



# Measurements of elliptic flow in forward and backward pseudo-rapidity in Au+Au Collisions at $\sqrt{s_{NN}}$ = 19.6 GeV in RHIC-STAR

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Hot Quarks 2025 @Hefei, China

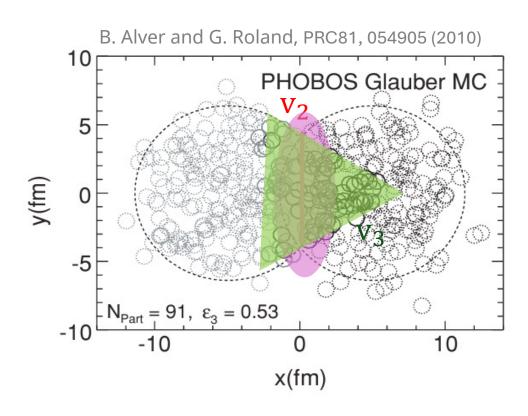
2025.05.11-17

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### Azimuthal anisotropy



#### Anisotropic flow:

- The initial spatial anisotropy is converted into momentum-space anisotropy through fluid expansion.
- The flow strength can be quantified by Fourier coefficient of azimuthal distribution of produced particles:

$$\frac{dN}{d(\phi - \Phi)} = \frac{1}{2\pi} \left\{ 1 + \sum_{n=1}^{\infty} 2\boldsymbol{\nu}_n \cos[n(\phi - \Phi)] \right\}$$

 $\Phi$ : reaction plane

- Elliptic flow  $(v_2)$ : caused by the pressure gradient coming from the initial almond shape.
- Triangular flow( $v_3$ ): originating from the initial triangular shape caused by event-by-event density fluctuation.

## Constraining power of flow on shear viscosity

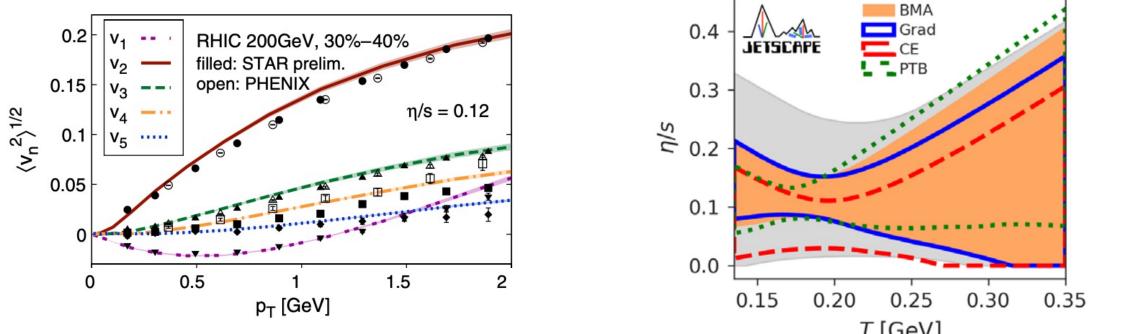
- **IP-Glasma + MUSIC** model shows good agreement with  $v_n$  with  $\eta/s = 0.12 0.20$ .
- Recent studies based on Bayesian analysis further constrain  $\eta/s(T)$ .

Charles et al, PRL 110, 012302 (2013)

#### IP-Glasma + MUSIC

#### **Bayesian analysis**

JE Bernhard, et al., Nat. Phys.15,1113(2019) D. Everett et al., PRL.126,242301(2021)(JETSCAPE)

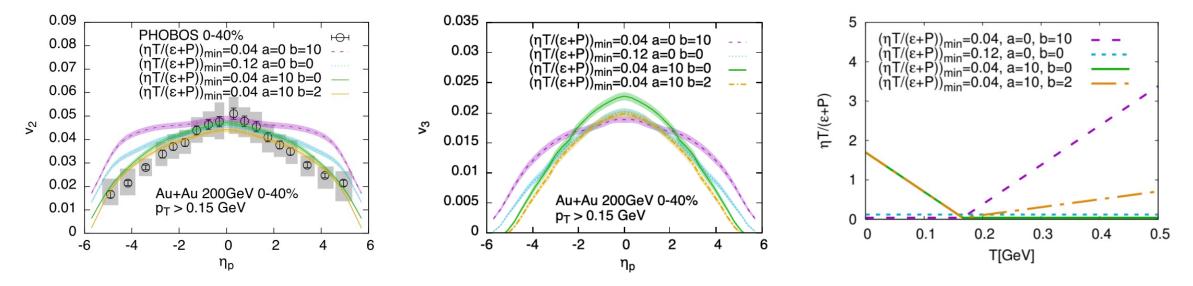


 Constraining the shear viscosity to entropy density ratio leads to understanding the transport properties of QGP.

•  $\eta/s \underset{5/13/25}{\text{still has a large uncertainty.}}$ 

### Constraining power of flow on shear viscosity

Denicol et.al, PRL 116, 212301 (2016)

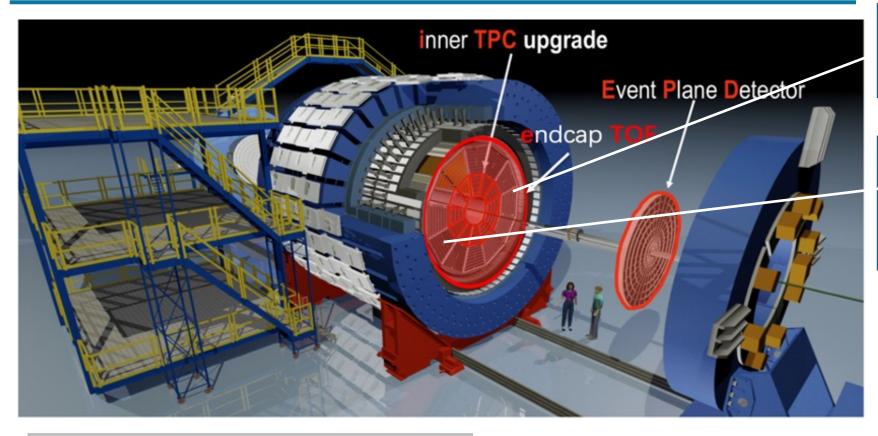


- According to the theoretical study, the η dependence of v<sub>2</sub> and v<sub>3</sub> is sensitive to T-dependence of η/s.
   PHOBOS v<sub>2</sub> exhibits a downward trend with considerable uncertainties, posing challenges for accurate determination.
- Acceptance in our forward detectors allows to capture spectators at lower energies, which remains under-explored.

#### **Final Goal**

• Precise measurements of  $\eta$  dependence of  $v_n$  to constrain the temperature-dependent shear viscosity  $\eta/s(T)$ .

## STAR experiment (BES-II 2019~)



#### **Beam Energy Scan II Program (BES-II)**

- For the search of the QCD critical point.
- Upgraded EPD (Event Plane Detector), iTPC (inner TPC), and forward detectors.
- Wider rapidity coverage and higher statistics than BES-I.
- BES-II energy:  $\sqrt{s_{NN}}$  = 3 GeV to 27 GeV (including FXT mode). 5/13/25 Hot Quarks 2025 (Moe Isshiki)

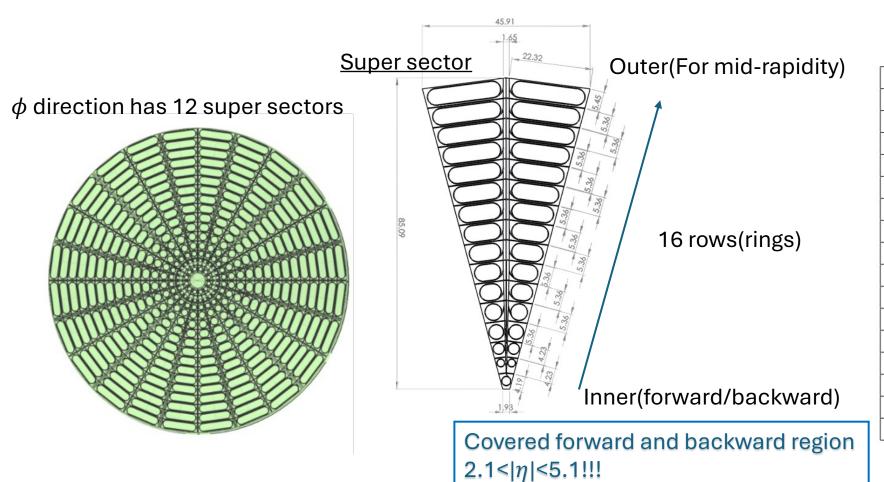
TOFTime of flight measurement ofcharged particles,  $|\eta| < 0.9$ 

# TPC+iTPC(inner readout upgrade)Measure dE/dx, reconstruct the<br/>tracks $|\eta| < 1.5, 0 < \phi < 2\pi$

**EPD** Provided the event planes covering 2.1<  $|\eta| < 5.1$ Scintillation detector

### **Event Plane detector**

- scintillation detector, so **TRACKING is not possible**.
- Particles lose energy in the EPD tile (ADC)



#### Covered $\eta$ acceptance

Row	$r_i$ (cm)	$r_f$ (cm)	$\eta_i$	$\eta_f$
1	4.6	9.0	5.09	4.42
2	9.0	13.4	4.42	4.03
3	13.4	17.8	4.03	3.74
4	17.8	23.33	3.74	3.47
5	23.33	28.86	3.47	3.26
6	28.86	34.39	3.26	3.08
7	34.39	39.92	3.08	2.94
8	39.92	45.45	2.94	2.81
9	45.45	50.98	2.81	2.69
10	50.98	56.51	2.69	2.59
11	56.51	62.05	2.59	2.50
12	62.05	67.58	2.50	2.41
13	67.58	73.11	2.41	2.34
14	73.11	78.64	2.34	2.27
15	78.64	84.17	2.27	2.20
16	84.17	89.70	2.20	2.14

#### Data set

- STAR Bes-II Au+Au 19.6 GeV (2019) ۲
- 260 billion events (after event cut) ۲
- Trigger: minimum bias ۲
- Pileup events: rejected based on the matched timing information of TOF and TPC ullet
- Centrality: determined by multiplicity  $\bullet$

parameters	peripheral. 80% centrality _ <sub>75-80 %</sub>	Central. 0%
$ V_r  \le 2.0 \text{ cm},  V_z  \le 40 \text{ cm},$ centrality 0-80%	10 <sup>7</sup> 10 <sup>6</sup> 10 <sup>5</sup>	Au+Au $\sqrt{s_{NN}}$ =19.6 GeV
$p_T$ >= 0.1 (GeV/c) , $ \eta $ <= 1.5	10 <sup>4</sup>	
0.3 <= nMip	10 <sup>2</sup>	
	10 10 10 10 200 $30$	00 400 500

Cut

cut

Event

Trackcut

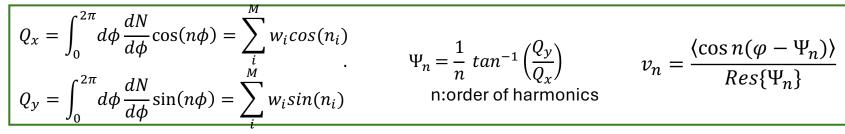
EPD cut

### Event plane determination

#### **Event plane definition**

A. M. Poskanzer and S. A. Voloshin. PRC.58,1671(1998)

This study used the event plane method to measure  $v_n$ ,



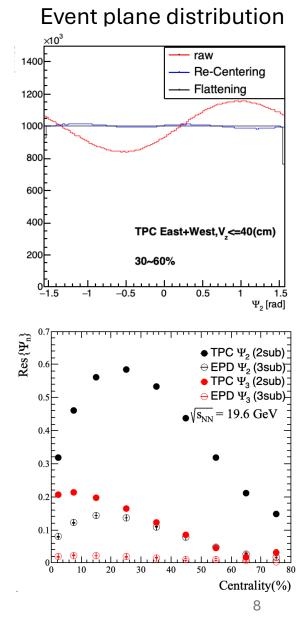
EPD event plane for  $v_n(|\eta| < 1.5)$ : outer 8 rings (participant region) TPC event plane for  $v_n(|\eta| > 2.1)$ : 0.5  $<|\eta| < 1.5$ 

#### **Event plane resolution**

3sub-event method: effective when the resolutions differ between 2 detectors.

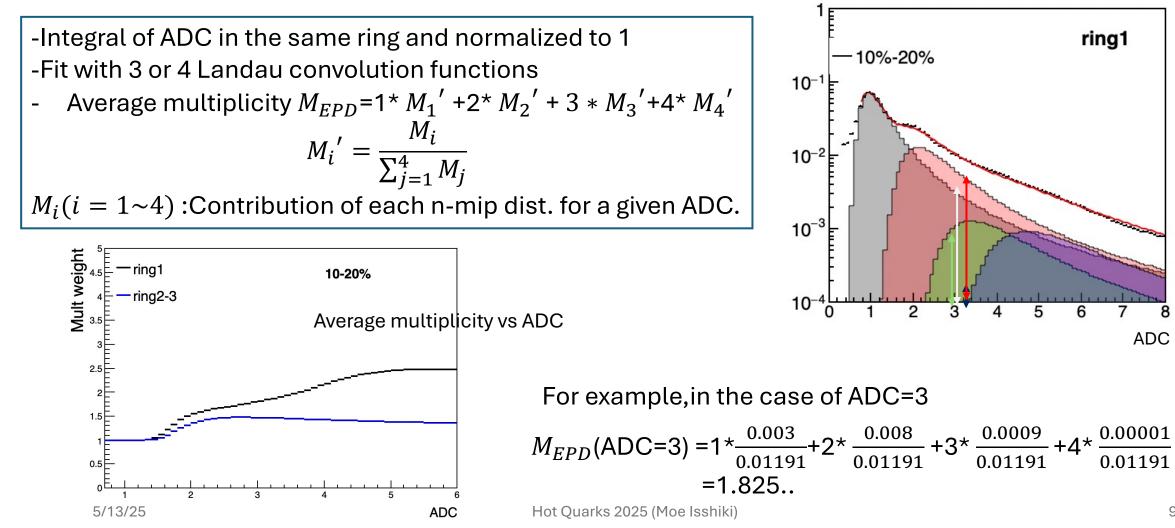
$$\operatorname{Res}\{\Psi_{nTPC(E+W)}\} = \sqrt{\frac{\left\langle \cos(n[\Psi_{n}^{EPDE} - \Psi_{n}^{TPC(E+W)}])\right\rangle \left(\cos(n[\Psi_{n}^{EPDW} - \Psi_{n}^{TPC(E+W)}])\right\rangle}{\left\langle \cos(n[\Psi_{n}^{EPDE} - \Psi_{n}^{EPDW}]\right\rangle}}$$

$$\operatorname{Res}\{\Psi_{nEPD(E+W)}\} = \sqrt{\frac{\left\langle \cos(n[\Psi_{n}^{TPCE} - \Psi_{n}^{EPD(E+W)}])\right\rangle \left(\cos(n[\Psi_{n}^{TPCW} - \Psi_{n}^{EPD(E+W)}])\right\rangle}{\left\langle \cos(n[\Psi_{n}^{TPCE} - \Psi_{n}^{TPCW}]\right\rangle}}$$
5/13/25 Hot Quarks 2025 (Moe Isshiki)



#### Determination of average multiplicity in EPD tile

- EPD: TRACKING is not possible.
- Estimate the average number of incident particles based on the energy loss value.



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### Correction: Detector material effect

#### What's the detector material effect?

Particles generated by collisions may interact with other materials and cause secondary particles.  $\rightarrow$ Need to estimate the effect caused by interactions with detector materials for <u>EPD</u>.

+

**<u>GEANT (STAR detector simulation)</u>** 

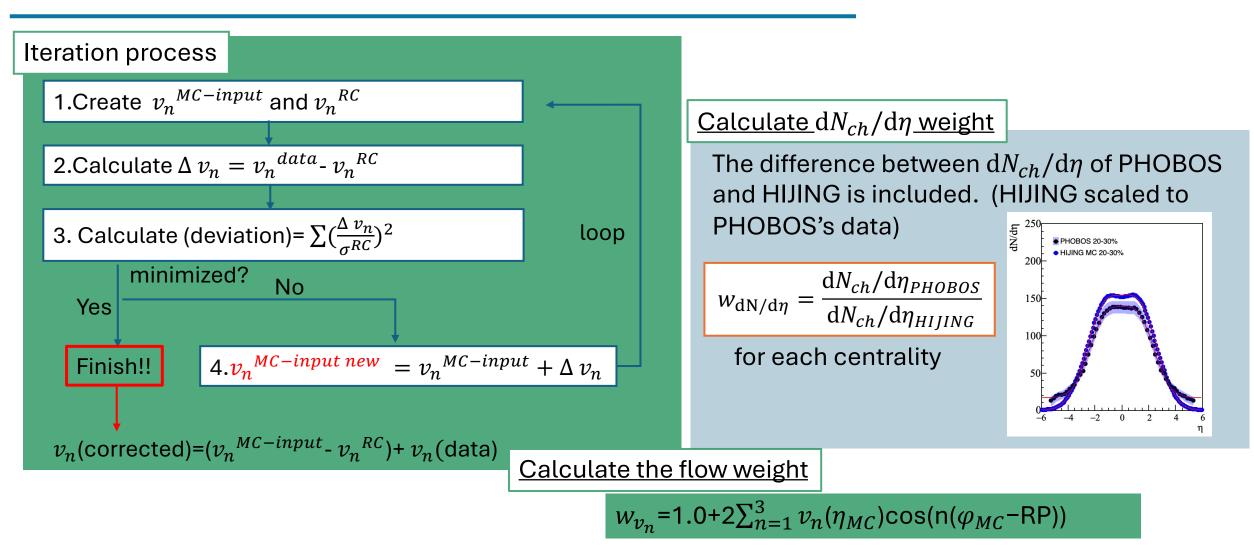
#### HIJING (Heavy Ion Jet INteration Generator)

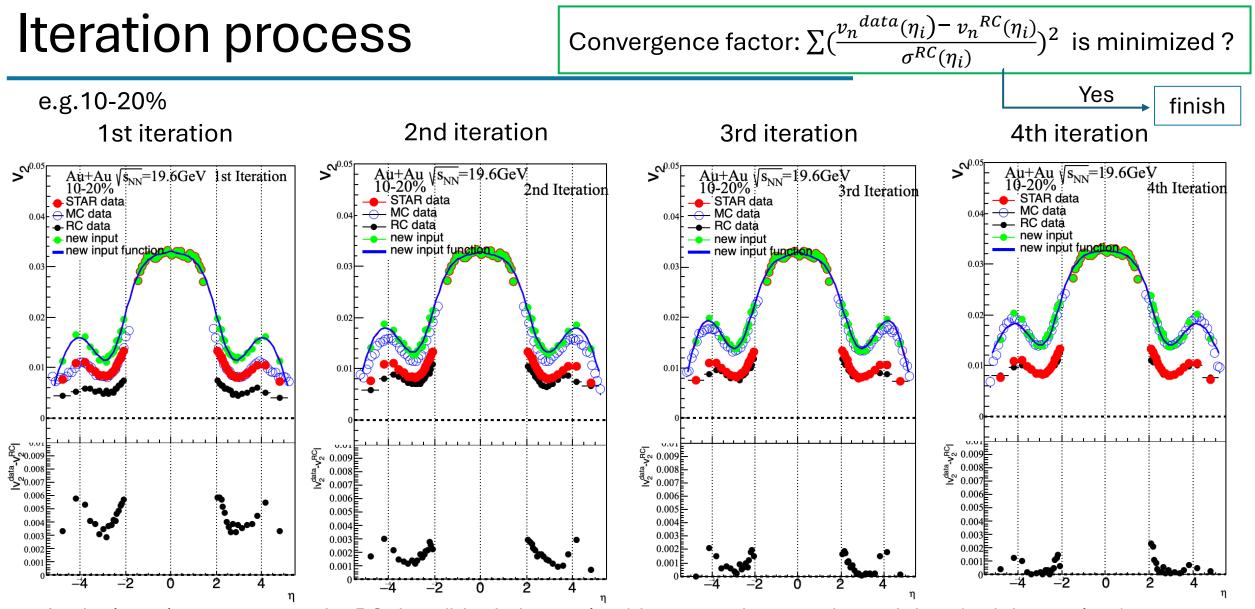
- Collision Energy:  $\sqrt{s_{NN}} = 19.6 \text{ GeV}$
- All Particles (including the neutral particles)

## generate the HIJING events + Reconstruction(RC) HC and RC $v_n$

#### Procedure of analysis

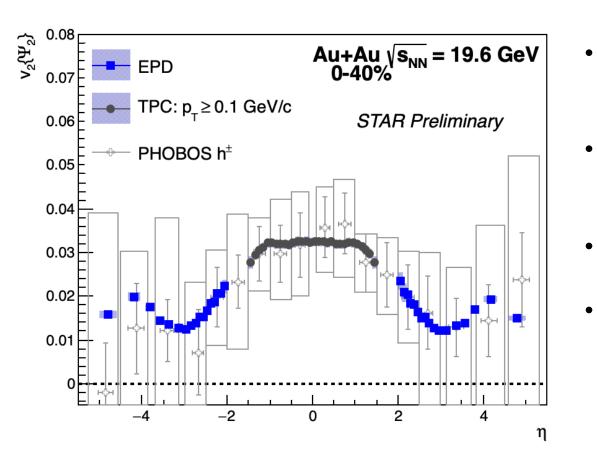
### **Procedure of Simulation**





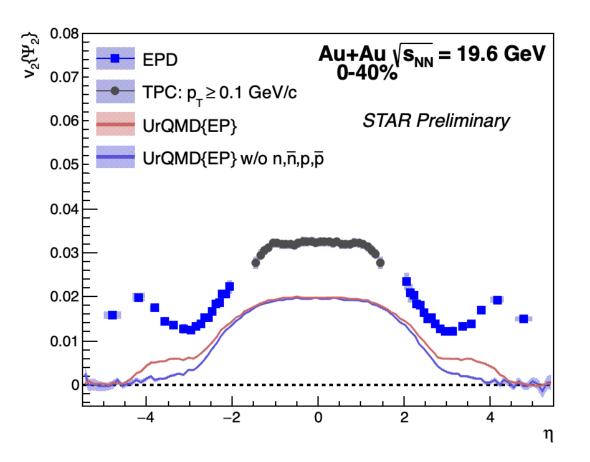
As the iterations progress, the RC data (black data points) become closer to the real data (red data points).

#### $v_2$ comparison with PHOBOS's result



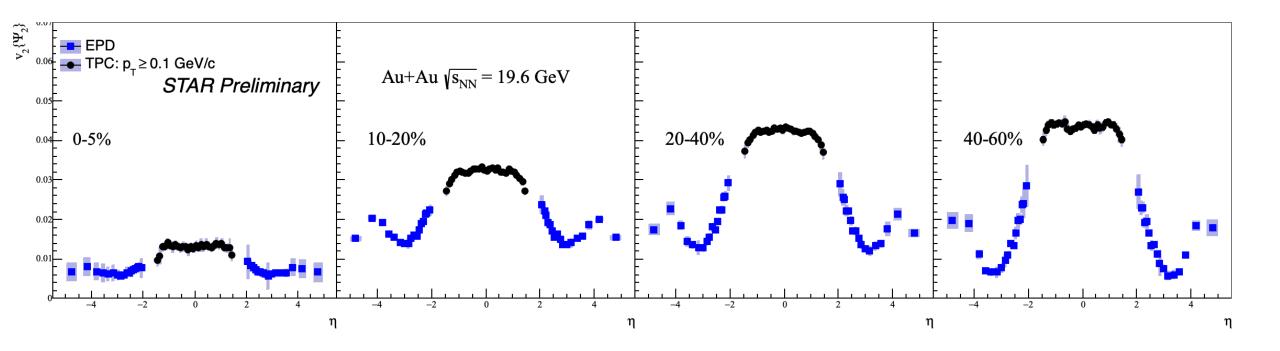
- High-precision measurements performed in the wide  $\eta$  region ( 2.1<  $|\eta|$  < 5.1).
- This result is consistent with PHOBOS's result, which has large uncertainty.
  - Interestingly, there are bump structures at  $|\eta| \sim 4$ .
  - The bump starts to appear at  $|\eta| \sim 3$ , which matches with the beam rapidity at 19.6 GeV ( $y_{beam}$ = 3) where spectator contribution starts to come in.

#### $v_2$ comparison with UrQMD model



- Although the absolute value differs, similar bump structures are seen in UrQMD, where spectators propagate throughout the simulation.
- UrQMD  $v_2$  without protons and neutrons (the main components of spectators) doesn't show the bump structure.

### Centrality dependence



- A clear trend of increasing  $v_2$  magnitude with centrality is observed.
- The bump increases in more peripheral collisions, likely due to spectators.

### Summary

#### Summary

- Elliptic flow  $v_2$  has been measured in the forward and backward regions using the EPD detector in BES-II Au+Au collisions at 19.6 GeV.
- $v_2$  is consistent with the PHOBOS result; however, our result shows significantly improved precision.
- The bump structure appears around the beam rapidity (spectator contribution).
- The trend of bump structure in UrQMD model is qualitatively consistent with our data, even though the absolute value differs.

#### Outlook

- Comparison with other theoretical models (e.g., hydrodynamic model) to constrain the temperature dependence of the shear viscosity.
- $v_3$  measurement is ongoing now.
- Study other collision energies to understand the energy dependence and effects from spectators  $(\sqrt{s_{NN}} = 7.7 \text{ GeV to } 200 \text{ GeV}).$

#### Thank you!!

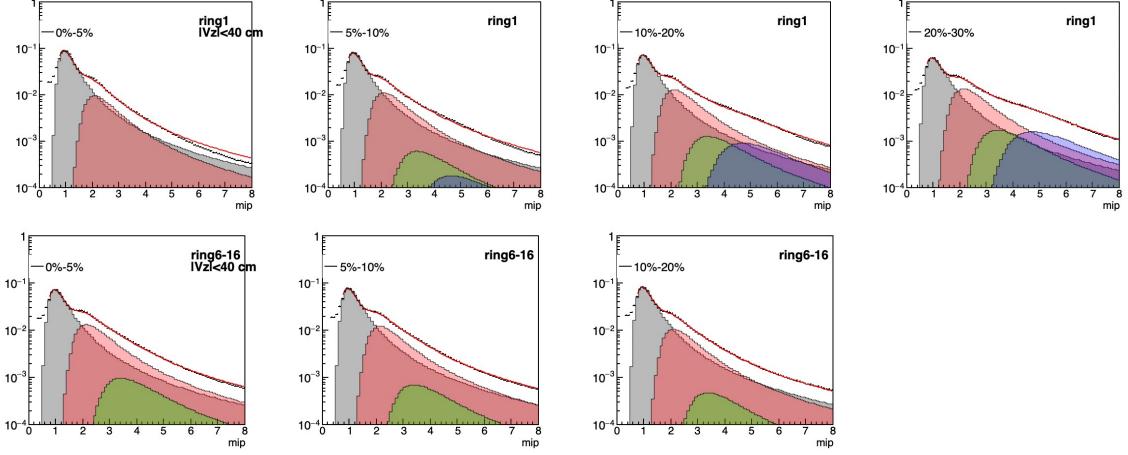
## Back up

### ADC distribution

-Multi landau fit was performed to calculate average multiplicity.

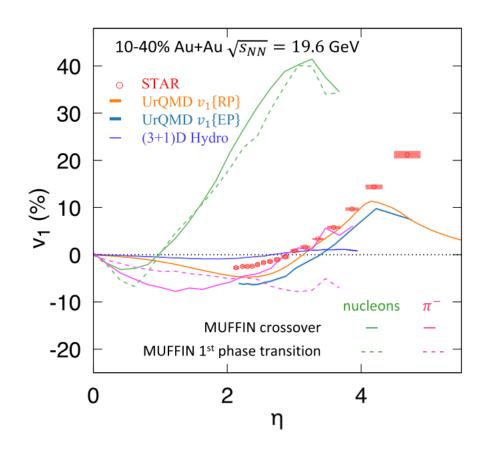
4 landau dist. convolution for ring1

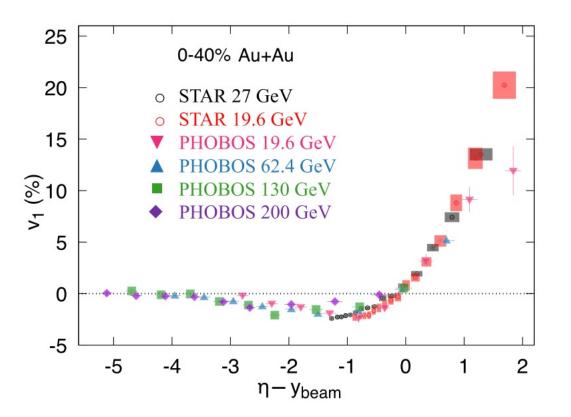
3 landau dist. convolution for ring 2-16



### $v_1$ result from STAR

STAR reported the  $v_1$  result in 2025.





M. I. Abdulhamid *et.al*, Phys. Rev. C **111**, 014906(STAR)