

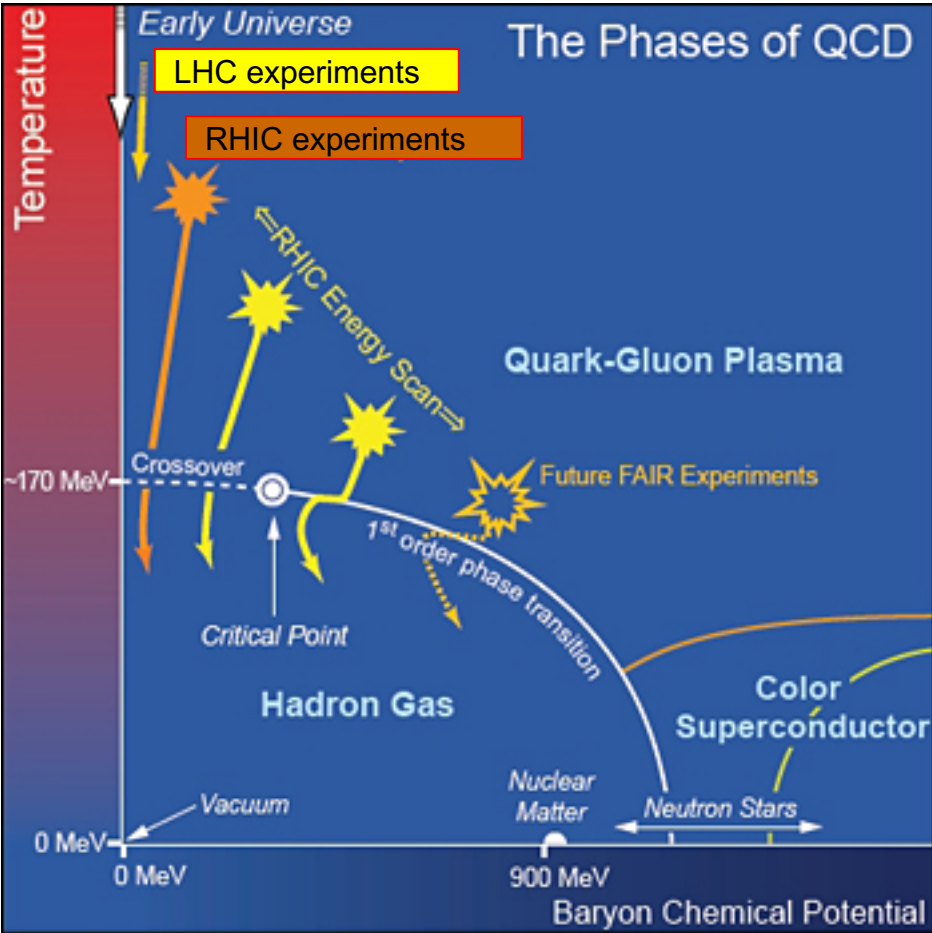
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# Search for the Chiral Magnetic Effect in Isobaric Collisions

## - A Status Report

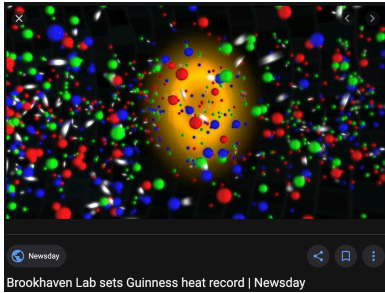
Aihong Tang (for the STAR Collaboration)  
Brookhaven National Lab



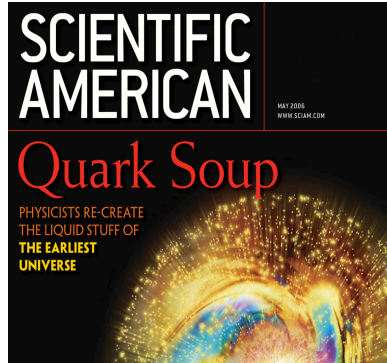


## 1. QCD phase diagram

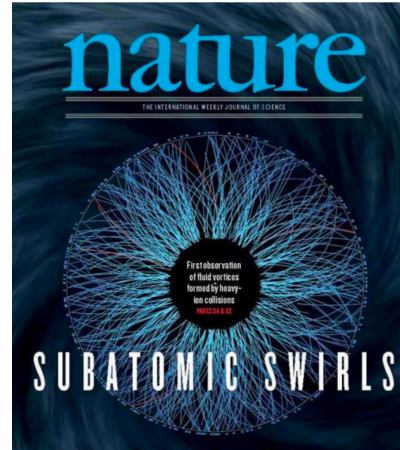
# Heavy Ion Missions at RHIC and the LHC



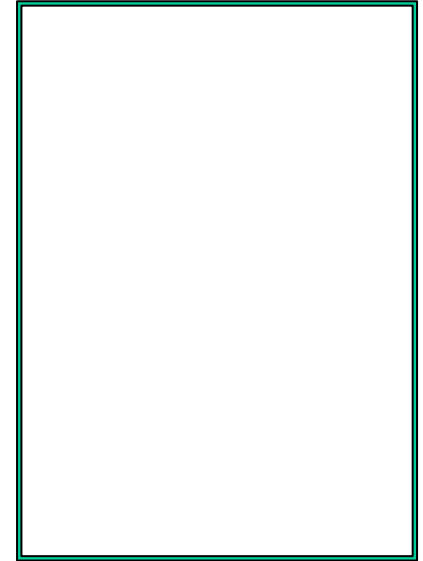
**Hottest**



**Least viscous**



**Most vortical**



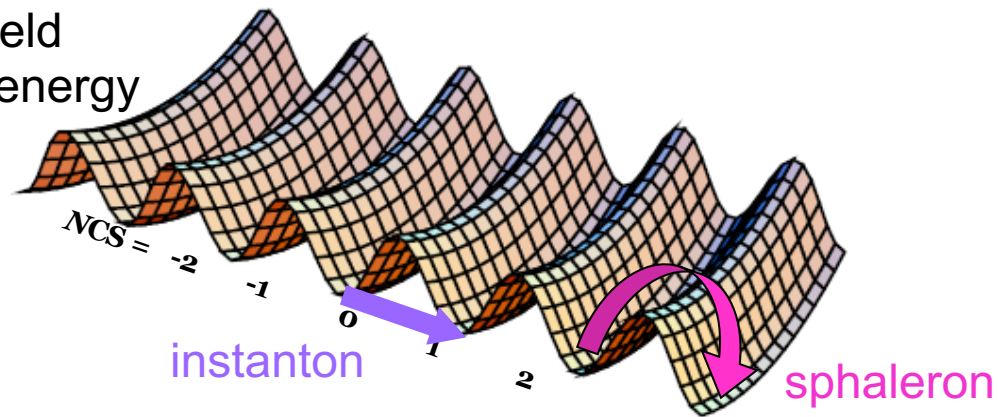
???

1. QCD phase diagram
2. Dynamic properties of QCD matter

Relativistic Heavy Ion Collisions : Excellent QCD test ground

# QCD Vacuum Transition

Gluonic field  
potential energy

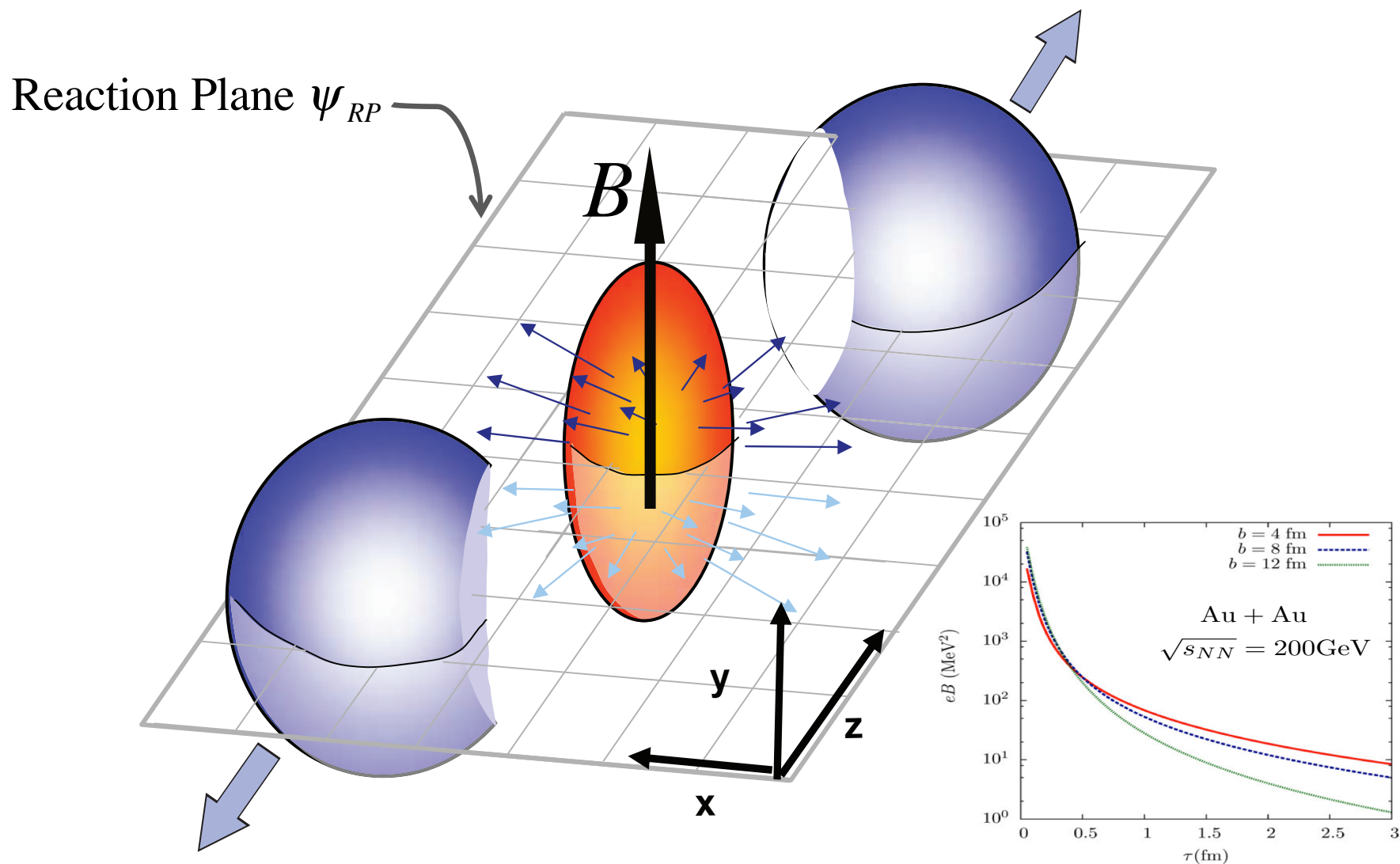


- QCD vacuum transition  $\rightarrow$  nonzero topological charge  $\rightarrow$  chirality imbalance

Local violations of fundamental symmetries  
of space-time in strong interaction



# The Setup of Heavy Ion Collision



D. Kharzeev, L. McLerran and H. Warringa. Nuclear Physics A 803, 227 (2008).

# The Setup of Heavy Ion Collision

Earth



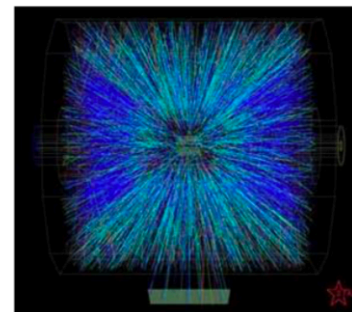
Magnet



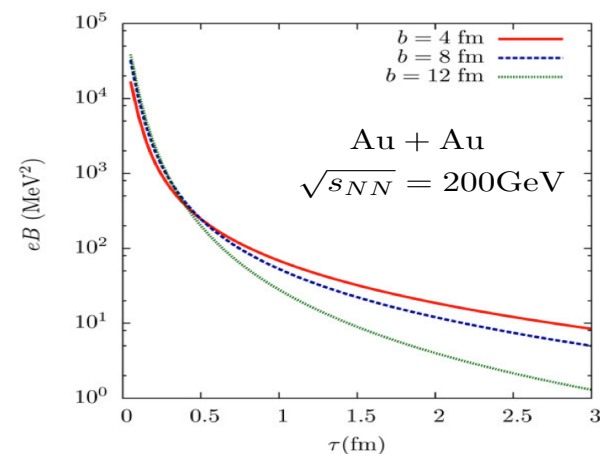
Neutron star



Heavy-ion collisions



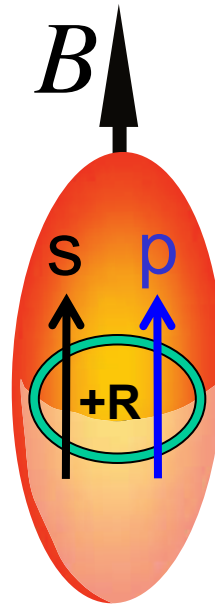
Plot courtesy of X. Huang



D. Kharzeev, L. McLerran and H. Warringa. Nuclear Physics A 803, 227 (2008).

# Chiral Magnetic Effect

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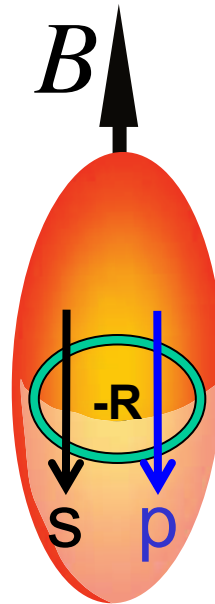
Positively charged particles moving upward.

$$\langle \vec{s} \rangle \propto (Qe) \vec{B}$$

$$\langle \vec{p} \rangle \propto \mu_A \langle \vec{s} \rangle$$

# Chiral Magnetic Effect

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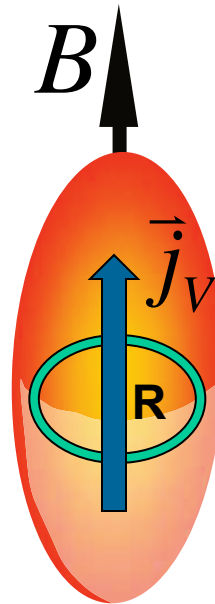
Negatively charged particles moving downward.

$$\langle \vec{s} \rangle \propto (Qe) \vec{B}$$

$$\langle \vec{p} \rangle \propto \mu_A \langle \vec{s} \rangle$$

# Chiral Magnetic Effect

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Charge current

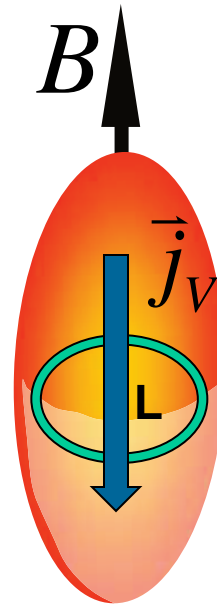
$$j_V = \frac{N_c e}{2\pi^2} \mu_A B$$

Chiral Magnetic Effect (CME)

Charge separation along B direction.

# Chiral Magnetic Effect

---



Charge current

$$j_V = \frac{N_c e}{2\pi^2} \mu_A B$$

Chiral Magnetic Effect (CME)

Charge separation along B direction.  
Dipole effect, flips event by event.

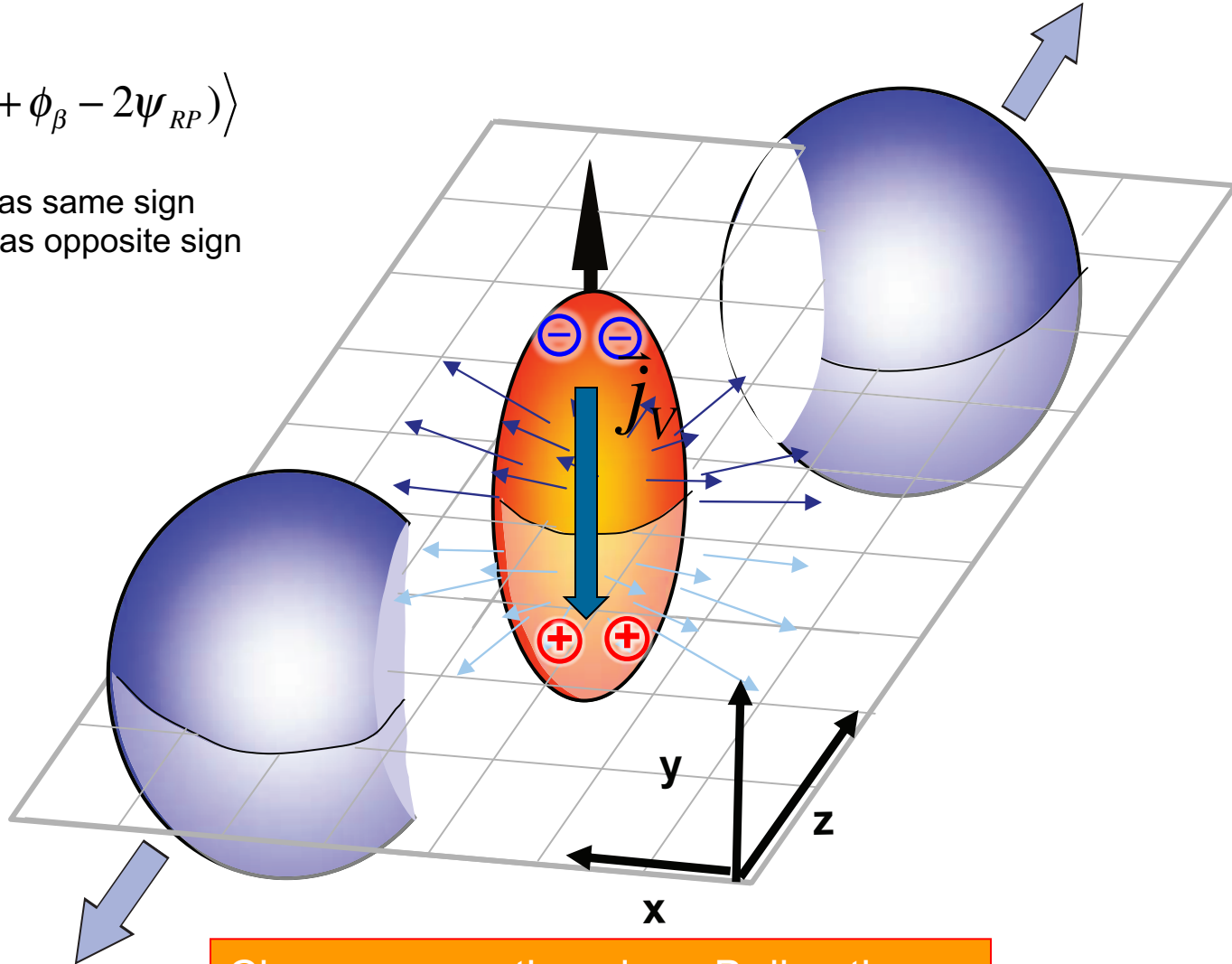
# Chiral Magnetic Effect

$$\gamma_{ss} < \gamma_{os}$$

$$\gamma \equiv \langle \cos(\phi_\alpha + \phi_\beta - 2\psi_{RP}) \rangle$$

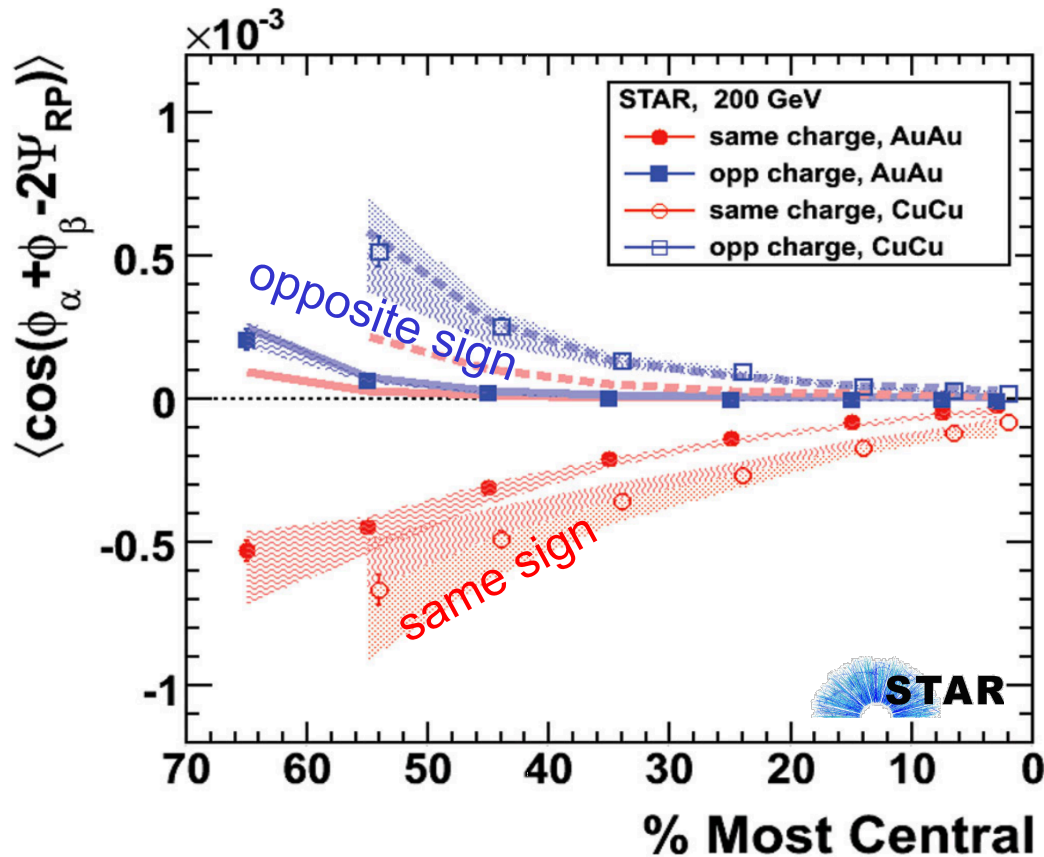
ss :  $\alpha$  and  $\beta$  has same sign

os :  $\alpha$  and  $\beta$  has opposite sign



Charge separation along B direction.  
Dipole effect, flips event by event.

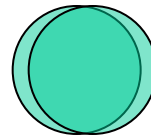
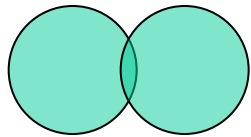
# Chiral Magnetic Effect



$$\gamma_{SS} < \gamma_{OS}$$

$$\gamma \equiv \langle \cos(\phi_\alpha + \phi_\beta - 2\psi_{RP}) \rangle$$

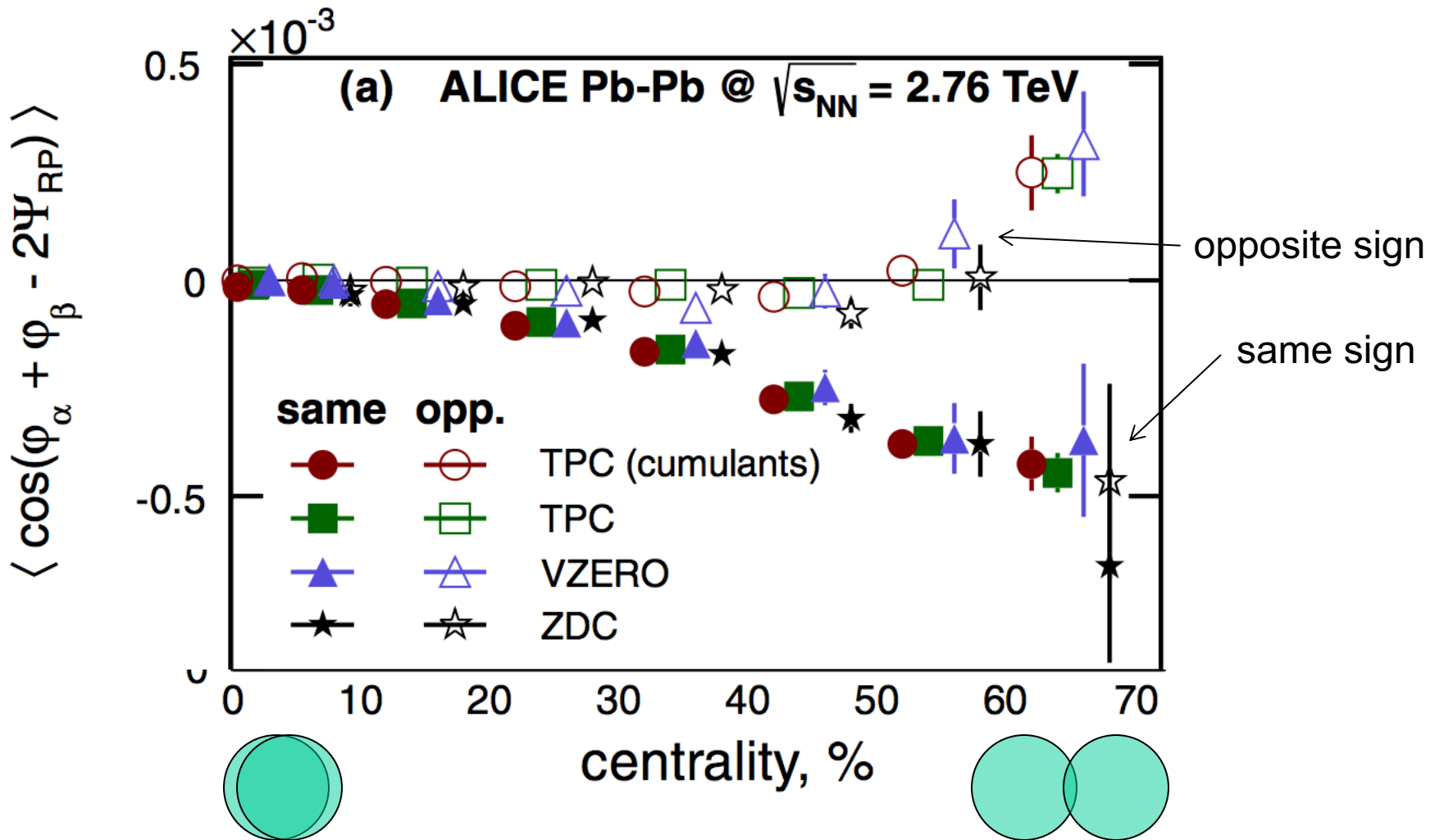
“A signal consistent with several expectations from the theory is detected.”



STAR PRL103 251601 (2009)  
STAR PRC 81 054908 (2010)



# Chiral Magnetic Effect



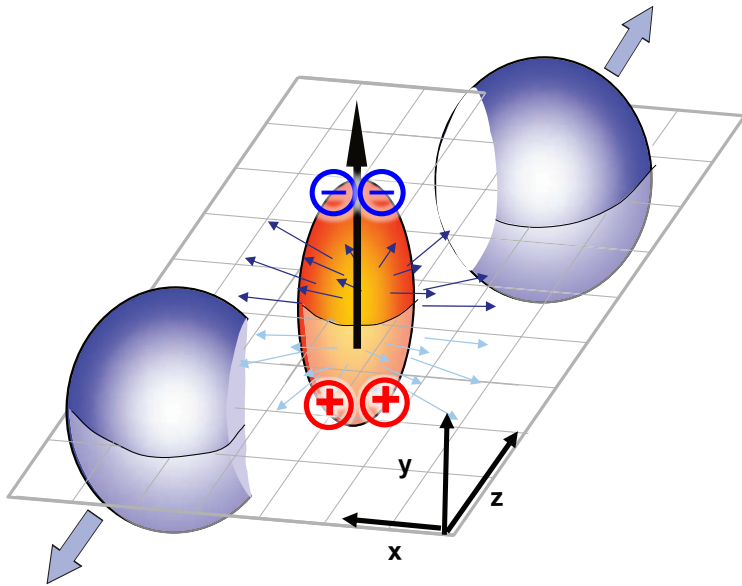
$$\gamma_{SS} < \gamma_{OS}$$

$$\gamma \equiv \langle \cos(\phi_\alpha + \phi_\beta - 2\Psi_{RP}) \rangle$$

ALICE PRL110 012301 (2013)

# Chiral Magnetic Effect : Backgrounds

Signal



Backgrounds:

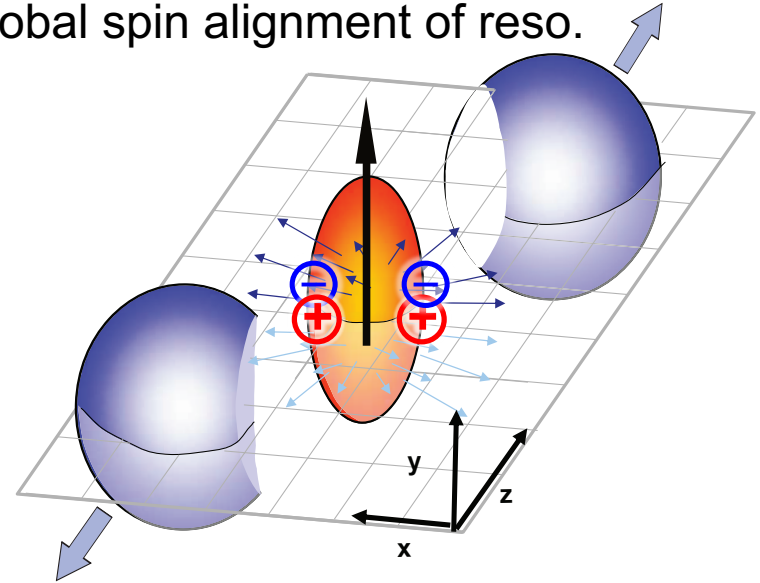
Total Momentum Conservation (TMC)

Local Charge Conservation (LCC)

Flowing cluster in plane

Global spin alignment of reso.

...



Flow boost collimates pairs more strongly in-plane than out of plane. Most background is driven by flow ( $v_2$ )

A. Bzdak, V. Koch, and J. Liao, Lect. Notes Phys. 871, 503 (2013)

S. Pratt, S. Shlichting and S. Gavin, PRC 84 024909 (2011)

S. Schlichting and S. Pratt, PRC 83 014913 (2011)

F. Wang PRC 81 064902 (2010)

A. Tang, Chin. Phys. C 44 No.5 054101 (2020)

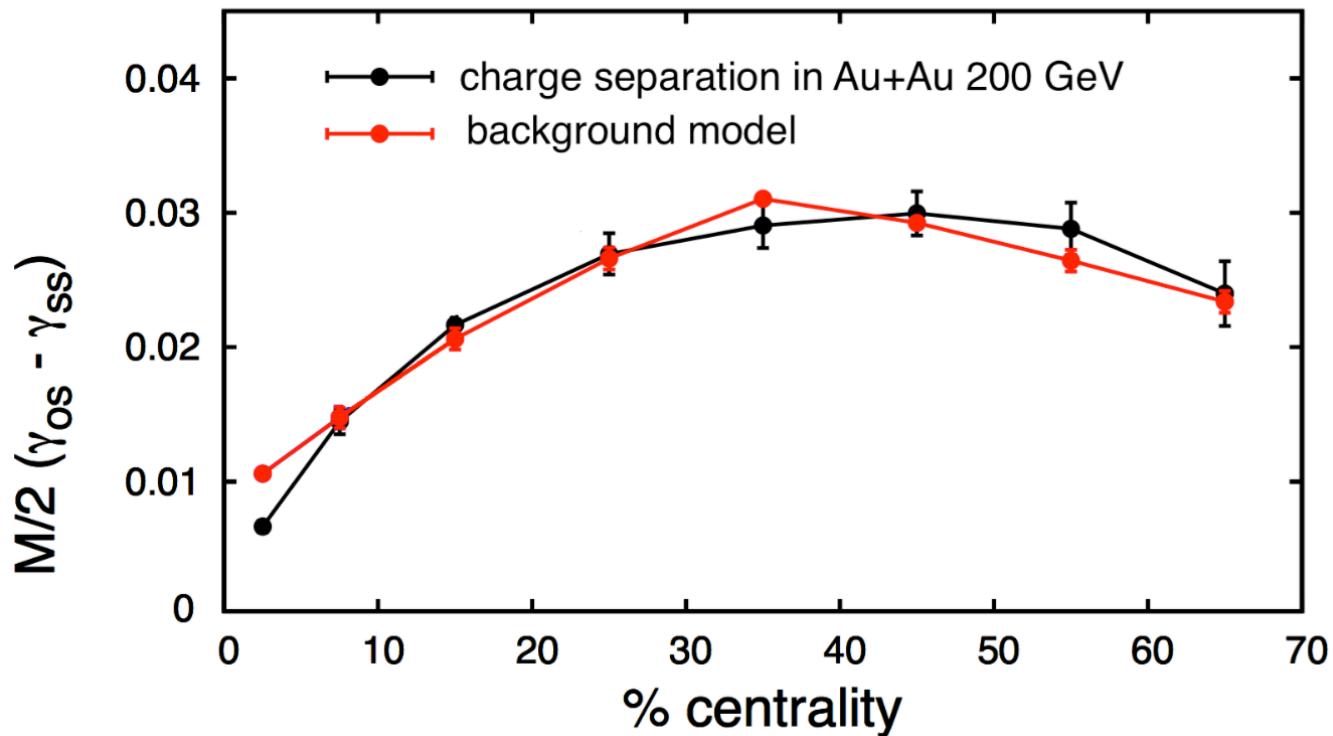
...

Aihong Tang

INT "Critically and Chirality", May 11-22 2020

# Chiral Magnetic Effect : Backgrounds

S. Schlichting and S. Pratt, Phys. Rev. C 83, 014913 (2011)



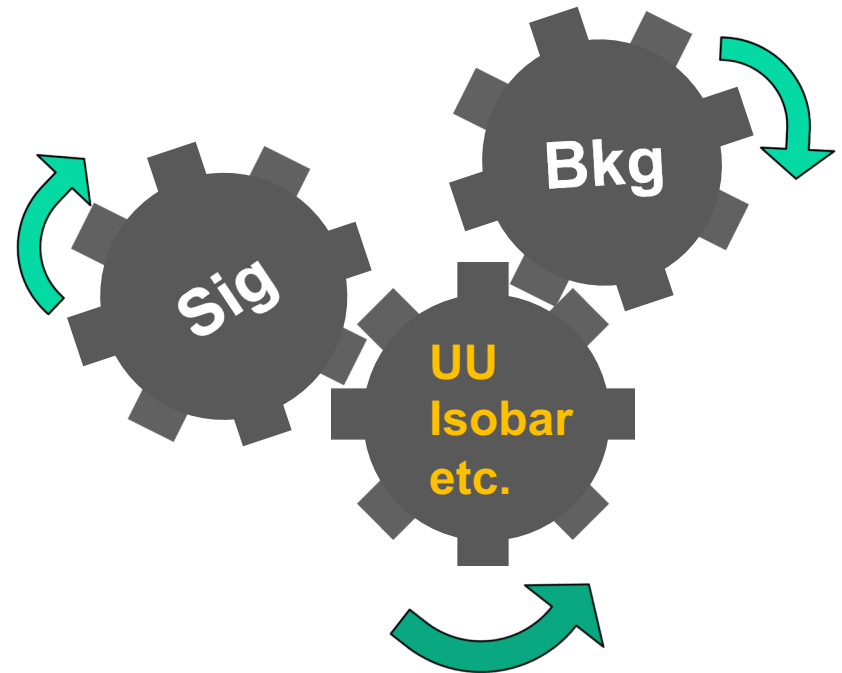
Backgrounds contributions could be comparable in magnitude to measured correlations

# Understanding the Background

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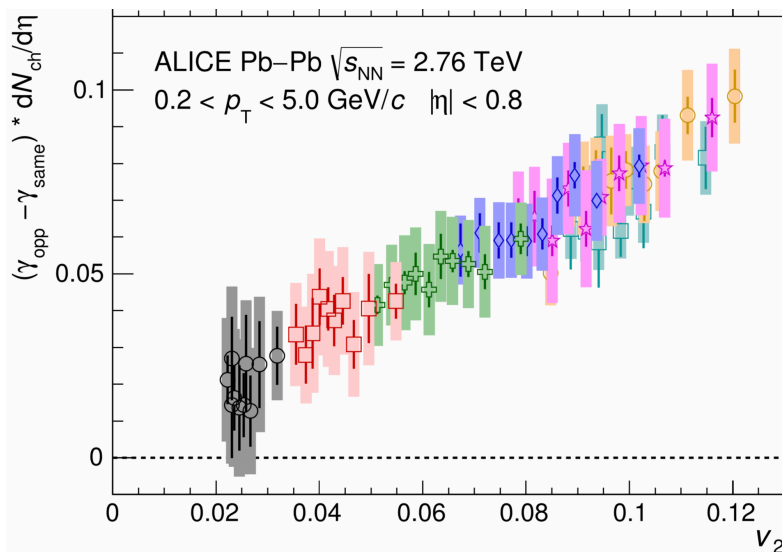
1. Turn off background ( event shape engineering,  $\Delta\gamma(M_{\text{inv}})$  ).
2. Turn off signal ( small system ).
3. Vary background only ( Ultra-central U+U vs. Au + Au ).
4. Vary signal only ( Isobaric collisions ).

Play with it and learn !



# Understanding the Background : Turn off Background

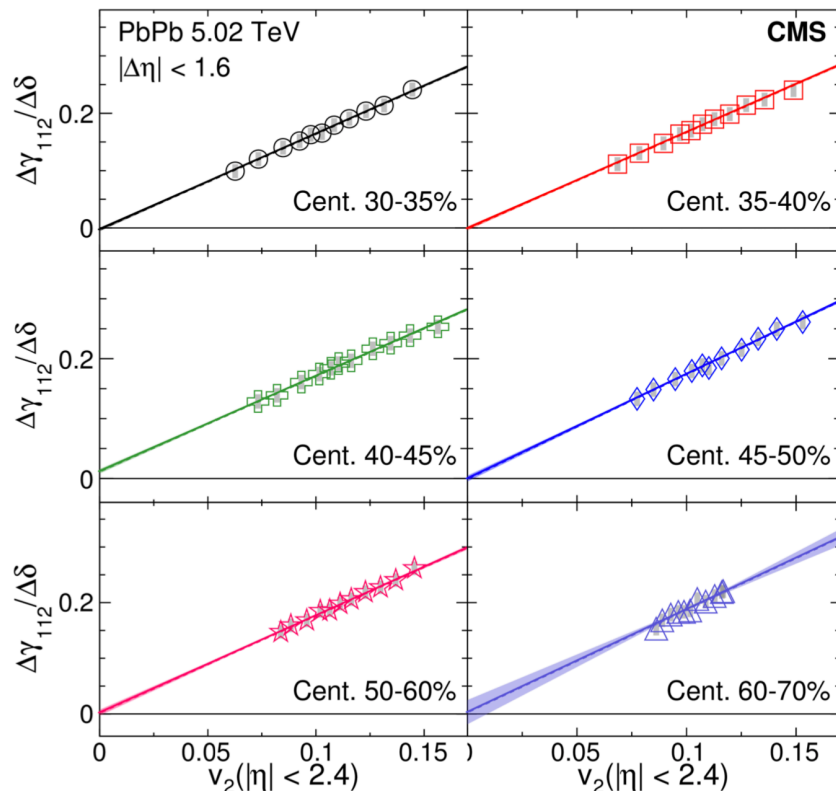
## Event Shape Engineering (ESE)



ALICE, Phys. Lett. B 777 151 (2018)

$$\Delta\gamma = \text{background} (\propto v_2) + \text{signal}$$

More ref. on ESE and EbyE  $v_2$  method :  
 J. Schukraft, A. Timmins and S. Voloshin, Phys. Lett. B 719 394 (2013)  
 STAR, Phys. Rev. C 89 44908 (2014)  
 F. Wen, J. Bryon, L. Wen and G. Wang. Chin. Phys. C 42 No. 1 014001 (2018)



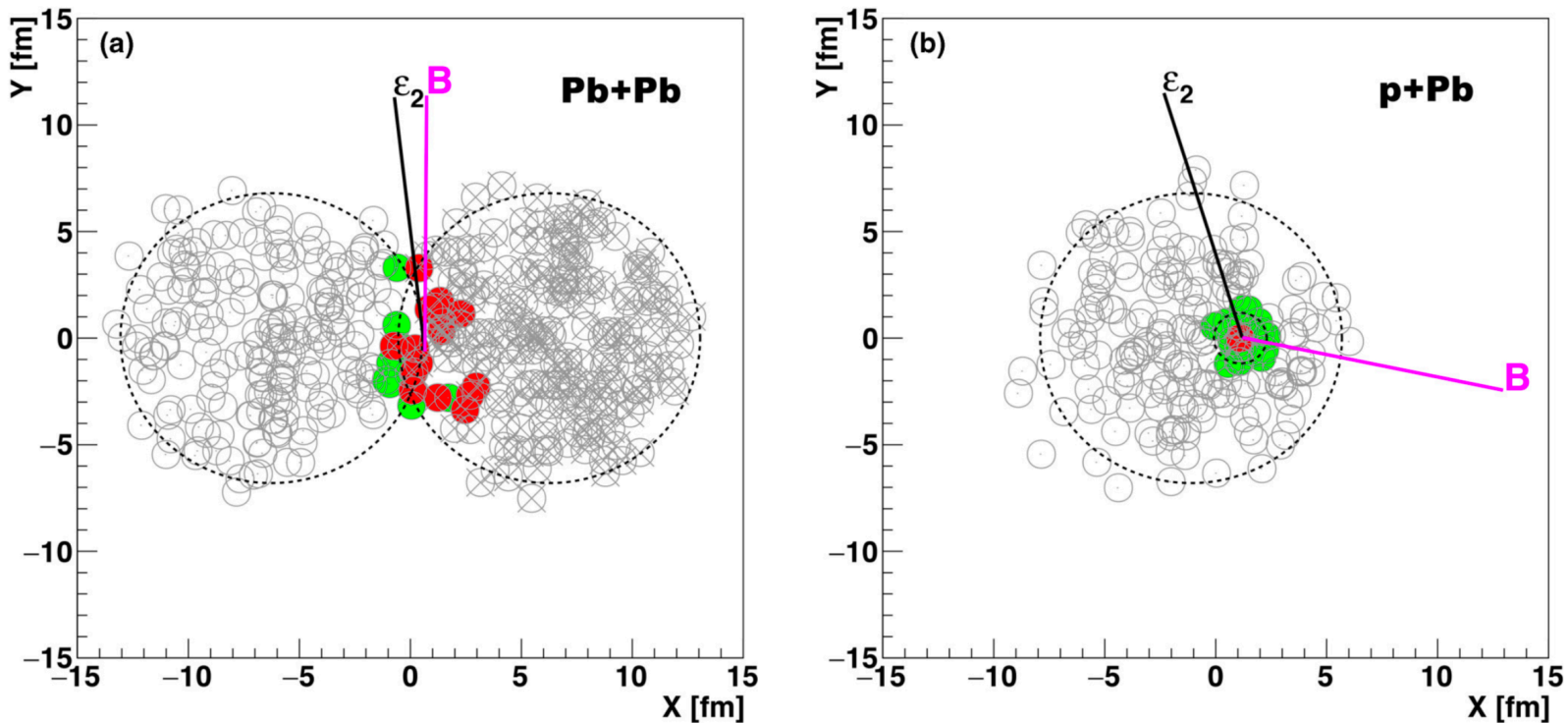
CMS, Phys. Rev. C 97 044912 (2018)

Caveat : long projection over empty  $v_2$  space.

Signal in  $\Delta\gamma$ , if exists, should be very small at LHC energies.

# Understanding the Background : Turn off Signal

## Small System

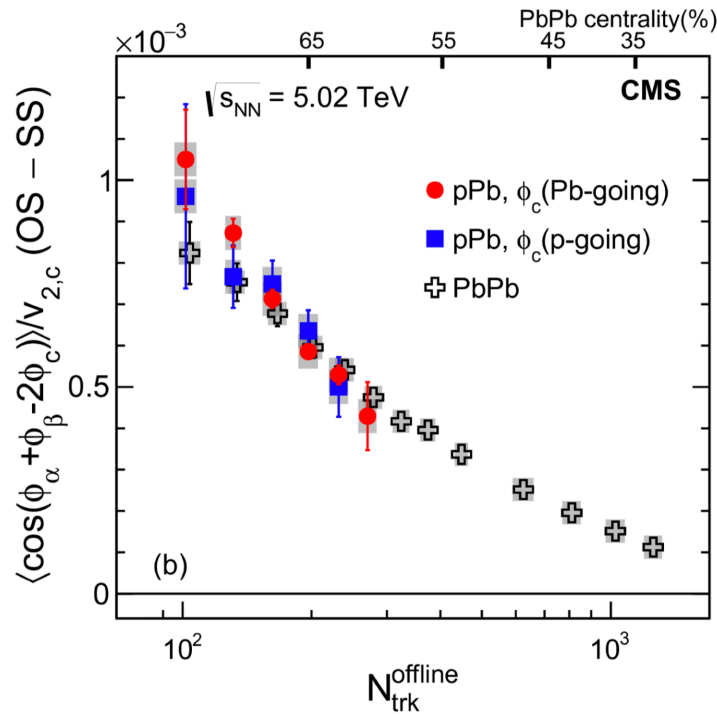


CMS, Phys. Rev. Lett. 118, 122301 (2017)  
R. Belmont and J. Nagle, Phys. Rev. C 96 024901 (2017)

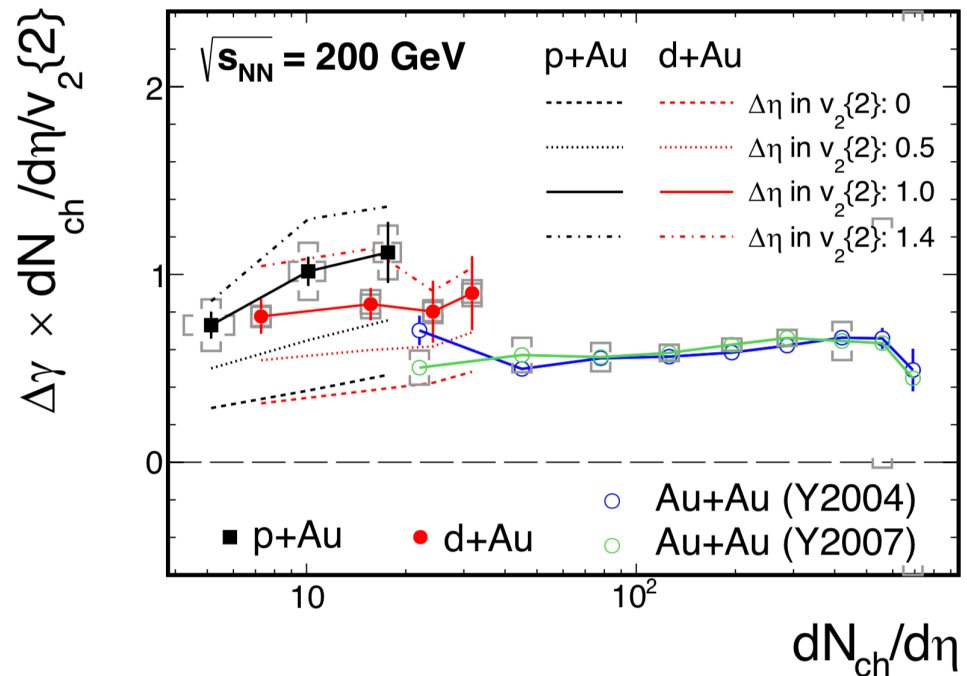
In small systems, B is not correlated with event plane, killing signal

# Understanding the Background : Turn off Signal

## Small System



CMS, Phys. Rev. Lett. 118, 122301 (2017)



STAR, Phys. Lett. B 798, 134975 (2019)

Similarity between large and small system at LHC suggests that background can dominate  $\Delta\gamma$ .

Open questions/issues remain :  
Residual B field in pPb at LHC ? [PRC 97 024905]  
RHIC and LHC have different  $\Delta\eta$  gap.  
Should RP independent background be the same between small and large system ?

# Understanding the Background : Vary Background

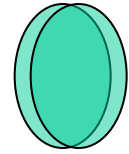
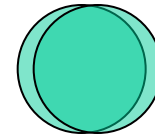
(Ultracentral collisions, **Au+Au vs U+U**)

Signal driver (B)

← **Fixed at  $\sim 0$**

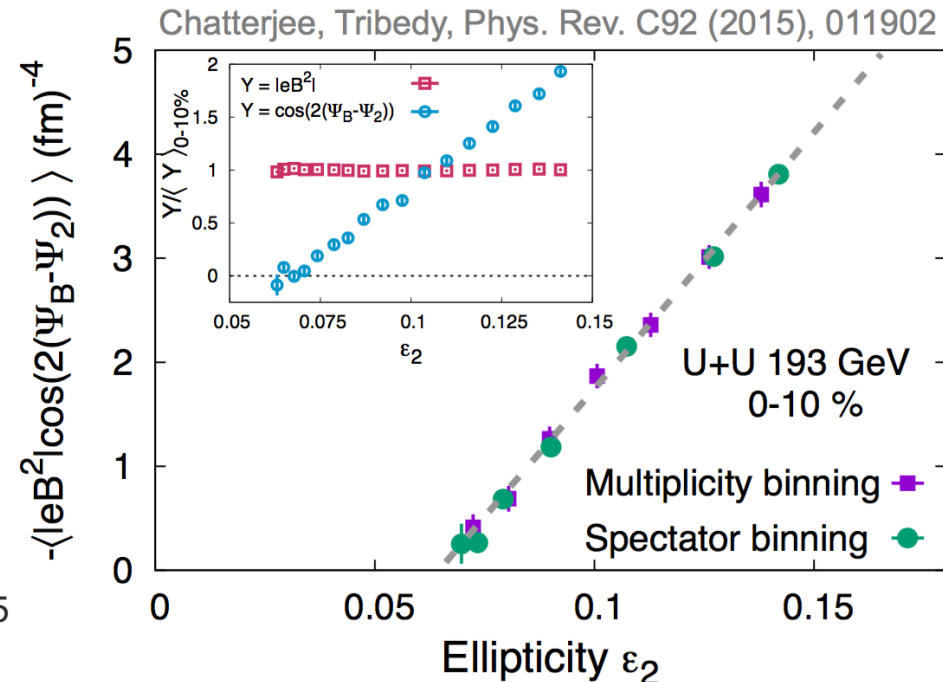
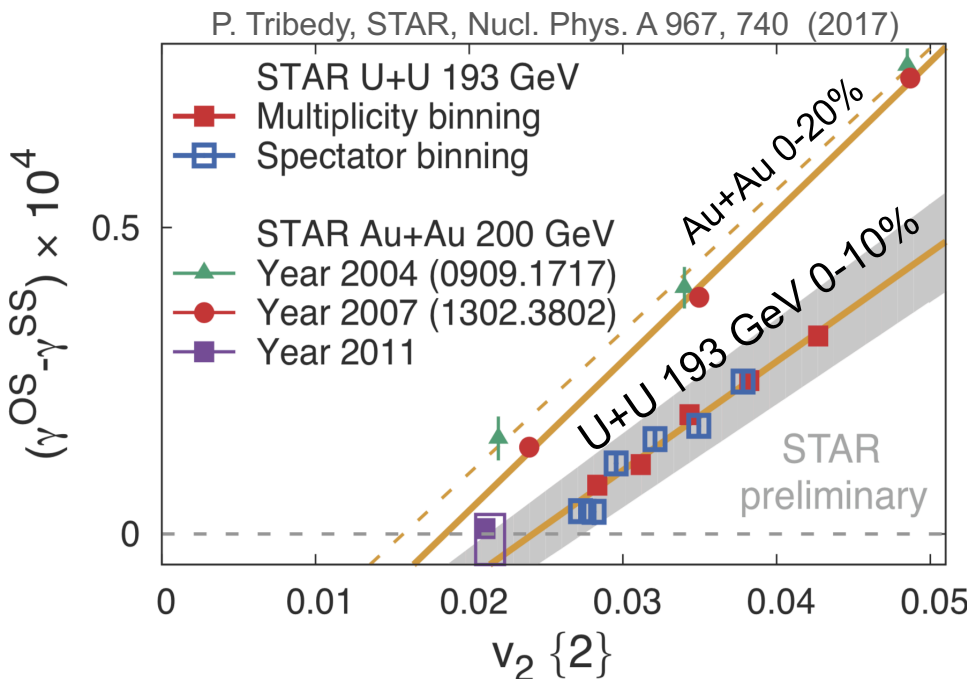
Background driver ( $v_2$ )

← **Vary**



Au+Au (ultracentral)  
 $\varepsilon_2 \sim 0$ ,  $B \sim 0$

U+U (ultracentral)  
 $\varepsilon_2 \neq 0$ ,  $B \sim 0$



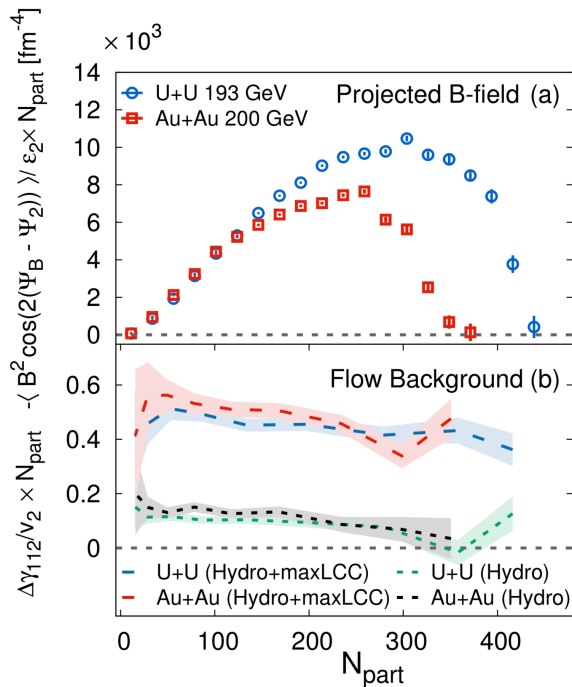
$(\gamma_{OS} - \gamma_{SS}) \sim 0$  while  $v_2$  is, 1) finite, and 2) larger in U+U than in Au+Au  
 $(\gamma_{OS} - \gamma_{SS})$  and the effective B has similar dependency on anisotropy

Not consistent with the existing understanding of background such as LCC

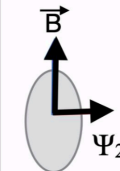
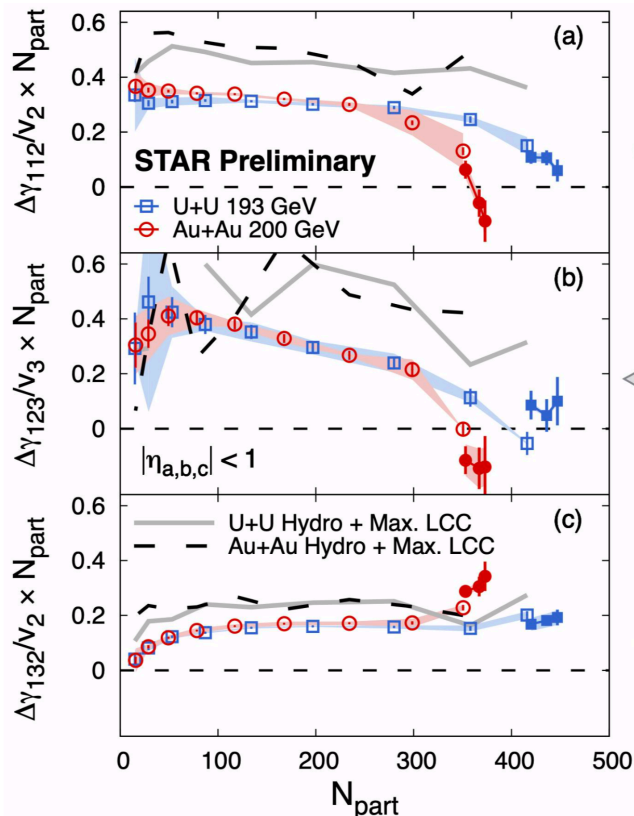
Note : The analysis depends on pseudorapidity separation of two particles.



# More on U + U collisions

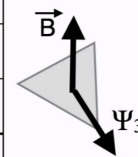


B-fields : U + U  $\neq$  Au + Au  
Backgrounds : U + U  $\sim$  Au + Au



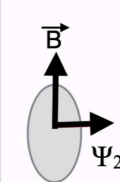
$$\gamma_{112} = \langle \cos(1\phi_\alpha + 1\phi_\beta - 2\Psi_2) \rangle$$

Sig + bkg.



$$\gamma_{123} = \langle \cos(1\phi_\alpha + 2\phi_\beta - 3\Psi_3) \rangle$$

Bkg. only



$$\gamma_{132} = \langle \cos(1\phi_\alpha + 3\phi_\beta - 2\Psi_2) \rangle$$

Bkg. dominant

P. Tribedy for STAR, WWND 2020

Also not consistent with signal-only explanation.  
Needs further investigations.

# Understanding the Background : Vary signal

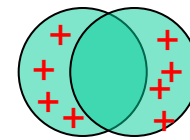
(Isobar collisions)

Signal driver (B)

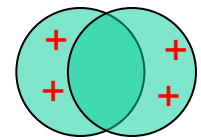
← **Vary**

Background driver ( $v_2$ )

← **Fixed**

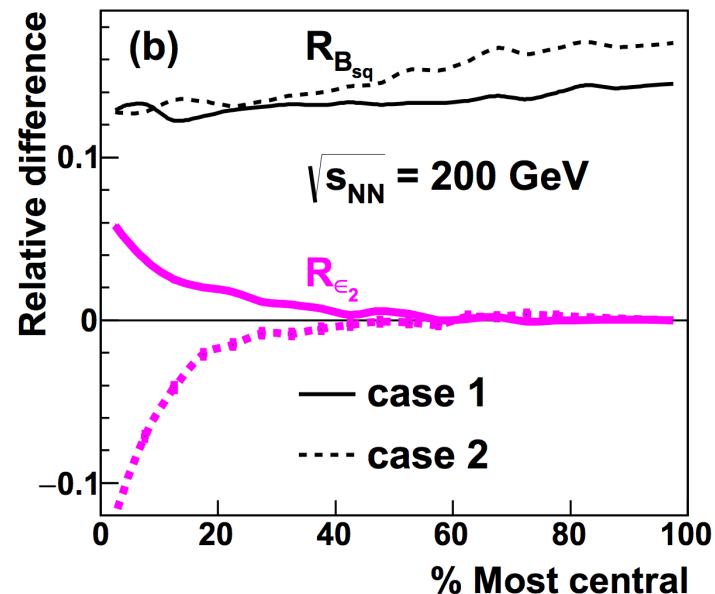
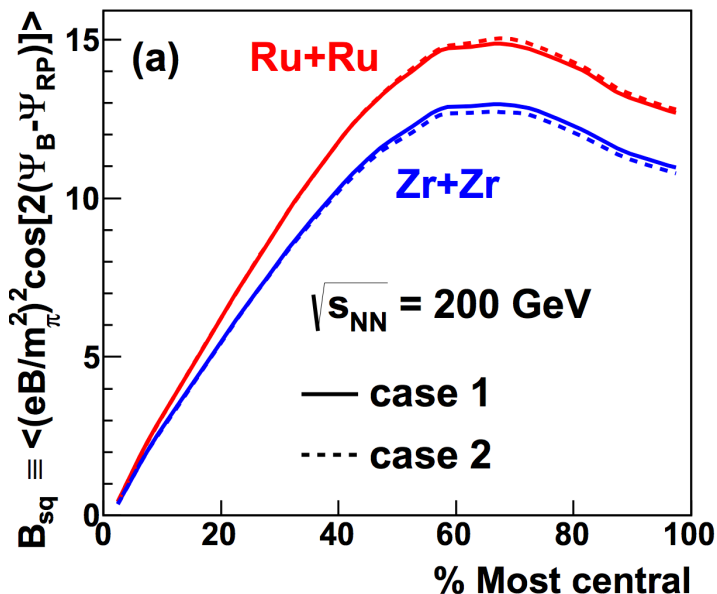


midcentral



$$\varepsilon_2 (\text{Ru}+\text{Ru}) \sim \varepsilon_2 (\text{Zr}+\text{Zr})$$

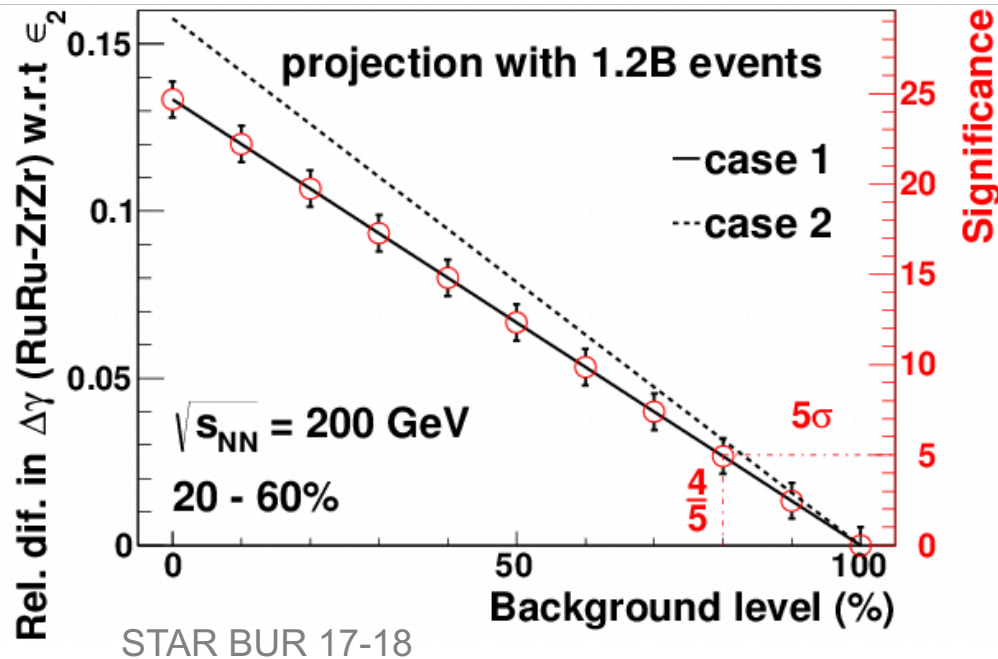
$$B (\text{Ru}+\text{Ru}) > B (\text{Zr}+\text{Zr})$$



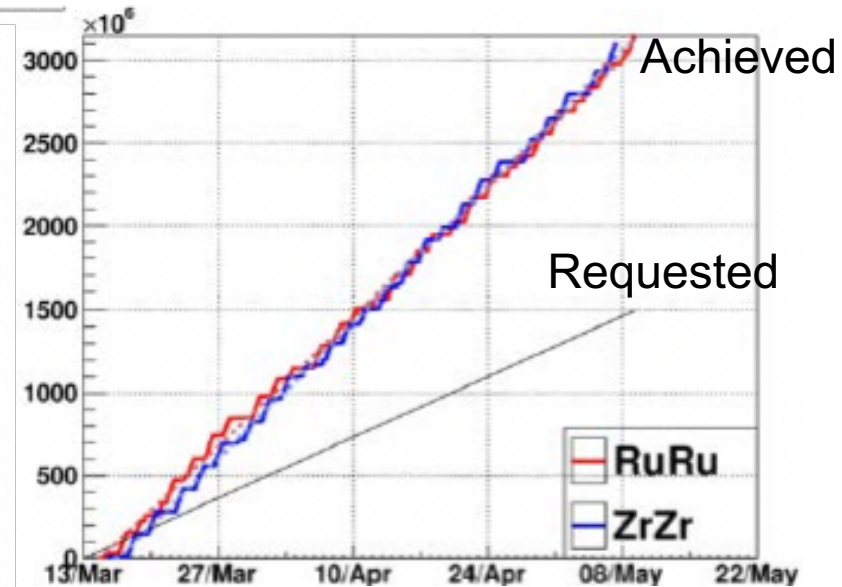
W. Deng, X-G. Huang G. Ma  
and G. Wang, PRC 94 041901 (2016)

Isobaric collisions : great potential for checking signal response.

# Isobar Data Taking at RHIC

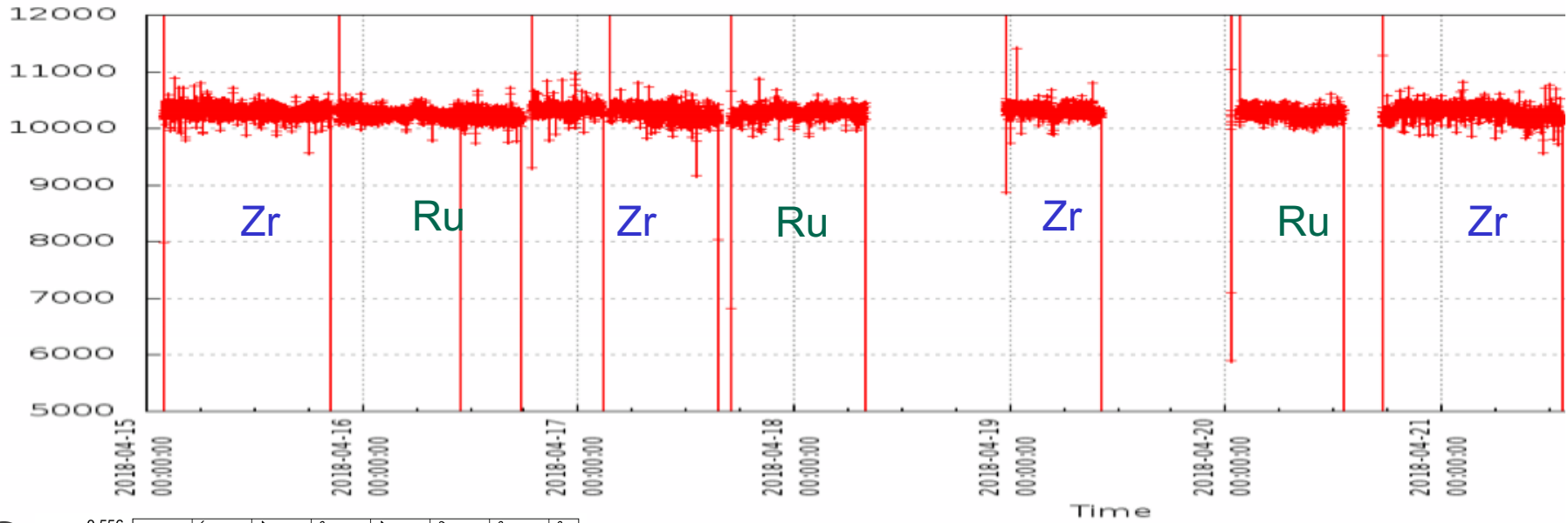


5 $\sigma$  difference in  $\Delta\gamma$  if bkg. is at ~80% level



Took more than requested.  
(3B each)

# Isobar Data Taking at RHIC

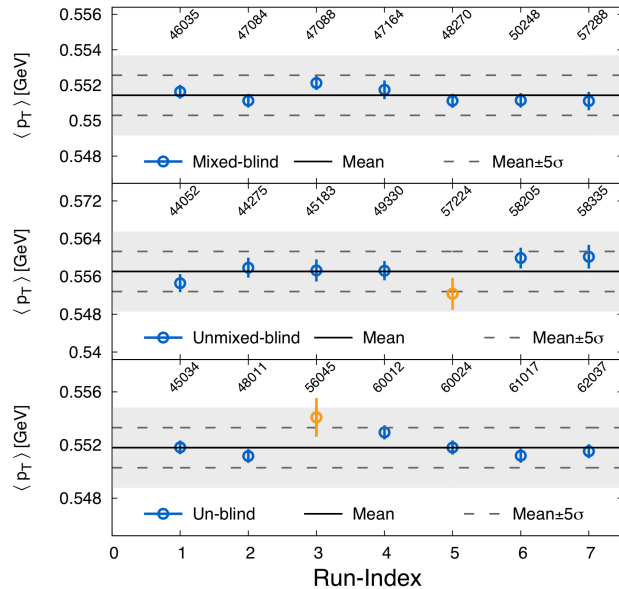


J-H Lee for STAR, RHIC/AGS Mtg 2018

Interleaved fills for isobar species to minimize systematics.

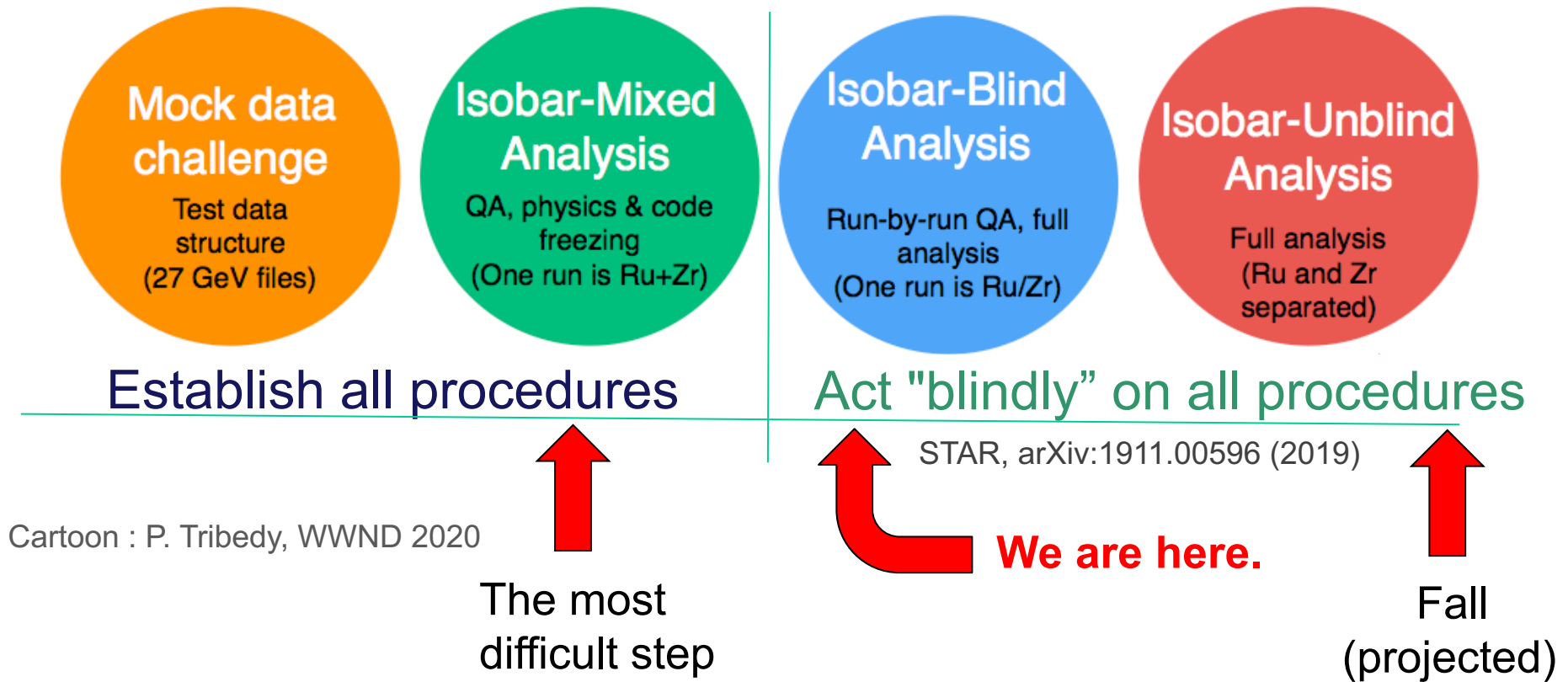
Good data set taken with very smooth running condition.

STAR, arXiv:1911.00596 (2019)





# Isobar Blind Analysis : Procedure

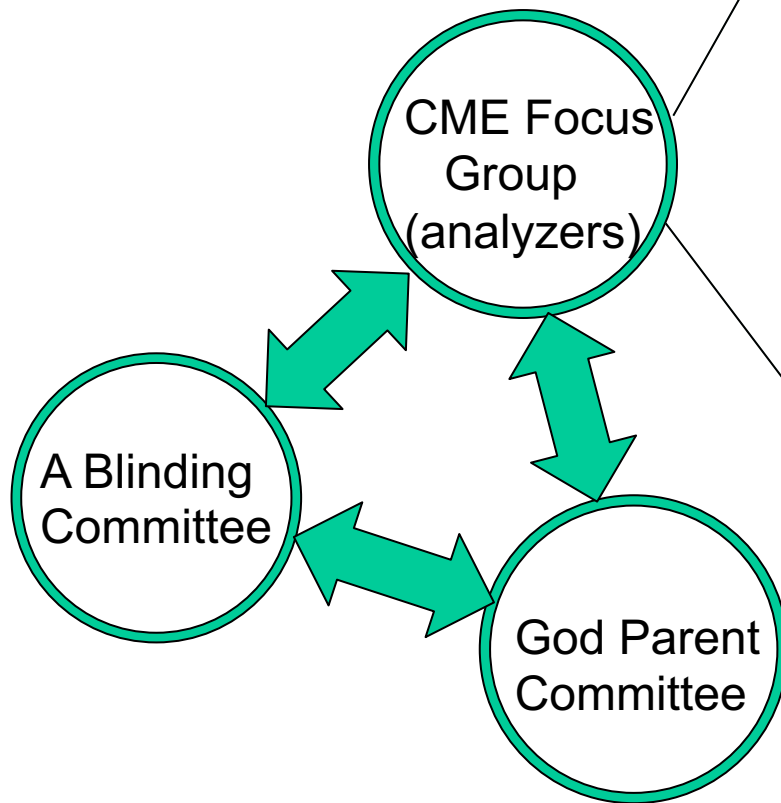


- Program Advisory Committee Recommendation:
  - The PAC strongly recommends that any STAR publication regarding CME observables should contain the result after unblinding and without any additional corrections applied after unblinding that are deemed necessary by STAR. If such additional corrections are needed, then a paper containing both the unblinded and post-unblinded results should be published for reference in papers reporting the isobar data.



# Isobar Analyses (5+1) in STAR

A large, collective effort



## Blind analyses (5 groups):

- $\Delta\gamma$ ,  $\Delta\delta$ , and  $\kappa$ .
- $\Delta\gamma$ ,  $\Delta\delta$ ,  $\Delta\gamma(\Delta\eta)$ .
- $\Delta\gamma$  in PP/SP,  $\Delta\gamma(M_{\text{inv}})$ .
- $\Delta\gamma$  in PP/SP.
- $R(\Delta S)$  Correlator.

## No-Blind analysis (1 group):

- Signed Balance Function.

## Challenges :

- Coordination and synchronization.  
(among groups, as well as between groups and committees).
- Unify procedures in common.
- Identify run-by-run abnormalities before hand without actual seeing them.  
.....

BNL, CCNU, Fudan, Huzhou,  
Purdue, SINAP, Stony Brook,  
Tsukuba, UCLA, UIC, Wayne State

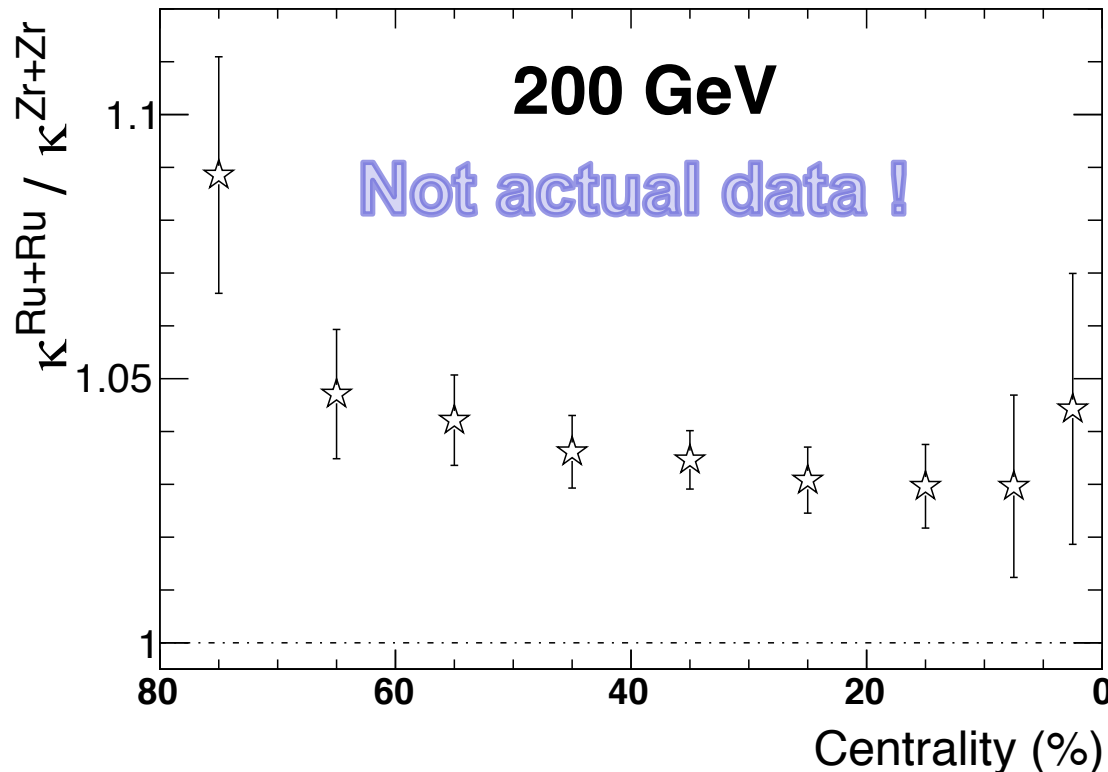
# Isobar Analysis (Grp 1) : $\Delta\gamma$ and $\Delta\delta$

$$\gamma = \langle \cos(\phi_\alpha + \phi_\beta - 2\Psi_{\text{RP}}) \rangle$$

$$\delta = \langle \cos(\phi_\alpha - \phi_\beta) \rangle$$

When there is  
pure background,

$$\Delta\gamma \sim \Delta\delta \cdot v_2,$$



motivates

$$\kappa \equiv \frac{\Delta\gamma}{\Delta\delta \cdot v_2}$$

**Case for CME :**

$$\kappa (\text{Ru} / \text{Zr}) > 1$$

Test isobar systems with pure  
background assumption.

S. Voloshin, Phys. Rev. C 70 057901

A. Bzdak, V. Koch, J. Liao Lect. Notes Phys. 871 503 (2013)

S. Shi, H. Zhang, D. Hou, and J. Liao arXiv : 1910.1401

+ many others



# Isobar Analysis (Grp 1) : $\Delta\gamma$ and $\Delta\delta$

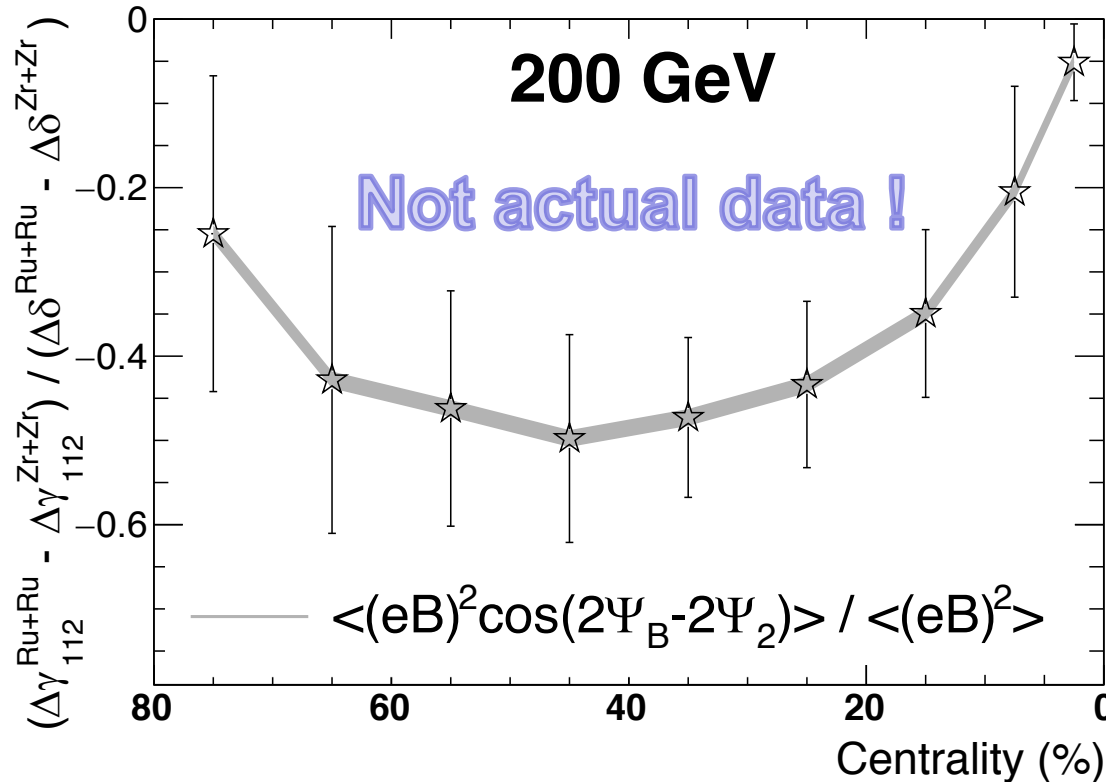
$$\gamma = \langle \cos(\phi_\alpha + \phi_\beta - 2\Psi_{RP}) \rangle$$

$$\delta = \langle \cos(\phi_\alpha - \phi_\beta) \rangle$$

When there is  
pure signal,

$$\Delta\gamma \sim 2\langle a_1^2 \cos(2\Psi_B - 2\Psi_{EP}) \rangle$$

$$\Delta\delta \sim -2\langle a_1^2 \rangle,$$



motivates  $\frac{\Delta\gamma\{\text{Ru-Zr}\}}{\Delta\delta\{\text{Ru-Zr}\}}$

$$\left. \frac{\Delta\gamma\{\text{Ru-Zr}\}}{\Delta\delta\{\text{Ru-Zr}\}} \right|_{v_2} \approx \langle \cos(2\Psi_B - 2\Psi_{EP}) \rangle^{[1]}$$

**Case for CME :**

$$\kappa (\text{Ru} / \text{Zr}) > 1$$

Check isobar systems with  
a relatively clean ratio

S. Voloshin, Phys. Rev. C 70 057901

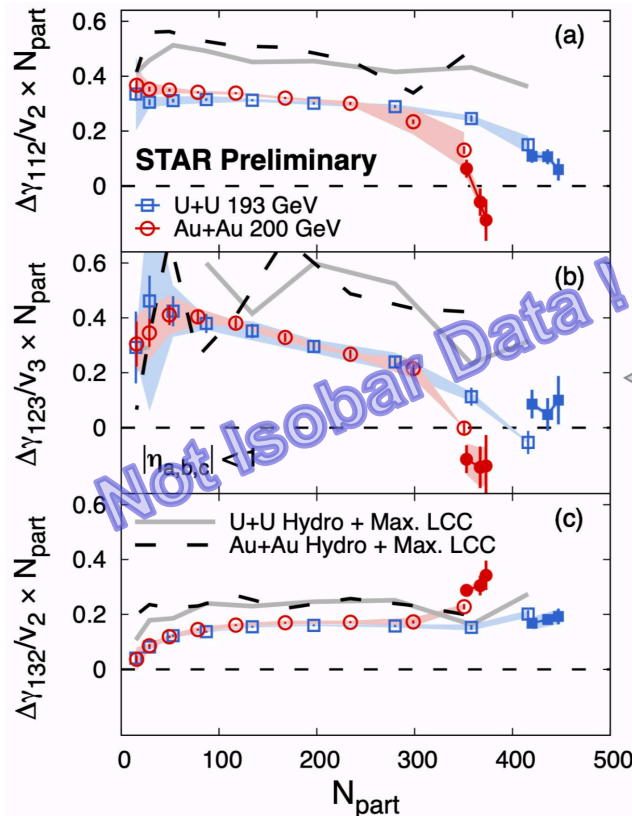
A. Bzdak, V. Koch, J. Liao Lect. Notes Phys. 871 503 (2013)

S. Shi, H. Zhang, D. Hou, and J. Liao arXiv : 1910.14010

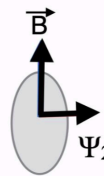
+ many others

<sup>[1]</sup>Assume same  $v_2$ , which can be attempted  
by event shape selection.

# Isobar Analysis (Grp 2) : $\Delta\gamma$ and $\Delta\delta$

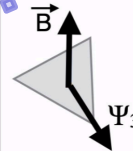


P. Tribedy for STAR, WWND 2020



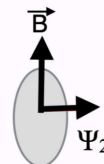
$$\gamma_{112} = \langle \cos(1\phi_\alpha + 1\phi_\beta - 2\Psi_2) \rangle$$

Sig + bkg.



$$\gamma_{123} = \langle \cos(1\phi_\alpha + 2\phi_\beta - 3\Psi_3) \rangle$$

Bkg. only



$$\gamma_{132} = \langle \cos(1\phi_\alpha + 3\phi_\beta - 2\Psi_2) \rangle$$

Bkg. dominant

In addition the  $\Delta\eta$  dependence of  $\Delta\gamma$  will be explored.

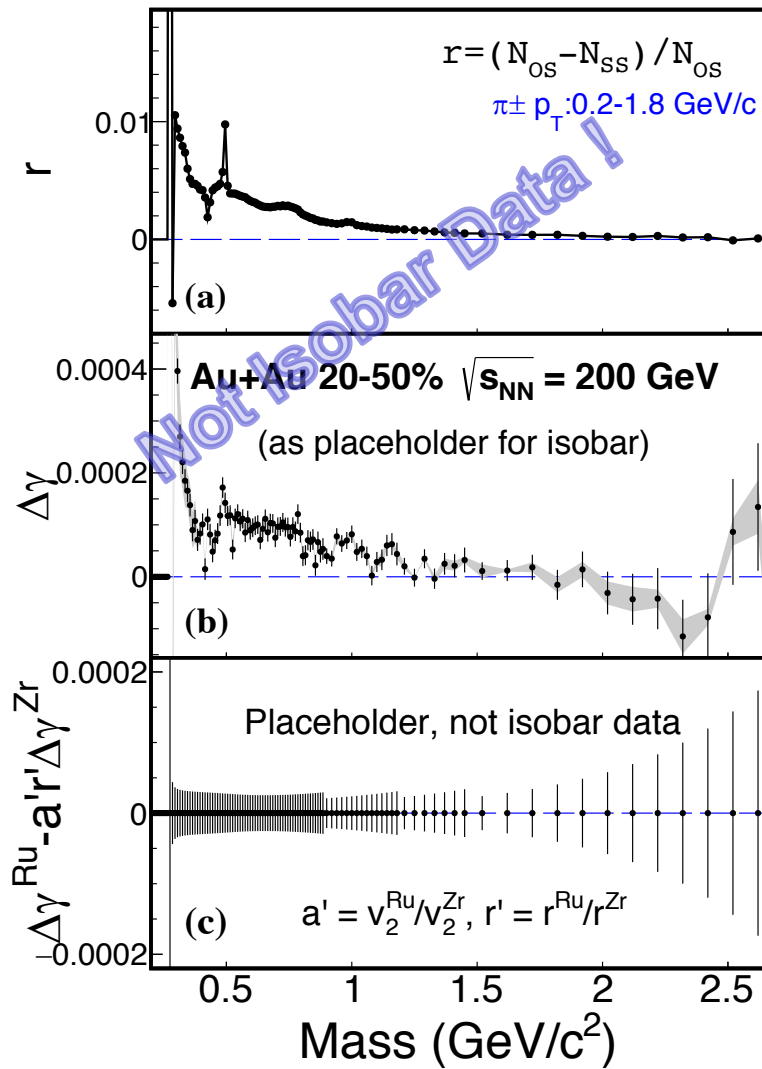
## Case for CME :

$$\Delta\gamma_{112}/V_2 (\text{Ru} / \text{Zr}) > 1$$

$$\Delta\gamma_{112}/V_2 (\text{Ru} / \text{Zr}) > \Delta\gamma_{123}/V_3 (\text{Ru} / \text{Zr})$$

$$\Delta\gamma_{112}/V_2 (\text{Ru} / \text{Zr}) > \Delta\delta (\text{Ru} / \text{Zr})$$

# Isobar Analysis (Grp 3) : $\Delta\gamma(M_{\text{inv}})$



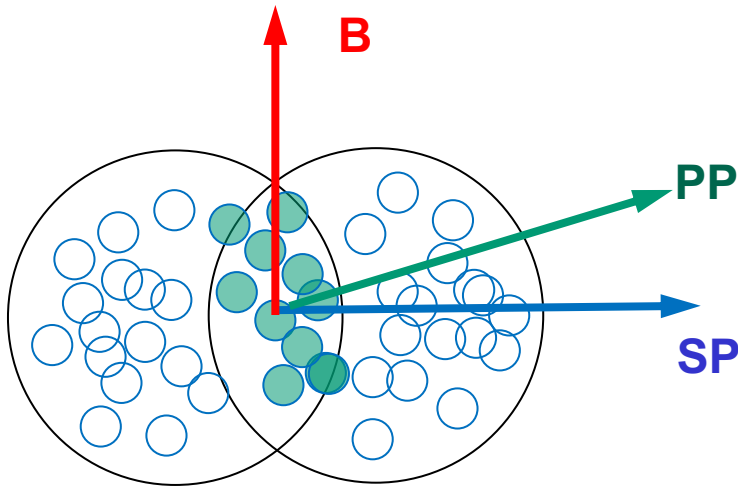
$$\Delta\gamma(m_{\text{inv}}) = r(m_{\text{inv}}) \langle \cos(\phi_\alpha + \phi_\beta - 2\phi_{\text{reso.}}) \rangle v_{2,\text{reso.}} + \Delta\gamma_{\text{CME}}$$

**Case for CME :**

$$\Delta\gamma^{\text{Ru}} - a'r'\Delta\gamma^{\text{Zr}} > 0$$

Different shape from inclusive  $\Delta\gamma$

# Isobar Analysis (Grp 3) : Participant Plane (PP) / Spectator Plane (SP)



$$f_{\text{CME}}^{\text{PP}} = \frac{\frac{\Delta\gamma\{\text{SP}\}}{\Delta\gamma\{\text{PP}\}}/a - 1}{1/a^2 - 1}$$

$$a = \langle \cos 2(\Psi_{\text{PP}} - \Psi_{\text{SP}}) \rangle \approx \frac{v_2^{\text{ZDC}}}{v_2^{\text{TPC}}} < 1$$

H-J. Xu et al., Chin. Phys. C 42 084103 (2018)

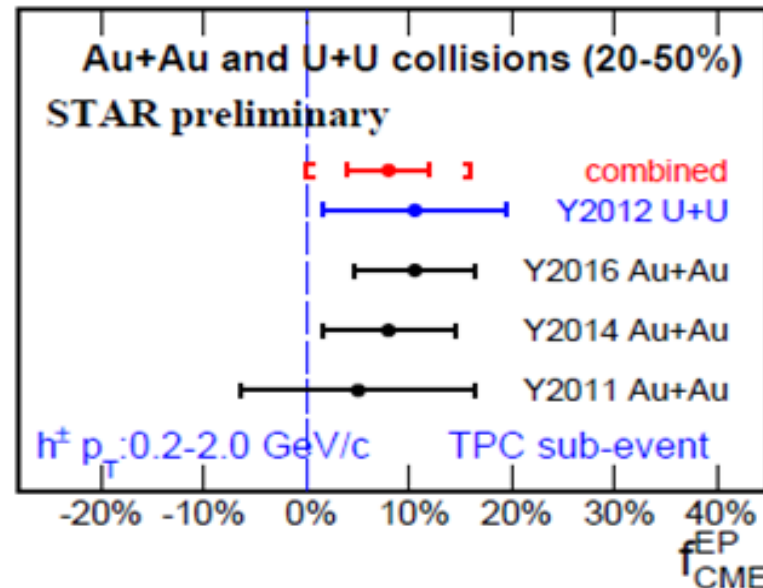
$$\Delta\gamma = \Delta\gamma^{\text{sig}} + \Delta\gamma^{\text{bg}}$$

PP : maximum background

$$\Delta\gamma^{\text{bg}}\{\text{PP}\} = \frac{\Delta\gamma^{\text{bg}}\{\text{SP}\}}{a}$$

SP : maximum signal

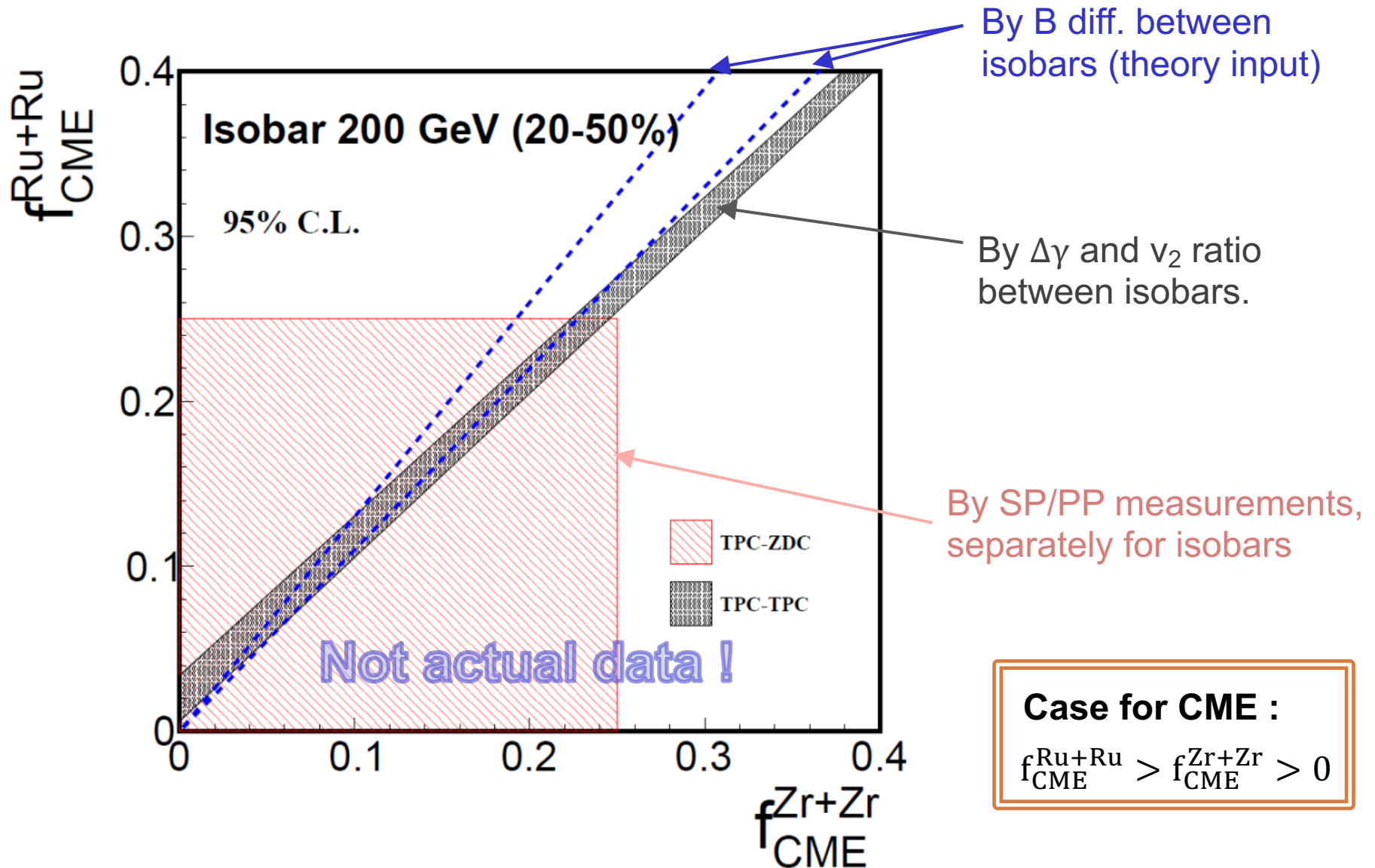
$$\Delta\gamma^{\text{sig}}\{\text{SP}\} = \frac{\Delta\gamma^{\text{sig}}\{\text{PP}\}}{a}$$



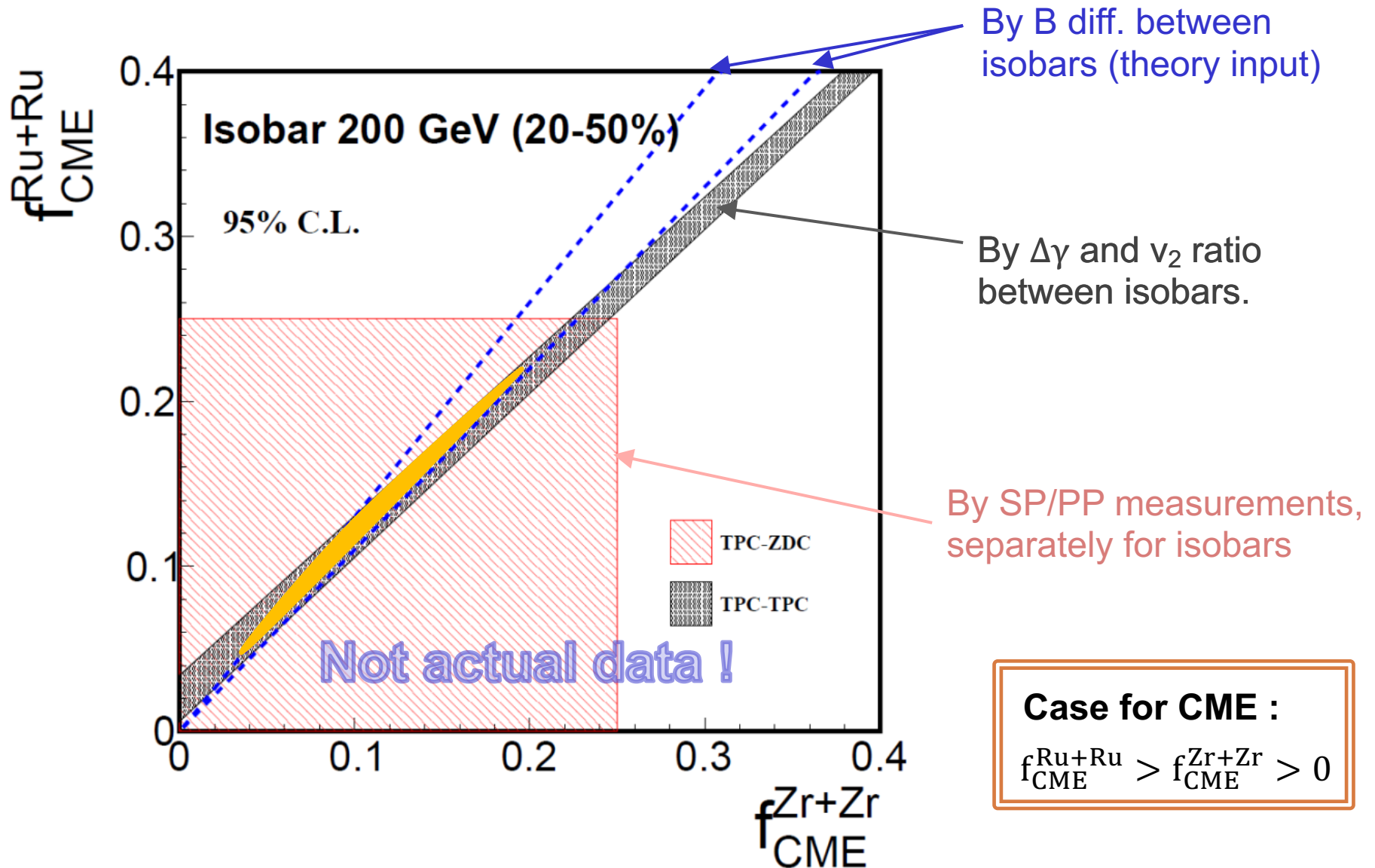
J. Zhao for STAR, QM 2019  
arXiv:2002.09410

Flow bkg is usually positively correlated with signal  $\Rightarrow$  Difficult to disentangle.  
This troublesome pattern can be broken by switching between PP and SP.

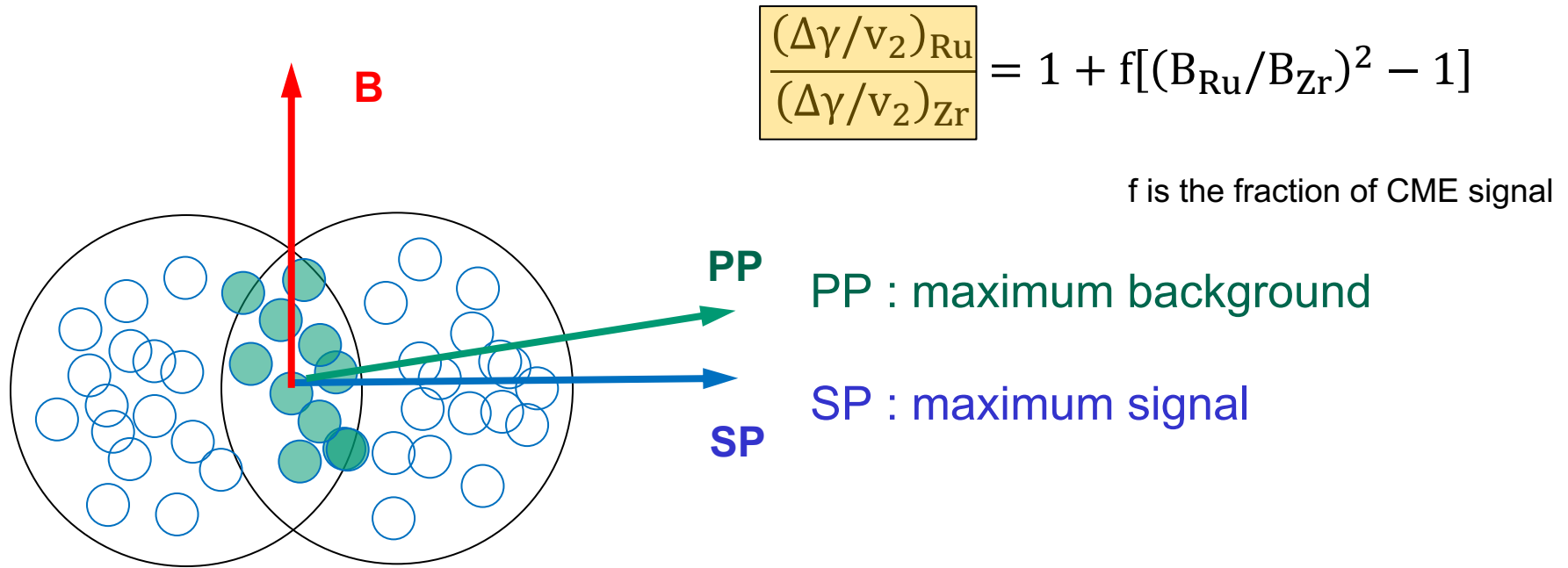
## Isobar Analysis (Grp 3) : Participant Plane (PP) / Spectator Plane (SP)



## Isobar Analysis (Grp 3) : Participant Plane (PP) / Spectator Plane (SP)



## Isobar Analysis (Grp 4) : Participant Plane (PP) / Spectator Plane (SP)



**Case for CME :  $\Delta\gamma/v_2 (\text{Ru} / \text{Zr}) > 1$**

Flow bkg is usually positively correlated with signal  $\Rightarrow$  Difficult to disentangle.  
This troublesome pattern can be broken by switching between PP and SP.

S. Voloshin, Phys. Rev. C 98 054911 (2018)

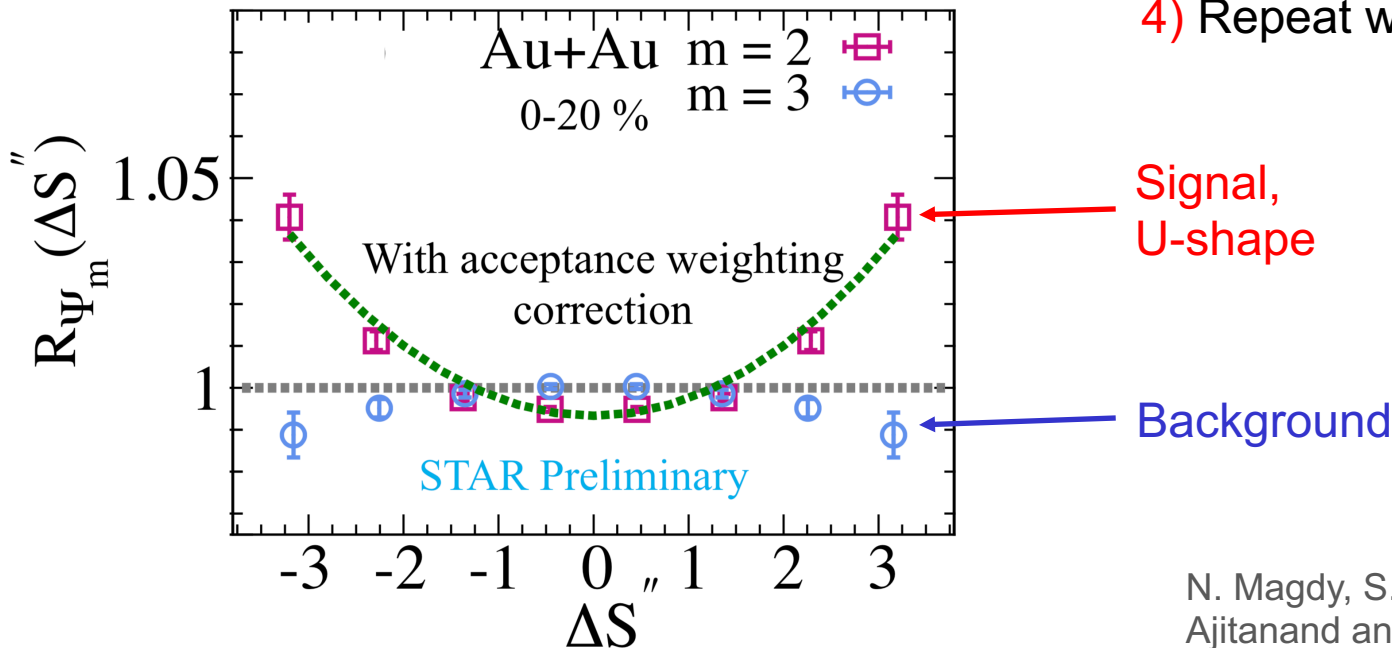
# Isobar Analysis (Grp 5) : R( $\Delta S$ ) Correlator

1) EbyE out-of-plane  $v_1$  difference between +/- charge  $\Delta S$ .

2) Removal of trivial contribution  $C(\Delta S) = \frac{N_{\text{real}}(\Delta S)}{N_{\text{shuffled}}(\Delta S)}$

3) Look for out-of-plane excess  $R(\Delta S) = \frac{C^\perp(\Delta S)}{C(\Delta S)}$

4) Repeat with  $\Psi_3$  EP for baseline.



N. Magdy, S. Shi, J. Liao, N. Ajitanand and R. Lacey Phys. Rev. C 97 061901 (2018)

N. Magdy for STAR, WWND 2019



# Isobar Analysis (Grp 5) : R( $\Delta S$ ) Correlator

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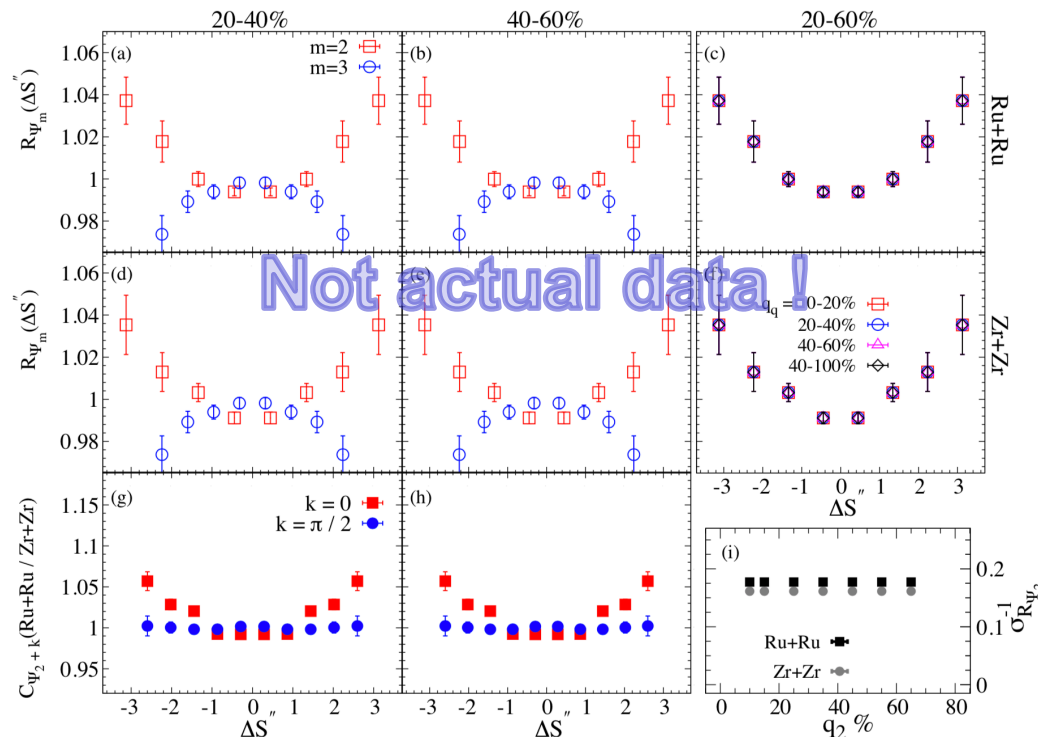
3) Look for out-of-plane excess  $R(\Delta S) = \frac{C^\perp(\Delta S)}{C(\Delta S)}$

4) Repeat with  $\Psi_3$  EP for baseline.

Intend to measure :

Sig. & bkg from each isobar.

Difference in sig. & bkg between isobars.



**Case for CME :**

**R (Ru / Zr) concave shape**

N. Magdy, S. Shi, J. Liao, N. Ajitanand and R. Lacey Phys. Rev. C 97 061901 (2018)

# Isobar Analysis (Grp 6) : Signed Balance Function

1) EbyE count +/- pair's momentum ordering in in- and out-of-plane direction.

2) EbyE count net-ordering  $\Delta B$ .

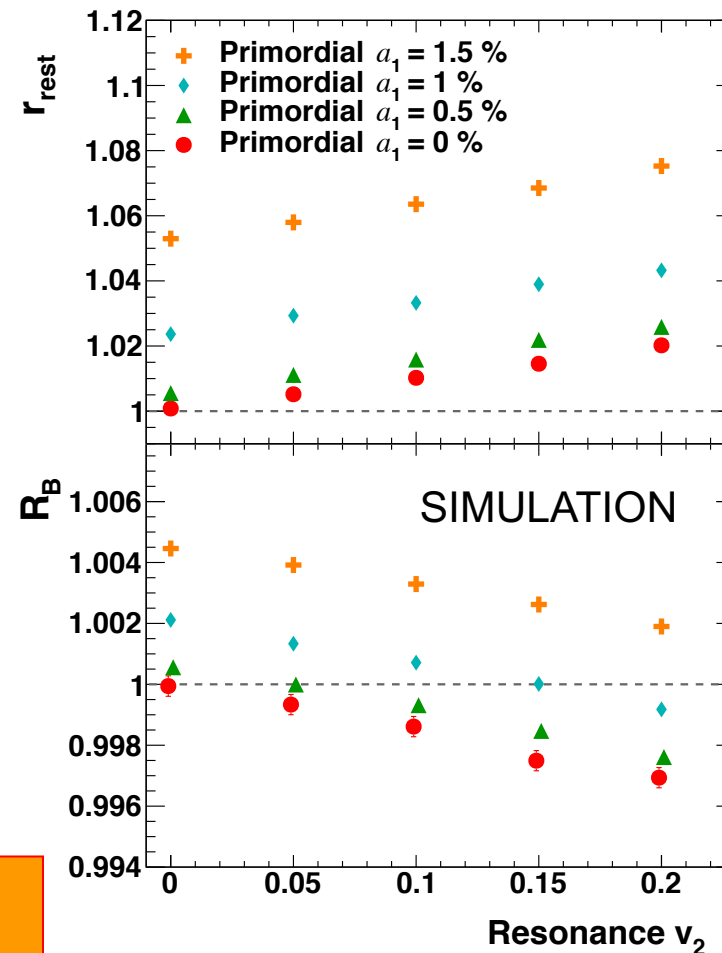
3) Excess  $\Delta B$  fluct. out-of-plane  $r = \frac{\sigma_{\Delta B}^{\perp}}{\sigma_{\Delta B}}$

4) Rest frame enhancement  $R_B = \frac{r_{\text{rest}}}{r_{\text{lab}}}$

Resonance daughters have no  $v_2$  in rest frame  $\Rightarrow r_{\text{rest}}$  tends to be closer to 1 than  $r_{\text{lab}}$ .



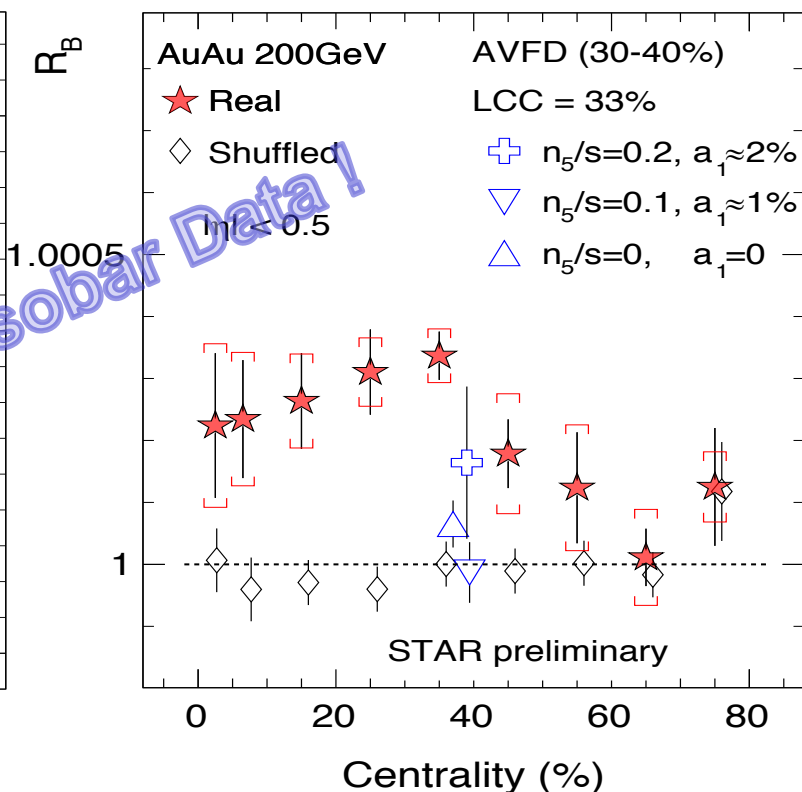
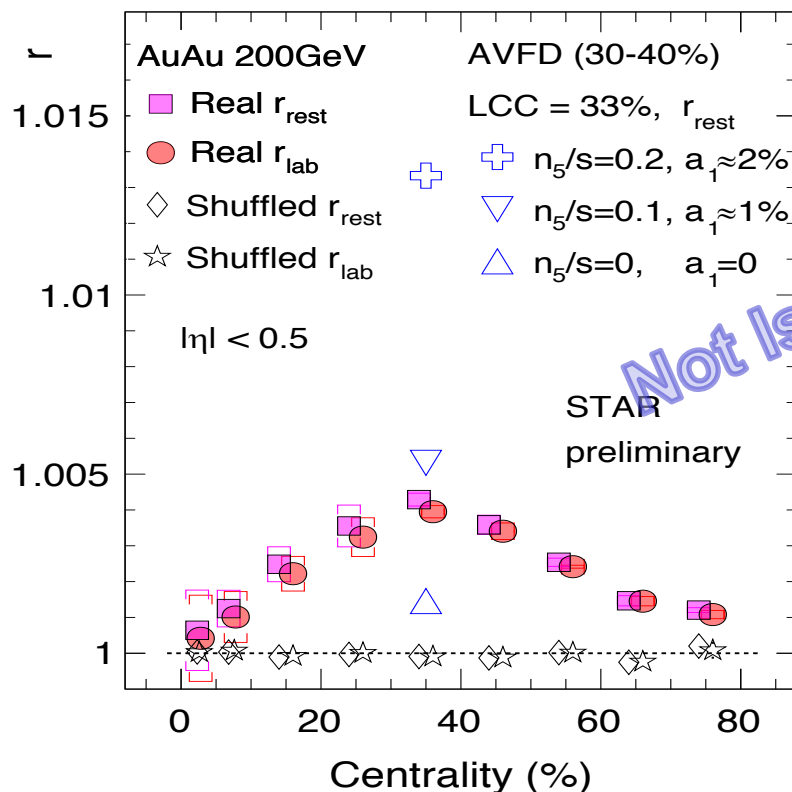
$r$  and  $R_B$  : Similar response to signal, opposite response to bkg.



A. Tang, Chin. Phys. C 44 No.5 054101 (2020)

Not participated in blind analyses

# Isobar Analysis (Grp 6) : Signed Balance Function



Y. Lin for STAR, QM 2019

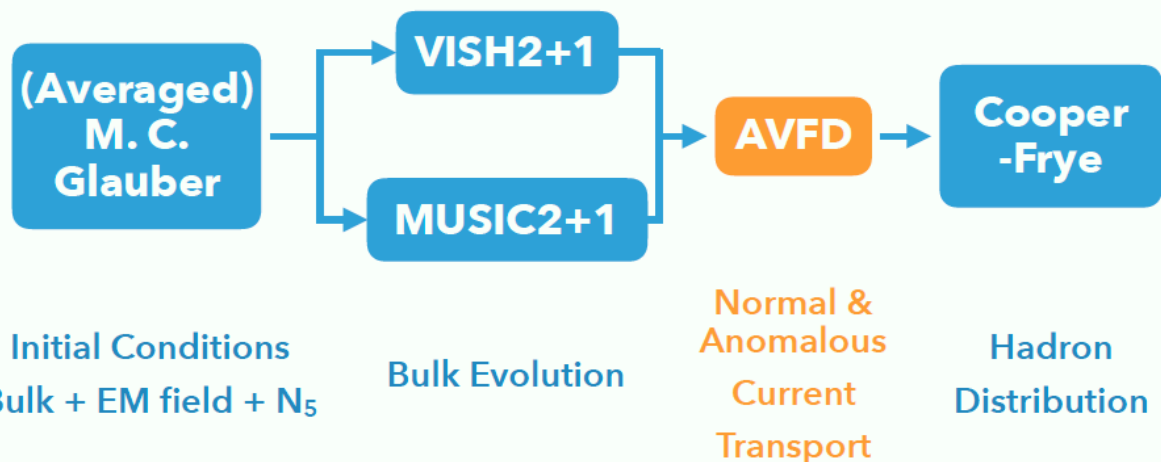
**Case for CME :**  $r > 1$  ,  $R_B > 1$   
 $r (Ru / Zr) > 1$   
 $R_B (Ru / Zr) > 1$

Not participated in blind analyses

# Observable's Response to Signal in AVFD

## AVFD Version Beta 1.0

The initial state fluctuations are fully accounted for by event-wise sampling for bulk entropy density and the fermion axial charge density ( $n_5/s$ ).



**Charge Asymmetry  
Correlation Measurement**

Background

Signal

**RuRu**

Background

Signal

**ZrZr**

	$n_5/s=0$	$n_5/s=0.10$	$n_5/s=0.2$
Ru	$a_1 = 0$	$a_1 \approx 0.75\%$	$a_1 \approx 1.49\%$
Zr	$a_1 = 0$	$a_1 \approx 0.69\%$	$a_1 \approx 1.38\%$

Three cases of AVFD evts are generated for each isobar.

**1<sup>st</sup> generation** (smooth initial condition, no bkg.):

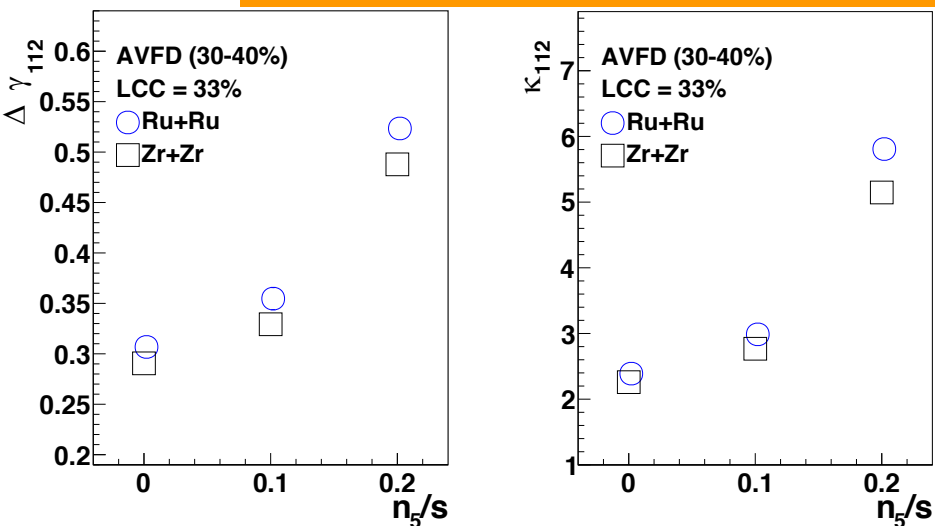
S. Shi, Y. Jiang, E. Lilleskov and J. Liao. Annals of Physics 394 50-72 (2018)

Y. Jiang, S. Shi, Y. Yin and J. Liao, Chin. Phys. C 42, No.1, 011001 (2018)

**2<sup>nd</sup> generation** (ebye initial cond., LCC and hadron cascade. Used in this work):

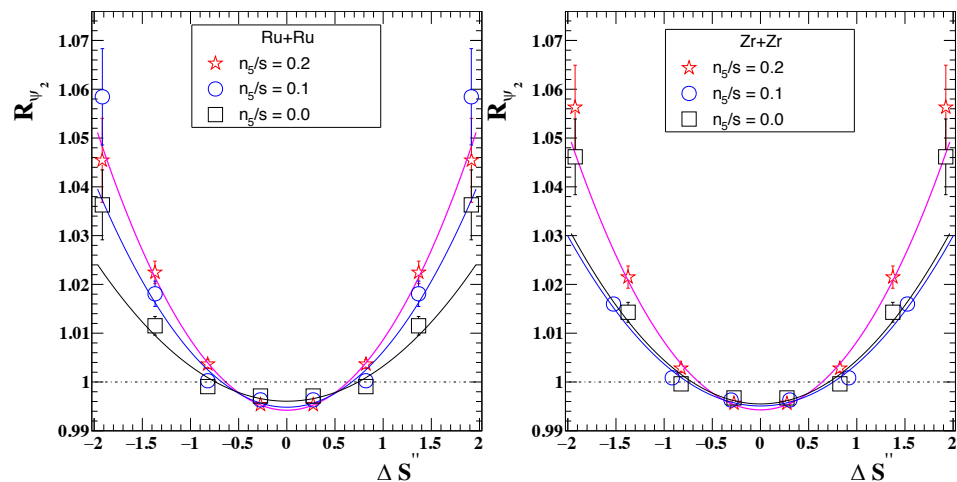
S. Shi, H. Zhang, D. Hou, and J. Liao. arXiv:1910.14010

# Observable's Response to Signal in AVFD



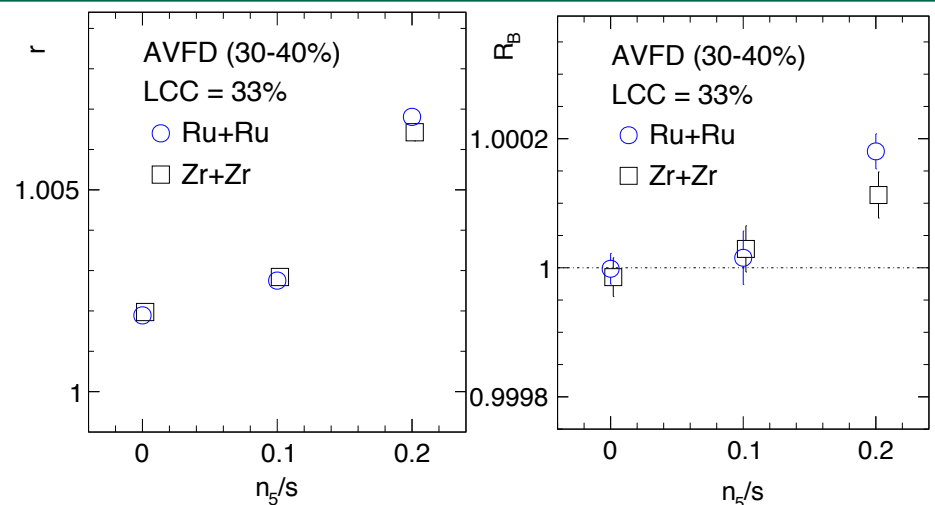
$\gamma$ -correlator

20 M evts



$R(\Delta S)$ -correlator

20 M evts



Signed Balance Function

50 M evts

Frozen codes are being tested with AVFD events (still accumulating/analyzing more evts).

**Working in progress**, independent of blind analyses.

- Connection between observables can be made with realistic CME model.
- Serve as guidance when comparing results across different methods.

# Summary

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The CME sub-field keeps evolving with new methods and innovations. Although a conclusion on the existence of CME cannot be made for now, a lot of progress have been made in terms of understanding the background and its interplay with signal.

STAR's blind analysis of isobar data has passed a critical step, expect relatively smooth sailing ahead.

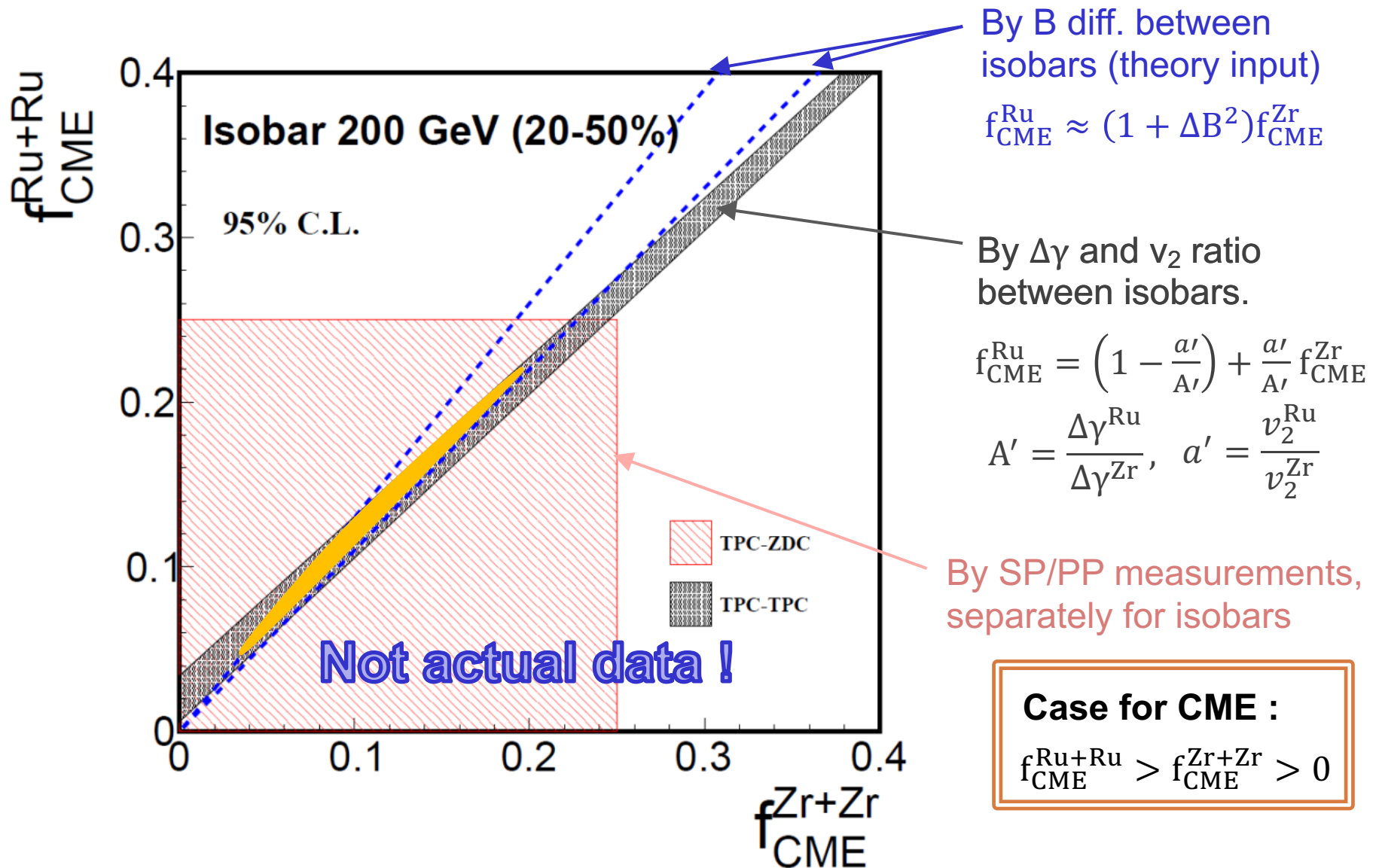
We are getting close, stay tuned !

# Backup Slides

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# Isobar Analysis (Grp 3) : Participant Plane (PP) / Spectator Plane (SP)





# Isobar Analysis (Grp +1) : Signed Balance Function

$$B_p(S) = \frac{N_{+-}(S) - N_{++}(S)}{N_+}$$

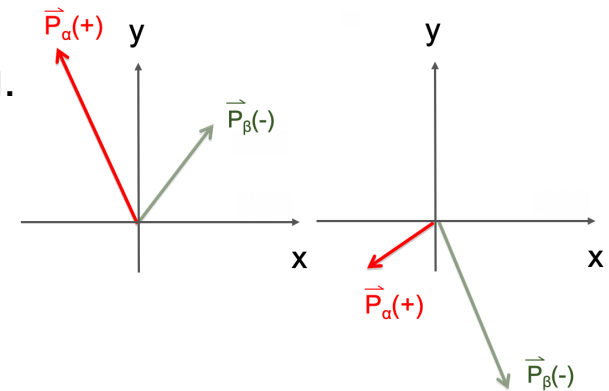
$$B_N(S) = \frac{N_{-+}(S) - N_{--}(S)}{N_-}$$

$$\delta B_p(\pm 1) = B_p(\pm 1) - B_N(\pm 1)$$

$$\Delta B = \delta B(+1) - \delta B(-1)$$

← 1) Pair's mtm ordering.  
( $S = \pm 1$ )

← 2) Count net ordering.



Two examples of  $\alpha$  leading  $\beta$  in  $p_y$   
( $S = +1$ )

$$r = \frac{\sigma_{\Delta B}^{\perp}}{\sigma_{\Delta B}}$$

← 3) Excess  $\Delta B$  fluct. out-of-plane.

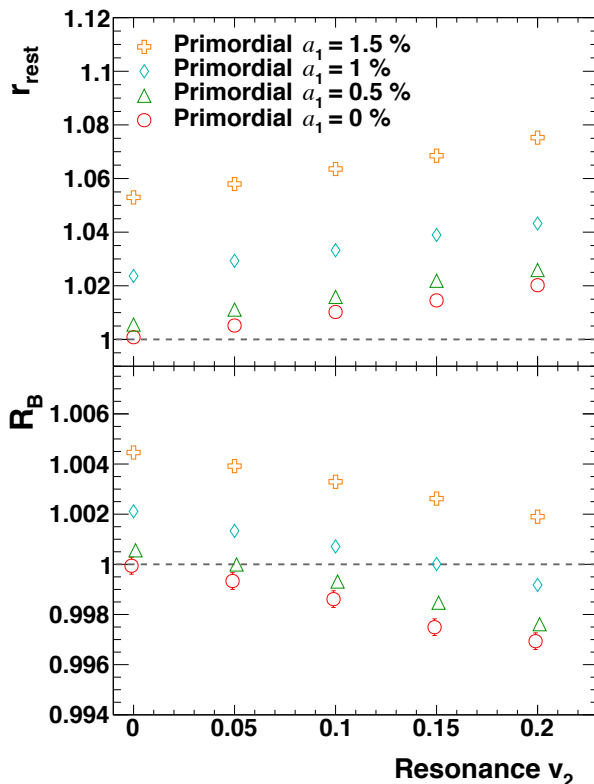
$$R_B = \frac{r_{\text{rest}}}{r_{\text{lab}}}$$

← 4) Rest frame enhancement.

← Resonance daughters have no  $v_2$  in rest frame  $\Rightarrow r_{\text{rest}}$  tends to closer to 1 than  $r_{\text{lab}}$ .

$r$  and  $R_B$  : Similar response to signal,  
opposite response to bkg.

Not participated in blind analyses



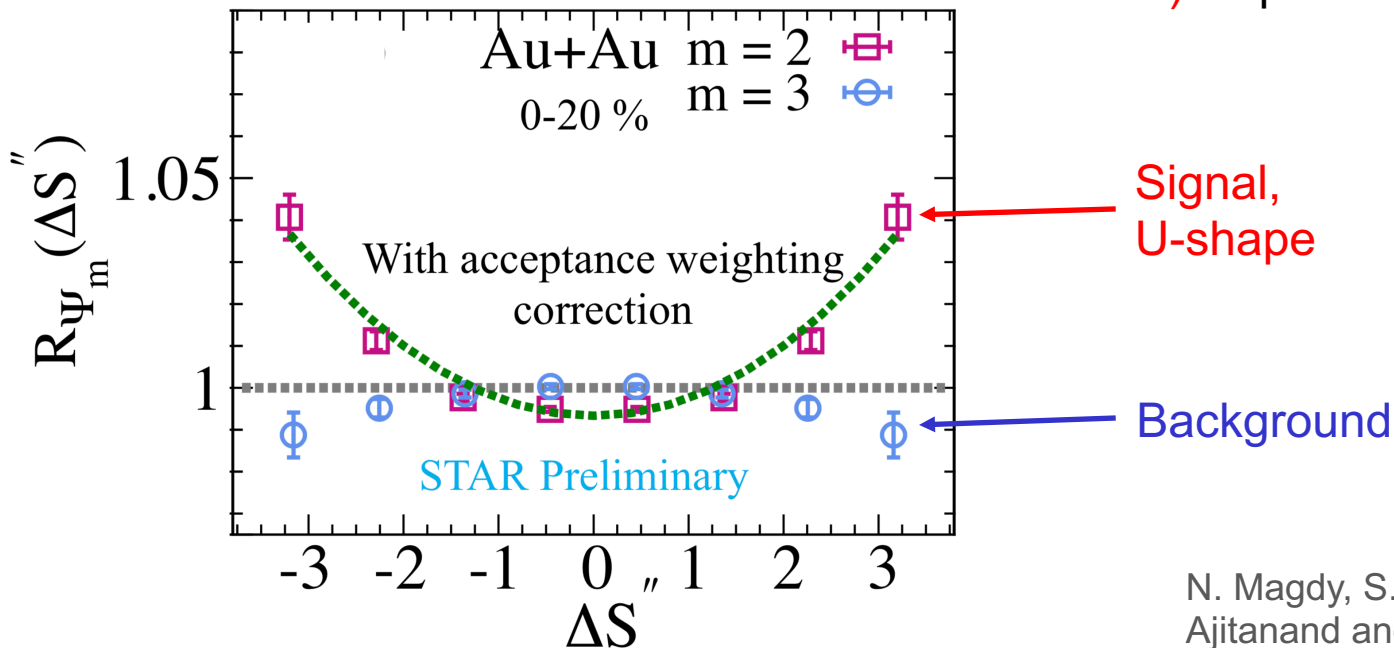
# Isobar Analysis (Grp 5) : R( $\Delta S$ ) Correlator

$$\Delta S = \langle \sin \Delta \phi \rangle_+ - \langle \sin \Delta \phi \rangle_- \quad \longleftarrow \quad 1) \text{ EbyE out-of-plane } v_1 \text{ diff.}$$

$$C(\Delta S) = \frac{N_{\text{real}}(\Delta S)}{N_{\text{shuffled}}(\Delta S)} \quad \longleftarrow \quad 2) \text{ Removal of trivial contribution.}$$

$$R(\Delta S) = \frac{C^\perp(\Delta S)}{C(\Delta S)} \quad \longleftarrow \quad 3) \text{ Look for out-of-plane excess.}$$

4) Repeat with  $\Psi_3$  EP for baseline.

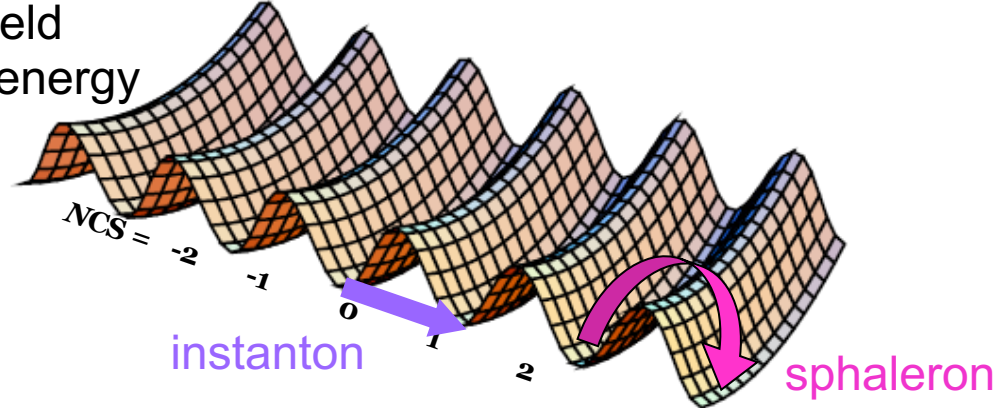


N. Magdy, S. Shi, J. Liao, N. Ajitanand and R. Lacey Phys. Rev. C 97 061901 (2018)

N. Magdy for STAR, WWND 2019

# QCD Vacuum Transition

Gluonic field  
potential energy



- Gluonic field energy is periodic in Chern–Simons number ( $N_{CS}$ ) direction
- Winding number:  $N_{CS}$  difference between initial and final states

$$Q_W = \frac{g^2}{32\pi^2} \int d^4x F_{\mu\nu}^a \tilde{F}_a^{\mu\nu}$$

- Axial current  $j_\mu^5$  : net handedness flow

$$\partial^\mu j_\mu^5 = 2 \sum_f m_f \langle \bar{\psi}_f i \gamma_5 \psi_f \rangle_A - \frac{N_f g^2}{16\pi^2} F_{\mu\nu}^a \tilde{F}_a^{\mu\nu}$$

- Nonzero topological charge generates chirality imbalance

$$N_L^f - N_R^f = 2Q_W, \quad Q_W \neq 0 \rightarrow \mu_A \neq 0$$

- QCD vacuum transition  $\rightarrow$  nonzero topological charge  $\rightarrow$  chirality imbalance