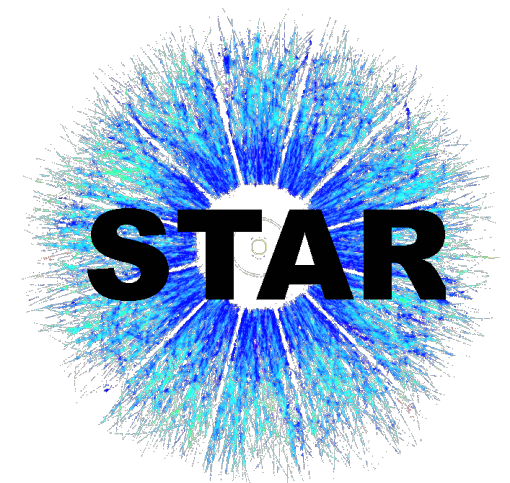


# $\phi$ spin alignment with respect to the global angular momentum reconstructed with the 1st-order event plane

Chensheng Zhou

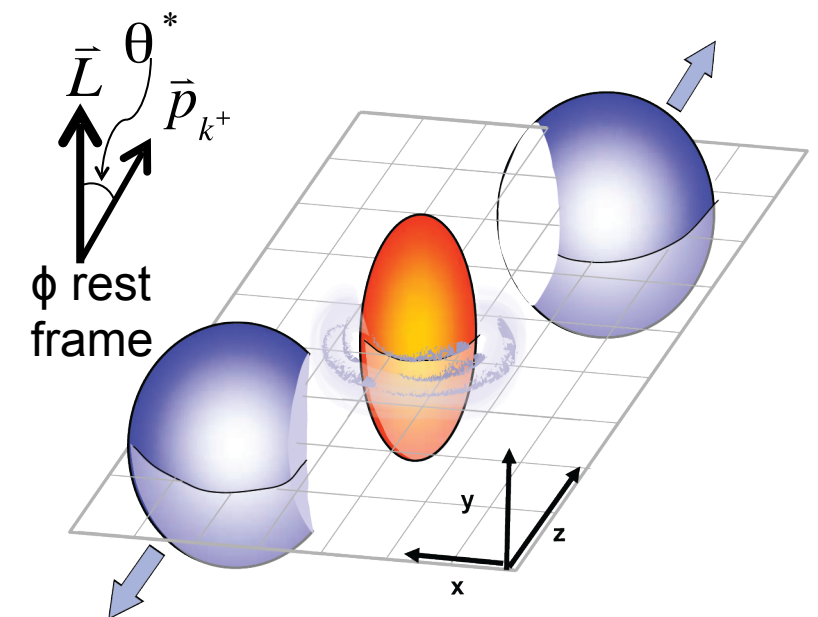
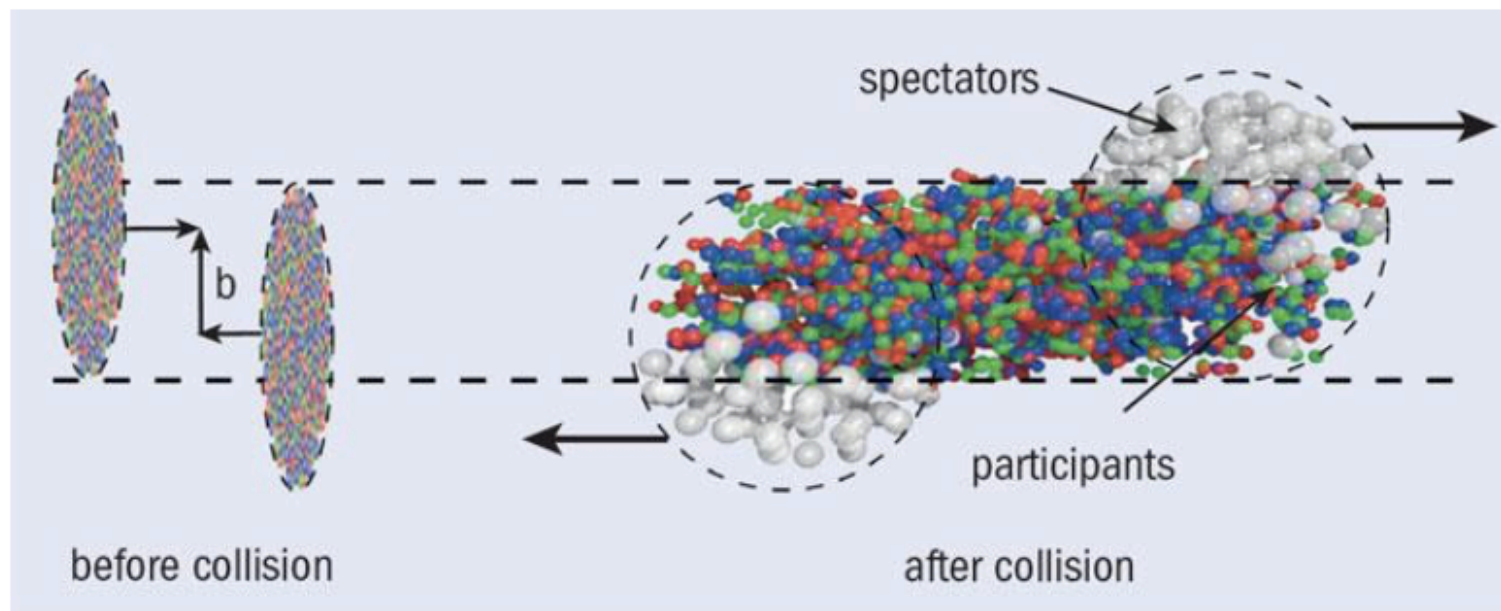
(For the STAR collaboration)  
Shanghai Institute of Applied Physics &  
Brookhaven National Laboratory



CPOD 2017  
Aug-09-2017, Stony Brook University

# Introduction

- Initial angular momentum  $\mathbf{L} \sim 10^3 \hbar$  in non-central heavy-ion collisions.
- Baryon stopping may transfer this angular momentum, in part, to the fireball.
- Due to vorticity and spin-orbit coupling,  $\phi$ -meson spin may align with  $\mathbf{L}$ .



# Spin alignment

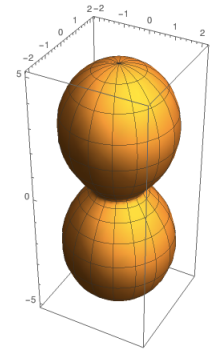
- Spin alignment can be determined from the angular distribution of the decay products\*:

$$\frac{dN}{d(\cos\theta^*)} = N_0 \times [(1 - \rho_{00}) + (3\rho_{00} - 1)\cos^2\theta^*]$$

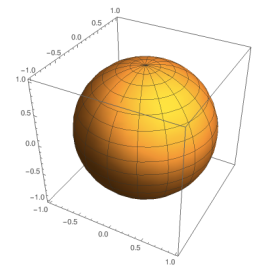
where  $N_0$  is the normalization and  $\theta^*$  is the angle between the polarization direction  $\mathbf{L}$  and the momentum direction of a daughter particle in the rest frame of the parent vector meson.

- A deviation of  $\rho_{00}$  from 1/3 signals net spin alignment.

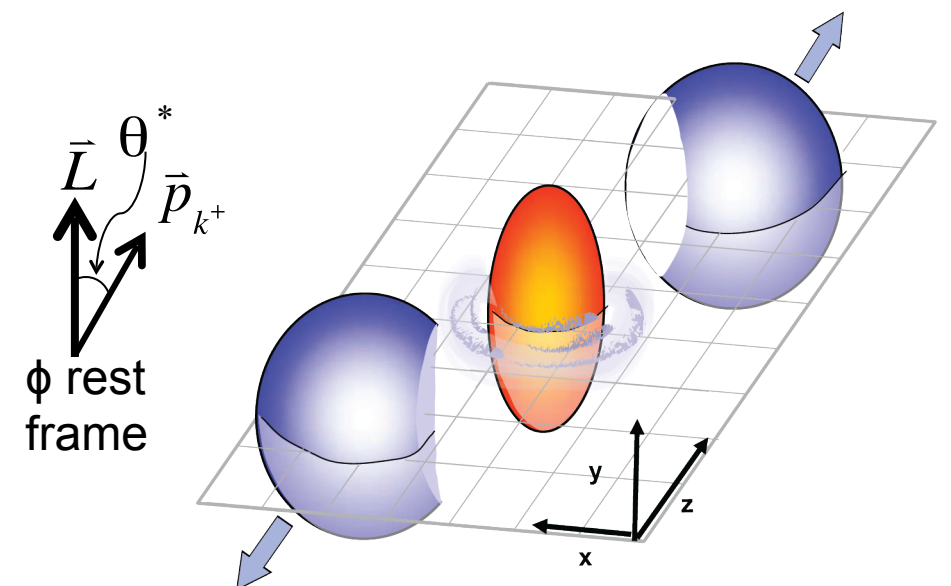
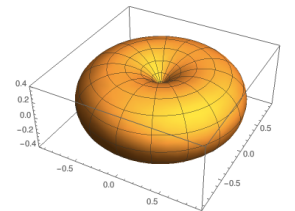
$\rho_{00} > 1/3$ :



$\rho_{00} = 1/3$ :



$\rho_{00} < 1/3$ :



\*K. Schilling et al., Nucl. Phys. B 15, 397 (1970)

# Hadronization scenarios

- Recombination of polarized quarks and anti-quarks in QGP likely dominates in the low  $p_T$  and central rapidity region.
- Fragmentation of polarized quarks  $q \rightarrow V + X$ , likely happens in the intermediate  $p_T$  and forward rapidity region. ( $V$  is the vector meson, which is  $\phi$  in our analysis)

$$\rho_{00}^{\phi(rec)} = \frac{1 - P_s^2}{3 + P_s^2}$$

Always smaller than 1/3

$$P_s = -\frac{\pi}{4} \frac{\mu p}{E(E + m_s)} \quad \text{is the global quark polarization}$$

$$P_{\bar{s}}^{frag} = -\beta P_s \quad \text{is the polarization of the anti-quark created in the fragmentation process}$$

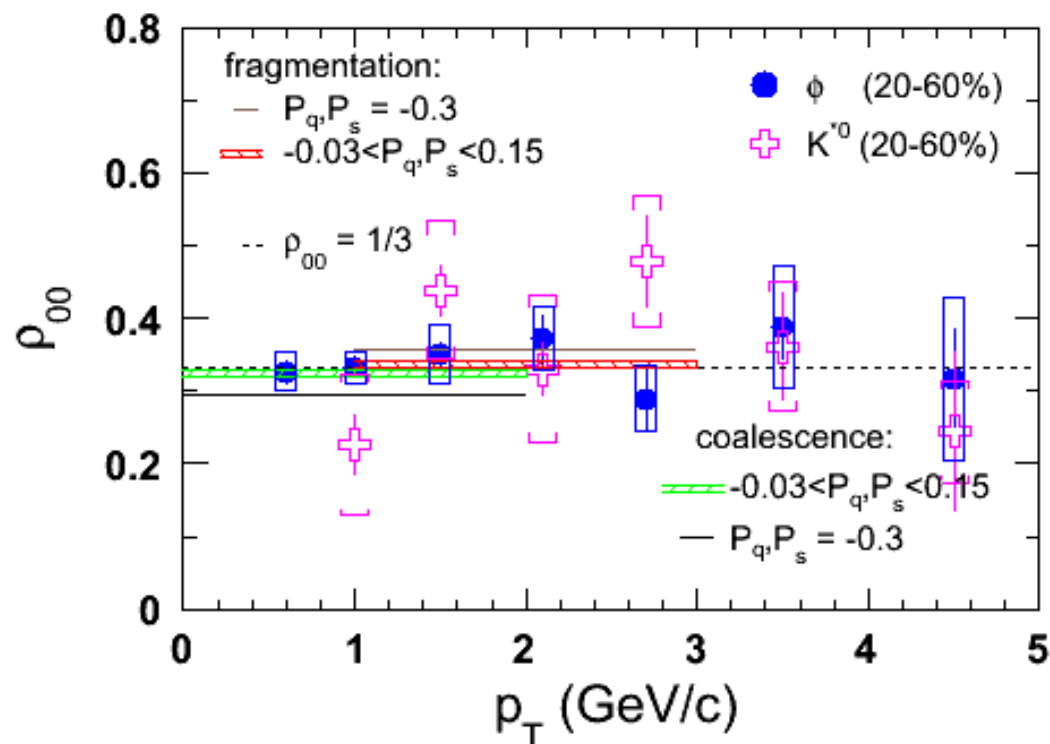
$$\rho_{00}^{\phi(frag)} = \frac{1 + \beta P_s^2}{3 - \beta P_s^2}$$

Always larger than 1/3



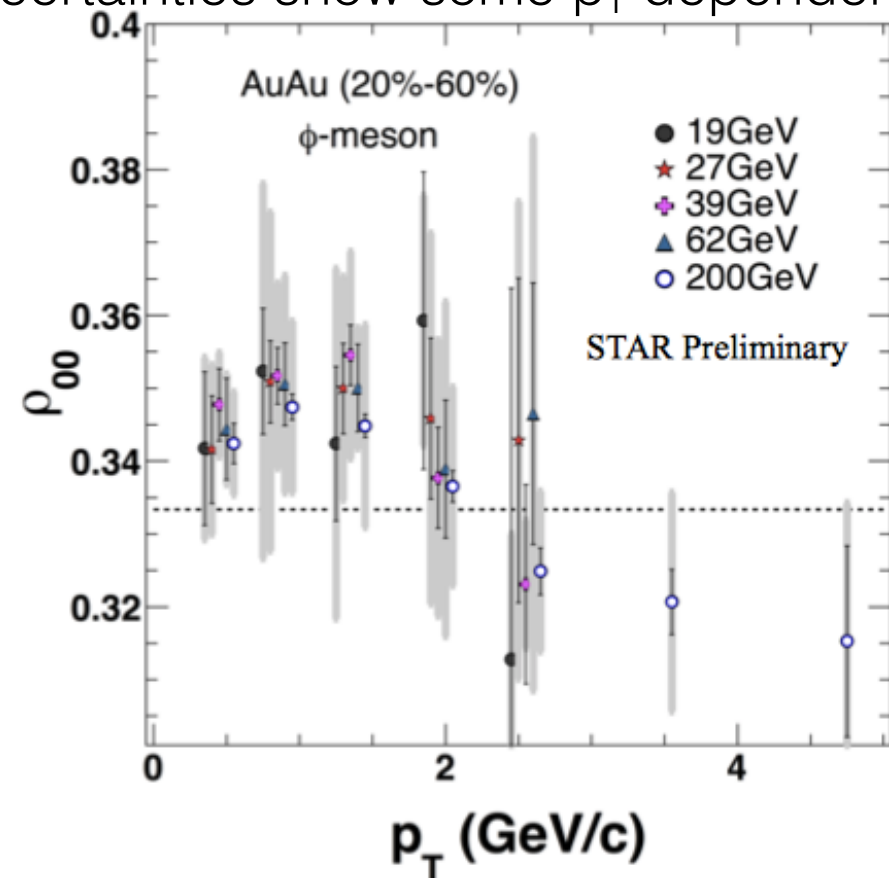
# STAR's previous results

- STAR has published results with data taken in year 2004.
- Updated results have been shown at QM2017 (Xu Sun's poster), with data taken in year 2010 & 2011.
- Both of the above use the 2nd-order event plane obtained from TPC. The published result is consistent with  $1/3$ ; New results with reduced uncertainties show some  $p_T$  dependence.



STAR's Published results

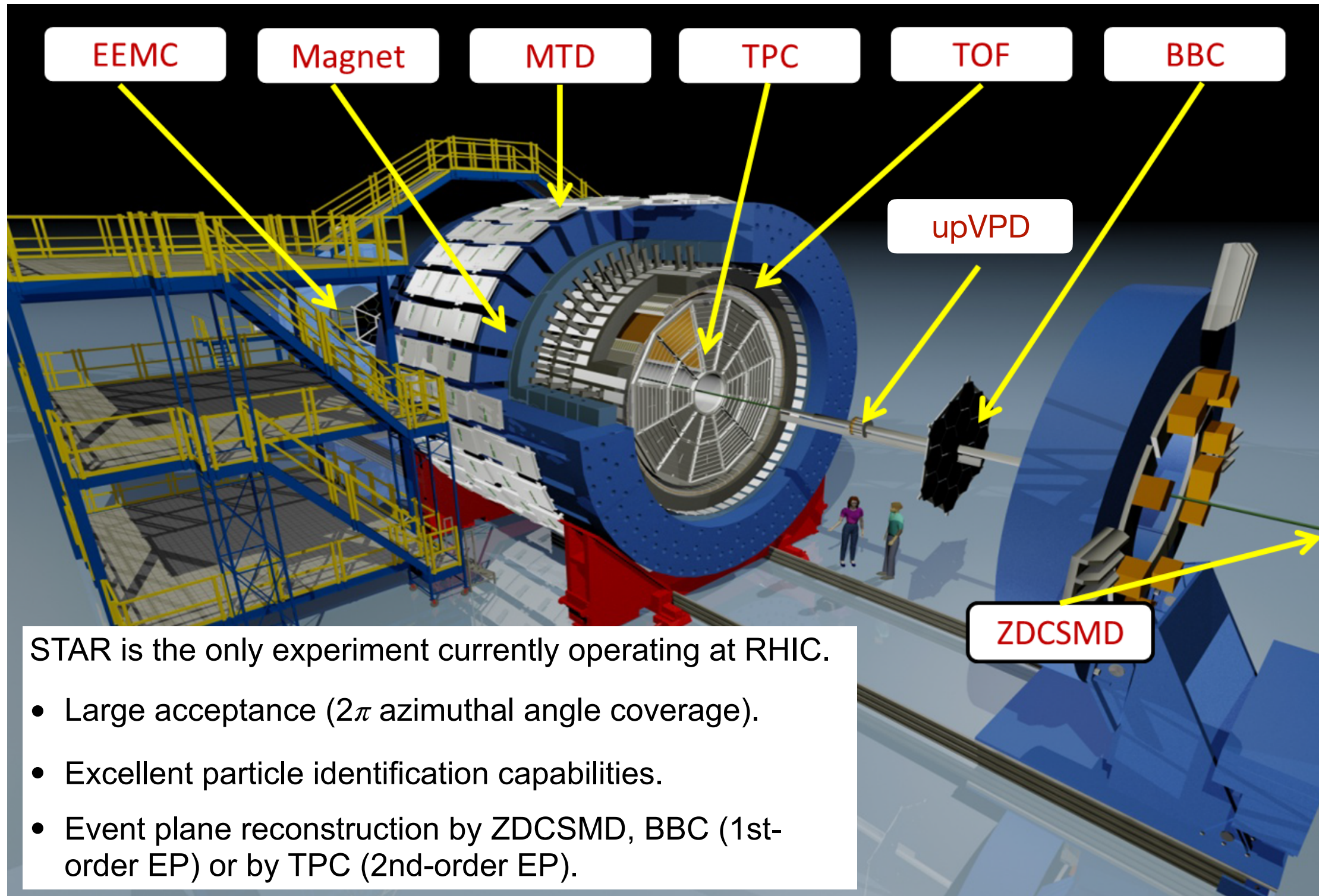
B.I.Abelev et al (STAR Collaboration), Phys. Rev. C77, 061902(R) (2008)



Xu Sun's QM2017 poster



# STAR detector



STAR is the only experiment currently operating at RHIC.

- Large acceptance ( $2\pi$  azimuthal angle coverage).
- Excellent particle identification capabilities.
- Event plane reconstruction by ZDCSMD, BBC (1st-order EP) or by TPC (2nd-order EP).



# Datasets and cuts

- Number of events:

- Au+Au 200 GeV ~ 500M
  - Au+Au 39 GeV ~ 100M
  - Au+Au 27 GeV ~ 30M
  - Au+Au 19.6 GeV ~ 10M
  - Au+Au 11.5 GeV ~ 3M

- Event cuts:

- $-30.0 < V_z < 30.0$  cm
  - $V_r < 2.0$  cm
  - $-3.0 < V_z - V_z^{\text{VPD}} < 3.0$  cm
  - Number ToF matched point  $> 3$
  - Minimum Bias Event
  - Bad runs are rejected

- Track cuts:

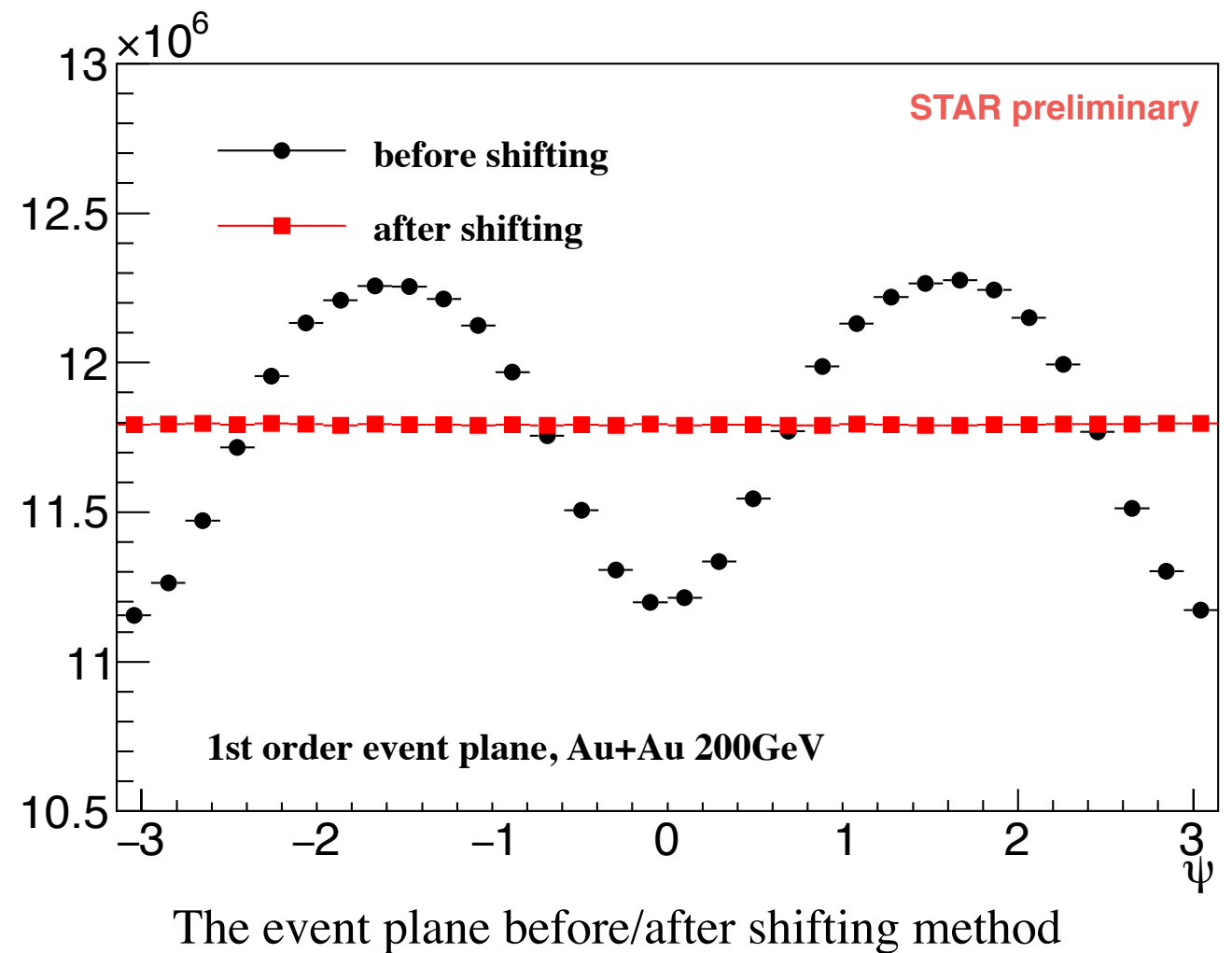
- $n_{\text{HitsFit}} > 15$
  - $n_{\text{HitsFit}}/n_{\text{HitsMax}} > 0.52$
  - $-1.0 < \eta < 1.0$
  - $dca < 2.0$  cm
  - $p_T > 0.1$  GeV/c
  - $p < 10$  GeV/c
  - invariant mass  $< 1.1$  GeV/c<sup>2</sup>

- Track PID:

Momentum(GeV/c)	With TOF	Without TOF
[0, 0.65]	$0.16 < m^2 < 0.36,  n_{\text{SigmaKaon}}  < 2.5$	$-1.5 < n_{\text{SigmaKaon}} < 2.5$
(0.65, 1.5)	$0.16 < m^2 < 0.36,  n_{\text{SigmaKaon}}  < 2.5$	—
[1.5, ∞)	$0.125 < m^2 < 0.36,  n_{\text{SigmaKaon}}  < 2.5$	—

# 1st order event plane

- In our analysis, the event plane is obtained from **ZDCSMD** (for 200 GeV data) or **BBC** (for low energy data) and flattened by shifting method\*. The flattening is applied for every 10 runs (about 60000 events in Au+Au 200 GeV collisions).



\*A. Poskanzer and S. Voloshin, PRC 58, 1671 (1998)

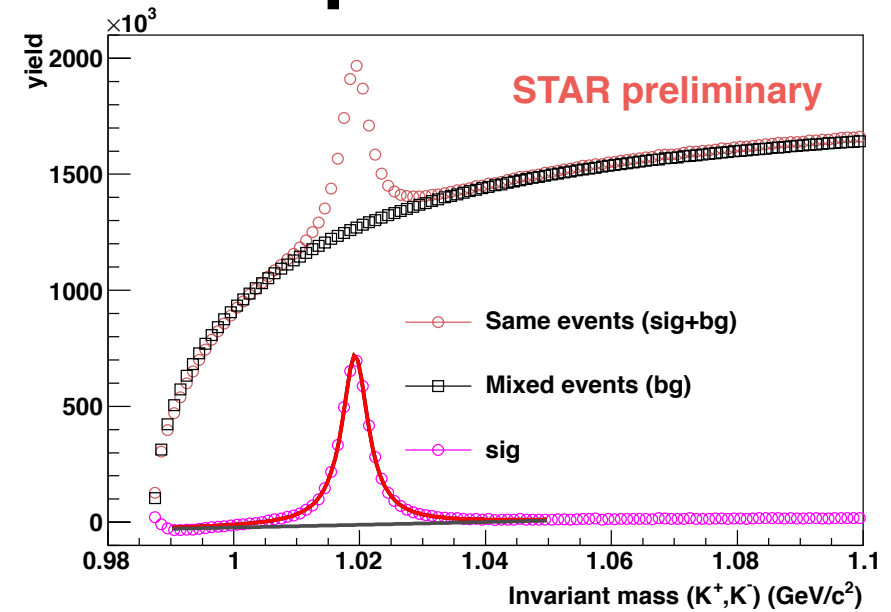


# Obtaining yields of $\phi$ meson

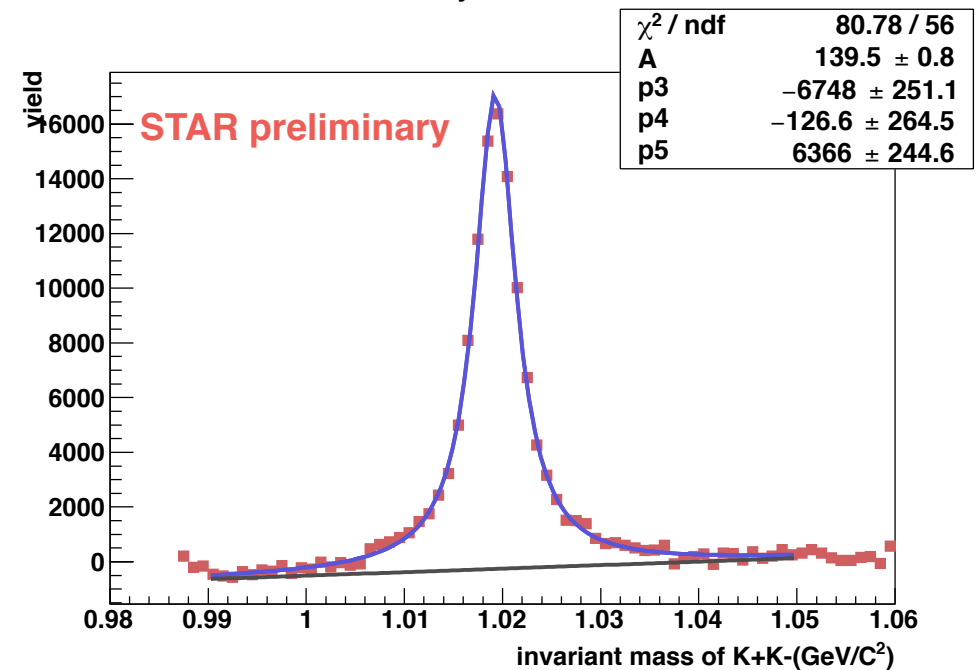
- The background is obtained using event mixing technique.
- The  $\phi$ -mesons signal is fitted with Briet-Wigner function and the 2nd order polynomial function for residual background to extract raw  $\phi$  meson yield:

$$BW(m_{inv}) = \frac{1}{2\pi} \frac{A\Gamma}{(m - m_\phi)^2 + (\Gamma/2)^2}$$

where  $\Gamma$  is the width of the distribution and  $A$  is the area of the distribution.  $A$  is the raw yield scaled by the bin width ( $= 0.001 \text{ GeV}/c^2$ ).



Fitting of all  $p_T$  &  $\cos\theta^*$  range.  
Centrality: 40~50%



Fitting of a single  $p_T$  &  $\cos\theta^*$  bin.  
Centrality: 40%-50%  $p_T$ : 1.2~1.8 GeV/C  $\cos\theta^*$ : -0.6~-0.4

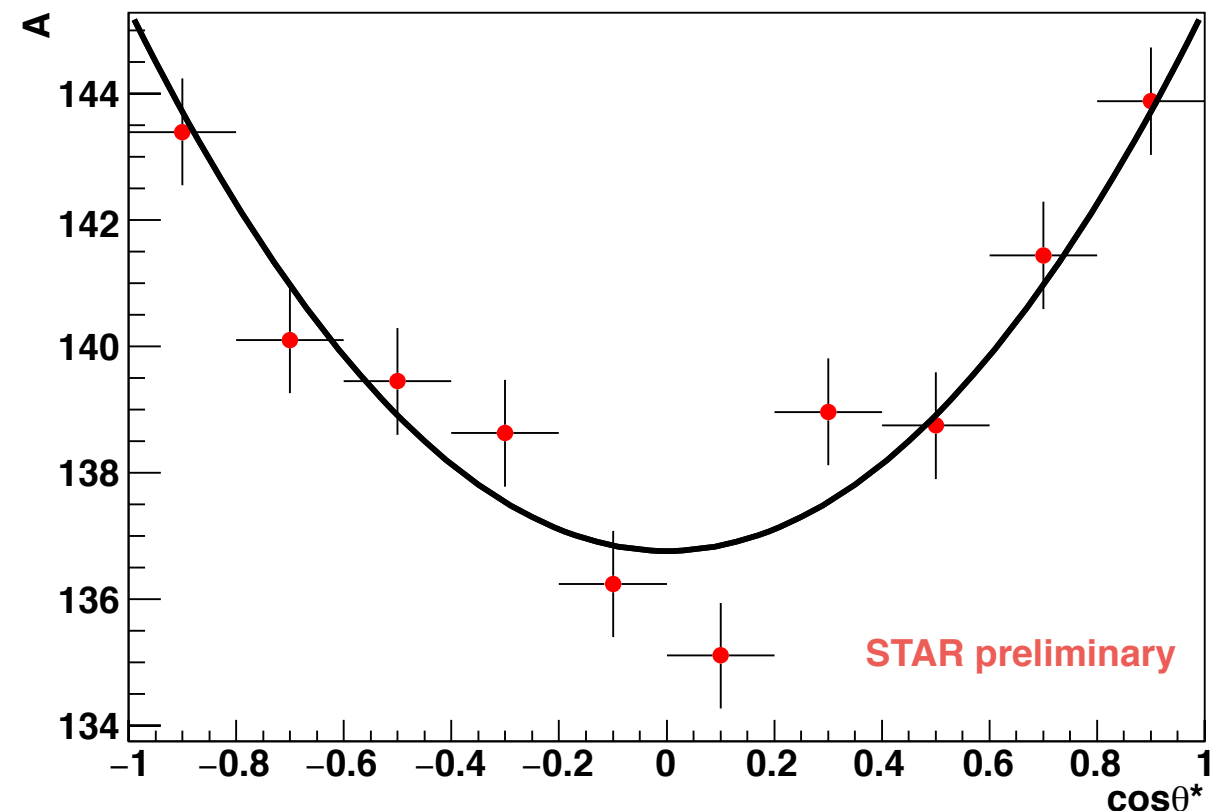
# Extracting observed $\rho_{00}$

- With yield of  $\phi$  for different bins, we can fit the yield distribution and obtain  $\rho_{00}$  using

$$\frac{dN}{d(\cos\theta^*)} = N_0 \times [(1 - \rho_{00}) + (3\rho_{00} - 1)\cos^2\theta^*]$$

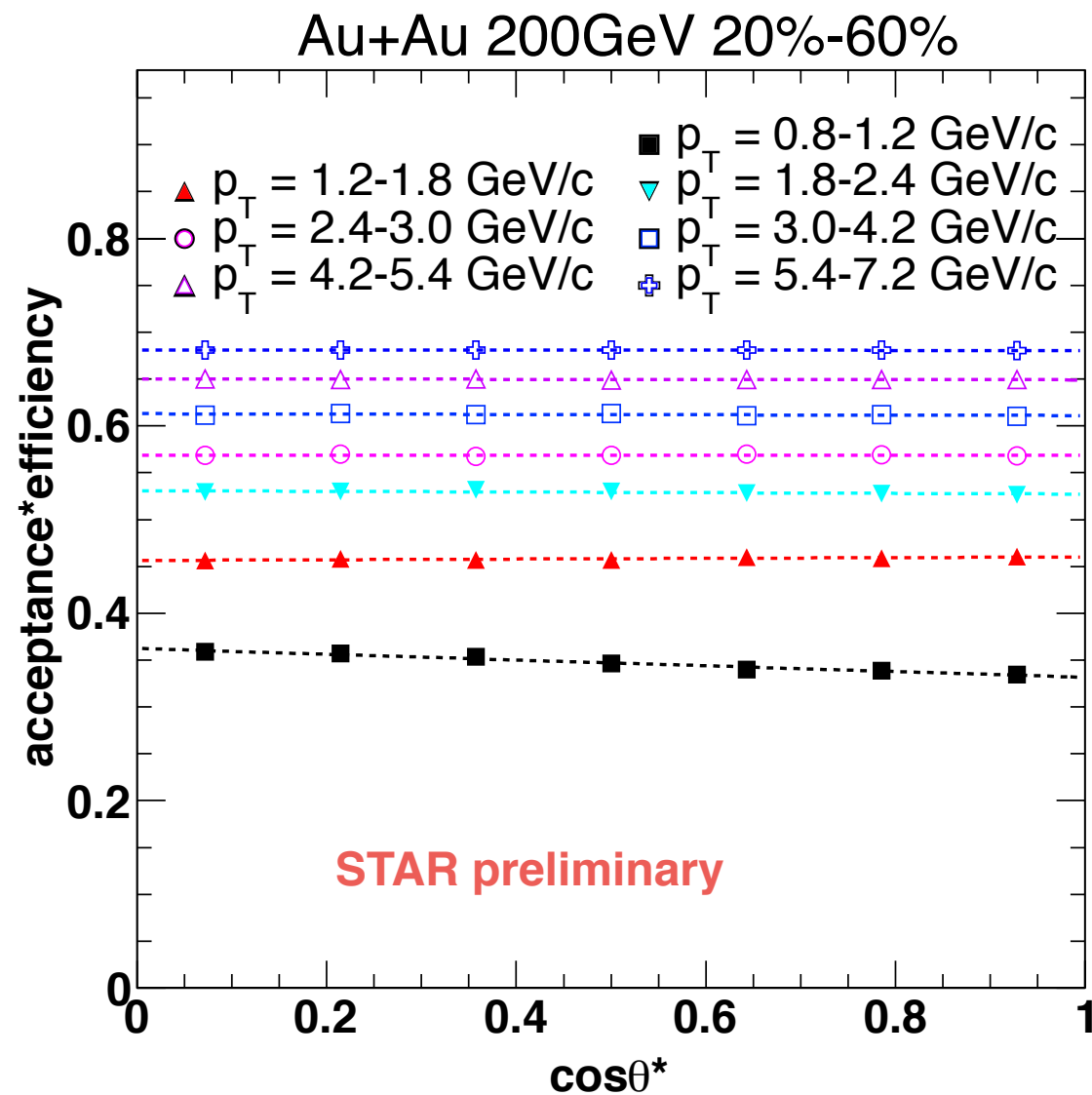
$\theta^*$  is the angle between the polarization direction  $\mathbf{L}$  and the momentum direction of a daughter particle in the rest frame of the parent vector meson.

- What we extracted here is the  $\rho_{00}$  before event plane resolution correction (observed  $\rho_{00}$ ).



Fitting of yield Vs  $\cos\theta^*$   
 Au+Au 200 GeV  
 Centrality: 40-50%  
 $p_T$ : 0.8~1.2 GeV/c

# Efficiency and acceptance



- $\phi$ -meson efficiency\*acceptance is calculated with  $K^+$  and  $K^-$  embedding data and shows very weak  $\cos\theta^*$  dependence, and the effect on  $\rho_{00}$  is negligible.

# Derivation of event plane resolution correction

- For spin = 1 particles, their daughter's angular distribution can be written in a general form as a function of  $\theta^*$  and  $\beta$  (the azimuthal angle w.r.t  $\mathbf{L}$ , see the picture at bottom right):

$$\frac{dN}{d\cos\theta^* d\beta} \propto 1 + A\cos^2\theta^* + B\sin^2\theta^* \cos 2\beta + C\sin 2\theta^* \cos\beta$$

- where

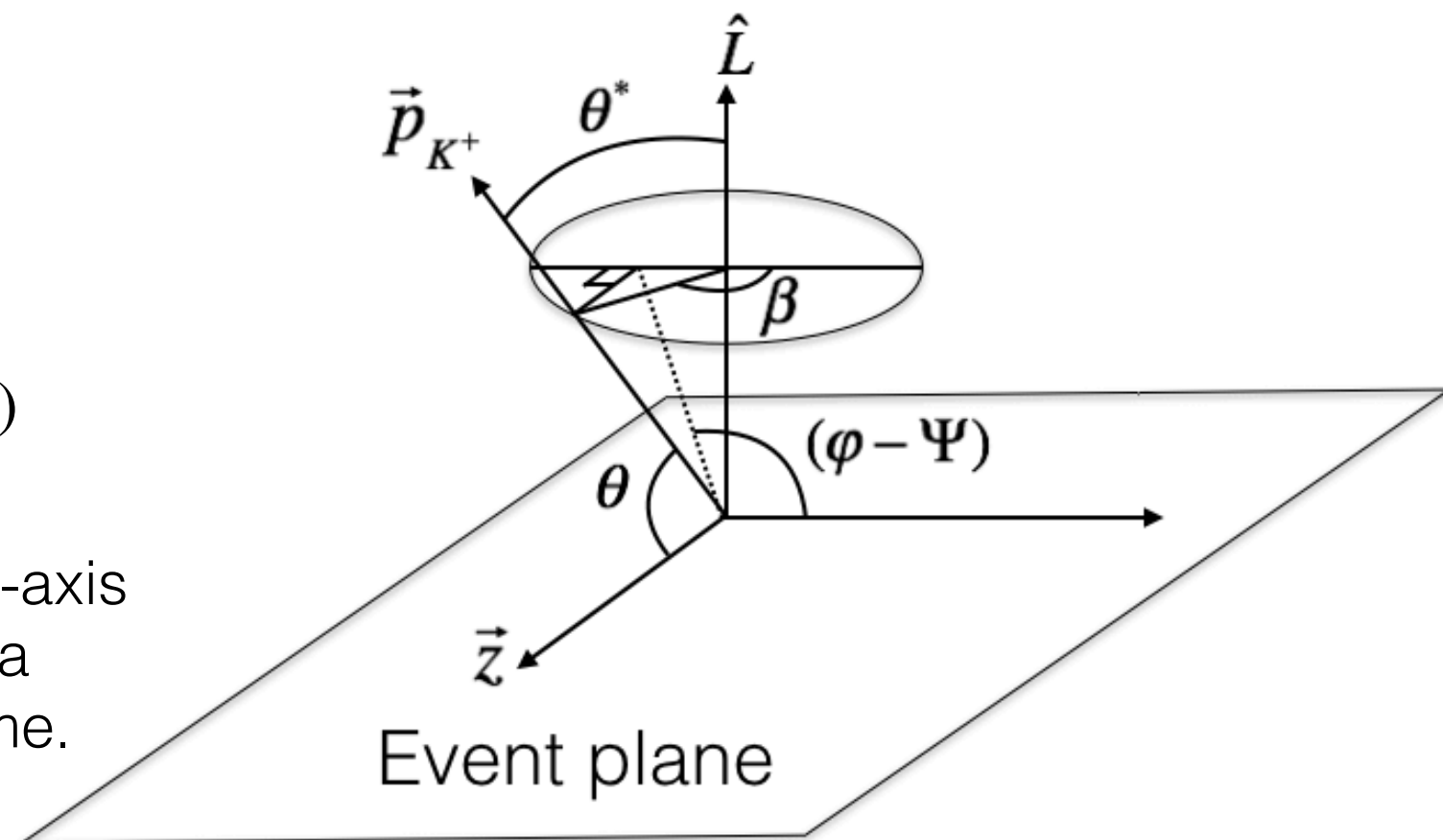
$$A = (3\rho_{00} - 1) / (1 - \rho_{00})$$

- We have

$$\cos\theta^* = \sin\theta \sin(\varphi - \psi)$$

$$\cos\theta = \sin\theta^* \sin\beta$$

where  $\theta$  is the angle between Z-axis and the momentum direction of a daughter particle in the rest frame.





# Derivation of event plane resolution correction

- The observed event plane  $\psi'$  may be different from the real event plane:

$$\psi' = \psi + \Delta$$

- The distribution of  $\Delta$  is supposed to follow an even function, so we can assume

$$\langle \cos 2\Delta \rangle = R, \quad \langle \sin 2\Delta \rangle = 0$$

- When  $\psi \rightarrow \psi'$ ,  $\theta^* \rightarrow \theta'^*$ ,  $\beta \rightarrow \beta'$ , we have

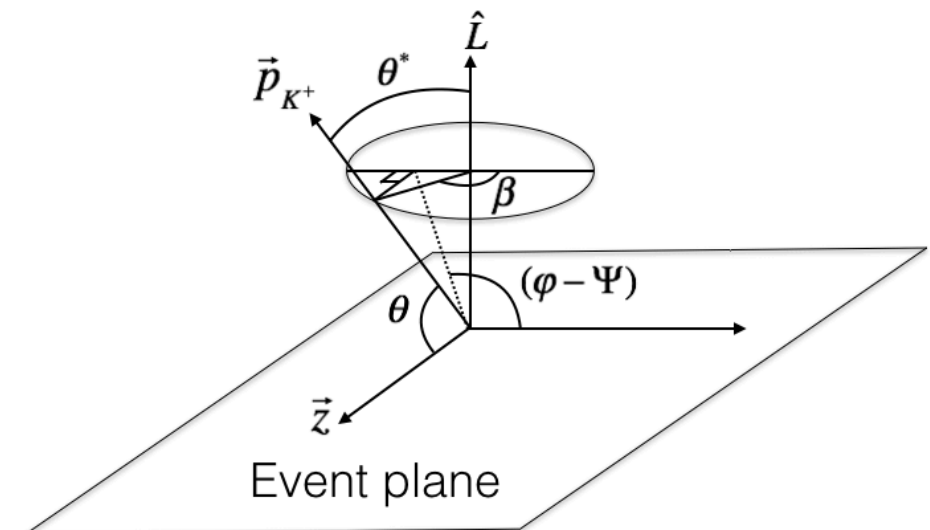
$$\begin{pmatrix} 1 \\ A \\ B \\ C \end{pmatrix} \rightarrow \begin{pmatrix} 1 \\ A' \\ B' \\ C' \end{pmatrix} = \begin{pmatrix} 1 \\ \frac{A(1+3R)+B(3-3R)}{4+A(1-R)+B(-1+R)} \\ \frac{A(1-R)+B(3+R)}{4+A(1-R)+B(-1+R)} \\ \frac{4 \cdot C \cdot R}{4+A(1-R)+B(-1+R)} \end{pmatrix}$$

- When  $B=0$ ,  $A' = \frac{A(1+3R)}{4+A(1-R)}$ ,  $\rho_{00}^{real} - \frac{1}{3} = \frac{4}{1+3R} (\rho_{00}^{obv} - \frac{1}{3})$

$$\frac{dN}{d \cos \theta^* d\beta} \propto 1 + A \cos^2 \theta^* + B \sin^2 \theta^* \cos 2\beta + C \sin 2\theta^* \cos \beta$$

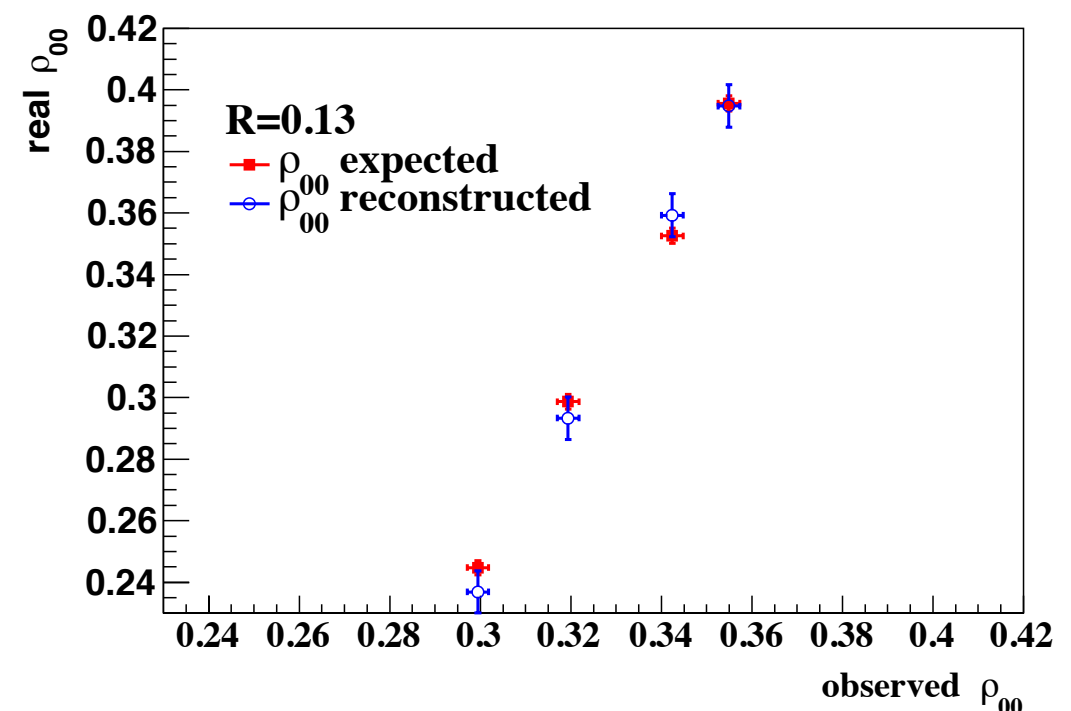
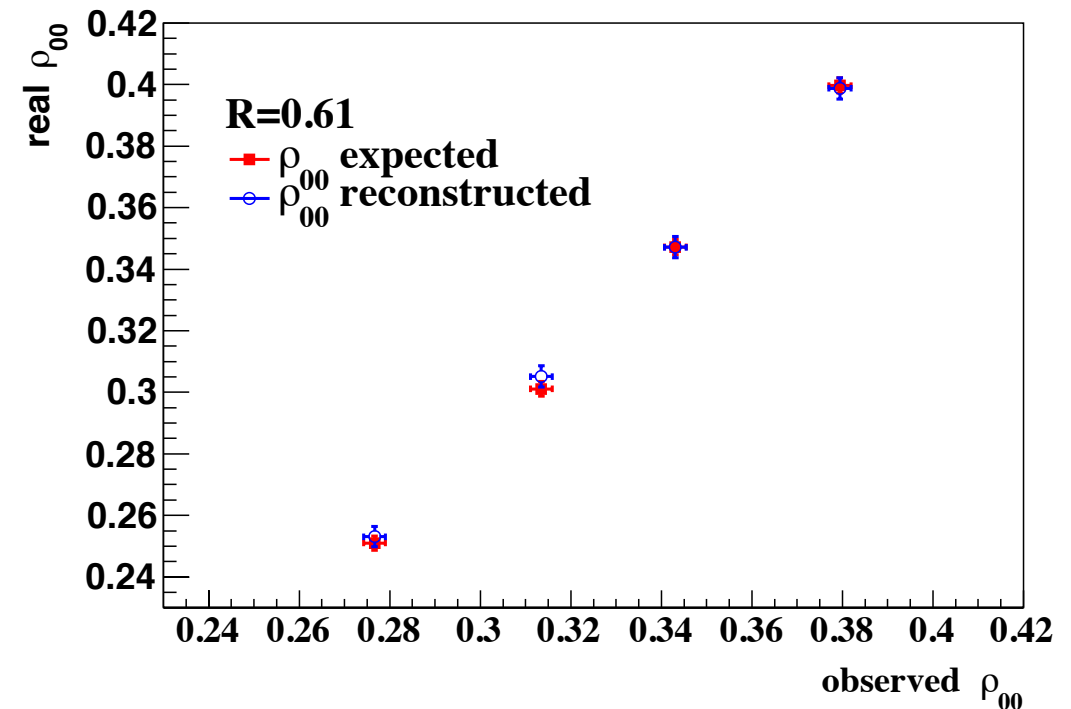
↓ rotate w.r.t z-axis by  $\Delta$

$$\frac{dN}{d \cos \theta'^* d\beta'} \propto 1 + A' \cos^2 \theta'^* + B' \sin^2 \theta'^* \cos 2\beta' + C' \sin 2\theta'^* \cos \beta'$$

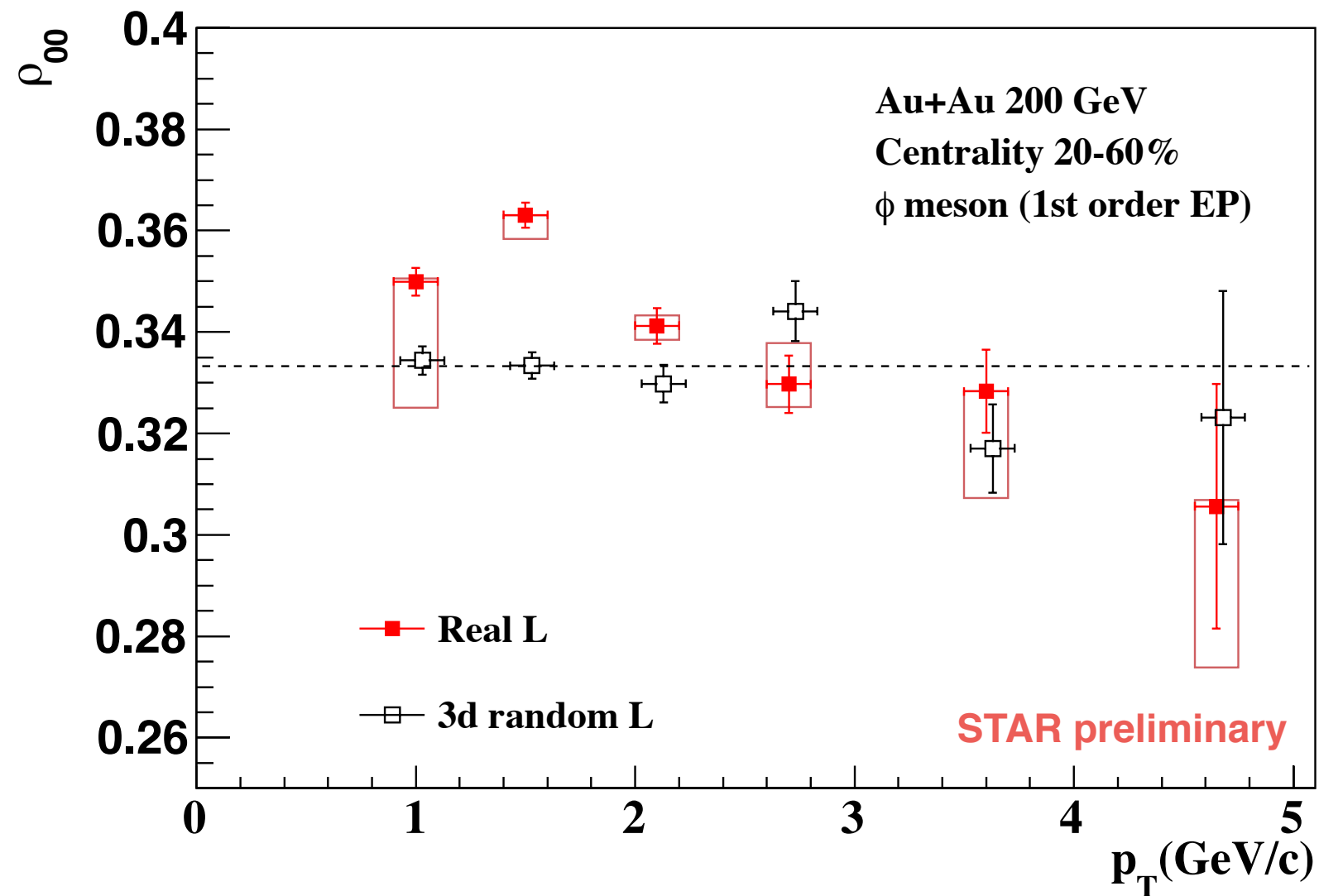


# Verify the resolution correction formula with simulations

- To test the formula of resolution correction, we generate Monte Carlo events by Pythia with  $\Delta$  following gaussian distributions.
- $\rho_{00}^{real}$  can be either obtained by fitting the yield with real event plane (without  $\Delta$ ), or by calculation with the correction formula we derived.
- The plots show the comparison of results between two methods. The correction works well even when the resolution is low.

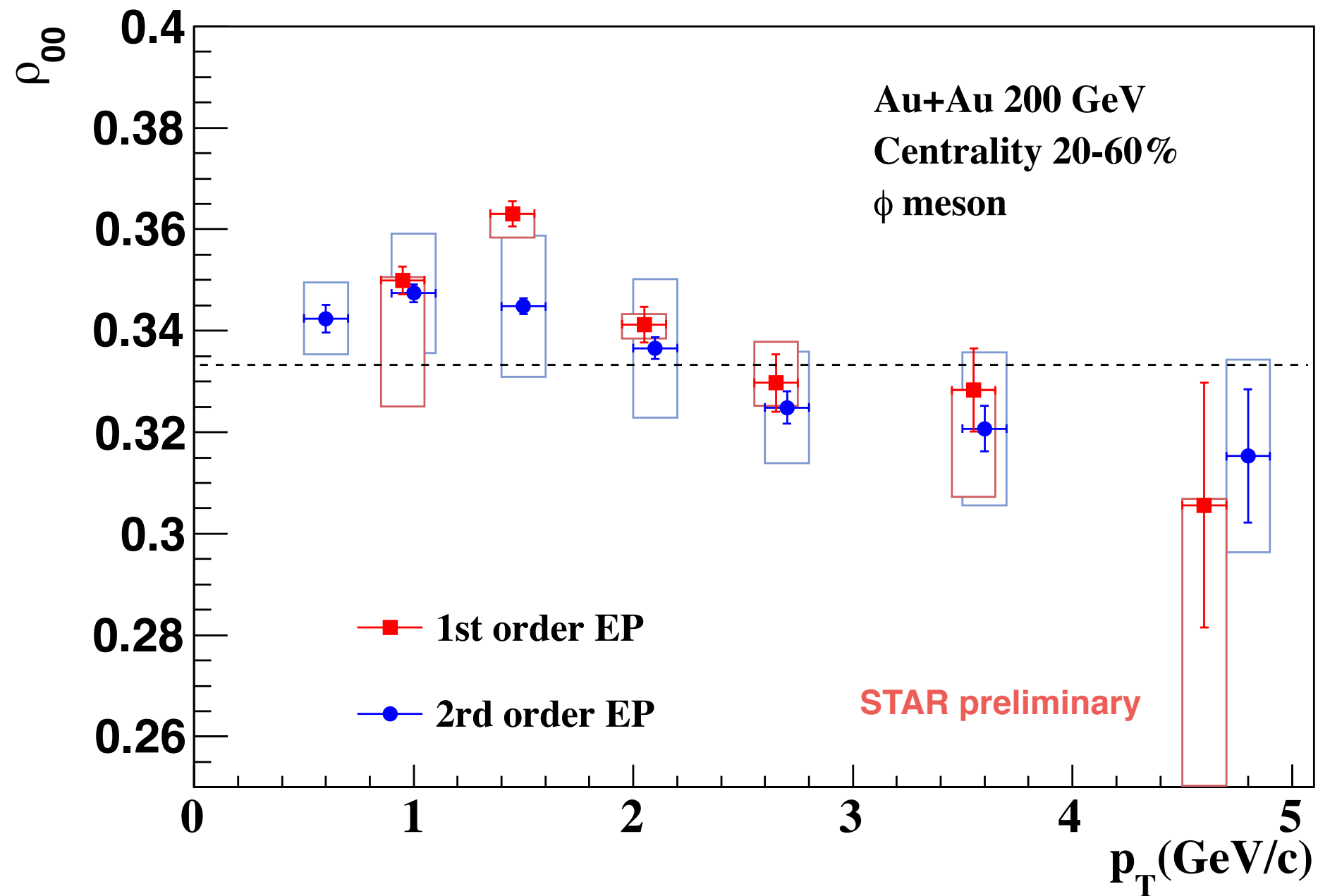


# $\rho_{00}$ VS. $p_T$



- Non-trivial  $p_T$  dependence is seen.  $6\sigma$  away from  $1/3$  at  $p_T = 1.5$  GeV/c.
- As a consistency check, the  $\rho_{00}$  is also studied with an  $\mathbf{L}$  direction randomized in 3d-space, which is at the expected value of  $1/3$ .

# 1st EP vs. 2nd EP



- To explain the difference at  $p_T \sim 1.5$  GeV/c, we need to consider the de-correlation between the two EPs.



# De-correlation between 1st and 2nd order event planes

- In the derivation of resolution, we have correction term  $R$  as:

$$R = \langle \cos 2\Delta \rangle$$

for 1st(2nd) order EP, the corresponding correction term becomes  $R_{1,2} = \langle \cos 2(\Psi_{1,2} - \Psi) \rangle$ ,

and for 2nd order EP with the consideration of de-correlation, the correction term can be written down as:

$$R_{12} = \langle \cos 2(\Psi_2 - \Psi_1 + \Psi_1 - \Psi) \rangle = D_{12} \cdot R_1,$$

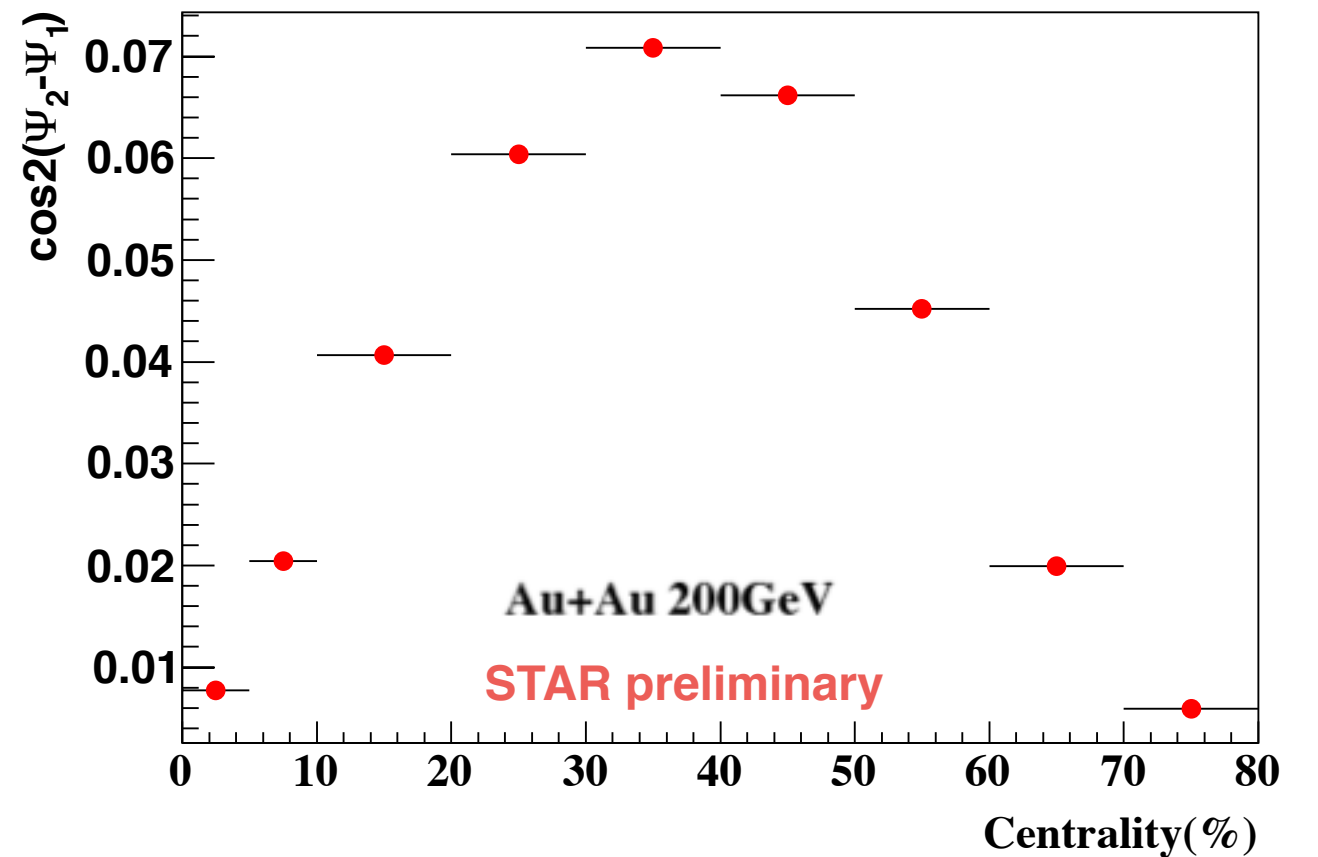
$$\text{where } D_{12} = \langle \cos 2(\Psi_2 - \Psi_1) \rangle$$

- Then we can take the corrected  $\rho_{00}$  from 1st order EP as real  $\rho_{00}$ , and use the resolution correction formula to recover 2nd order EP result:

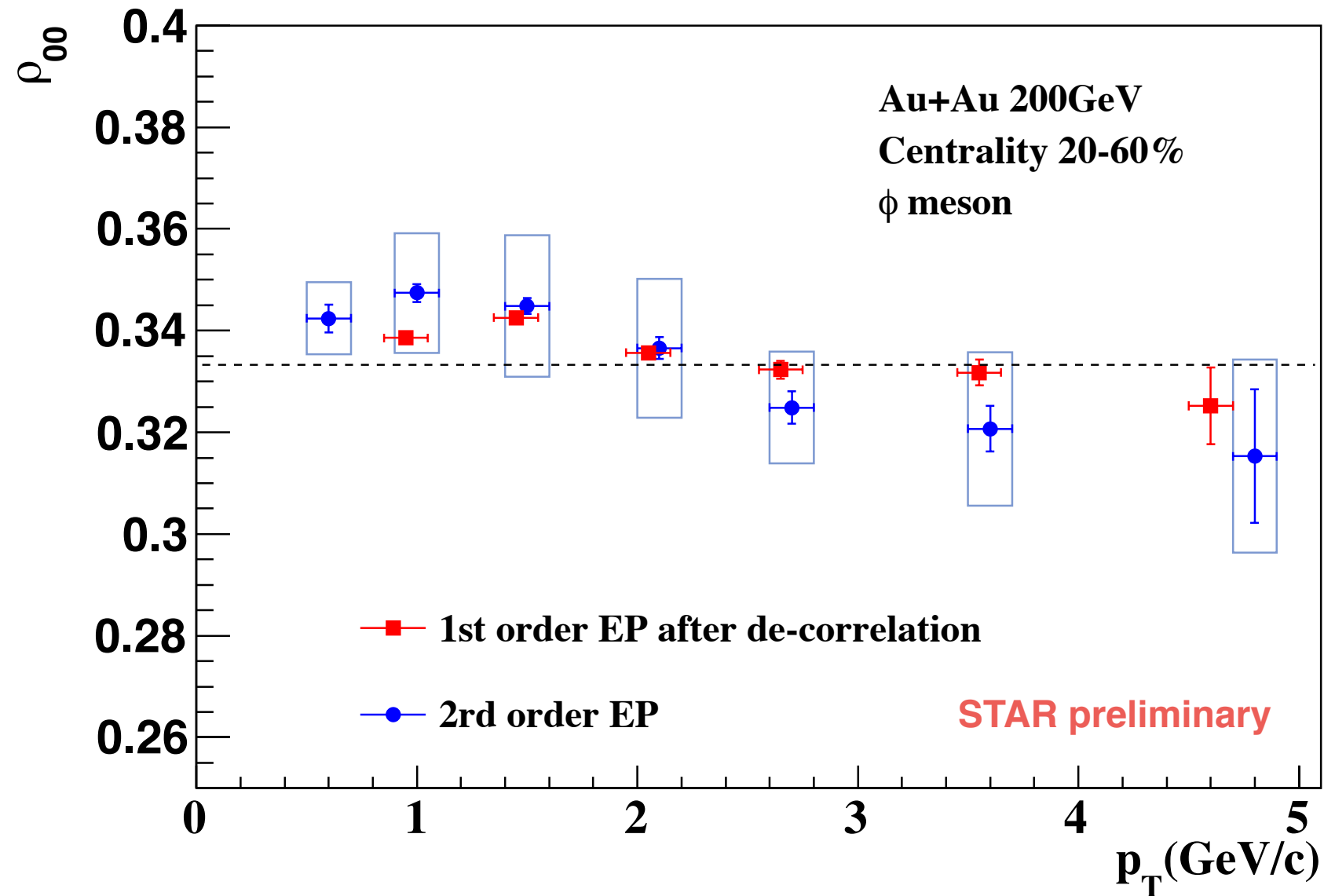
$$\rho_{\text{obv}}^{2\text{nd}} - \frac{1}{3} = \frac{1 + 3R_2}{4} \left( \rho_{00}^{2\text{nd}} - \frac{1}{3} \right)$$

$$\rho_{\text{obv}}^{2\text{nd}} - \frac{1}{3} = \frac{1 + 3D_{12} \cdot R_1}{4} \left( \rho_{00}^{1\text{st}} - \frac{1}{3} \right)$$

$$\Rightarrow \rho_{00}^{2\text{nd}} - \frac{1}{3} = \frac{1 + 3D_{12} \cdot R_1}{1 + 3R_2} \left( \rho_{00}^{1\text{st}} - \frac{1}{3} \right)$$



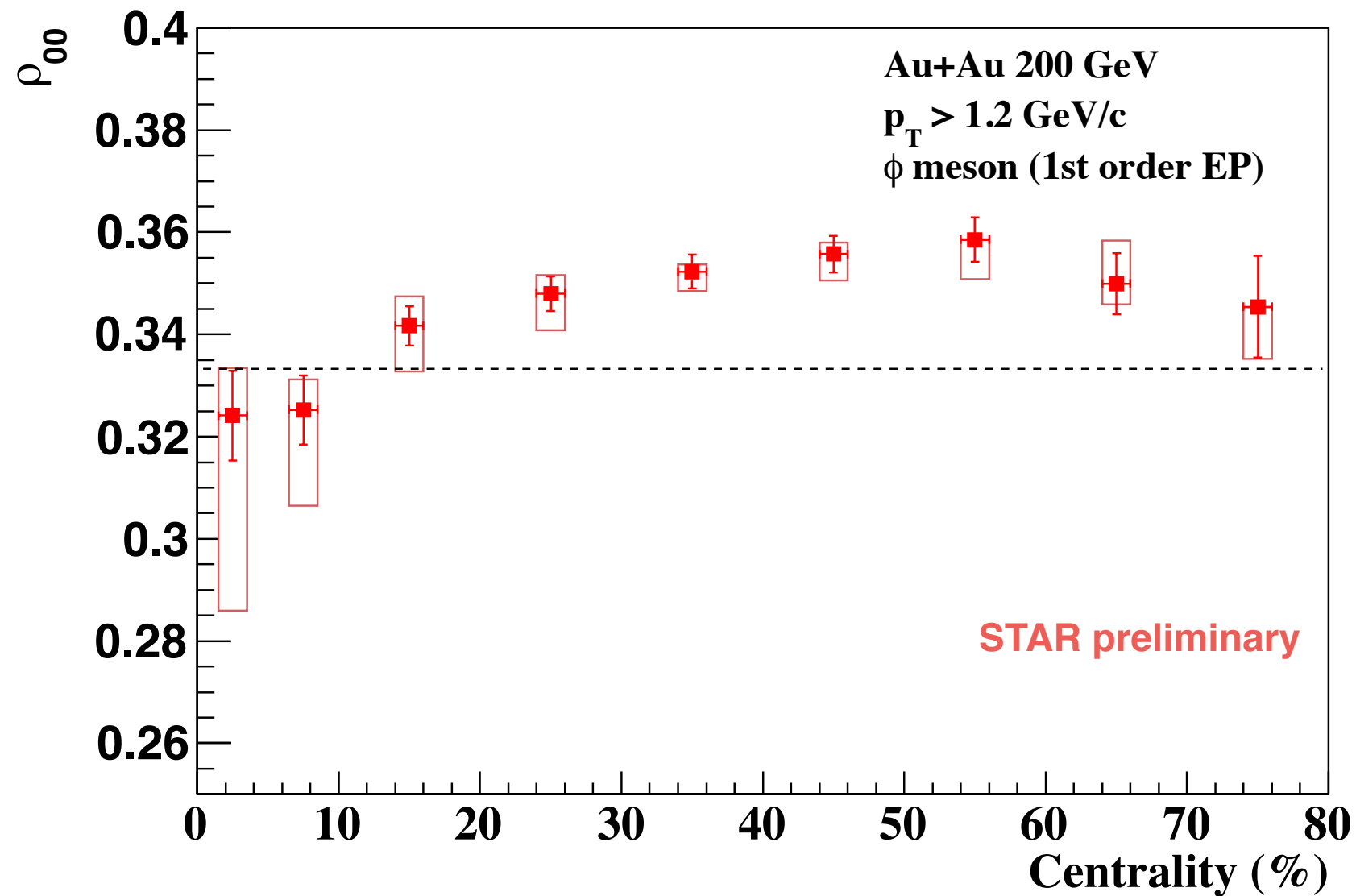
# De-correlation results



- The de-correlation between 1st and 2nd-order events plane explains part of the difference.
- The remaining difference may be due to  $B \neq 0$  in the angular distribution (or other physics origin?):

$$\frac{dN}{d\cos\theta^* d\beta} \propto 1 + A\cos^2\theta^* + B\sin^2\theta^* \cos 2\beta + C\sin 2\theta^* \cos\beta$$

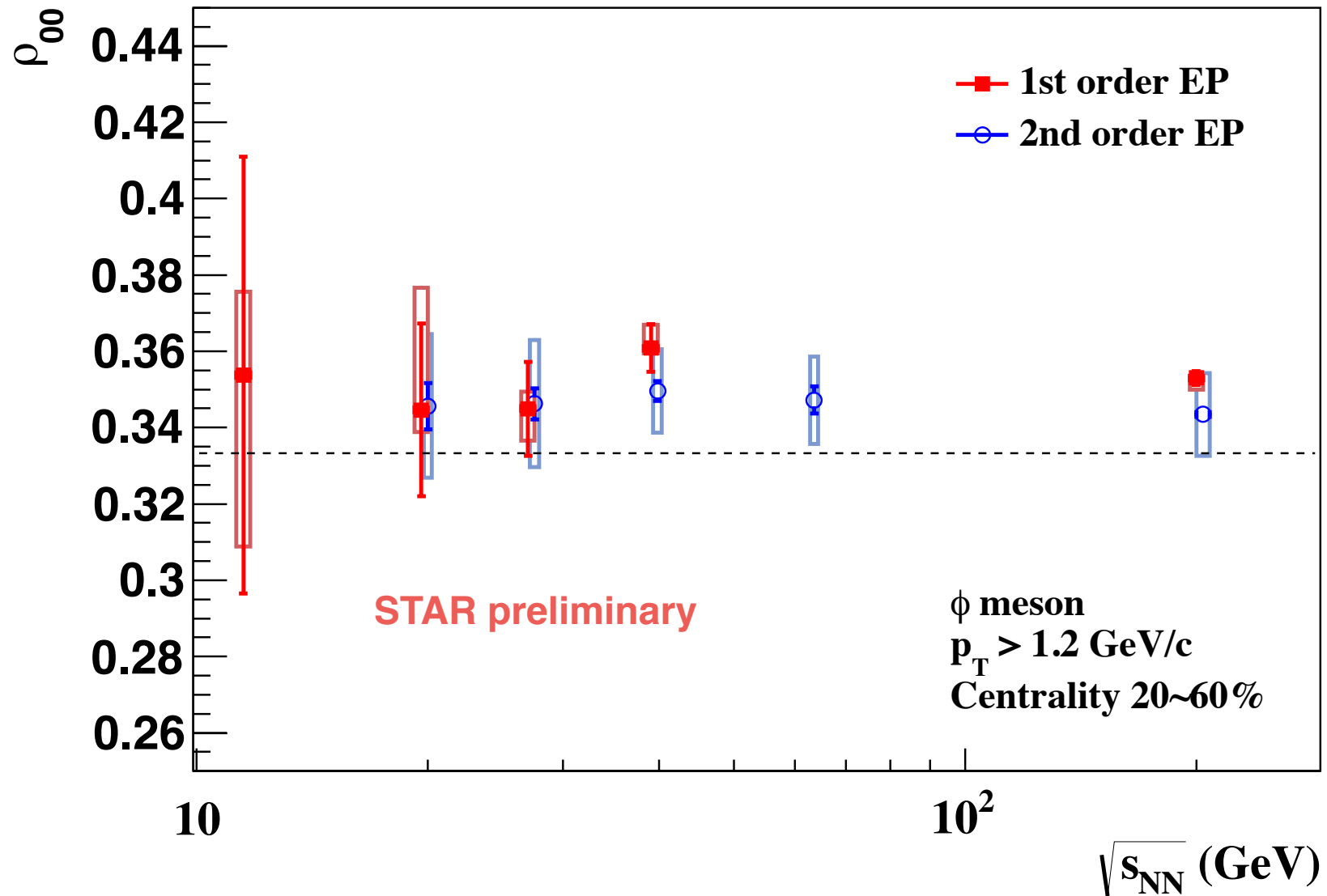
# $\rho_{00}$ vs. centrality



- $\rho_{00}$  are around 1/3 at most central collisions.
- For non-central collisions,  $\rho_{00}$  are significantly higher than 1/3, supporting the fragmentation scenario?

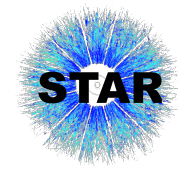


# $\rho_{00}$ vs. energy



- $\rho_{00}$  are significantly higher than  $1/3$  at 39 and 200 GeV.



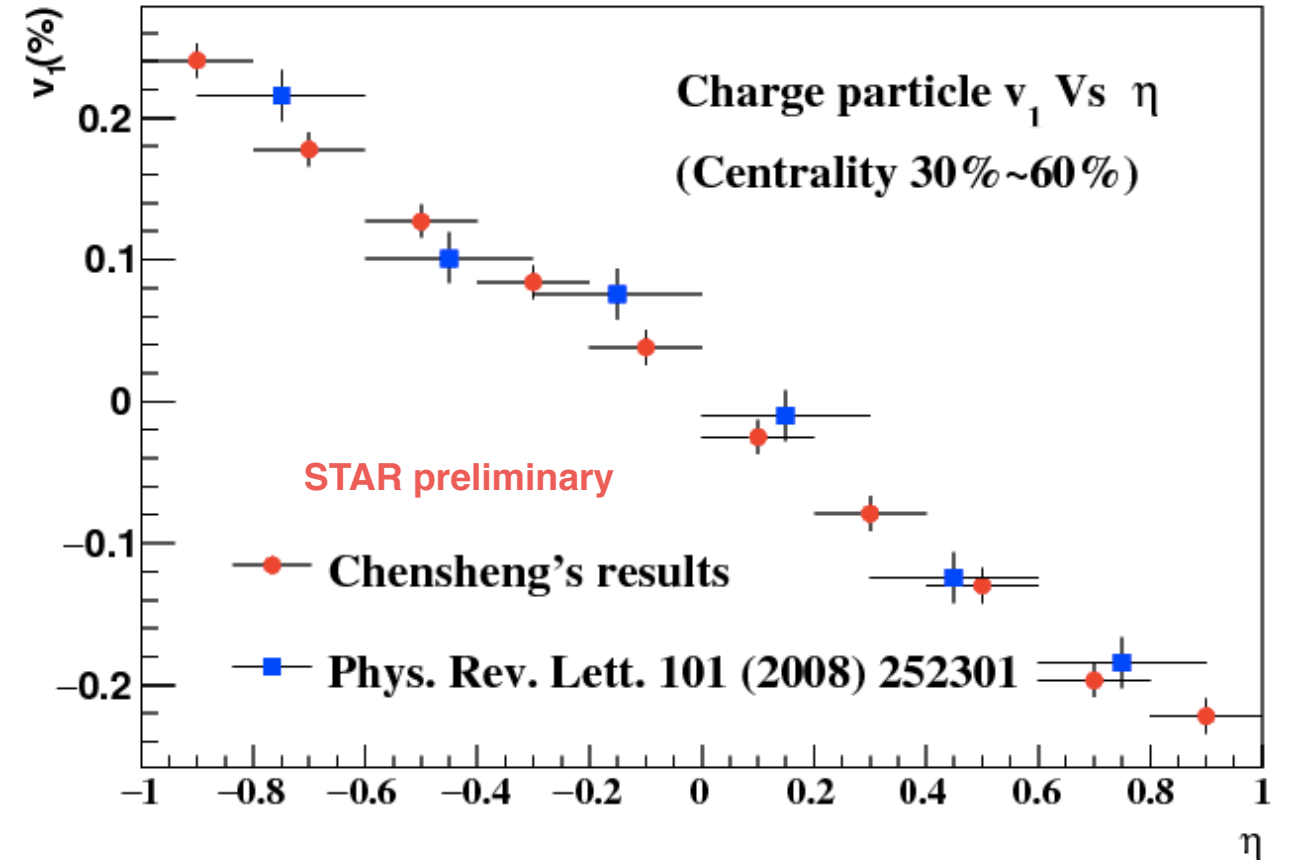
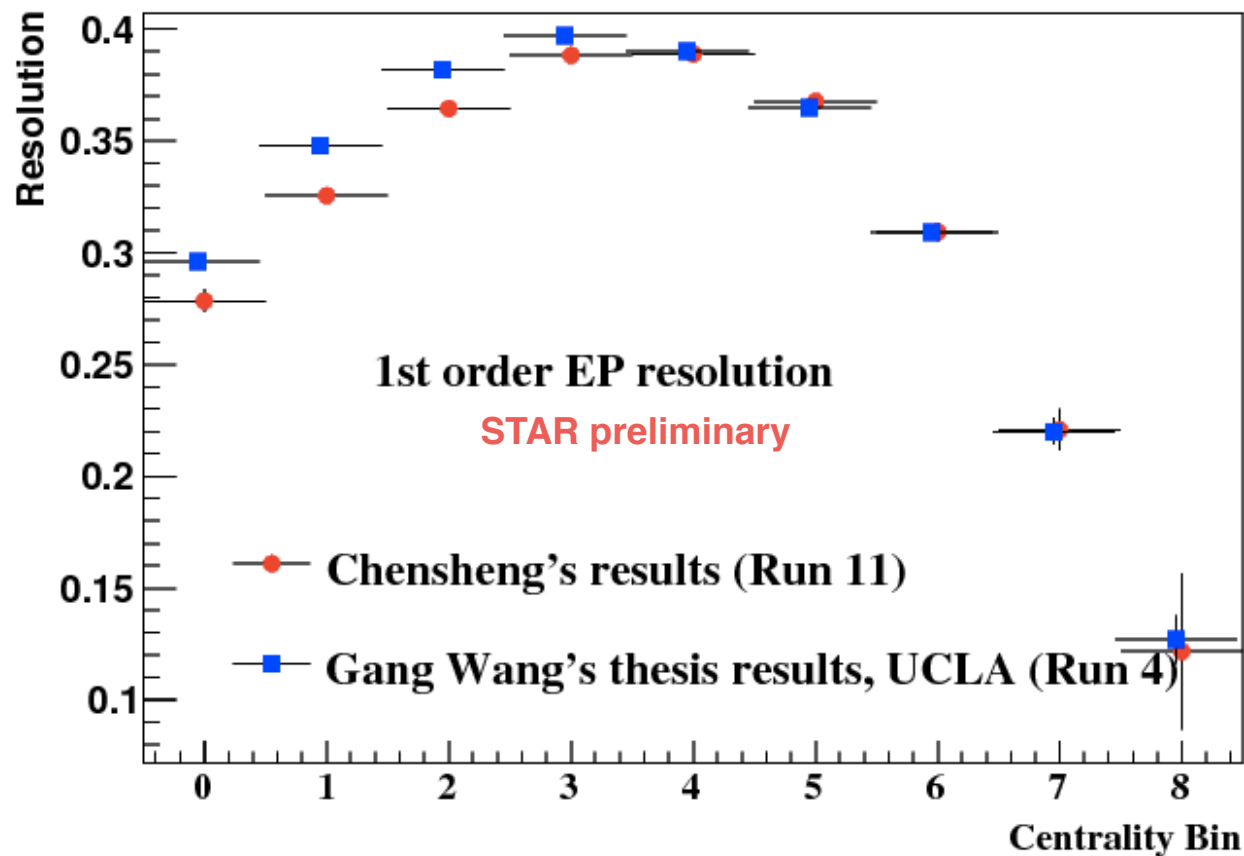


# Summary

- Non-trivial dependence of  $\rho_{00}$  as a function of  $p_T$  and centrality has been observed with 1st-order event plane. At 200 GeV the measured  $\rho_{00}$  is  $> 1/3$  at  $p_T \sim 1.5$  GeV/c in non-central collisions.
- For  $\rho_{00}$  integrated from  $p_T > 1.2$  GeV/c, the deviation from  $1/3$  is found to be significant at 39 and 200 GeV.
- This is the first time  $\rho_{00} > 1/3$  being observed in heavy ion collisions. Vorticity induced by initial global angular moments and particle production from quark fragmentation are possible sources that might contribute to the new observation.

# Backups

# Comparing charged particle $v_1$



1st order event plane resolution  
 Gang's thesis results : Run 4, Au-Au 200GeV  
 Our analysis: Run 11, Au-Au 200GeV

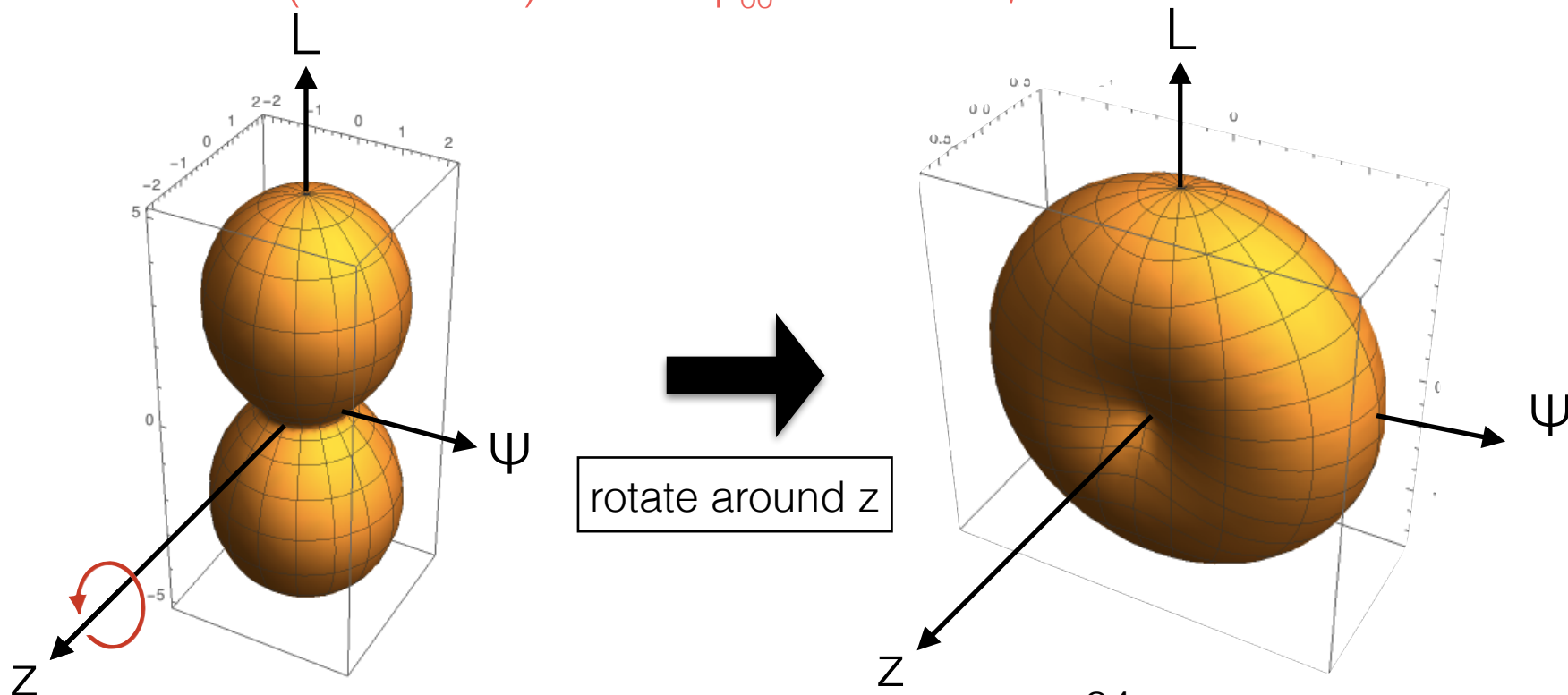
Charged particle  $v_1$  vs Eta

# What to expect when using random event plane

- Recall the formula for resolution correction:

$$\rho_{00}^{real} - \frac{1}{3} = \frac{4}{1+3R} \left( \rho_{00}^{obv} - \frac{1}{3} \right)$$

- For random event plane,  $\mathbf{L}$  is random in the transverse plane, and  $R=0$ . Only when the real  $\rho_{00}$  is  $1/3$ , the observed  $\rho_{00}$  from random event plane will become  $1/3$ . Putting it in simple words, **an irregular shape won't become a ball when rotated around a fixed axis** (z in this case). So the observed random plane result will be closer to  $\rho_{00} = 1/3$ , but hardly to be right at  $1/3$ . With the resolution correction formula ( $R=0$ ), we can still obtain the real  $\rho_{00}$ .
- Only when  $\mathbf{L}$  can take any direction in space (not confined to the transverse plane), it becomes truly random (3d-random) and the  $\rho_{00}$  becomes  $1/3$ .



Rotation around z axis will not necessarily make a round shape (strictly speaking, not make a flat distribution in  $\cos\theta^*$ )





# $\rho_{00}$ vs. $p_T$ (Au+Au 39 GeV)

