

# Strangeness Production in O+O Collisions at $\sqrt{s_{NN}} = 200$ GeV

Iris Ponce for the STAR Collaboration  
Yale University  
CPOD 2024



U.S. DEPARTMENT OF  
**ENERGY**

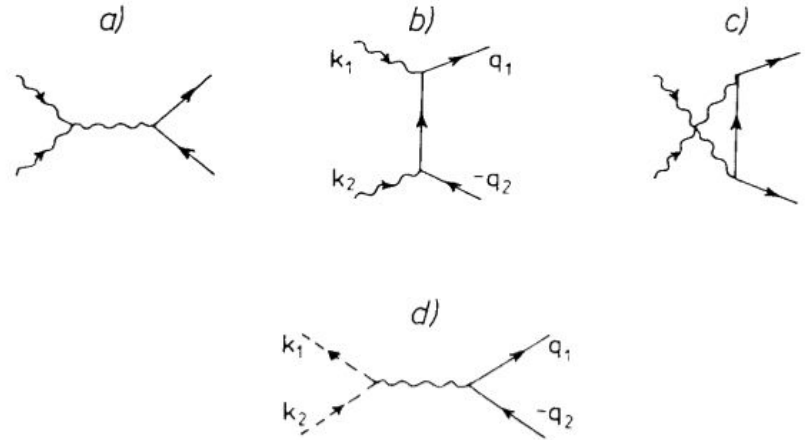
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# Strangeness Enhancement and the QGP

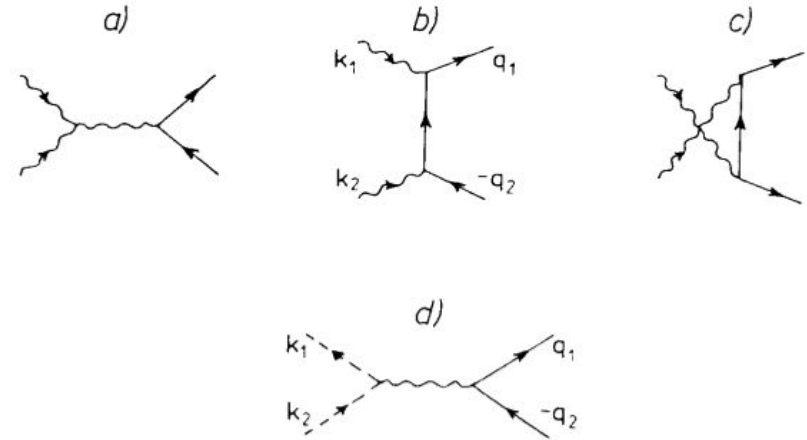
- Strangeness enhancement was one of the first observables predicted as a signature of the QGP.



P. Koch, B. Müller and J. Rafelski, "Strangeness in Relativistic Heavy Ion Collisions," Phys. Rep. 142, 167 (1986).

# Strangeness Enhancement and the QGP

- Strangeness enhancement was one of the first observables predicted as a signature of the QGP.
- The thermal production of  $s\bar{s}$  pairs is favorable in the QGP since the  $s\bar{s}$  masses are close to the QGP transition temperature  $\sim 156$  MeV.
  - $2 \times M_s \sim 192$  MeV
  - There are abundant thermal gluons in the QGP medium.
- Multi-strange ( $\Xi^\pm, \Omega^\pm$ ) hadrons are more sensitive to the existence of QGP.

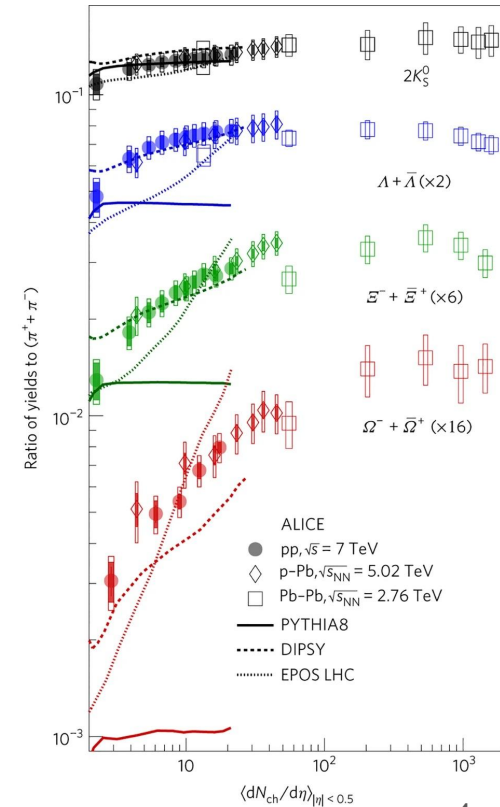


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# Motivation

- A smooth increase in the ratio of strange hadron production to the pion yield as a function of multiplicity has been found in various collision systems (p+p, p+A, A+A).

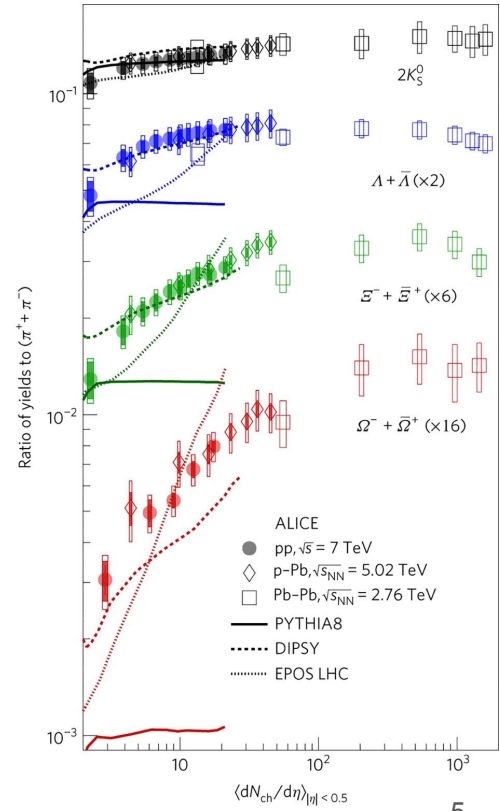
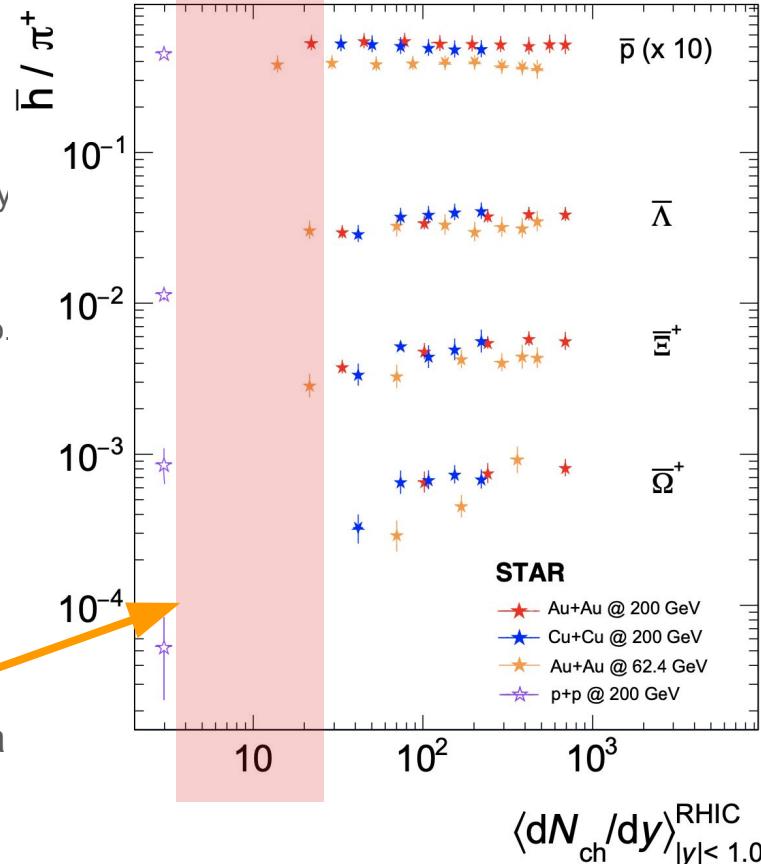




[STAR Collaboration. Phys. Rev. Lett. 98. 062301 \(2007\)](#)  
[STAR Collaboration. Phys. Rev. C 77, 044908 \(2008\)](#)  
[STAR Collaboration. Phys. Rev. C 83. 024901 \(2011\)](#)  
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  - STAR has reproduced this ratio.



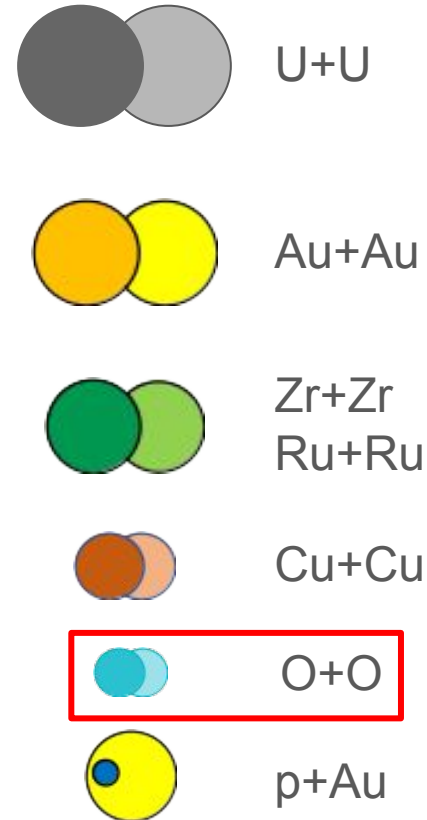
However, there is a notable data gap in the low multiplicity region



# Motivation

- A smooth increase in the ratio of strange hadron production to the pion yield as a function of multiplicity has been found in various collision systems (p+p, p+A, A+A).
  - STAR has reproduced this ratio.
- Oxygen is one of the smallest ions used at RHIC.
  - Fill in the hyperon to pion ratio in the low multiplicity gap
  - Allows a more straightforward geometry mapping with centrality than those asymmetric small system collisions like He+Au, or d+Au

Some of RHIC's collision systems



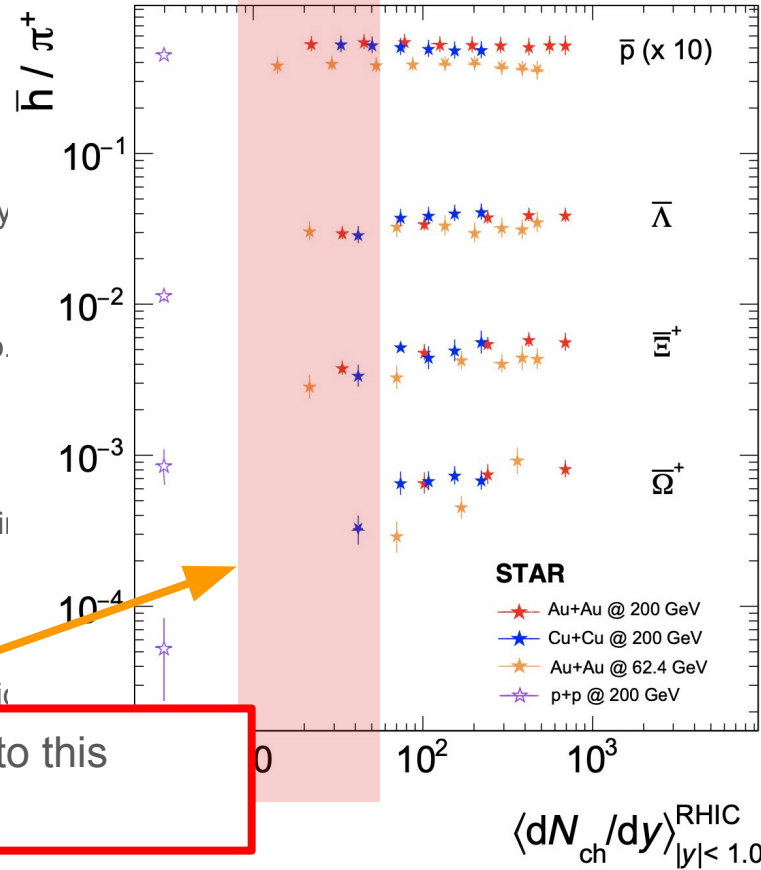
Increasing System Size



[STAR Collaboration. Phys. Rev. Lett. 98. 062301 \(2007\)](#)  
[STAR Collaboration. Phys. Rev. C 77. 044908 \(2008\)](#)  
[STAR Collaboration. Phys. Rev. C 83. 024901 \(2011\)](#)  
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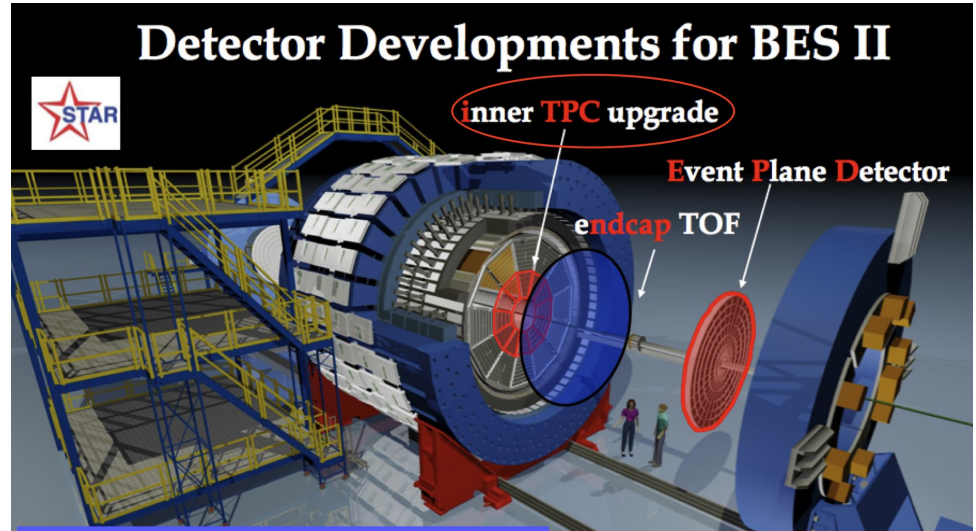
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O+O's multiplicity can extend to this unexplored region

# O+O Run Information at STAR

- The Solenoidal Tracker at RHIC (STAR) has been operating since 2000.
- From 2018 on, STAR had two detector upgrades: iTPC and eTOF



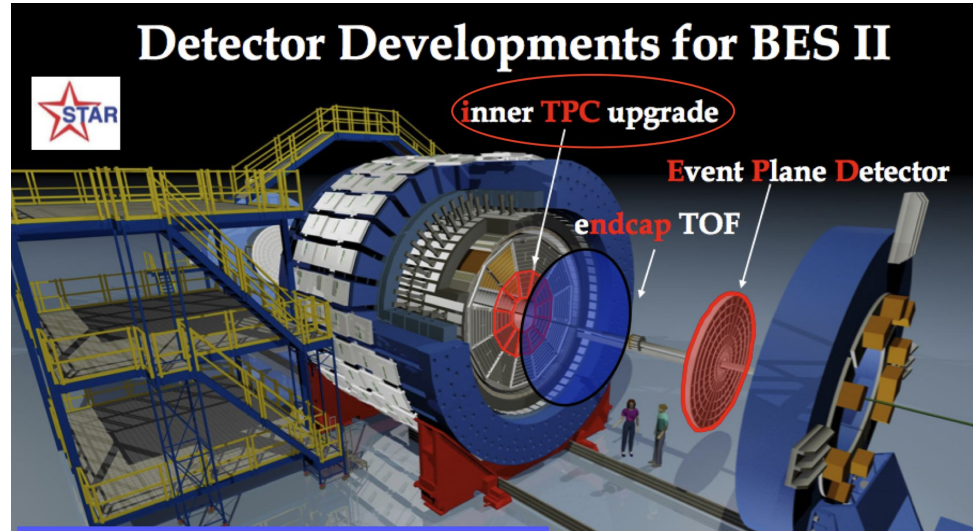
Picture: Alex & Maria Schmah

[Q. Xu. \(STAR Collaboration\). 8th Workshop on Hadron Physics \(2016\)](#)



# O+O Run Information at STAR

- The Solenoidal Tracker at RHIC (STAR) has been operating since 2000.
- From 2018 on, STAR had two detector upgrades: iTPC and eTOF
  - Improved coverage:
    - | $\eta$ | < 1.5 from | $\eta$ | < 1.0
    - Lower pT coverage 125 MeV => 60 MeV
- There are ~650M O+O minimum bias events total.
  - 1/4 of the O+O run was taken with the magnetic field reversed.
    - Testing calibration and TPC distortions



Picture: Alex & Maria Schmah

[Q. Xu. \(STAR Collaboration\). 8th Workshop on Hadron Physics \(2016\)](#)

# Flow Measurements in O+O collisions shown in QM2023

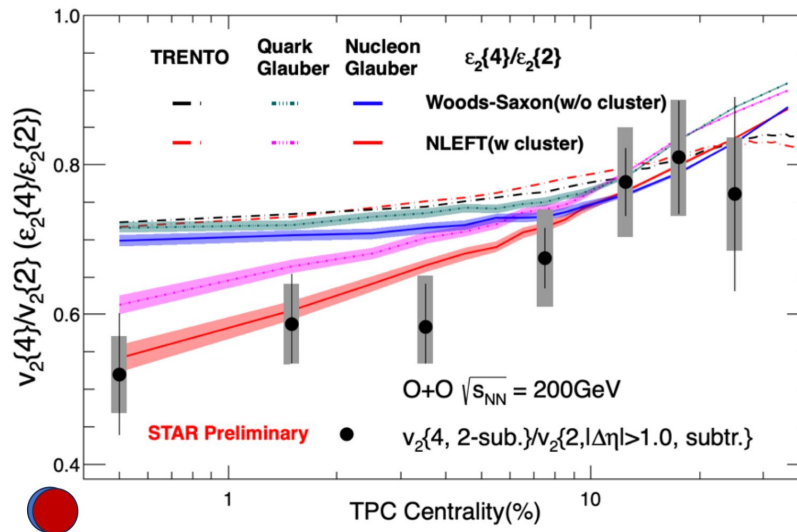
## Bulk Results:

Similar  $N_{\text{part}}$  to  ${}^3\text{He}\text{-Au}$

$$v_2(\text{O+O}) < v_2(\text{d+Au}) \approx v_2({}^3\text{He+Au})$$

$$v_3(\text{O+O}) \approx v_3(\text{d+Au}) \approx v_3({}^3\text{He+Au})$$

- $v_n/\varepsilon_n$  are similar between O+O and  ${}^3\text{He+Au}$ , within a quark Glauber model
- $v_2\{4\}/v_2\{2\}$  show clear decrease in ultra-central collisions, consistent with  $\varepsilon_2\{4\}/\varepsilon_2\{2\}$ , indicating enhanced fluctuations due to possible many-nucleon correlations.



Quark Glauber:  
PRC **94**, 024914 (2016)  
TRENTO:  
PRC.92.011901(2015)  
Calculated by Giuliano

[S. Huang \(STAR Collaboration\). QM2023](#)



# Particles To Be Reconstructed

I am interested in reconstructing particles with s-quarks, as listed below.

Particle	Strangeness	Mass (MeV)	Decay Mode	Branching Ratio
$\phi(1020)$	0	$1,019.461 \pm 0.020$	$K^+ K^-$	49.5 %
$K_s^0$	$\pm 1$	$497.611 \pm 0.013$	$\pi^+ \pi^-$	69.20 %
$\Lambda$	-1	$1,115.683 \pm 0.006$	$p \pi^-$	64.1 %
$\Xi^-$	-2	$1,321.71 \pm 0.07$	$\Lambda \pi^-$	99.887%
$\Omega^-$	-3	$1,672.45 \pm 0.29$	$\Lambda K^-$	67.8%

- This presentation will focus on  $\Lambda$ 's.
- The  $\Xi^-$ ,  $\Omega^-$ ,  $\phi$ , and  $K_s^0$  results will follow soon.

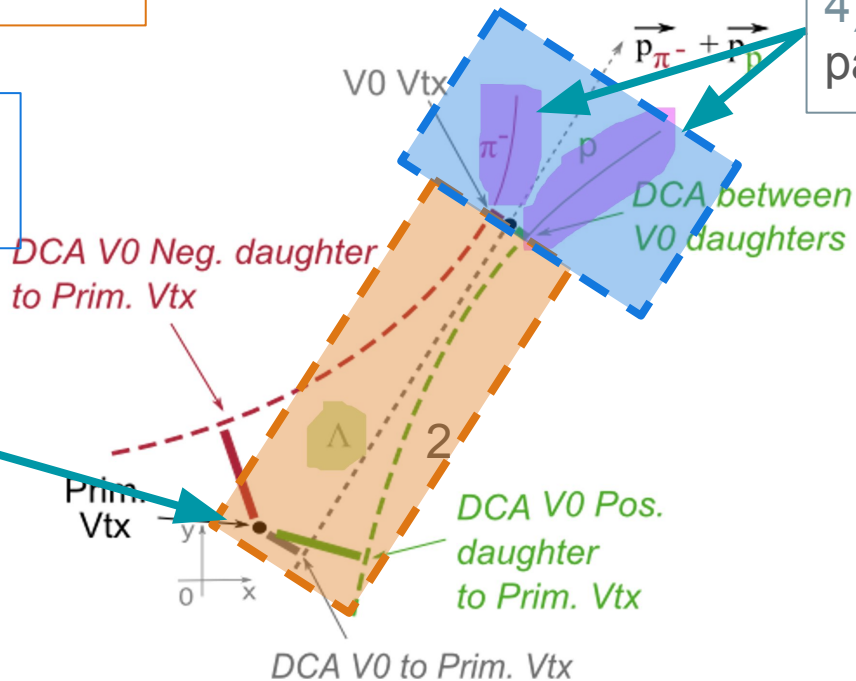
# Reconstruction Details and Topological Cuts

1) The distance to primary vertex

2) The distance between daughter particles

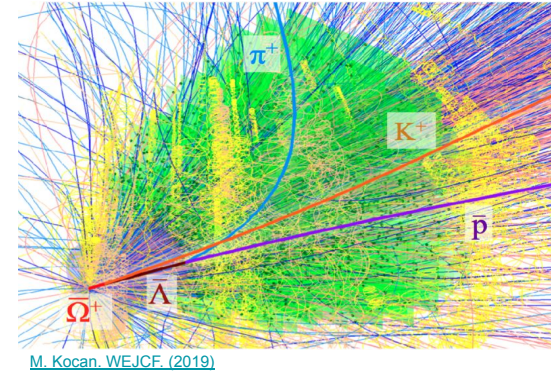
3) Vertex location

4)  $\chi^2$  cuts on particle fits



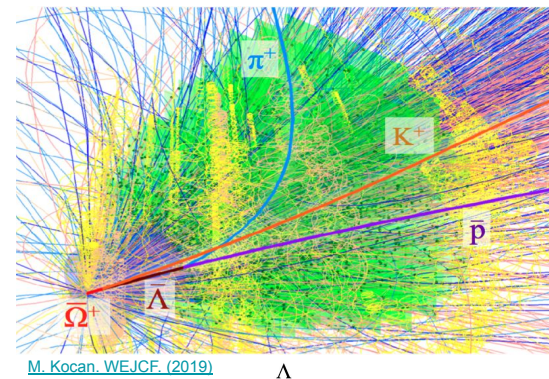
# Reconstructing Lambdas and Signal Extraction

- Using Kalman Filter Particle (KF Particle) reconstruction algorithm.
  - Standard reconstruction for decayed particles.
  - Initially developed for other heavy ion experiments but was adapted in 2018 for STAR.



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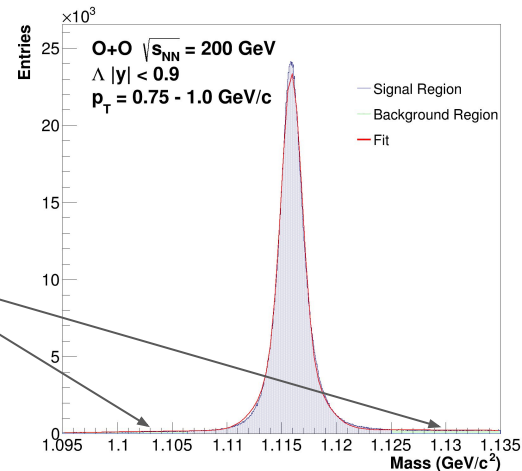


## For the $\Lambda$ Signal Extraction:

- The signal (without background subtraction) region is  $[\mu-3\sigma, \mu+3\sigma]$ , and the background region is  $[0$  to  $\mu-3\sigma$ ,  $\mu+3\sigma$  to  $X_{max}]$ .
- Fitting function: 2nd poly + double Gauss function.

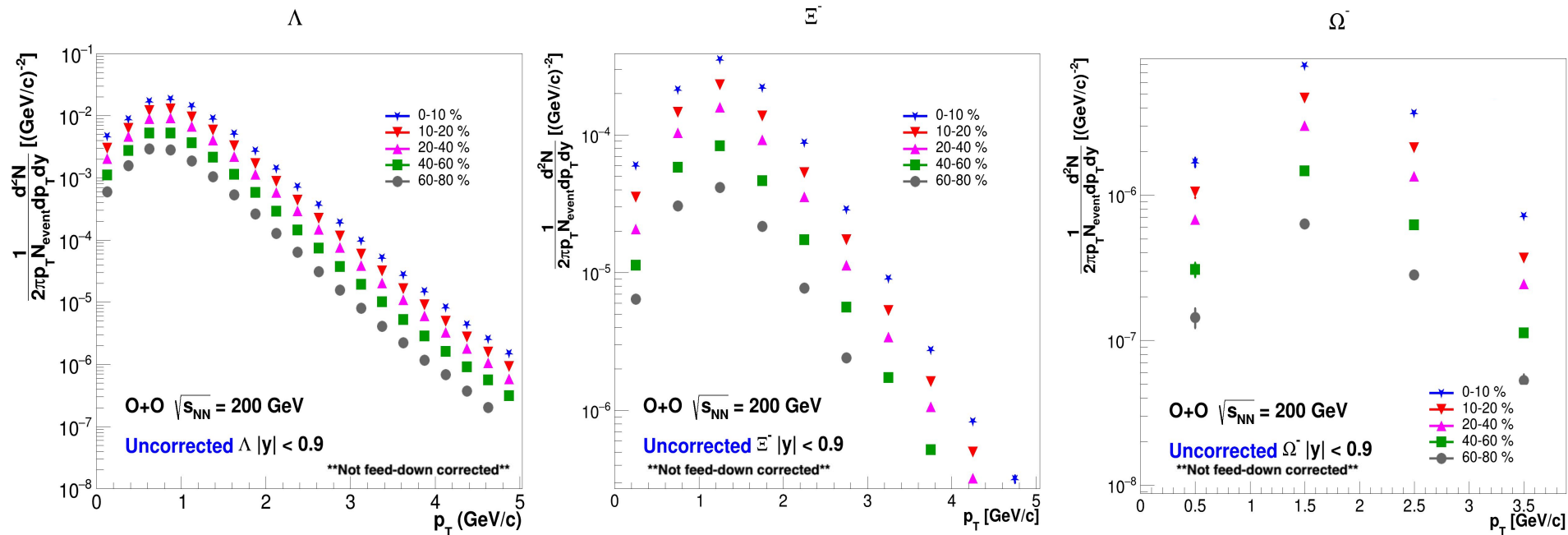
The blue region is the signal w.o background subtraction.

The green region is the background region.



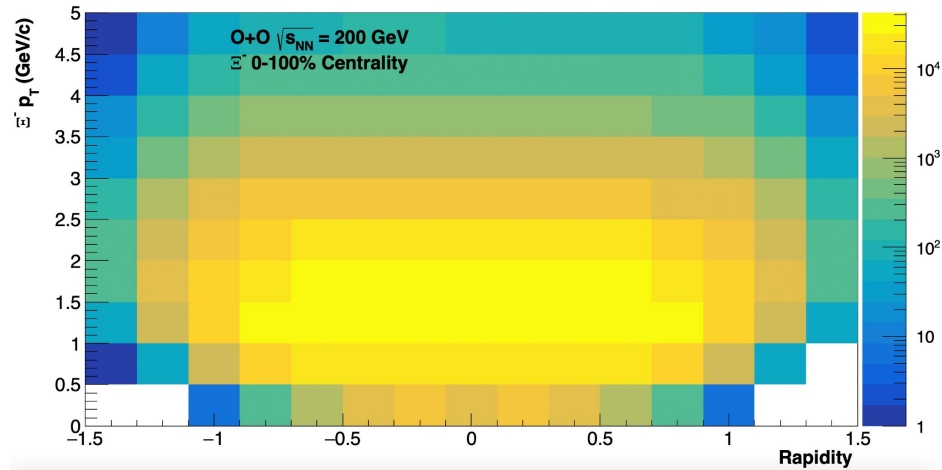
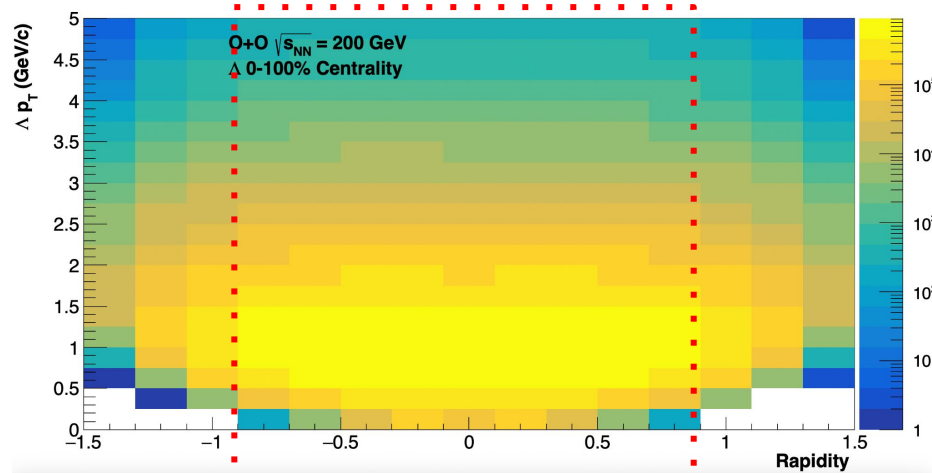


# Uncorrected Raw Spectra for Hyperions in O+O



- The large statistics, improved  $p_T$  and rapidity coverage enables STAR to have good statistics for multi-strange hadrons.
- There is good coverage through 0 - 80% centralities.

# What does our rapidity coverage looks like?



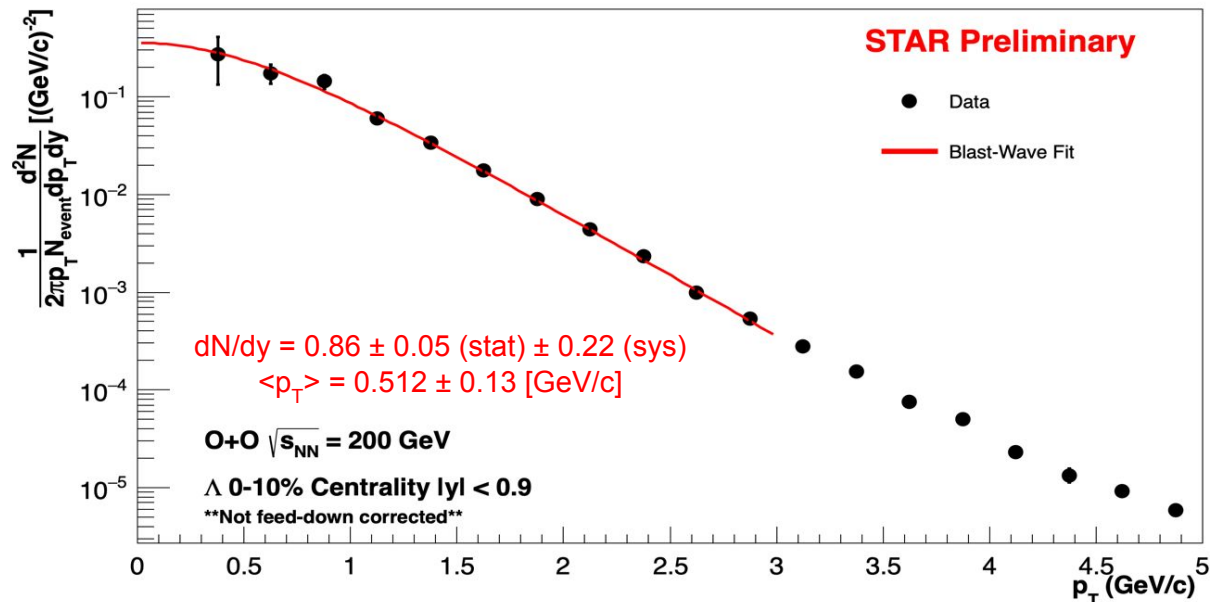
The iTPC provides extended coverage.

The rest of the talk will focus on this phase-space.

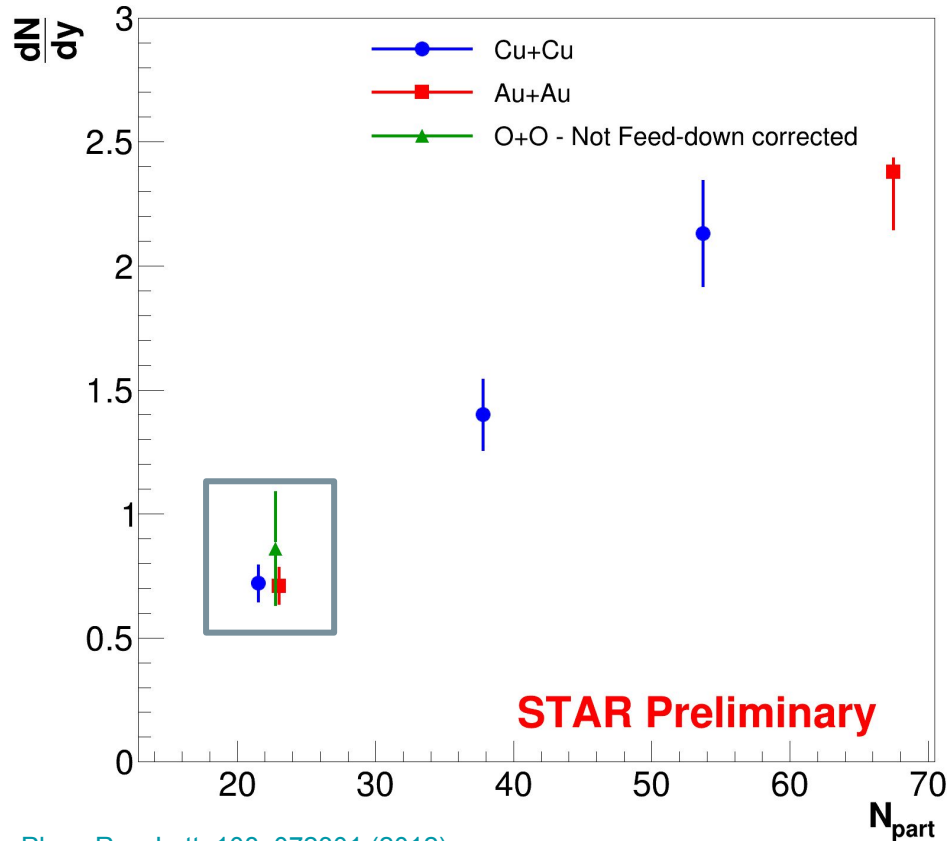


# Corrected $p_T$ spectrum for $\Lambda$ 's in O+O

- The  $p_T$  spectra is calculated from the  $\Lambda$ 's invariant mass distributions for the different momenta.
- The reconstruction efficiency is calculated using monte carlo which is embedded in real data and then propagated through the detector simulation.
- The  $\Lambda$  spectra is the average of both magnetic field configurations.



# Comparing the O+O yield to similar Collision Systems



Most central O+O collisions have a similar  $N_{part}$  as peripheral Au+Au collisions.

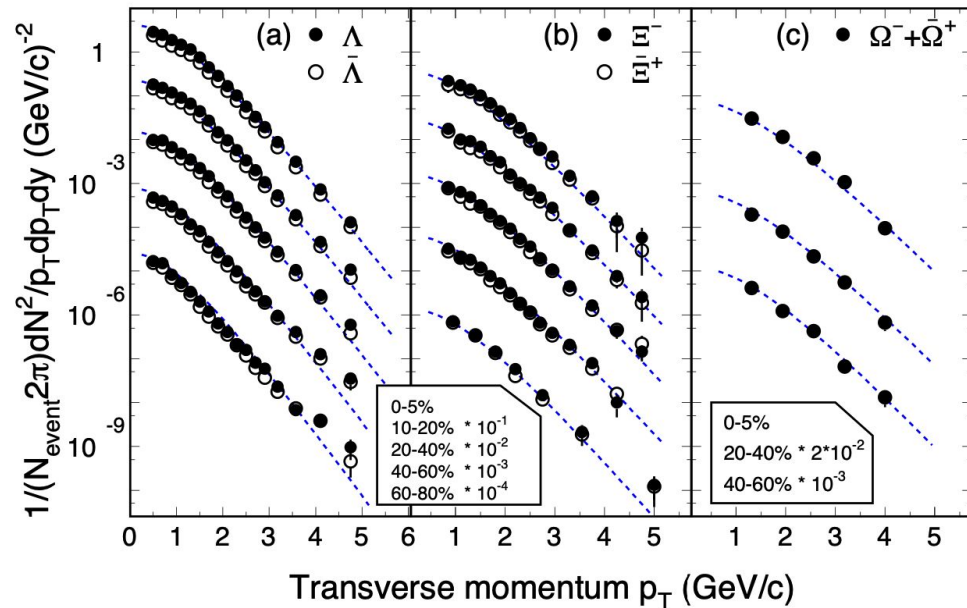
Integrating the  $\Lambda$   $p_T$  spectra from 0 to  $\infty$  the yield ( $dN/dy$ ) is  $0.86 \pm 0.05 \pm 0.22$

\*\*O+O yield is not feed-down corrected.

# Next Steps for Analysis

- Extend the analysis to other hyperons.
  - The raw  $p_T$  spectra have been made but is pending the corrections.
- Calculate the yields from corrected spectra.
- Calculate the pion yield.
- Apply feed-down corrections to spectra for yield calculations.
  - Compute the pion/hyperon ratio in the low multiplicity region
- Use thermal model for freeze-out parameter (e.g.  $\mu_B$ ,  $T_{ch}$ ) extraction.

Transverse momenta distribution for Au+Au at  $\sqrt{s_{NN}} = 200$  GeV



[STAR Collaboration. Phys. Rev. Lett. 98, 062301 \(2007\)](https://arxiv.org/abs/0705.3802)

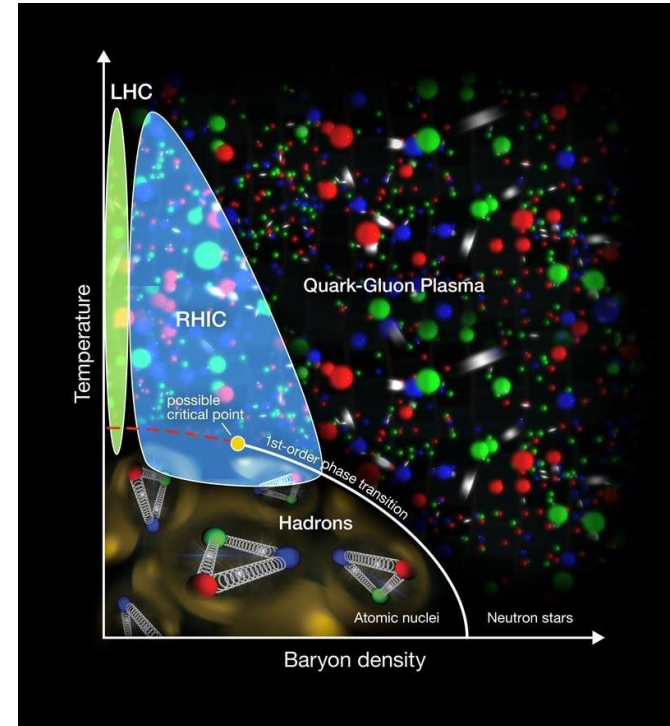


# Conclusions

- The O+O at  $\sqrt{s_{NN}} = 200$  GeV is a newer data set for STAR.
- The O+O dataset can fill in the gaps in the low multiplicity regions in the ratio of strange hadron production to the pion yield for the STAR data.
- I presented the first yield calculation for  $\Lambda$ 's in the 0-10% centrality region for O+O.
- With the great statistics there will be interesting results for the near future!

# Some of STAR's other strangeness results at CPOD

- [Y. Zhou](#) presented measurements of  $K_S^0$ ,  $\Lambda$ ,  $\Xi^-$  production at  $\sqrt{s_{NN}} = 3 - 4.5$  GeV in Au + Au collisions.
  - Soon there will be more measurements from BESII too.
- [Y. Leung](#) presented on hypernuclei production at  $\sqrt{s_{NN}} = 3 - 27$  GeV in Au+Au.
- Plus several other analysis!
- Covering different phase-space of the QCD diagram!



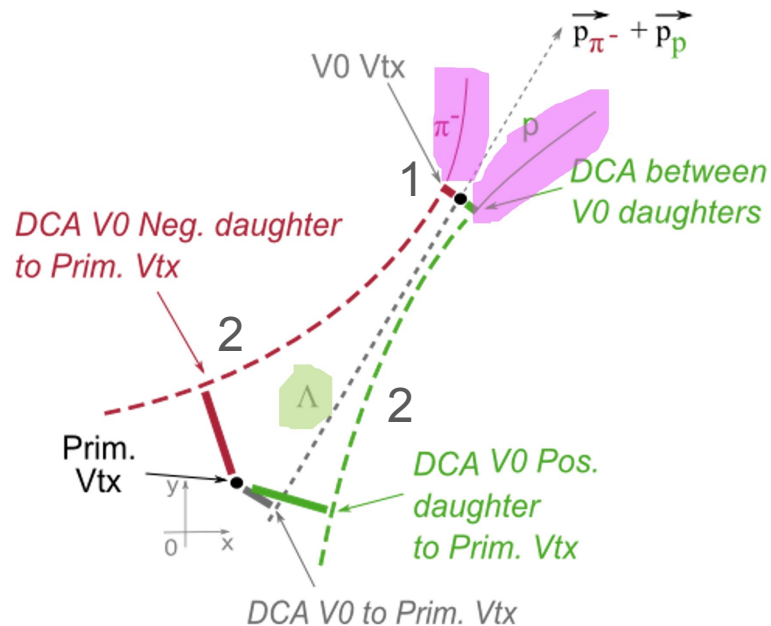
<https://www.bnl.gov/newsroom/news.php?a=121072>



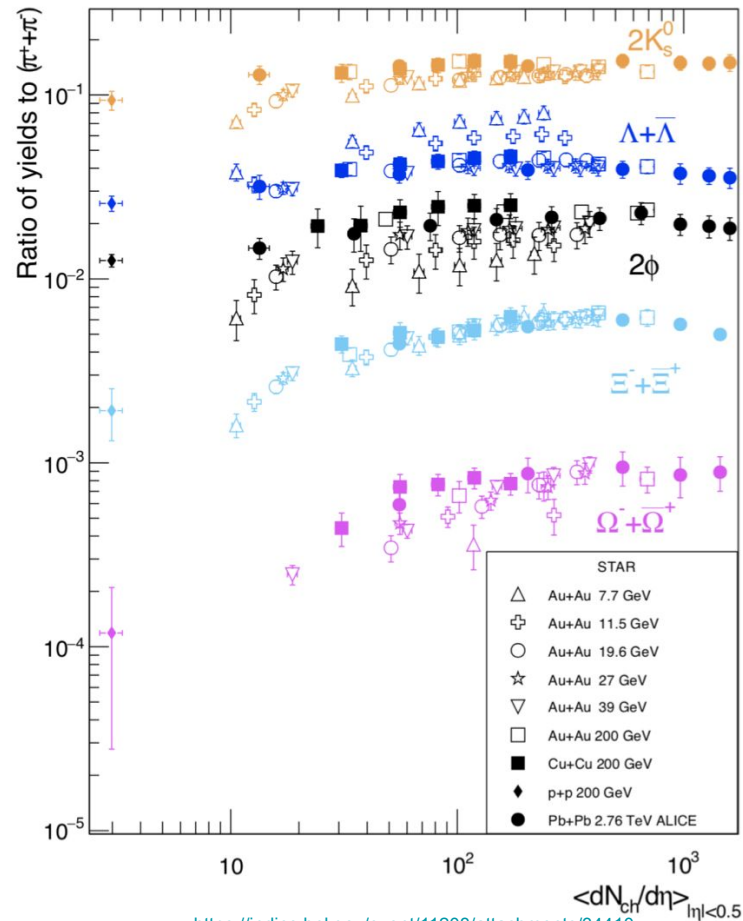
# Backup

# Reconstruction Details and Topological Cuts

- 1) **MaxDistanceBetweenParticlesCut** (DCA between daughters): 5 cm
- 2) **LCut** (distance to primary vertex): > 1 cm
- 3) **Chi2Cut2D** (cut on  $\chi^2$  of the particle fit): > 20
- 4) **ChiPrimaryCut** (cut on  $\chi^2$  of the tracks to the PV to divide tracks into primary and secondary) : > 3.
- 5) **ChiPrimaryCut2D** (cut on  $\chi^2$  of the track to the PV): > 3.
- 6) **LdLCut2D** (cut on the distance to PV normalized on the error): > 3
- 7) **Vz** < | 145 | cm
- 8) **Vr** < 2 cm
- 9) **nHitsFit** > 15



# Full spectra with BES yields



[https://indico.bnl.gov/event/11208/attachments/34410/55818/zhu\\_BNL\\_nuclear\\_seminar\\_2021.pdf](https://indico.bnl.gov/event/11208/attachments/34410/55818/zhu_BNL_nuclear_seminar_2021.pdf)