses presented here concentrate on LMR physics

STAR in unique position to study energy dependent dielectron production and study/confirm medium consequences on spectra w.r.t to their energy dependence

# 

#### **Datasets, Electron-ID & Conversion Rejection**



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Mee (GeV/c2)

### **Background Subtraction Methods**

 $e^+/e^-$  created in pairs

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 $\Rightarrow$  unlike-sign BG is geometric mean of the like-sign BGs independent of primary probability/multiplicity distribution

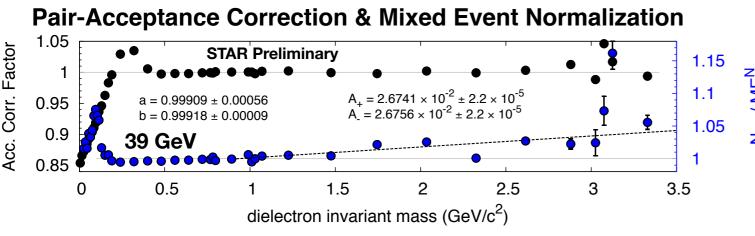
 $\langle \mathrm{BG}_{+-}\rangle = 2\sqrt{\langle \mathrm{BG}_{++}\rangle \langle \mathrm{BG}_{--}\rangle}$ 

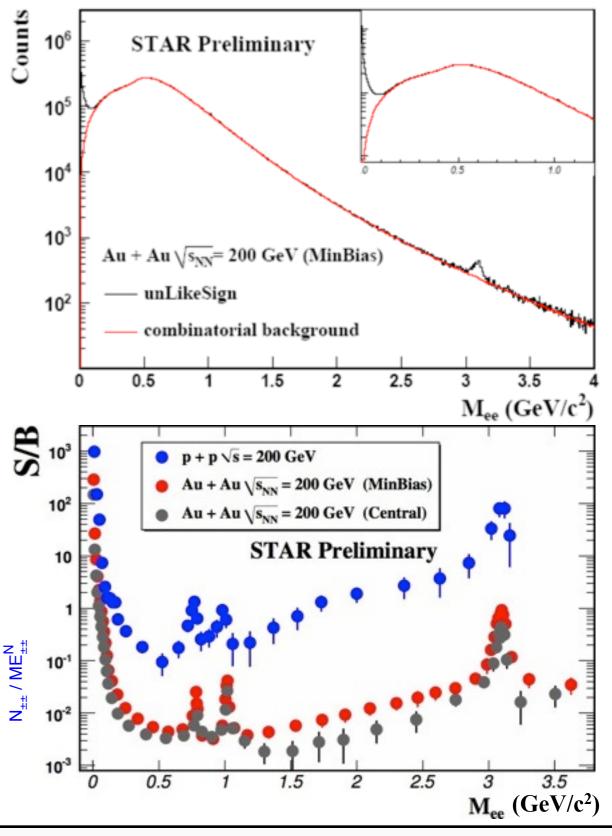
#### I) Like-Sign Same Event Method

- All like-sign pairs of one event are combined and averaged.
- Method reproduces the background from all correlated sources.
- Acceptance difference of like-sign to unlike-sign pairs is corrected using the ME Technique.

#### 2) Unlike-Sign Mixed Event Method

- Charges from two different events within same event class are combined (event vertex, reference multiplicity & event plane).
- Method describes uncorrelated BG only.





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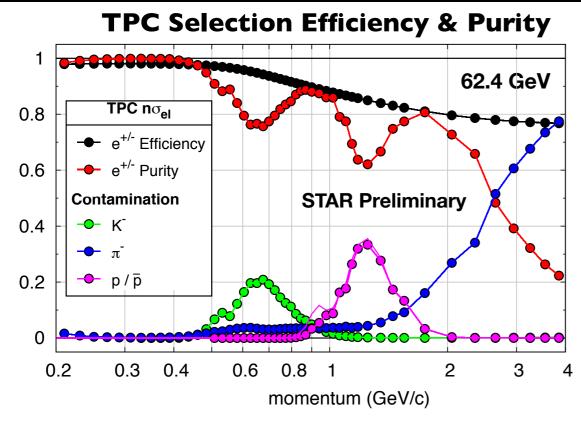
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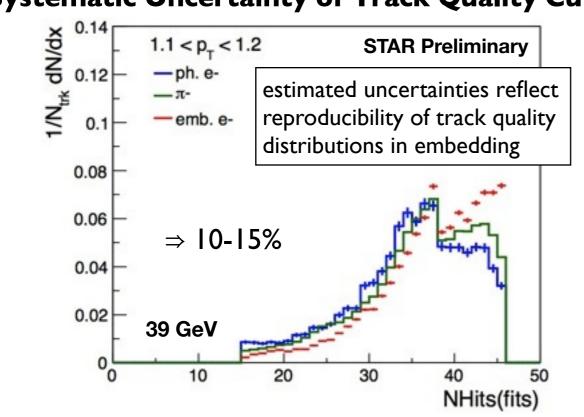


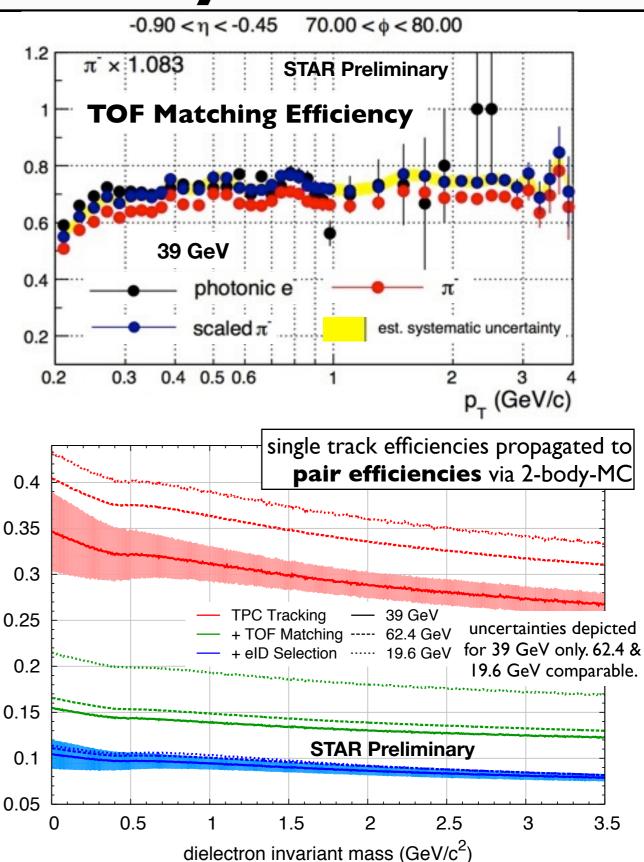
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## **Efficiency Correction**



#### Systematic Uncertainty of Track Quality Cuts





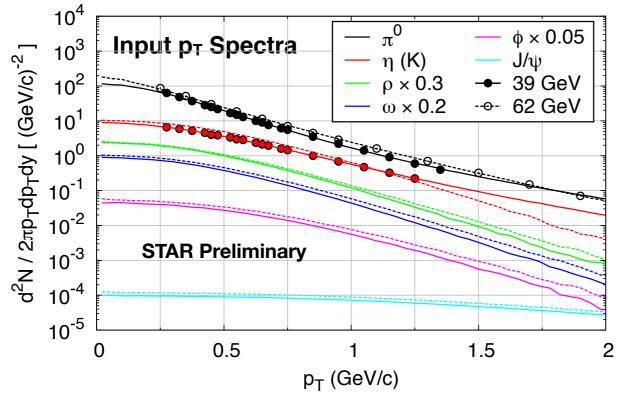
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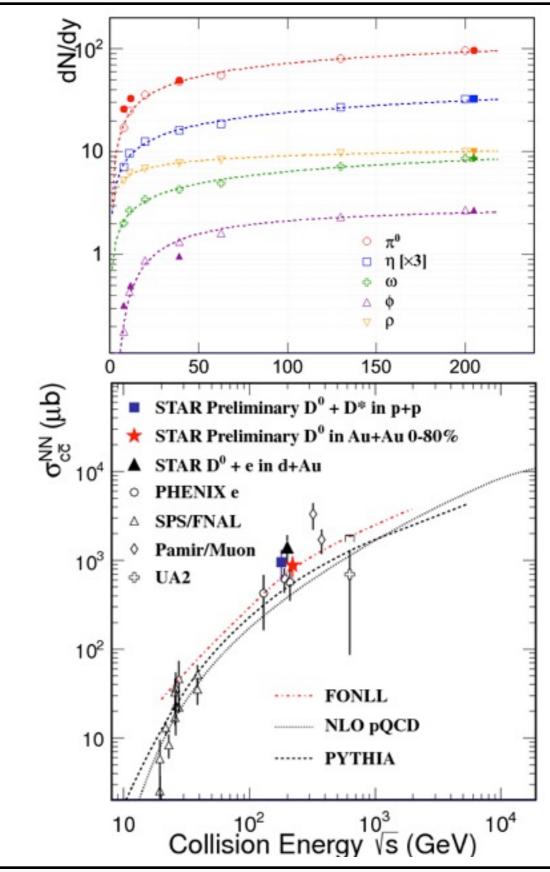
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### **Cocktail Simulation**

- flat η [-1,1] & φ [0,2π], Kroll-Wada for Dalitz decays & according form-factors from measurements (PDG)
- AuAu@19.6 GeV:
  - Tsallis fits to meson spectra from SPS PbPb@17.3GeV
  - + meson/ $\pi^0$  ratio from SPS &  $\pi^0$  yield from STAR
  - Conversion included via full STAR GEANT simulation
- AuAu@39 & 62.4 GeV:
  - +  $\pi^0 p_T$  spectra from  $\pi^{+/-}$  @STAR, K spectra used for  $\eta$
  - Unknown pT distributions taken from AMPT
  - According yields extrapolated from 200 GeV based on AMPT's  $\sqrt{s}$ -dependence
  - + conversion rejected via  $\phi_{\rm V}$  cut
- Contributions due to correlated charmed decays simulated using PYTHIA and scaled to Au+Au by N<sub>bin</sub>



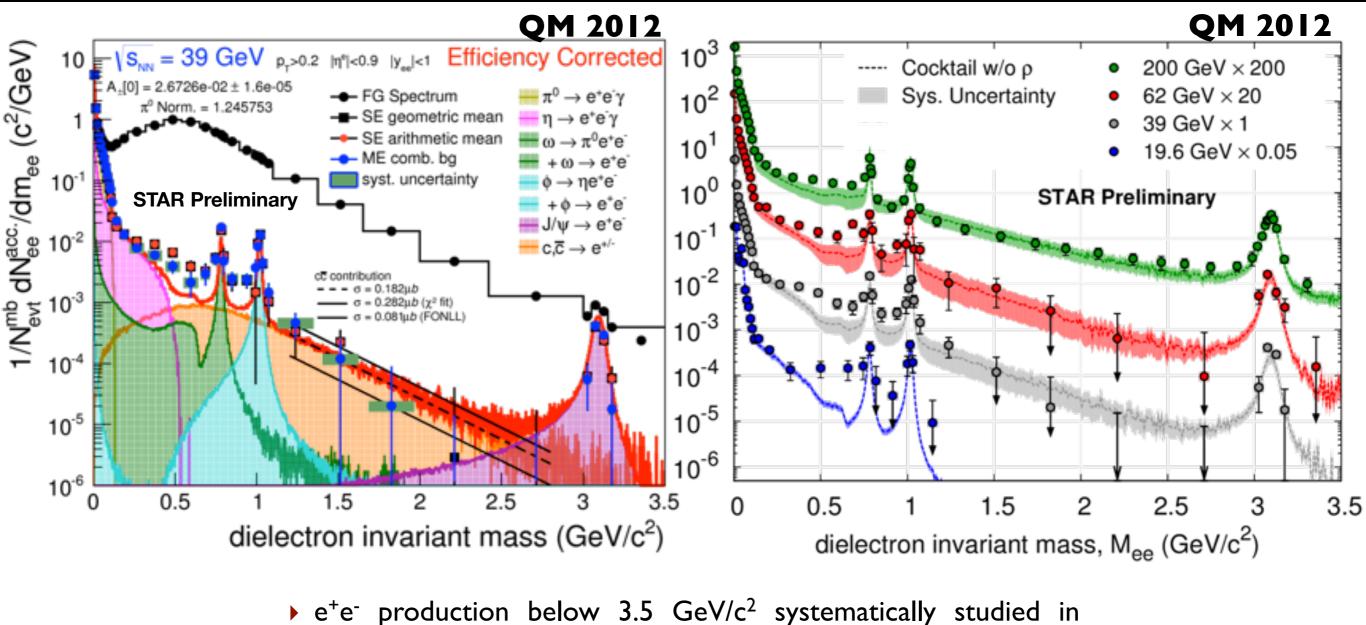


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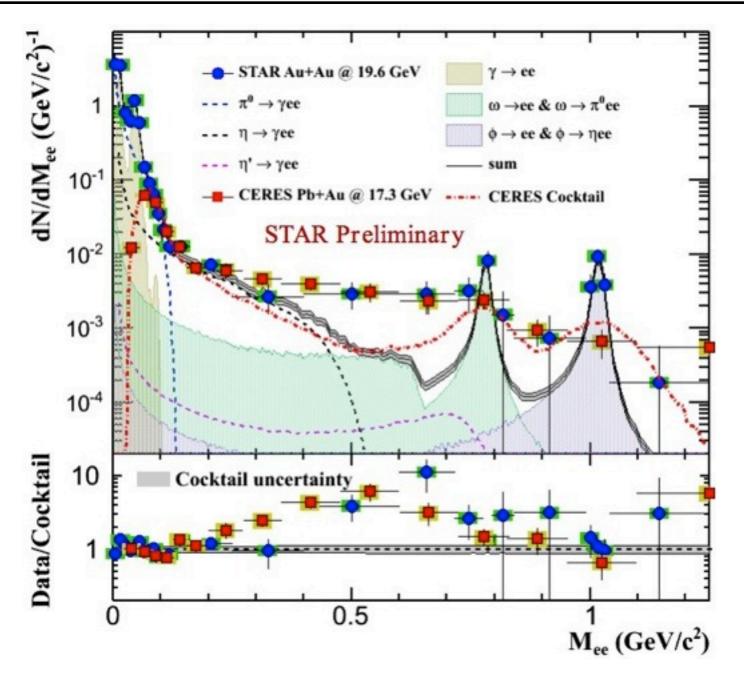
## **BES Dielectron Spectra**

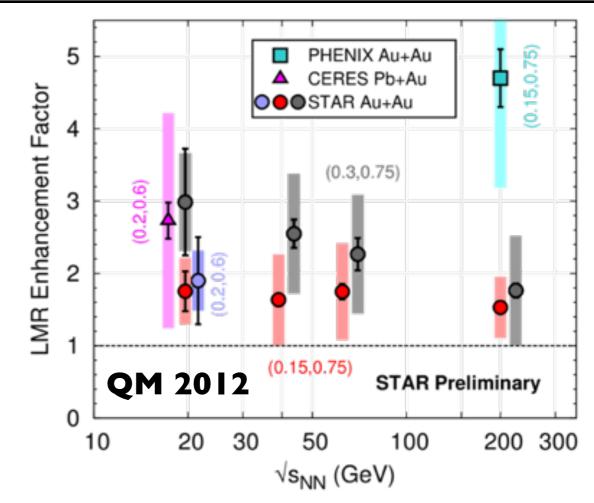


- $e^+e^-$  production below 3.5 GeV/c<sup>2</sup> systematically studied in STAR from  $\sqrt{s_{NN}} = 19.6$  GeV up to top RHIC energy.
- correlated charm adjusted to observed dielectron yield FONLL predictions are used as lower and χ<sup>2</sup> fits to the IMR data as upper limits
- $\blacktriangleright$  vacuum- $\rho$  does not account for the excess yield in the LMR



### LMR Enhancement

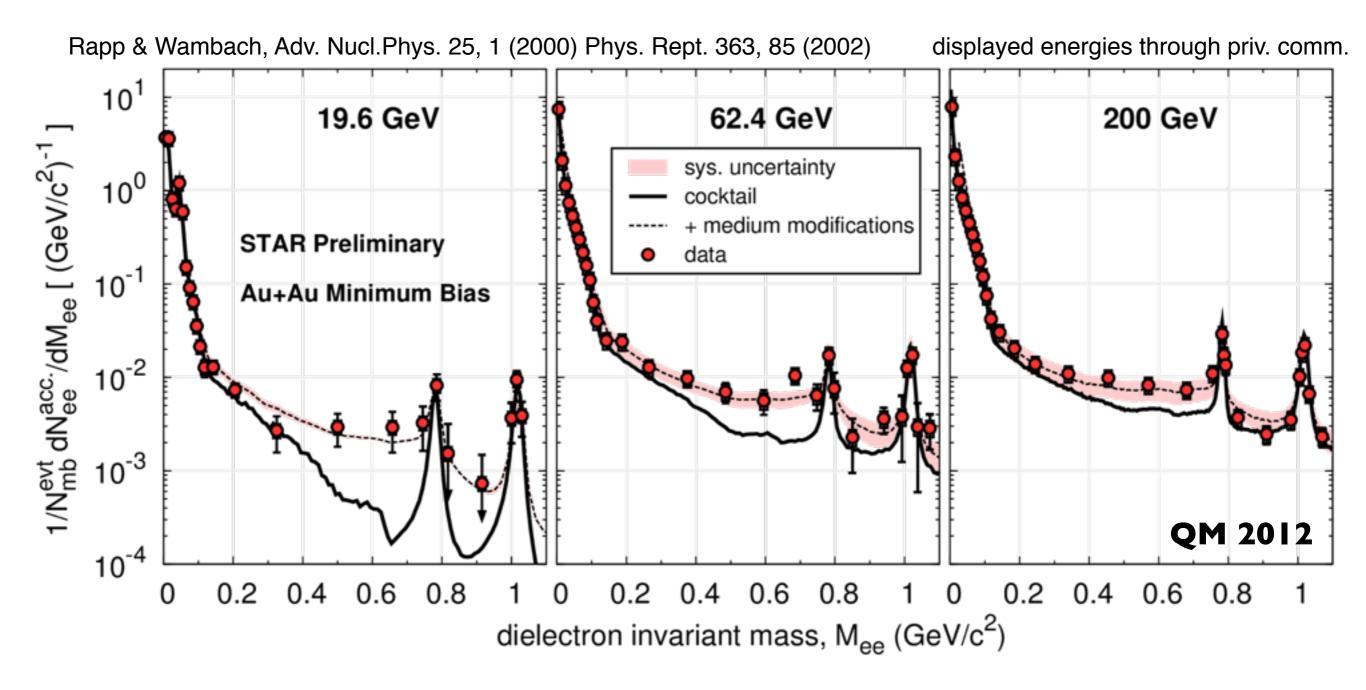




- visible LMR excess over hadronic cocktail observed for all energies. (excl. p→e<sup>+</sup>e<sup>-</sup>)
- systematic measurement of the LMR enhancement factor (agreement with CERES result) [Mee-dep. energy overlay see backup]
- LMR enhancement at 19.6 GeV comparable with CERES at 17.3 GeV (note: different experimental acceptances)
- increasing enhancement with decreasing energy w.r.t. the cocktail?
   "any energy dep. in X-Factor might be physics directly related to dielectrons from earlier creation times due to ρ<sub>B</sub><sup>tot</sup>~const" Z. Xu

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### **Energy Dependence of In-Medium Mods**



 Within systematic uncertainties, in-medium modifications of the ρ spectral function consistently describe the LMR enhancement from SPS to top RHIC energies.

### Summary, Outlook & Work in Progress

• Dielectron spectra from Au+Au collisions measured in STAR at  $\sqrt{s_{NN}} = 19.6, 39, 62.4 \& 200$ GeV and compared to cocktail calculations.

- LMR excess yield can be accounted for by inmedium modifications to the ρ spectral function across a wide range of energies.
- enhancement increasing with decreasing energy w.r.t. the cocktail?
- measurements will provide comprehensive data for the better understanding of the LMR enhancement (p<sub>T</sub>, centrality and energy dependence)

#### work in progress

- complete BES data set
- $p_T$  spectra for  $M_{ee}$  regions
- detailed systematic uncertainty studies
- cocktail improvements

#### outlook

- IMR: Charm continuum contribution and its possible in-medium modification need better understanding in Au+Au to possibly access QGP radiation in the future
  - $\Rightarrow$  study energy dependence of initial temperature
  - $\Rightarrow$  STAR HFT & MTD upgrades

### Thank you for your attention



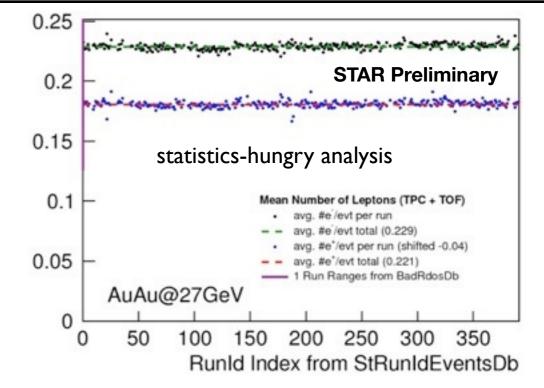
## **BACKUP SLIDES**

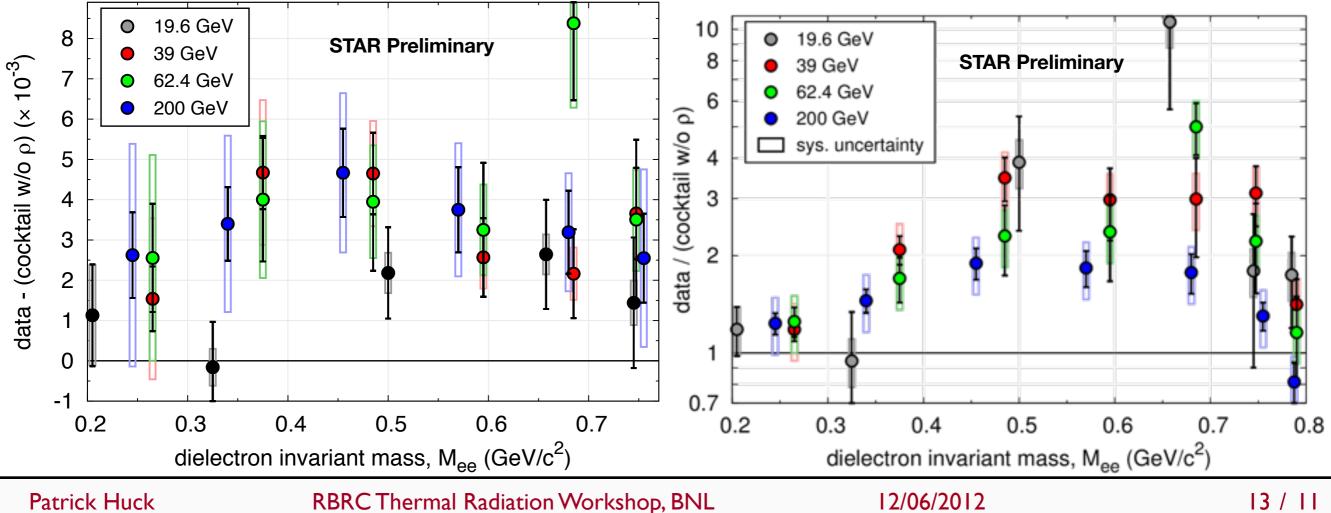
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• absence of baryonic resonances with  $\phi N$  decay channels due to OZI-rule \*  $*\overline{s}$  annihilation into excitation energy strongly suppressed

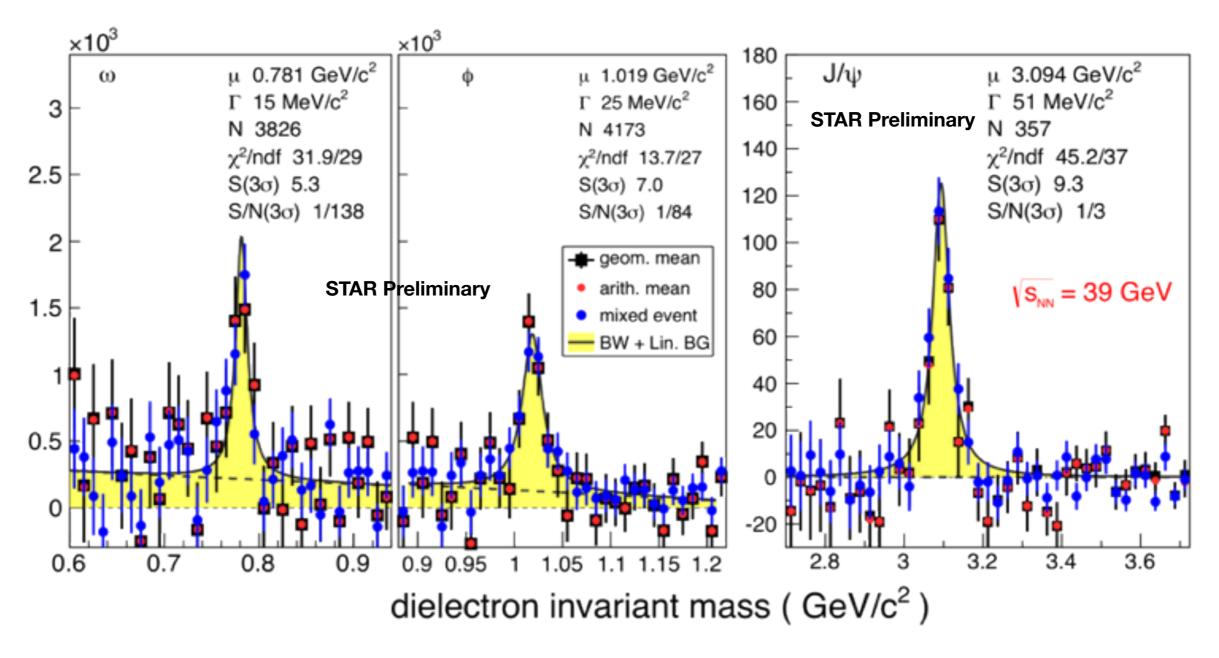
√s <sub>NN</sub> (GeV)	Vector Meson Yields (30% uncertainty assigned)					$\sigma^{car{c}}_{pp}~({ m mb})$	۸۲coll
	$\pi^{0\dagger}$	η †	ω	φ	J/ψ	± Sys.	$N_{ m bin}^{ m coll}$
39	57	9.37	4.42	1.39	4.8 × 10 <sup>-4</sup>	0.19 ± 0.11	243
62.4	72.9	11.4	5.38	1.79	1.2 × 10 <sup>-3</sup>	$0.40 \pm 0.25$	253





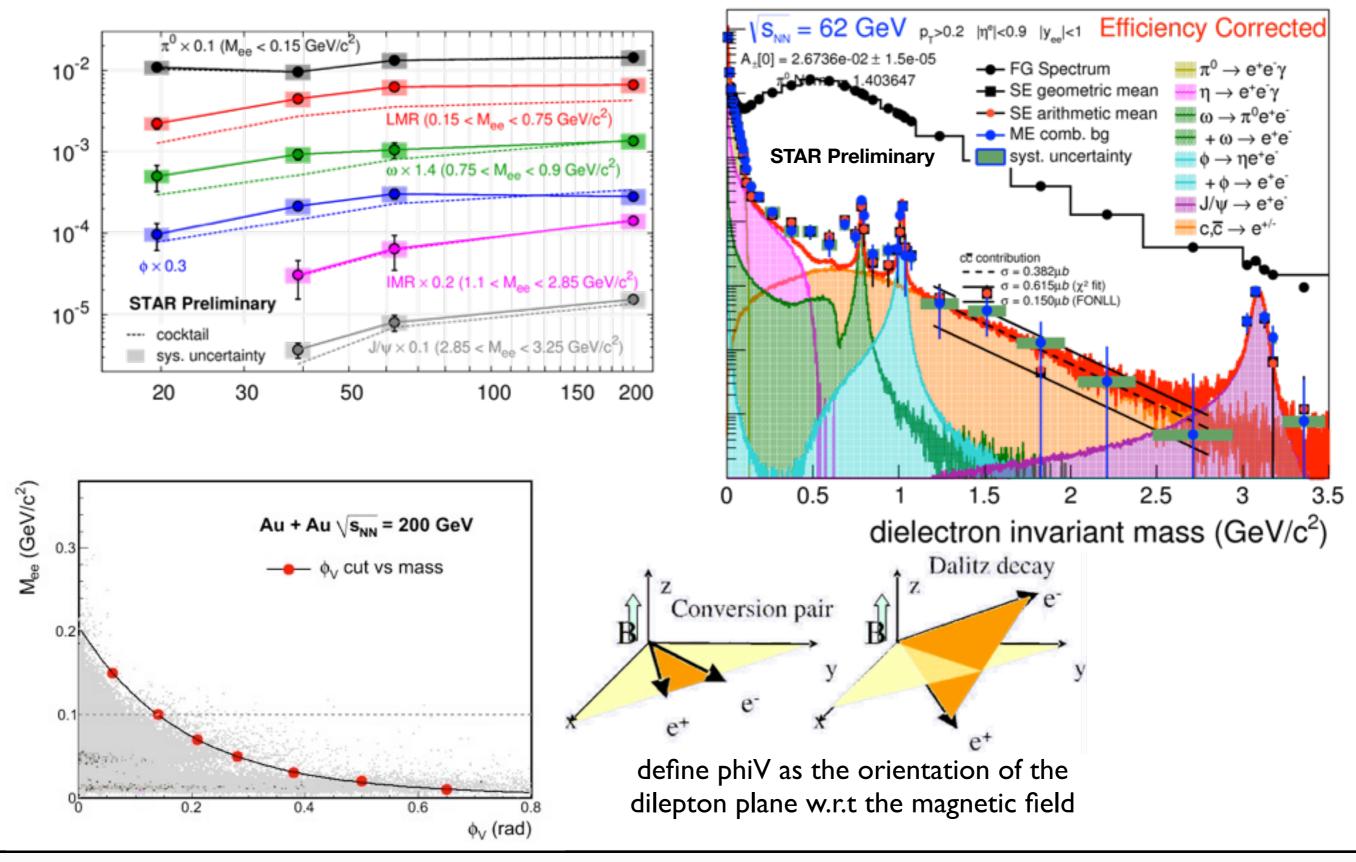
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## STAR 🛠 Vector Meson Signals & S/N



- $\rho/\omega$  region exhibiting a S/N ratio of ~1/100 1/250
- background subtraction crucial
- prominent vector meson signals after background subtraction









#### Cocktail w/ Vacuum-Rho

