# Directed flow and Elliptic flow in 22.4 GeV CuCu collisions 

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STAR Detector

## Outline

- Introduction
- Data-set and Cuts
- Event Plane Reconstruction
- Results
- Summary



## Introduction

Anisotropic flow: Anisotropy of the azimuthal distribution of particles with respect to the reaction plane: Fourier expansion of the particle's azimuthal distribution with respect to the reaction plane is given by:

$$
\mathrm{E} \frac{\mathrm{~d}^{3} \mathrm{~N}}{\mathrm{~d}^{3} \mathrm{p}}=\frac{1}{2 \pi} \frac{\mathrm{~d}^{2} \mathrm{~N}}{p_{\mathrm{t}} \mathrm{~d} \mathrm{p}_{\mathrm{t}} \mathrm{dy}}\left(1+\sum_{\mathrm{n}=1}^{\infty} 2 \mathrm{v}_{\mathrm{n}} \cos \left(\mathrm{n}\left(\phi-\Psi_{\mathrm{r}}\right)\right)\right)
$$



Coordinate-
Space
Anisotropy

$$
\begin{aligned}
& \mathrm{v}_{\mathrm{n}}=\left\langle\cos \left(\mathrm{n}\left(\phi-\Psi_{r}\right)\right)\right\rangle \\
& \phi=\tan ^{-1}\left(\frac{\mathrm{p}_{\mathrm{x}}}{\mathrm{p}_{\mathrm{y}}}\right)
\end{aligned}
$$



## Data: 22 GeV CuCu Collisions( year 2005)

The Cu+Cu 22 GeV Dataset was taken during run 5 from Mar 22 -24th

Cuts for event:
$|\mathrm{Vx}|,|\mathrm{Vy}|<1.0$
Sqrt(vx2+vy2)<1.0
|Vz|<30.0
Mult>10
BBCAdc<120
Centrality 0-60\%

Cuts for track:
Pt(TPC\&FTPC) 0.15,2.0
FitPts(TPC) 15,50
FitPts(FTPC) 5,11
DCA global(FTPC) <1.0
$\mathrm{Eta}(\mathrm{TPC})<|1.2|$
$\mathrm{Eta}(\mathrm{FTPC})<|2.5-4.0|$


## Event Plane



## BBC Event Plane

We have used psi weight method and shift method [J.Barrette et al Phys.Rev.C56(1997)3254] to make the event plane distribution flat.



shift correction method

## Directed Flow v1(n) at 22GeV CuCu

We use east BBC event plane to calculate v1 in west and west BBC event plane to calculate v1 at east in FTPC eta[2.5-4.0] coverage region and full event plane in TPC eta[1.2-1.2] coverage


BBC EP Resolution~15\%


Asymmetry is within stat. error

## Comparison with Other Energies



Comparison with STAR

Elliptic Flow $\mathrm{v}_{2}(\mathrm{pt})$ at 22 GeV CuCu Collisions


Elliptic Flow $\mathrm{v}_{2}(\eta)$ at 22 GeV CuCu Collisions
0-60\% Centrality
0-40\% Centrality
arXiv:nucl-ex/0701054v1



## Comparison with Other Energies

Comparing with PHOBOS(CuCu)

$$
\text { Phys. Rev. Lett. 98, } 242302 \text { (2007) }
$$



## Summary

- We present the first measurement of charged particle directed flow $\mathrm{v}_{1}(\mathrm{\eta})$ in CuCu collisions at 22.4 GeV using BBC event plane.
- Directed flow $\mathrm{v}_{1}(\mathrm{\eta})$ in 22.4 GeV CuCu is similar to 19.6 GeV AuAu collisions and we observe nice scaling behavior for different energies.
- We present the first STAR results for v2( f$)$ and $\mathrm{v} 2(\mathrm{pt})$ with TPC Event plane at 22.4 GeV CuCu.
- Elliptic flow v2(pt) in 22.4 GeV CuCu collisions follows the similar trend with CuCu 200 GeV collisions.
- Elliptic flow v2( $\eta$ ) in 22.4 GeV CuCu collision is consistent with the PHOBOS result at 22.4 GeV CuCu and we observe the incident energy dependence of $\mathrm{v2}(\mathrm{\eta})$.


## Back Up Slides

## BBC Geometry

STAR Beam-Beam Counter Schematic


- Nearly circular Geometrical Shape.
- About 3.7 m from the IR center.
- $2.2<|\mathrm{D}|<5.2$.
- 4 rings with 36 tiles on each side.
- 18 small tiles in inner two rings.
- 18 big tiles in outer two rings.
- 24 PMTs read the 36 tiles/side; linked tiles are mostly in outer rings.
- Find the 1st-order Event Plane at Low Beam Energies, following the mode of the ZDC- SMD at higher energies
- Less non-flow than other methods based on FTPC/TPC


## Gain correction

We have made channel-by-channel gain correction assuming that each channel at the same radius from the center of the beam pipe should be hit by same number of particles.

Channel by channel gain correction

BBCAdc East


Frl Sep 11 09:49:56 2009

## BBCAdc West



Frl Sep 11 09:49:01 2009

## BBC Event Plane

$$
\begin{aligned}
& \mathrm{Q} \cos \Psi_{0}=\mathrm{X}=\sum \mathrm{w}_{\mathrm{i}} \cos \Phi_{\mathrm{i}} \\
& \mathrm{Q} \sin \psi_{0}=\mathrm{Y}=\sum \mathrm{w}_{\mathrm{i}} \sin \phi_{\mathrm{i}} \\
& \tan \left(\psi_{0}\right)=\frac{\sum \sum_{\mathrm{i}} \mathrm{w} \sin \left(\varphi_{\mathrm{i}}\right)}{\sum \mathrm{w}_{\mathrm{i}} \cos \left(\varphi_{\mathrm{i}}\right)} \quad \mathrm{W}_{\mathrm{i}}=\frac{\mathrm{BBCadc}_{\mathrm{i}}}{\sum \mathrm{BBCadc}_{\mathrm{i}}} \\
& \mathrm{v}_{1}=\frac{\mathrm{v}_{\mathrm{obs}}}{\operatorname{Res}} \quad \mathrm{~V}_{\mathrm{obs}}=\left\langle\cos \left(\theta-\Psi_{0}\right)\right\rangle \\
& \operatorname{Res}=\left\langle\cos \left(\psi_{0}-\psi_{\mathrm{R}}\right\rangle=\mathrm{C} \sqrt{\left\langle\cos \left(\psi^{\mathrm{E}}-\psi^{\mathrm{w}}\right)\right\rangle}\right.
\end{aligned}
$$

$\Psi_{0}$ : Event plane angle
$\Phi_{i}$ : Azimuthal angle of center of the BBC tile. Where multiple tiles are connected to same $\mathrm{PMT}, \Phi_{i}$ corresponds to the center of gravity of the tile combination.
$\theta$ :azimuthal angle of the particle.
C : Conversion factor [A.M. Poskanzerer and S.A.Voloshin,1998 ]

## Shift Correction Formula:

$$
\begin{array}{r}
\Psi_{1}^{\prime}=\Psi_{1}+\sum_{n} \frac{2}{n}\left(-\left\langle\sin n \Psi_{1}\right\rangle \cos n \Psi_{1}+\left\langle\cos n \Psi_{1}\right\rangle \sin n \Psi_{1}\right) \\
\text { J.Barrette et al. Phys. Rev. C56 (1997)3254 }
\end{array}
$$

## Raw BBC Adc Distribution



BBC Adc Distribution after BBCAdc cut (Events with BBCAdc<120)
































BBC Adc distribution after gain correction


