

1 CME search at STAR

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6 **Abstract.** The hot and dense medium produced in relativistic heavy-ion col-
7 lisions has been conjectured to be accompanied by an axial charge asymmetry
8 that may lead to a separation of electric charges in the direction of the extremely
9 strong magnetic field, also known as the Chiral Magnetic Effect (CME) [1–
10 3]. The measurement of azimuthal correlator ($\Delta\gamma$) with respect to the spectator
11 plane [5], gives us an opportunity to measure the possible CME fraction beyond
12 the flow background. Preliminary results using this approach with combined
13 Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV and U+U at $\sqrt{s_{NN}} = 193$ GeV show
14 f_{CME} at $(8\pm 4\pm 8)\%$. Meanwhile, the observability of CME has been conjectured
15 to be dependent on $\sqrt{s_{NN}}$ due to changes in the lifetime of the magnetic field,
16 the strengths of CME signal and non-CME background. At lower energies,
17 the Event Plane Detector (EPD) installed in the year 2018 provides a unique
18 capability for CME search. The background scenario test at Au+Au $\sqrt{s_{NN}} =$
19 27 GeV by using $\Delta\gamma$ with respect to TPC and the new installed EPD shows a
20 consistency with no-CME scenario in the current statistics. The method of the
21 ongoing isobar blind analysis, and the latest sensitivity check with the event-
22 by-event AVFD model on the different observables between Ru+Ru and Zr+Zr
23 are also briefly discussed.

24 1 Introduction

25 Symmetry and its breaking are parts of the most fundamental laws in the universe. In rel-
26 ativistic heavy-ion collisions, accompanied with the produced hot and dense medium, one
27 conjecture is there could be a symmetry breaking leading to the difference in the number of
28 right-handed and left-handed quarks [1, 2]. This imbalance leads to a separation of electric
29 charge in the direction of the strong magnetic field ($B \sim 10^{14}$ Tesla), produced by the protons
30 in the colliding heavy ions [3]. This phenomenon is known as the Chiral Magnetic Effect
31 (CME). STAR (Solenoidal Tracker at RHIC) experiment is currently the only operational ex-
32 periment at the Relativistic Heavy Ion Collider (RHIC), has been contributing to the CME
33 search over many years. RHIC has collided various ion species, at many different energies
34 over the past 20 years. In heavy-ion collisions, CME is expected to lead to charge separation
35 across the reaction plane determined by the impact parameter and the collision direction. The
36 reaction plane is correlated to the direction of the magnetic field [4]. To measure the charge
37 separation quantitatively, one uses the well known observable known as the gamma correlator
38 (γ).

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39 To reduce the charge-independent correlation backgrounds, like from global momentum
40 conservation, we usually use the $\Delta\gamma$, which measures the difference between the correlations
41 from the opposite charge and the same charge. There are still many background sources that
42 have also been identified to contribute to this observable, such as the local charge conserva-
43 tion and the resonance decays. We can write the $\Delta\gamma$ into 2 parts, one from the background
44 ($\Delta\gamma^{bkg}$), another from the CME signal($\Delta\gamma^{sig}$). In principle, the $\Delta\gamma$ can be measured with re-
45 spect to any plane [5, 6]: e.g. using the participate plane, which is estimated by the produced
46 particles (Ψ_{PP}); or the spectator plane, which is estimated by the spectators neutrons (Ψ_{SP}).

47 2 At top RHIC energies

48 At the top RHIC energies, we use the Au+Au collision at 200 GeV (form year 2011, 2014,
49 2016) and U+U collision at 193 GeV to search for the CME. As we mentioned before, if
50 we use the particles measured with Time Projection Chamber (TPC) to reconstruct the event
51 plane (Ψ_{EP}) as a proxy of the Ψ_{PP} , and use the Zero Degree Calorimeters (ZDC) to recon-
52 struct the spectator plane as a proxy of the Ψ_{SP} . The $\Delta\gamma$ with respect to Ψ_{PP}^{TPC} and Ψ_{SP}^{ZDC}
53 contain different fractions of the CME signal and background. If we assume the the back-
54 ground is proportional to the elliptic flow(v_2), we will have the following relations:

$$55 \Delta\gamma(\Psi_{TPC}) = \Delta\gamma^{bkg}(\Psi_{TPC}) + \Delta\gamma^{sig}(\Psi_{TPC}) \quad (1)$$

$$56 \Delta\gamma(\Psi_{ZDC}) = \Delta\gamma^{bkg}(\Psi_{ZDC}) + \Delta\gamma^{sig}(\Psi_{ZDC}) \quad (2)$$

$$57 \Delta\gamma^{bkg}(\Psi_{TPC})/\Delta\gamma^{bkg}(\Psi_{ZDC}) = v_2(\Psi_{TPC})/v_2(\Psi_{ZDC}) \quad (3)$$

$$58 \Delta\gamma^{sig}(\Psi_{TPC})/\Delta\gamma^{sig}(\Psi_{ZDC}) = v_2(\Psi_{ZDC})/v_2(\Psi_{TPC}) \quad (4)$$

59 With four known variables: $\Delta\gamma(\Psi_{TPC})$, $\Delta\gamma(\Psi_{ZDC})$, $v_2(\Psi_{TPC})$, $v_2(\Psi_{ZDC})$, and four unknwn ones
60 $\Delta\gamma^{bkg}(\Psi_{TPC})$, $\Delta\gamma^{bkg}(\Psi_{ZDC})$, $\Delta\gamma^{sig}(\Psi_{TPC})$, $\Delta\gamma^{sig}(\Psi_{ZDC})$, we can solve the above equations.
61 Then we can estimate f_{CME} which represents the fraction of CME signal in the observable
62 measured using TPC defined as:

$$63 f_{CME} = \Delta\gamma^{sig}(\Psi_{TPC})/\Delta\gamma(\Psi_{TPC}). \quad (5)$$

64 Figure 1 (left) the measurement of the f_{CME} for different run years in 20% – 50% centrality.
65 The Au+Au and U+U combined result shows that the faction is $8 \pm 4 \pm 8\%$. The systematic
66 uncertainties in this measurement are assessed by varying the track quality cuts and the η gap.

70 3 At Lower Energy

71 At lower energies, the Event Plane Detector (EPD) installed in the year 2018 provides a
72 unique capability for CME search. EPD covers a pseudorapidity range of 2.1 to 5.1 on either
73 side of the TPC. For the $\Delta\gamma$ observable, we choose particles of interest carrying charges
74 from the TPC, and measure the event plane from the EPD. At 27 GeV, the beam rapidity
75 ($Y_{beam} = 3.4$) falls at the center of EPD acceptance. Due to the kinematics, the inner EPD
76 ($\eta > Y_{beam}$) detects a lot of protons, coming from spectators, stopped protons or beam
77 fragments and they all have large directed flow which determines first-order event plane
78 well. The outer EPD ($\eta < Y_{beam}$) detects mostly produced particles. Produced particles have
79 large elliptic flow so we have a very well defined second order event plane. So with EPD
80 and 27 GeV collisions we have a way to, at the same time, measure charge separation w.r.to
81 planes of large directed flow due to spectator protons and w.r.to elliptic flow plane due to
82 participants.

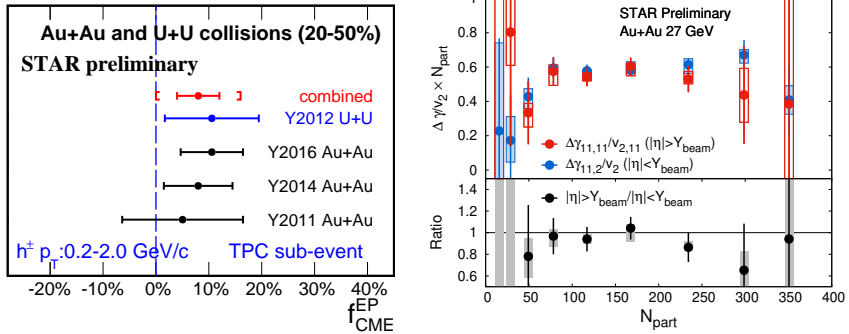


Figure 1. (Left) The fraction of the possible CME signal(f_{CME}) in $\Delta\gamma$ with respect to TPC and ZDC event planes for Au+Au collision at 200 GeV and U+U collision at 193 GeV, combined result shows $8 \pm 4 \pm 8\%$. (Right) $\Delta\gamma$ scaled by v_2 with respect to TPC-EPD-inner and TPC-EPD-outer planes to the no-CME scenario for Au+Au collision at 27 GeV.

In the background only scenario, if the background is proportional to v_2 , the ratio of $\Delta\gamma/v_2$ with respect to different planes (Ψ_A, Ψ_B, Ψ_C) will be the same, in other words we expect:

$$\Delta\gamma/v_2(\Psi_A) = \Delta\gamma/v_2(\Psi_B) = \Delta\gamma/v_2(\Psi_C) = \dots \quad (6)$$

Therefore, in our case we measured the $\Delta\gamma$ and v_2 with respect to the TPC-EPD-inner 1st order plane ($\eta > Y_{beam}$) and the $\Delta\gamma$ and v_2 with respect to the TPC-EPD-outer 2nd order plane ($\eta < Y_{beam}$). Figure 1(right) shows the $\Delta\gamma/v_2 \times N_{part}$ with respect to different planes. The bottom panel shows the double ratio. The results show, within the uncertainty, the ratio is consistent with unity. No significant deviation from Eq. 6 is observed.

4 Methods for blind analysis of isobar data

To better control the signal and the background in CME observables isobar program is introduced. To remove unconscious human induced biases we implement the isobar blind analysis. The $\Delta\gamma$ can be written into the signal ($\Delta\gamma^{sig}$), v_2 related background ($\Delta\gamma^{flow-bkg}$), and the non-flow driven background ($\Delta\gamma^{non-flow-bkg}$). In an ideal scenario if we can select two different systems that produce similar non-flow effects, similar flow backgrounds, then the residual difference in charge separation can be attributed to signal – isobar collisions were originally proposed with such expectations [7]. In practice with the two isobar species $^{96}_{44}\text{Ru} + ^{96}_{44}\text{Ru}$ and $^{96}_{40}\text{Zr} + ^{96}_{40}\text{Zr}$, according to model calculations one can achieve up to 4% difference in flow background [8–11], similar non-flow determined by the multiplicity difference, and about B-field square difference of the order of 10 – 18% that is relevant for a signal difference. If one can keep the systematics under control, predictions show, with two billion events collected for each species, we can achieve 5σ significance given the CME fraction in observables is at 14% level [8].

The isobar collisions are proposed to perform an experimental test of CME with the best possible control of signal and background as compared to all previous measurements. To minimize the run-to-run variation of the detector response due to the acceptance loss, efficiency changes, or the luminosity variation, we switch the collisions of two species frequently during data taking. To minimize the unconscious bias by the analysts we: 1) use the 27 GeV data

112 which is collected in the same year to perform a closure test of our analysis, 2) mix the Ru
113 and Zr events to freeze the analysis methods, observables, codes, and the event selection cri-
114 teria, 3) perform run-by-run quality assurance using blind data where species information is
115 hidden and, 4) finally perform the full analysis after unblinding the species information. The
116 details of this analysis method can be found in Ref. [12]. One final cross check of the frozen
117 code and response of different observables to CME signals has been recently performed using
118 an event-by-event Anomalous-Viscous Fluid Dynamics (AVFD) simulation [13].

119 **5 Summary**

120 At the top RHIC energy, we measured the possible CME fraction beyond the flow background
121 by measuring the $\Delta\gamma$ observable with respect to TPC and ZDC planes. At Au+Au 200 GeV
122 and U+U 193 GeV collisions, the combined results indicate the CME fraction is $(8 \pm 4 \pm 8)\%$
123 in data for 20% – 50% centrality.

124 At lower energy, a test of background scenario was performed using Au+Au 27 GeV data
125 by measuring $\Delta\gamma$ with respect to TPC and the new installed EPD systems. The results indicate
126 no significant deviation from background scenario with the uncertainty of our measurement.

127 Meanwhile, for the ongoing analysis of the isobar collisions, a blind analysis is imple-
128 mented to minimize unconscious bias. The sensitivity of different CME observables was
129 checked using the event-by-event AVFD model. The results from the isobar blind analysis
130 will soon be released by the STAR collaboration.

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