CPOD2022 - Workshop on Critical Point and Onset of Deconfinement

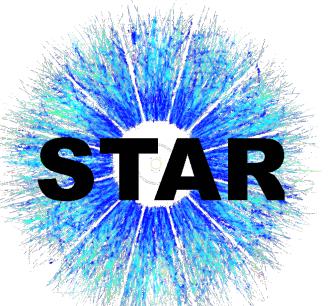
Light Hypernuclei Measurements in Au+Au Collisions from STAR

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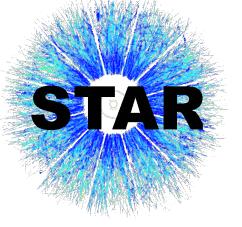




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Outline



- Introduction
- Hypernuclei measurements in STAR BES-II
 - Internal structure

Branching ratios, lifetimes

Production mechanism

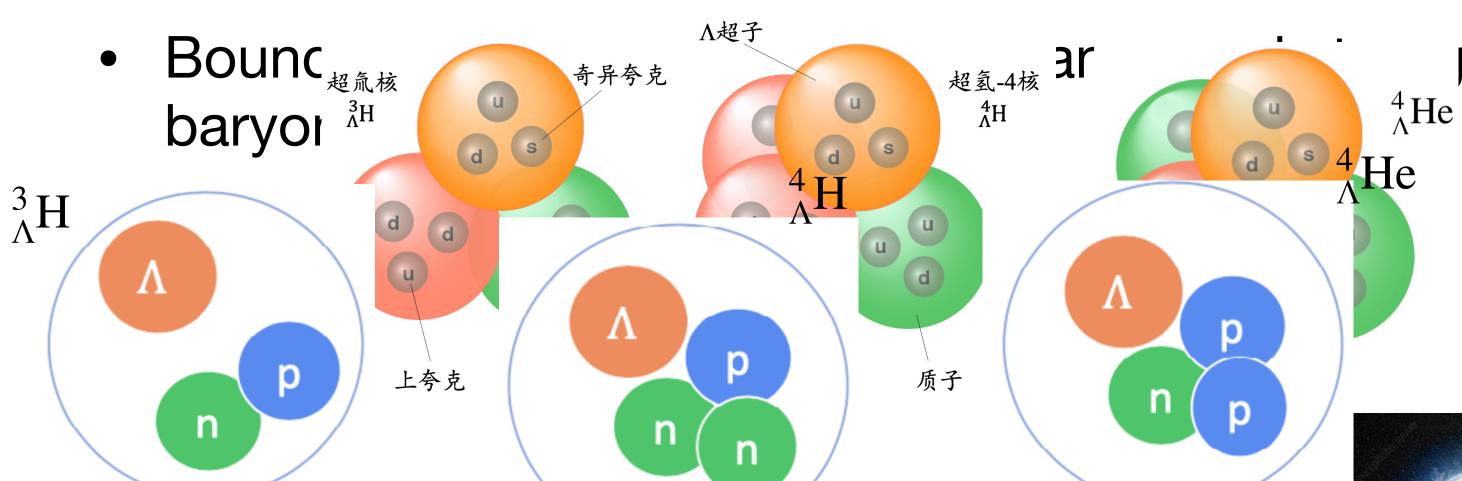
Yields, particle ratios, directed flow

Summary and outlook

Introduction: what and why



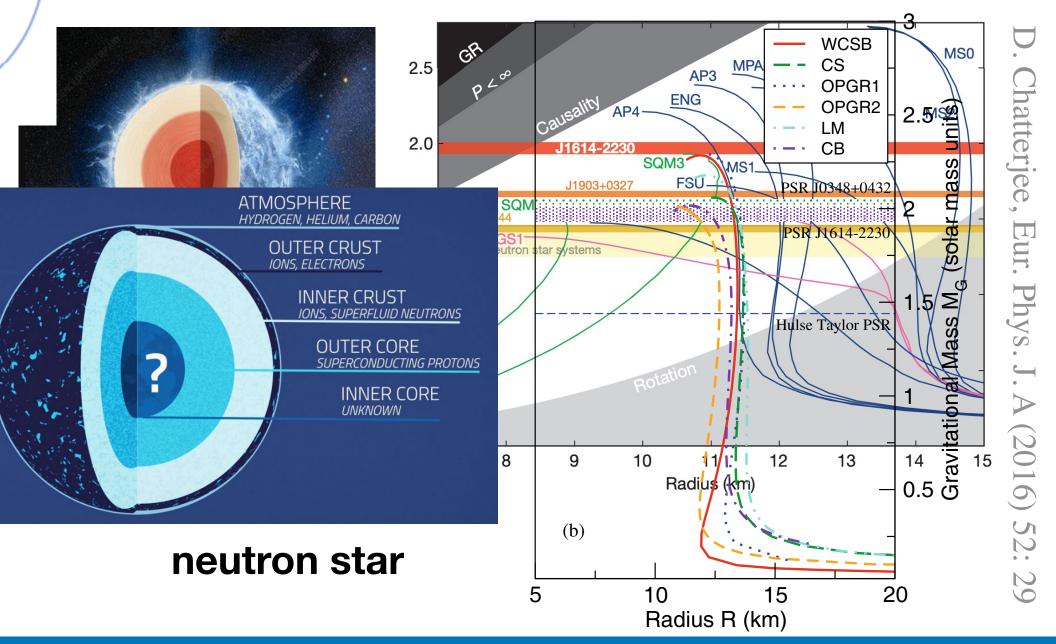
What are hypernuclei?



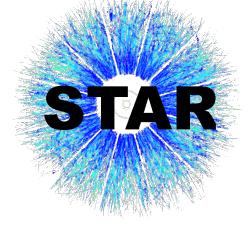


Marian Danysz (right) and Jerzy Pniewski (left) discovered hypernuclei in 1952

- Why hypernuclei?
 - Probe hyperon-nucleon (Y-N) interaction
 - Strangeness in high density nuclear matter
 - Equation-of-State (EoS) of neutron star

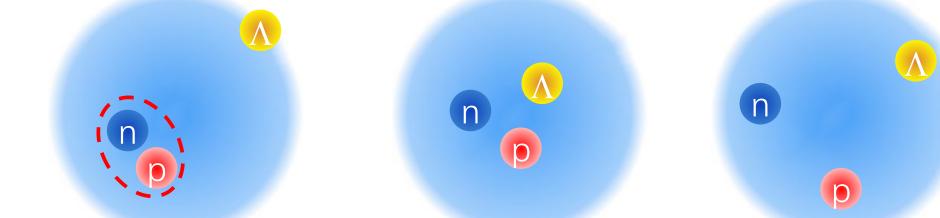


Introduction: how



• Experimentally, we can make measurements related to:

1. Internal structure



• Lifetime, binding energy, branching ratios etc.

Understanding hypernuclei structure can provide insights to the Y-N interaction

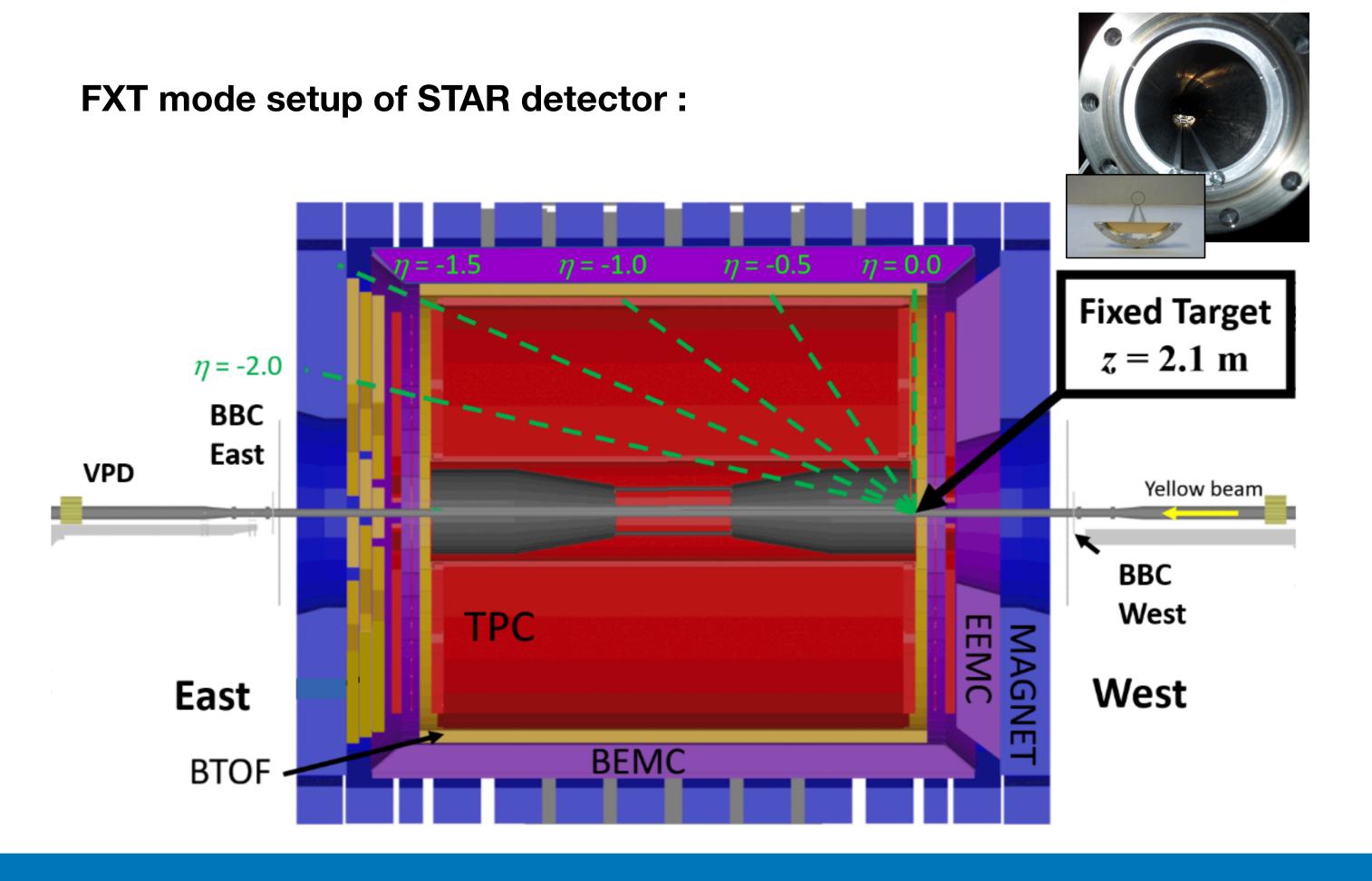
- 2. Production mechanism
 - Spectra, collectivity etc.

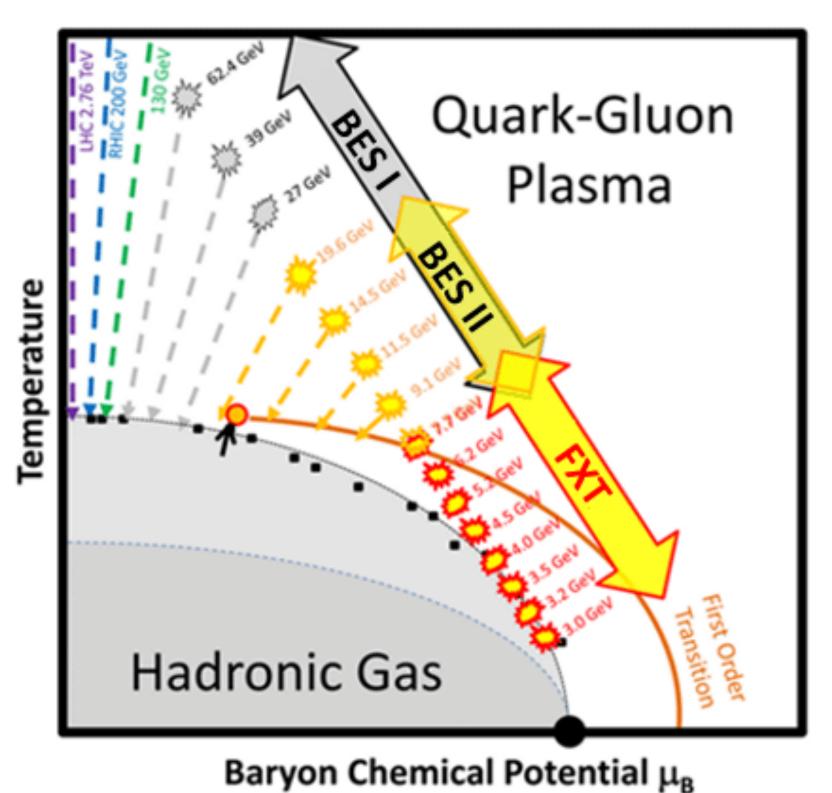
The formation of hypernuclei in violent heavy-ion collisions is not well understood

Introduction: RHIC BES program



• During the BES-II program, STAR utilized the fixed-target (FXT) setup, which extends the energy reach below $\sqrt{s_{NN}}$ = 7.7 GeV, down to 3.0 GeV

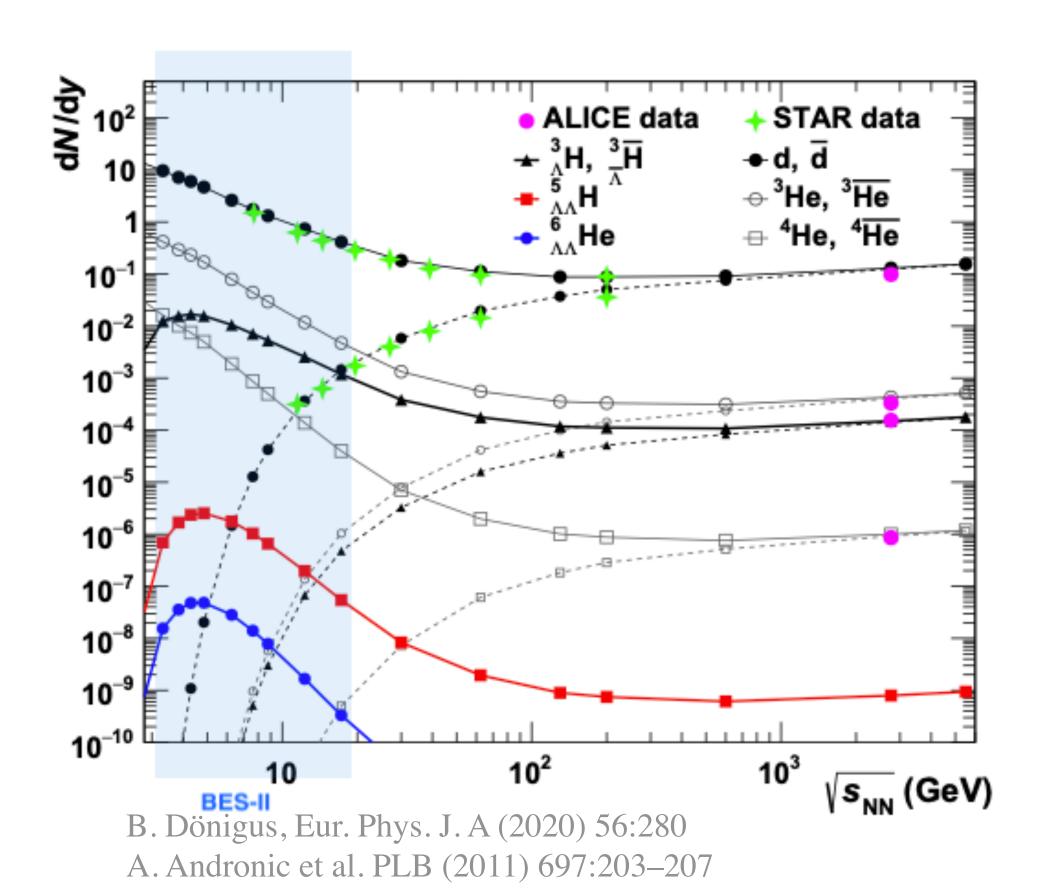




Introduction: hypernuclei and STAR BES-II

STAR

 Hypernuclei measurements are scarce in heavy-ion collision experiments



 At low beam energies, hypernuclei production is expected to be enhanced due to high baryon density

List of BES-II datasets:

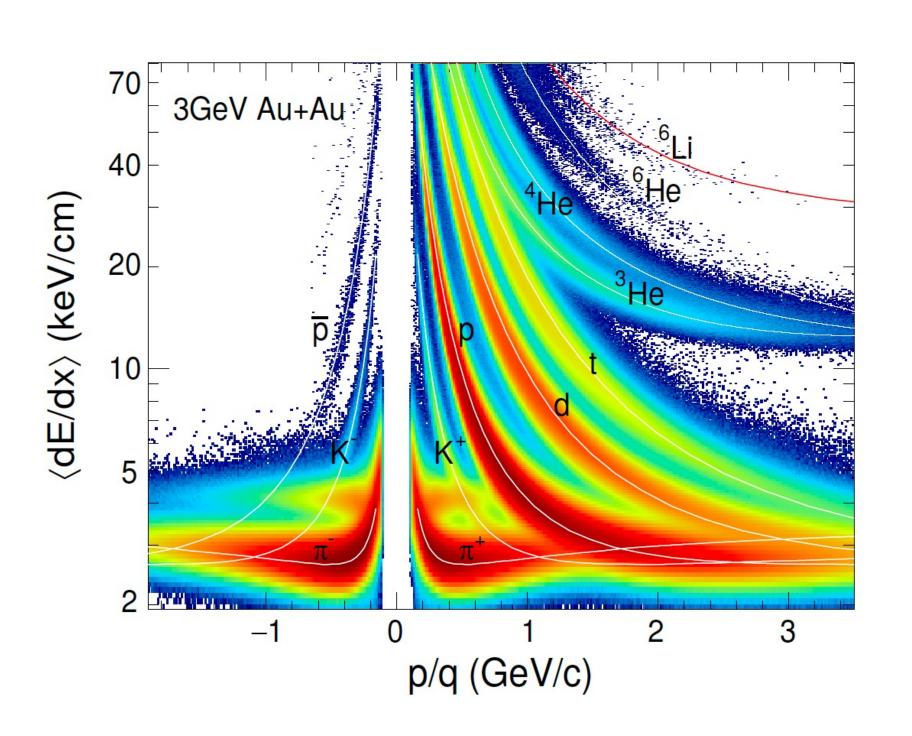
- Datasets with large statistics taken during BES-II
 - → BES-II is a great opportunity to study hypernuclei production

Year	$\sqrt{s_{NN}}$ [GeV]	Events
2018	27	555 M
	3.0	258 M
	<u>7.2</u>	155 M
2019	19.6	478 M
	14.6	324 M
	<u>3.9</u>	53 M
	<u>3.2</u>	201 M
	<u>7.7</u>	51 M
2020	11.5	235 M
	<u>7.7</u>	113 M
	<u>4.5</u>	108 M
	<u>6.2</u>	118 M
	<u>5.2</u>	103 M
	<u>3.9</u>	117 M
	<u>3.5</u>	116 M
	9.2	162 M
	<u>7.2</u>	317 M
2021	7.7	101 M
	<u>3.0</u>	2103 M
	<u>9.2</u>	54 M
	<u>11.5</u>	52 M
	<u>13.7</u>	51 M
	17.3	256 M
	<u>7.2</u>	89 M

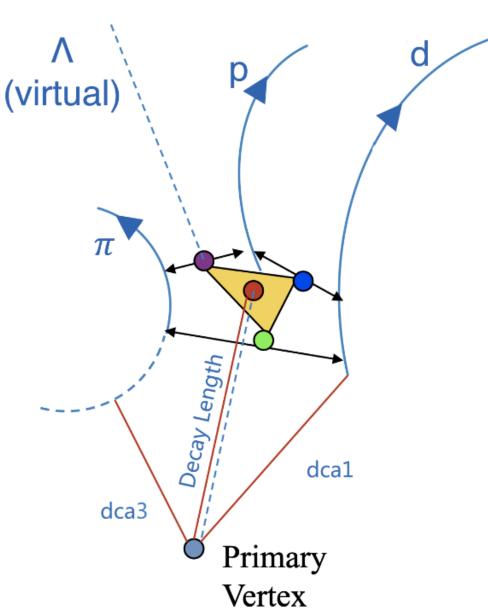


fication and hypernuclei reconstruction

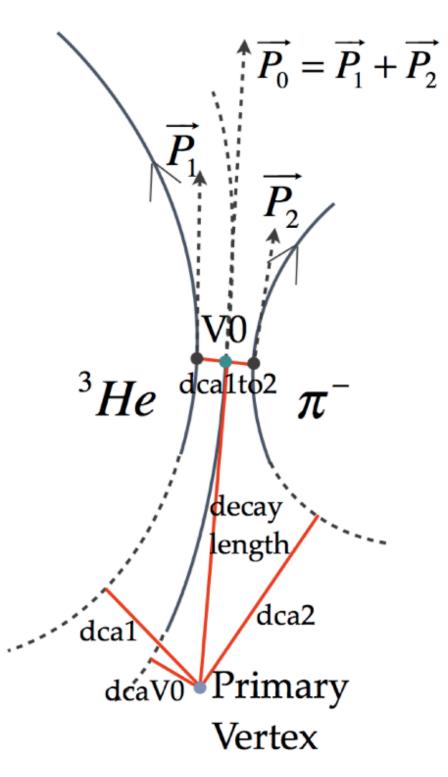




$^{4}\text{He} \rightarrow ^{3}\text{He} + p + \pi^{-}$:







- Particle identification from energy loss measurement using TPC
- KF particle package^[1] is used for signal reconstruction
- Hypernuclei reconstructed via their weak decay channels:

$$^{3}_{\Lambda}H \rightarrow ^{3}He + \pi^{-}$$
 $^{3}_{\Lambda}H \rightarrow d + p + \pi^{-}$ $^{4}_{\Lambda}H \rightarrow ^{4}He + \pi^{-}$ $^{4}_{\Lambda}He \rightarrow ^{3}He + p + \pi^{-}$

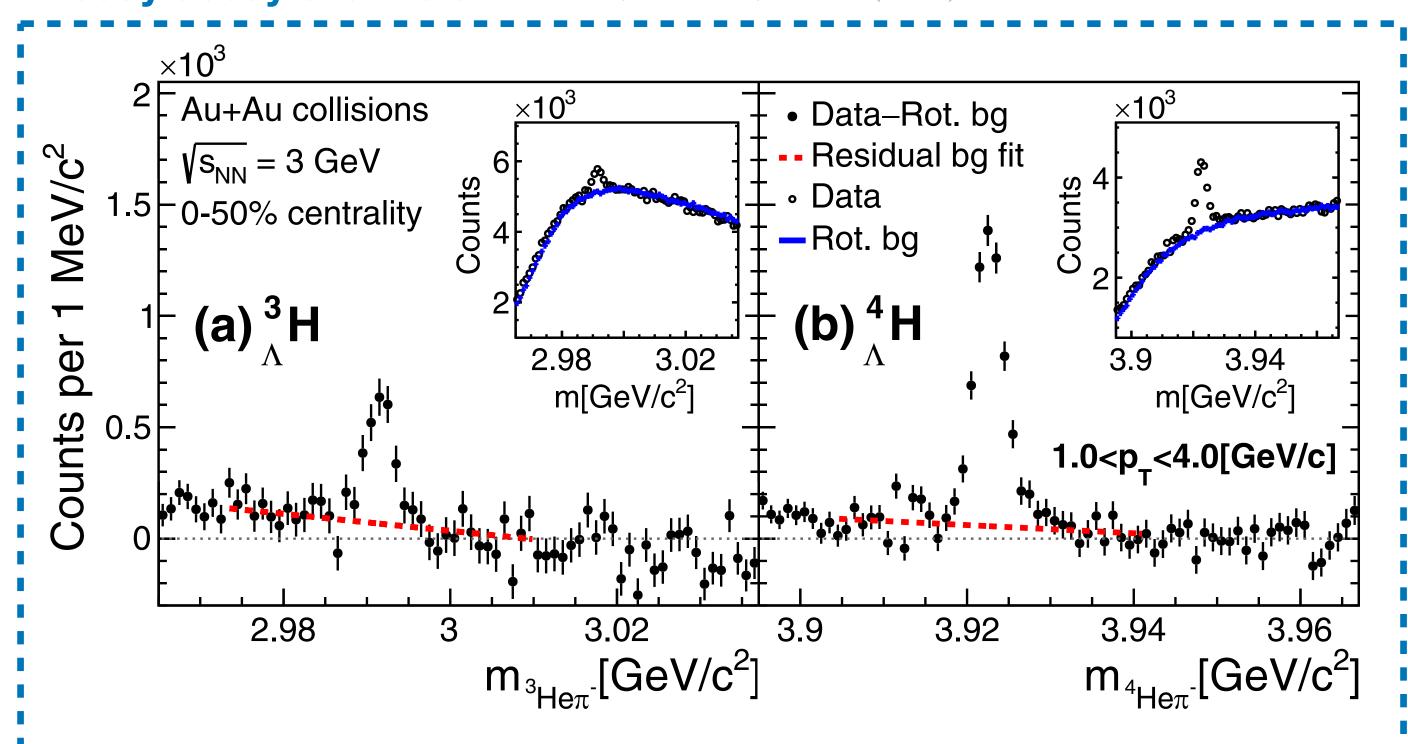
[1]Zyzak M, Kisel I, Senger P. Online selection of short-lived particles on many-core computer architectures in the CBM experiment at FAIR[R]. Collaboration FAIR: CBM, 2016.

Hypernuclei signal reconstruction

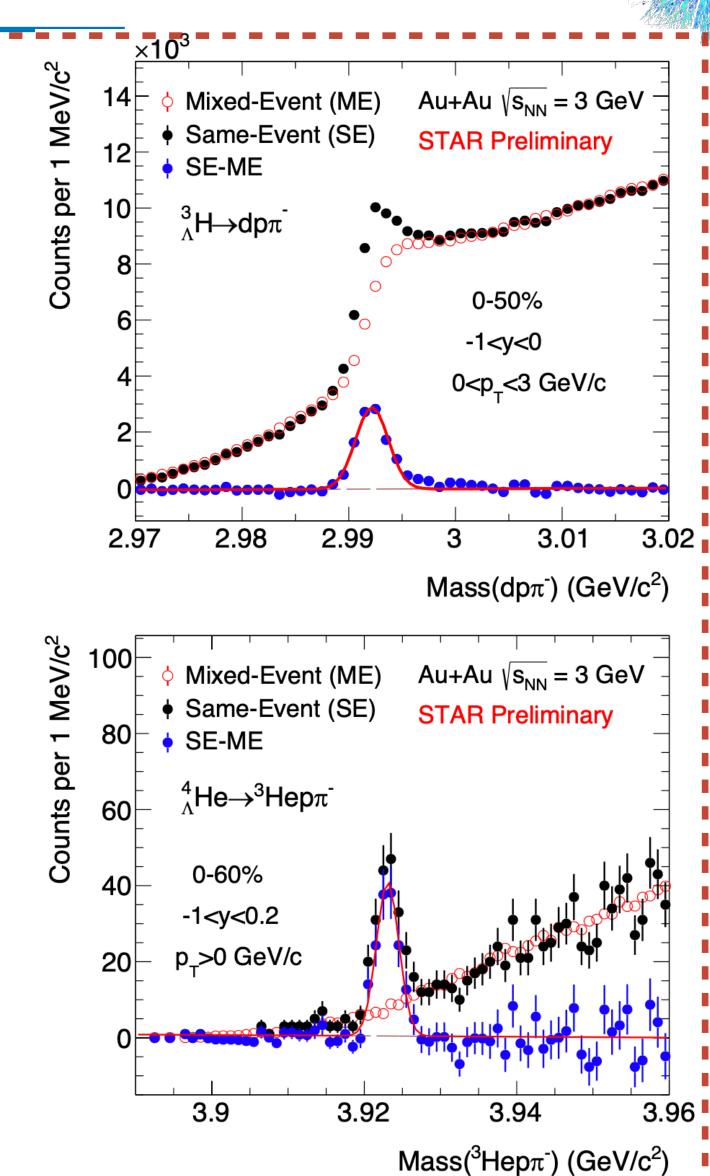


2-body decay channels: STAR, PRL 128, 202301(2022)

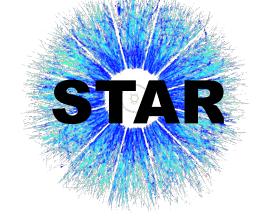
3-body decay channels:



- Combinatorial background estimated via:
 - Rotating pion tracks for 2-body decay channels
 - Event mixing for 3-body decay channels

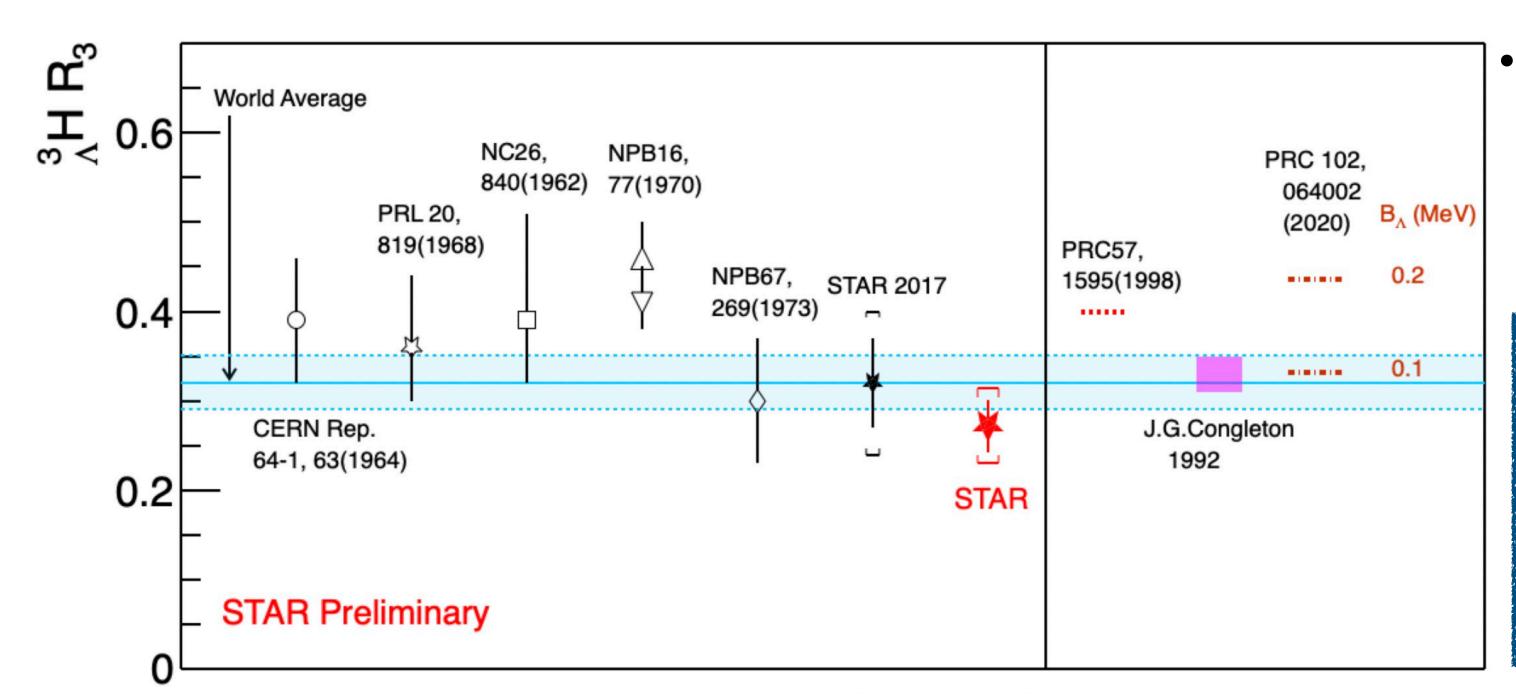


$^3_\Lambda H$ branching ratio R_3



Relative branching ratio:
$$R_3 = \frac{B \cdot R \cdot ({}^3_{\Lambda}H \rightarrow {}^3He\pi^-)}{B \cdot R \cdot ({}^3_{\Lambda}H \rightarrow {}^3He\pi^-) + B \cdot R \cdot ({}^3_{\Lambda}H \rightarrow dp\pi^-)}$$

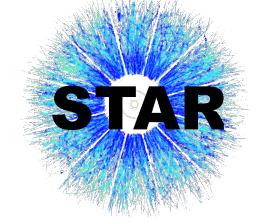
F. Hildenbrand et al. PRC 102, 064002 (2020)



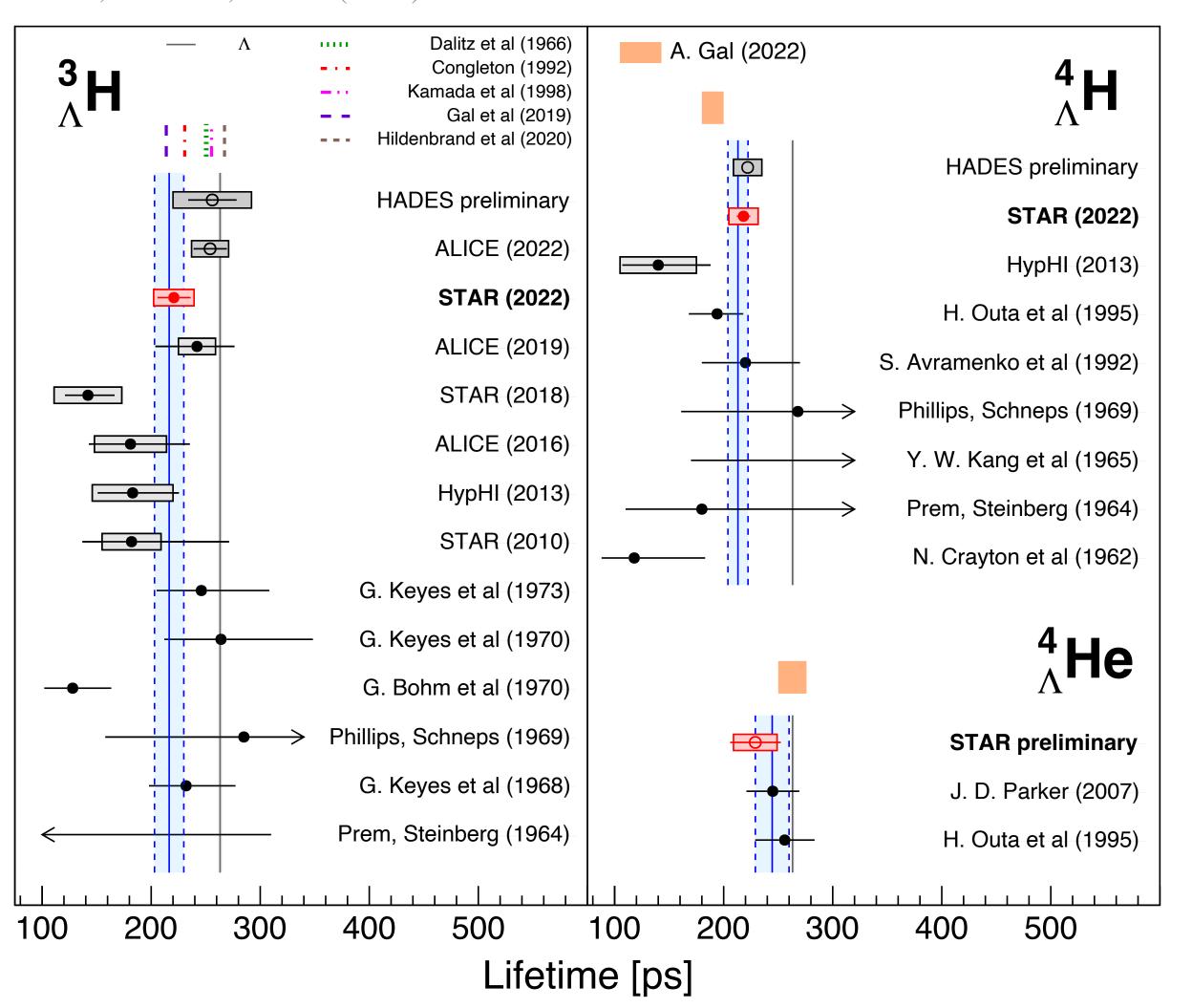
- Recent calculation shows that R_3 may be sensitive to the binding energy (B_Λ) of $^3_\Lambda H$
 - $B_{\Lambda} \rightarrow$ provide constraints to Y-N interaction
- Using $\sqrt{s_{NN}}$ = 3.0 GeV data:
 - $R_3 = 0.272 \pm 0.030(stat.) \pm 0.042(syst.)$
 - Updated world average R_3 (0.32 \pm 0.03) is consistent with theoretical models assuming B_Λ ~ 0.1 MeV

- ullet Improved precision on R_3
 - Stronger constraints on absolute B.R.s and hypertriton internal structure models

$^3_{\Lambda}$ H, $^4_{\Lambda}$ H and $^4_{\Lambda}$ He lifetimes



STAR, PRL 128, 202301(2022)



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Using \sqrt{s_{NN}} = 3.0 GeV and 7.2 GeV datasets:

^3_{\Lambda}H: \tau = 221 \pm 15 (\text{stat.}) \pm 19 (\text{syst.}) [\text{ps}]

^4_{\Lambda}H: \tau = 218 \pm 6 (\text{stat.}) \pm 13 (\text{syst.}) [\text{ps}]

^4_{\Lambda}He: \tau = 229 \pm 23 (\text{stat.}) \pm 20 (\text{syst.}) [\text{ps}]
```

- Lifetimes of light hypernuclei ${}^3_\Lambda H$, ${}^4_\Lambda H$ and ${}^4_\Lambda He$ are shorter than that of free Λ (with 1.8 σ , 3.0 σ , 1.1 σ respectively)
- Consistent with former measurements (within 2.5 σ for $^3_{\Lambda}$ H, $^4_{\Lambda}$ H)
- $au_{^3\mathrm{H}}$: consistent with calculation including pion FSI $^{[1]}$ and calculation with Λd 2-body picture $^{[2]}$ within 1 σ
- $au_{^4\mathrm{H}}$ and $au_{^4\mathrm{He}}$: consistent with expectations from isospin rule

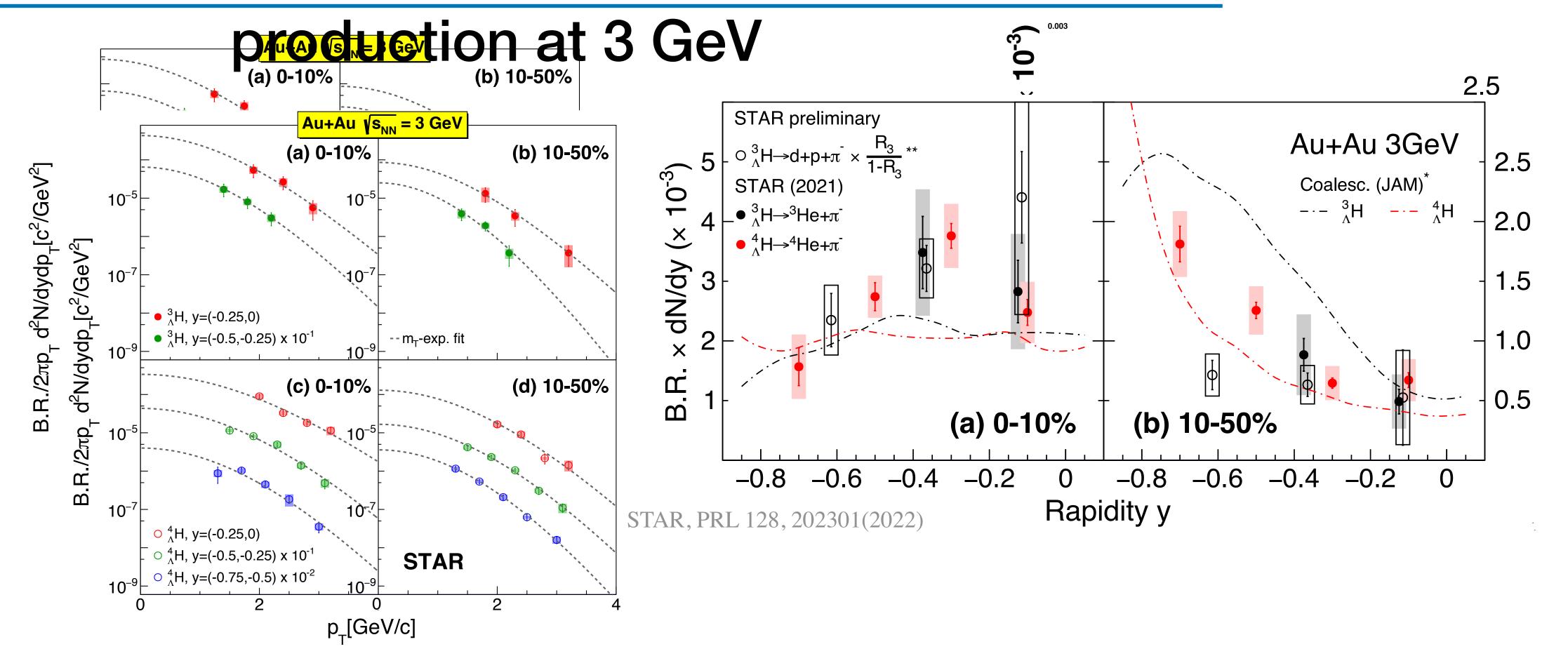
 $^3_{\Lambda}$ H, $^4_{\Lambda}$ H results with improved precision

→ Provide tighter constraints on models.

[1]A. Gal and H. Garcilazo, PLB 791, 48 (2019) [2]J.G. Congleton, J. Phys. G 18, 339 (1992)

Hypernuclei3production at 3 GeV

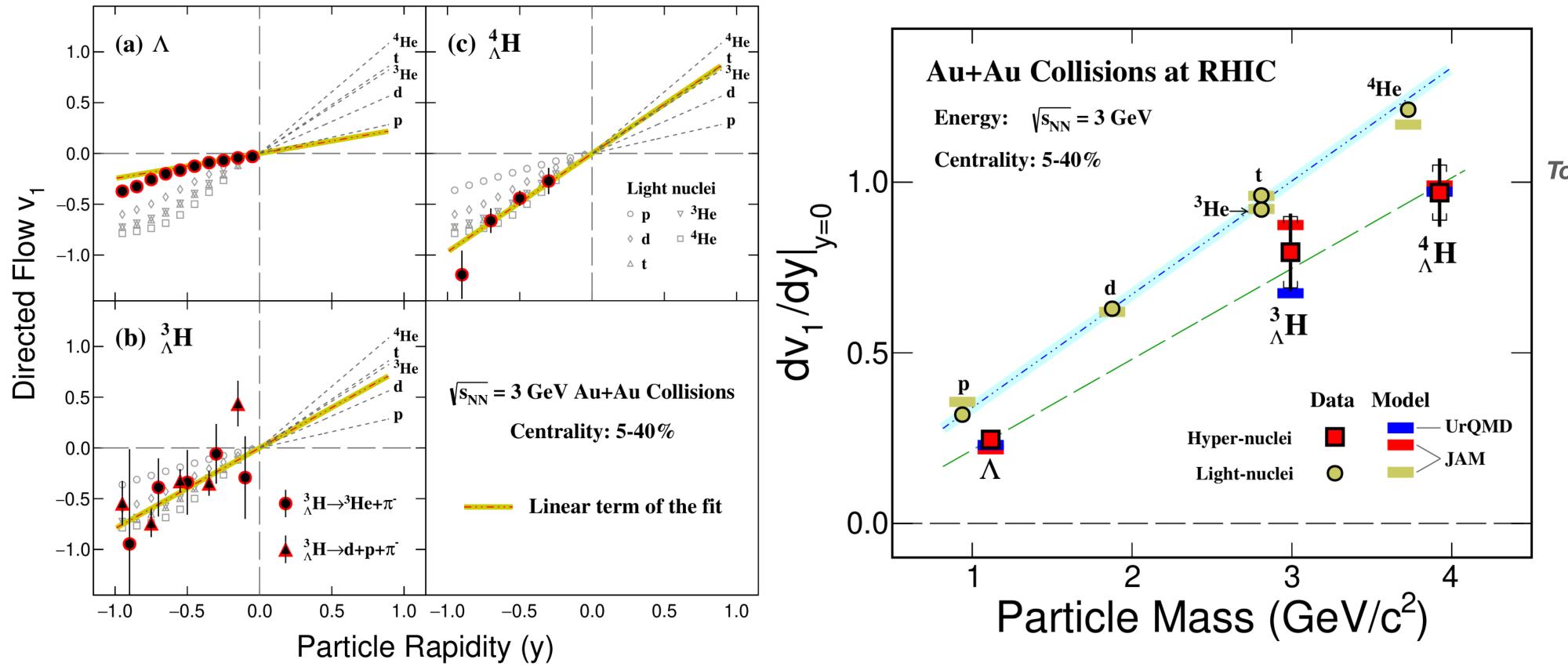




- Different trends in the $^4_\Lambda H$ rapidity distribution in central (0-10%) and mid-central (10-50%) collisions at $\sqrt{s_{NN}}$ = 3.0 GeV
 - ullet Transport model (JAM) with coalescence approximately reproduces trends of $^4_\Lambda {
 m H}$ rapidity distributions seen in data

$^3_{\Lambda} H$ and $^4_{\Lambda} H$ directed flow at 3 GeV

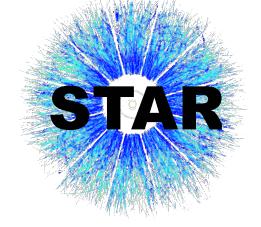


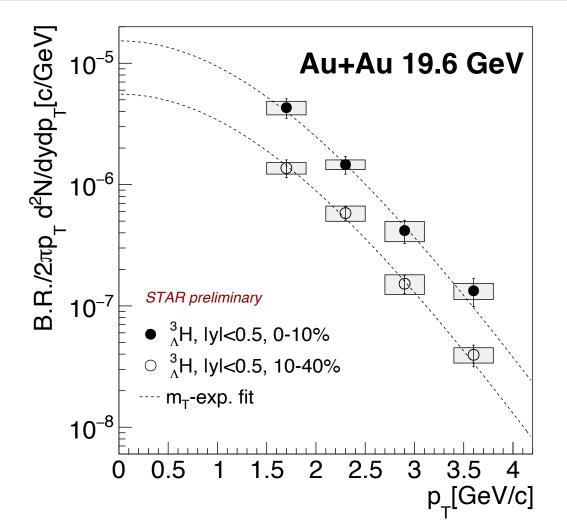


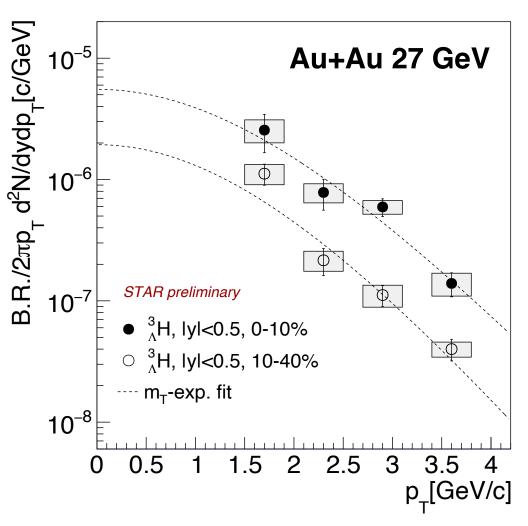
To be submitted to arXiv soon

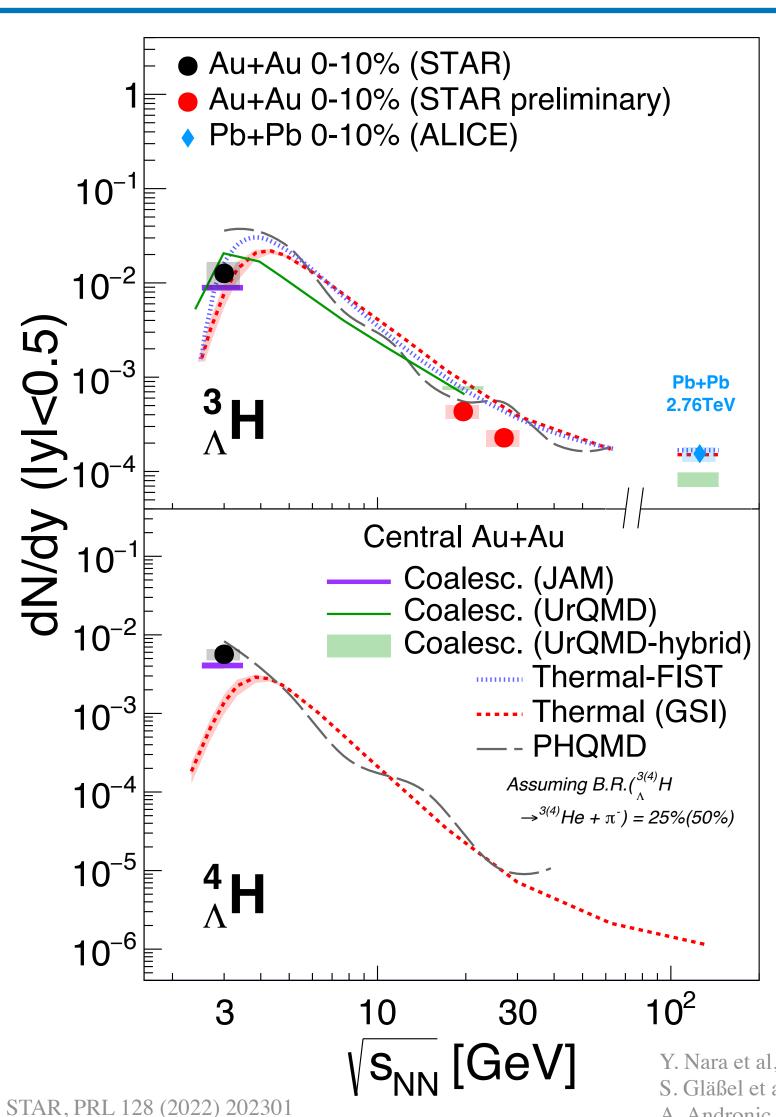
- First measurements of $^3_\Lambda H$ and $^4_\Lambda H$ directed flow (v₁) in 5-40% central Au+Au collisions at 3 GeV
- v_1 slopes of ${}^3_{\Lambda}H$ and ${}^4_{\Lambda}H$ follow mass number scaling.
 - → Imply coalescence process to be the dominant formation mechanism for hypernuclei in heavy-ion collisions

Energy dependence of hypernuclei production in heavy-ion collisions









ALICE, PLB 754 (2016) 360

- $^3_\Lambda H$ yield at mid-rapidity increases from 2.76 TeV to 3 GeV
 - Driven by increase in baryon density at low energies
- Thermal(GSI), Coalescence(UrQMD), Thermal-FIST and PHQMD reproduce the trend

For Au+Au @ 3 GeV

- Coalescence(JAM) with tuned coalescence parameters can describe data
- PHQMD describes ${}^4_{\Lambda}H$, but slightly overestimates ${}^3_{\Lambda}H$

Provide first constraints for hypernuclei production models in the high-baryon-density region

Y. Nara et al, PRC 61 (1999) 024901 (JAM)

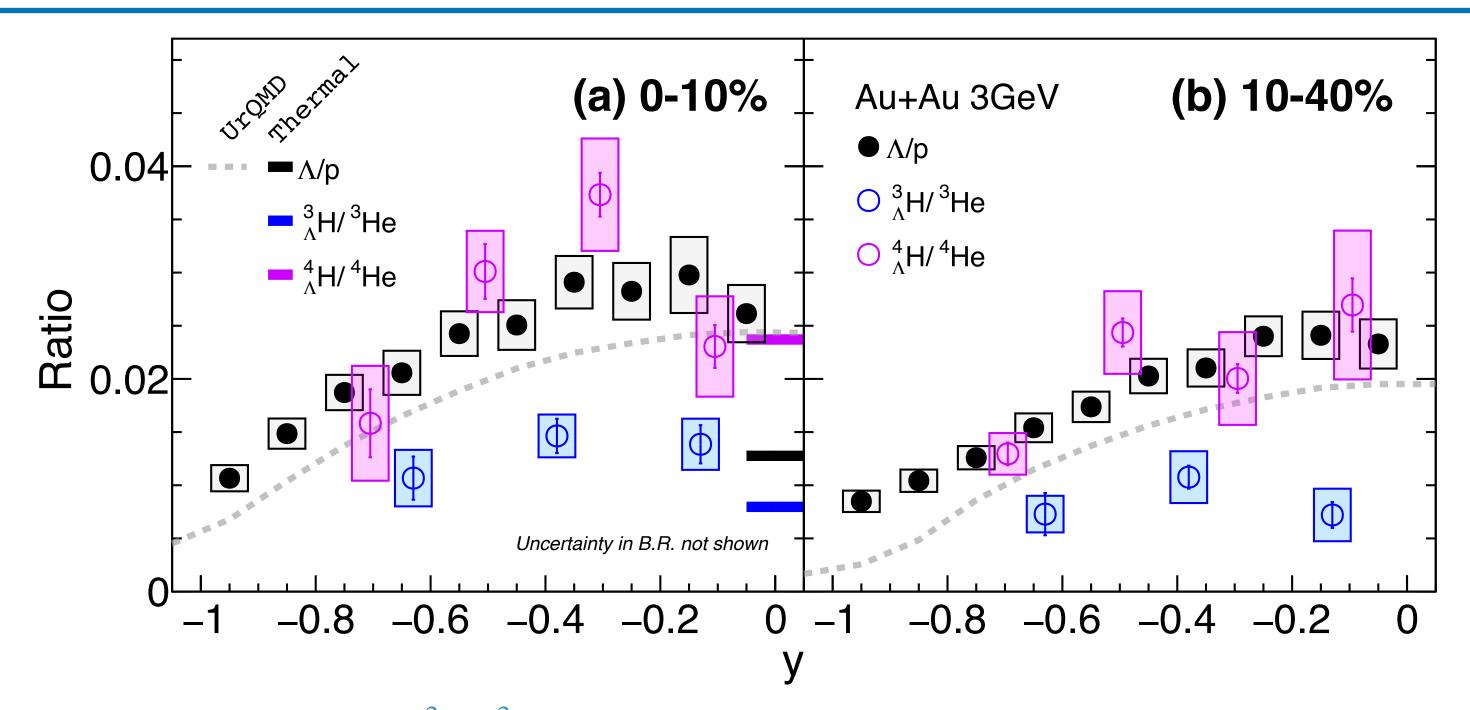
S. Gläßel et al, arXiv: 2106,14839 (PHQMD)

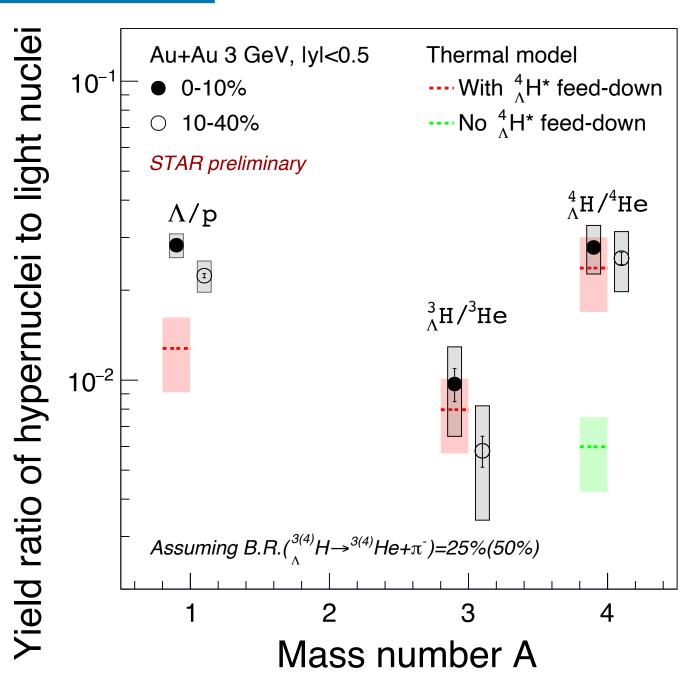
A. Andronic et al, PLB 697 (2011) 203 (Thermal (GSI))

T. Reichert, J. Steinheimer et al, arXiv:2210.11876(2022) (UrQMD, Thermal-FIST)

Hyper-toughthenees ratios



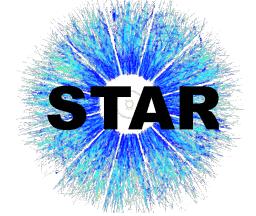




- Suppression of $^3_{\Lambda} \text{H}/^3 \text{He}$ yield ratios compared to that of Λ/p
 - Observed at both 0-10% and 10-40% centrality in Au+Au collisions at 3 GeV.
- The $^4_{\Lambda} \text{H}/^4 \text{He}$ yield ratios are comparable to that of Λ/p
- Thermal model calculations including excited $^4_\Lambda H^*$ feed-down show a similar trend
 - Feed-down from excited state enhances $^4_\Lambda H$ production
 - Support creation of excited A=4 hypernuclei in heavy-ion collisions

A. Andronic et al, PLB 697 (2011) 203 (Thermal model)

S_{3,4} at 3 GeV



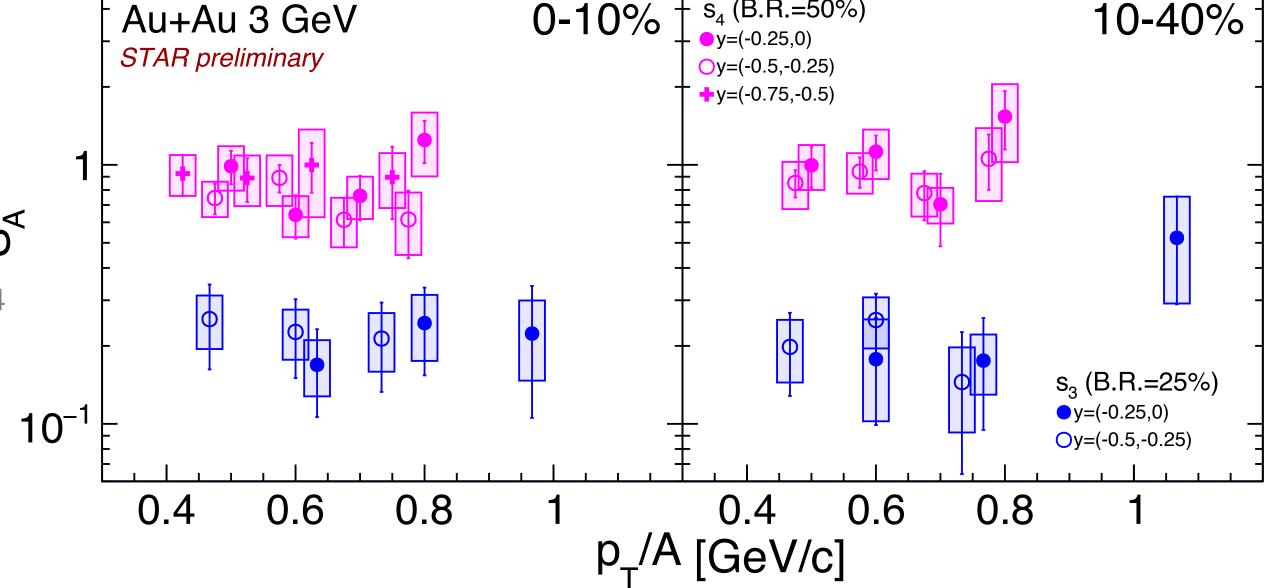
ullet Strangeness population factor S_A

 Relative suppression of hypernuclei production compared to light nuclei production

$$S_{A} = \frac{{}^{A}_{\Lambda}H}{{}^{A}_{He} \times \frac{\Lambda}{p}} = \frac{B_{A}({}^{A}_{\Lambda}H)(p_{T})}{B_{A}({}^{A}_{He})(p_{T})}$$

$$S.Zhang, PLB 684(2010)224$$

- B_A: Coalescence parameters
- Expect ~1 if no suppression



$$S_3$$
< 1 \rightarrow relative suppression of $^3_{\Lambda}H$ to 3He

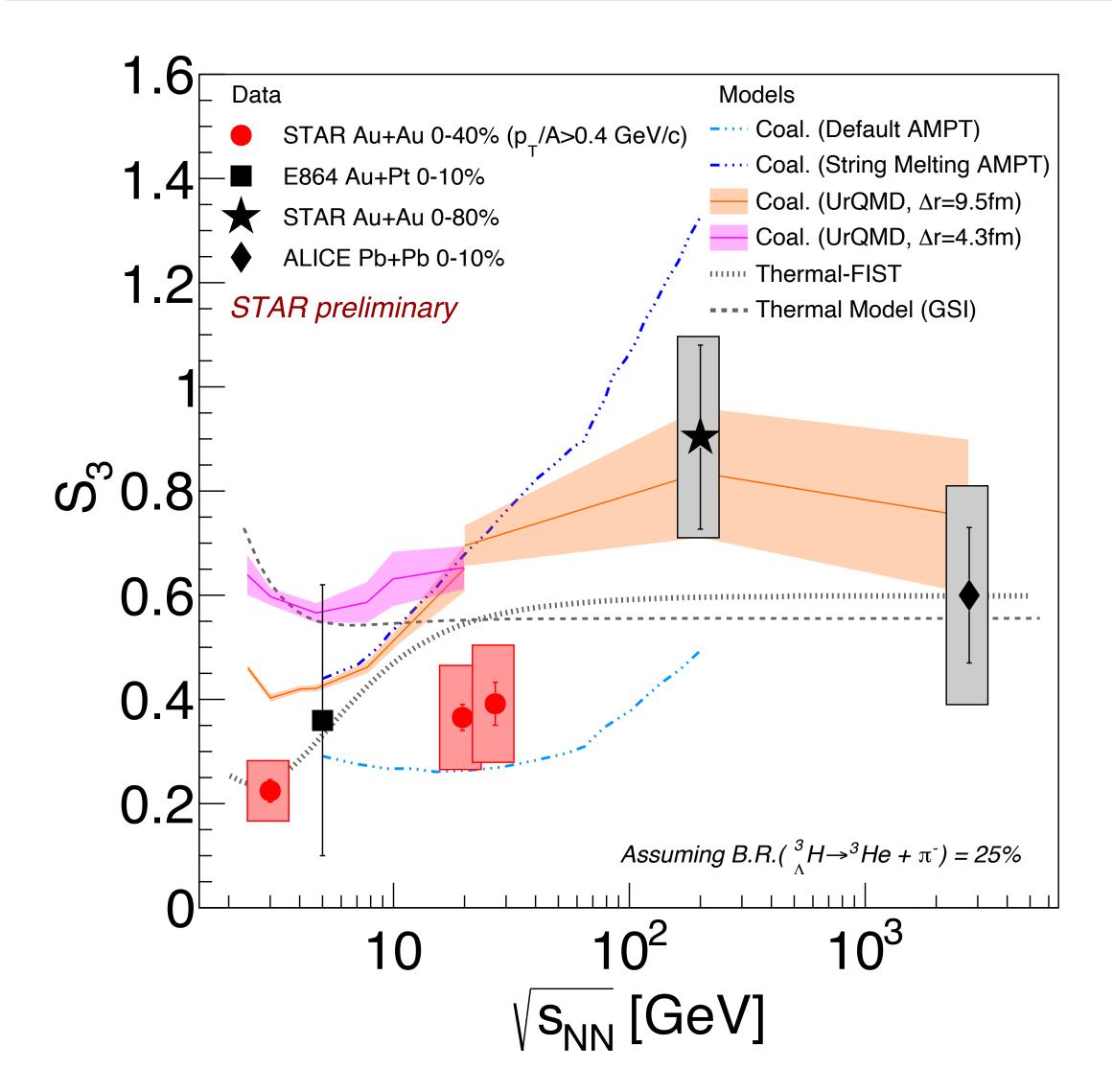
 $S_4 > S_3 \rightarrow \text{enhanced } ^4_\Lambda H$ production due to feed-down from excited state

No obvious kinematic and centrality dependence of $\boldsymbol{S}_{\boldsymbol{A}}$ is observed at 3 GeV.

ightarrow Coalescence parameter B_A of $^A_\Lambda H$ and $^A_H H$ e follows similar tendency versus p_T , rapidity and centrality.

Energy dependence of S₃





STAR, Science 328 (2010) 58 ALICE, PLB 754 (2016) 360 E864, PRC 70 (2004) 024902 NA49, J.Phys.Conf.Ser.110(2008)032010

- A. Andronic et al, PLB 697 (2011) 203 (Thermal (GSI))
- S. Zhang, PLB 684(2010)224 (Coal.+AMPT)
- T. Reichert, J. Steinheimer et al, arXiv:2210.11876(2022) (UrQMD, Thermal-FIST)

- Data shows a hint of an increasing trend from $\sqrt{s_{NN}}$ = 3.0 GeV to 2.76 TeV
- For coalescence models, the energy dependence is sensitive to the source radius (Δr)
- Thermal-FIST describes the \mathbf{S}_3 data reasonably well

Summary



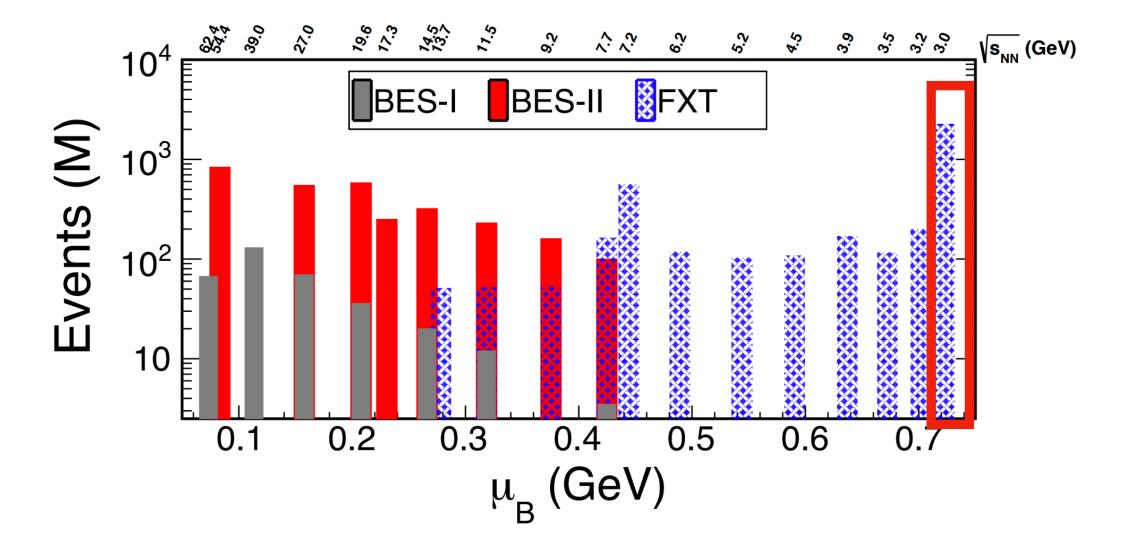
Presented measurements on hypernuclei production in the high-baryon-density region with high statistical precision using STAR data

- Hypernuclei structure
 - ${}^3_{\Lambda}H$, ${}^4_{\Lambda}H$ lifetimes and R_3 of ${}^3_{\Lambda}H$ measured with improved precision
 - Strong constraints on hyperon-nucleon interaction models
- Hypernuclei production in heavy-ion collisions
 - ${}^3_{\Lambda} H$, ${}^4_{\Lambda} H$ production yields at 3.0, 19.6 and 27 GeV
 - Coalescence models approximately describe the trends of $^4_\Lambda H$ rapidity distribution
 - S_3 and S_4 show weak centrality/kinematic dependence
 - Energy dependence of $^3_\Lambda H$, $^4_\Lambda H$ yields and S_3 compared with models are shown
 - Provide constraints to hypernuclei production models
 - $^3_{\Lambda}{
 m H}$ and $^4_{\Lambda}{
 m H}$ collectivity ${
 m v}_1$
 - v₁ slopes follow mass number scaling -> Support coalescence picture

Outlook

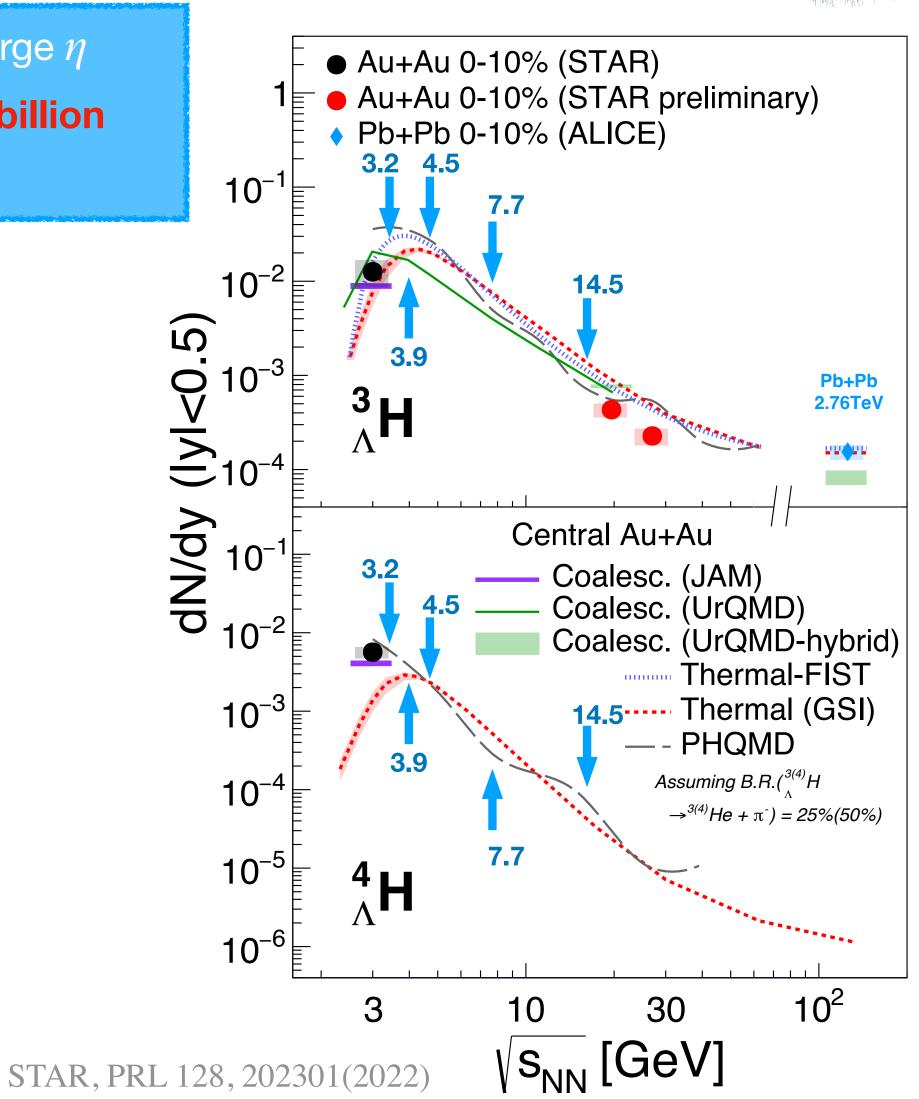


- 1. iTPC and eToF fully installed in 2019 \rightarrow improve η acceptance and PID at large η
- 2. High statistics data in STAR BES-II $\sqrt{s_{NN}}$ = 3.0 54.4 GeV, especially the **2 billion** events collected at 3 GeV in 2021 \rightarrow larger statistics, higher precision

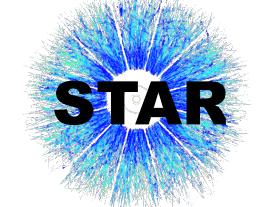


- Precision measurements on hypernuclei properties
- Energy dependence study of hypernuclei yields
- Search for double Λ hypernuclei

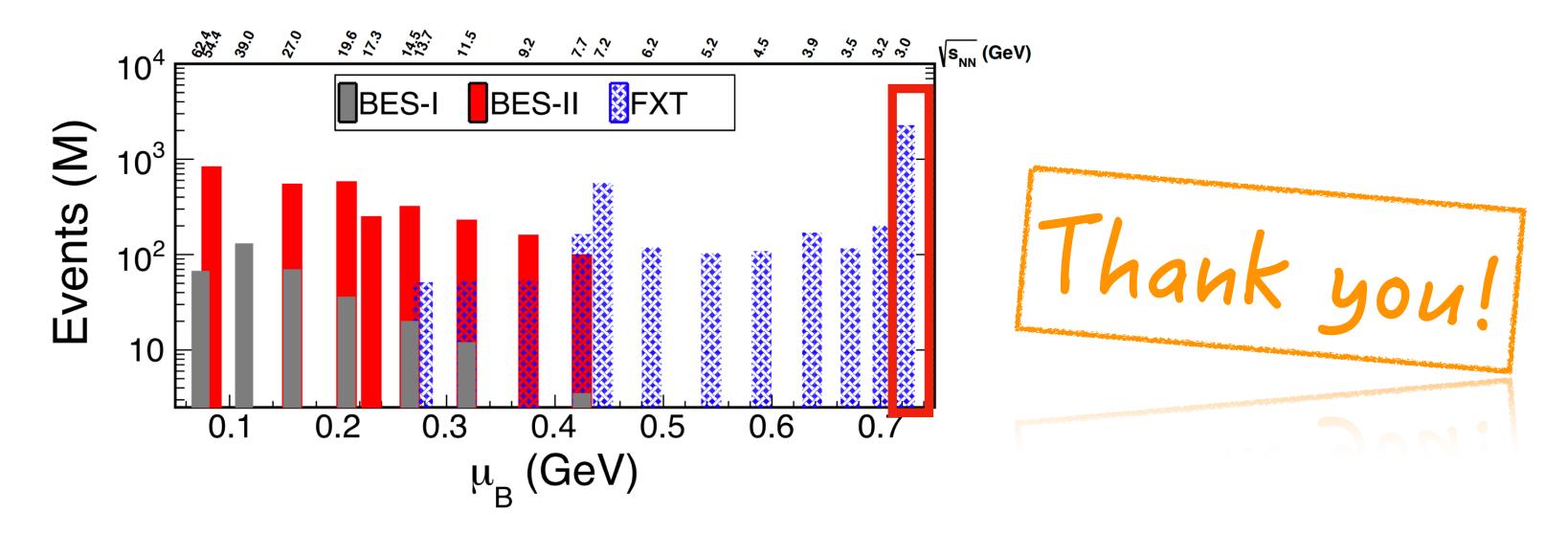
• e.g.
$$^4_{\Lambda\Lambda} \text{He} \rightarrow ^4_{\Lambda} \text{He} \pi$$
, $^5_{\Lambda\Lambda} \text{He} \rightarrow ^5_{\Lambda} \text{He} \pi$



Outlook

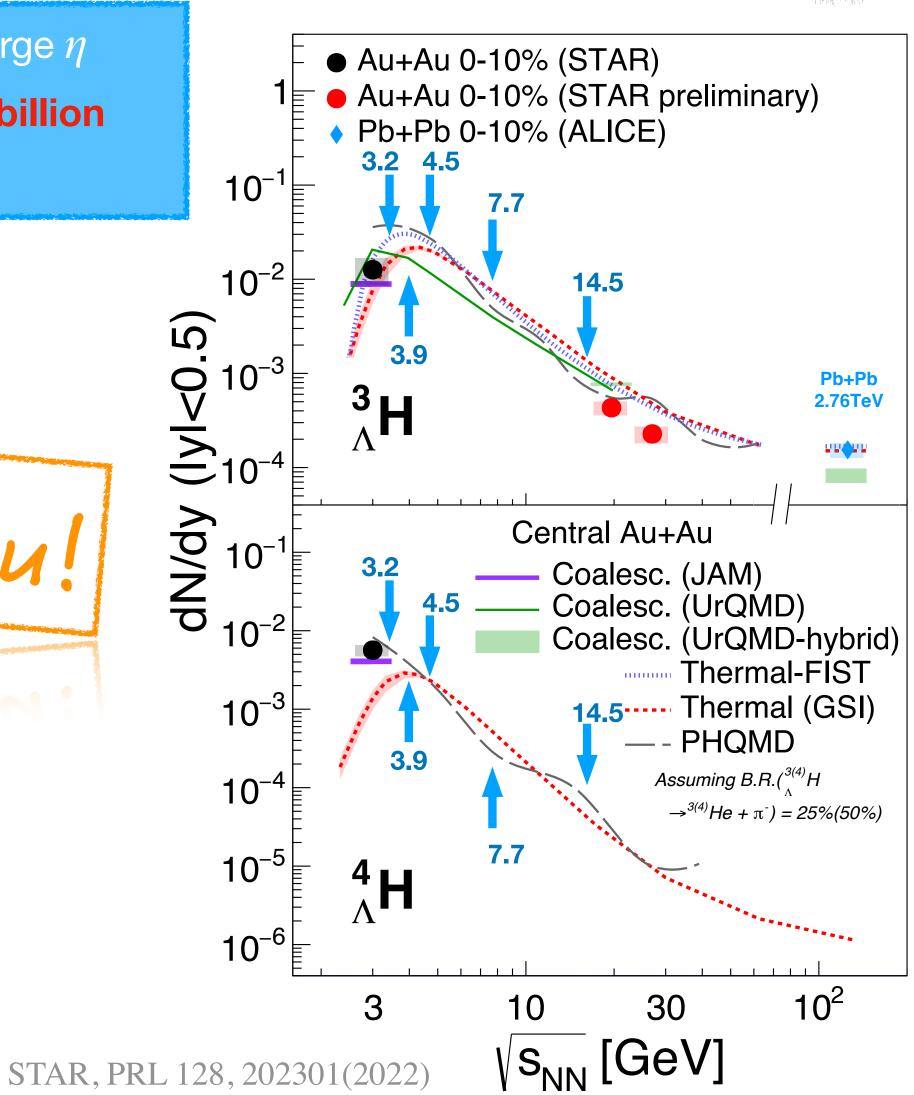


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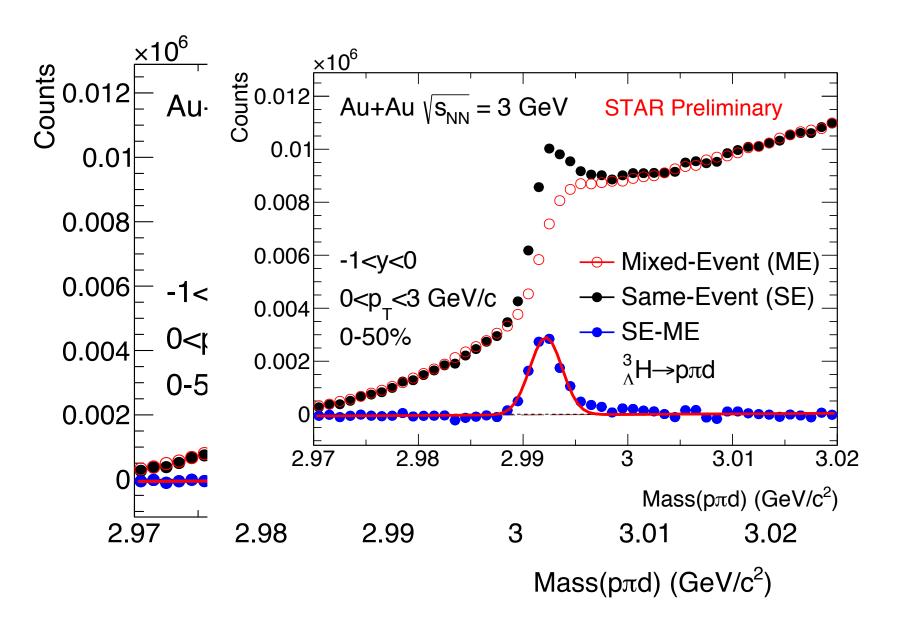
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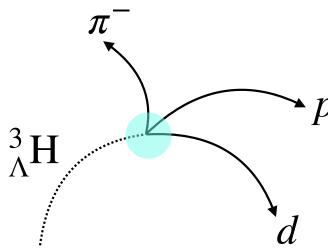


Backup slides

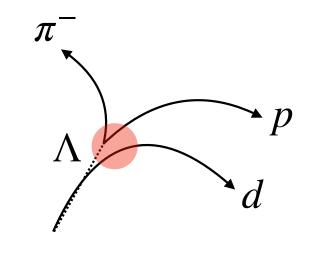
³H 3-body signal



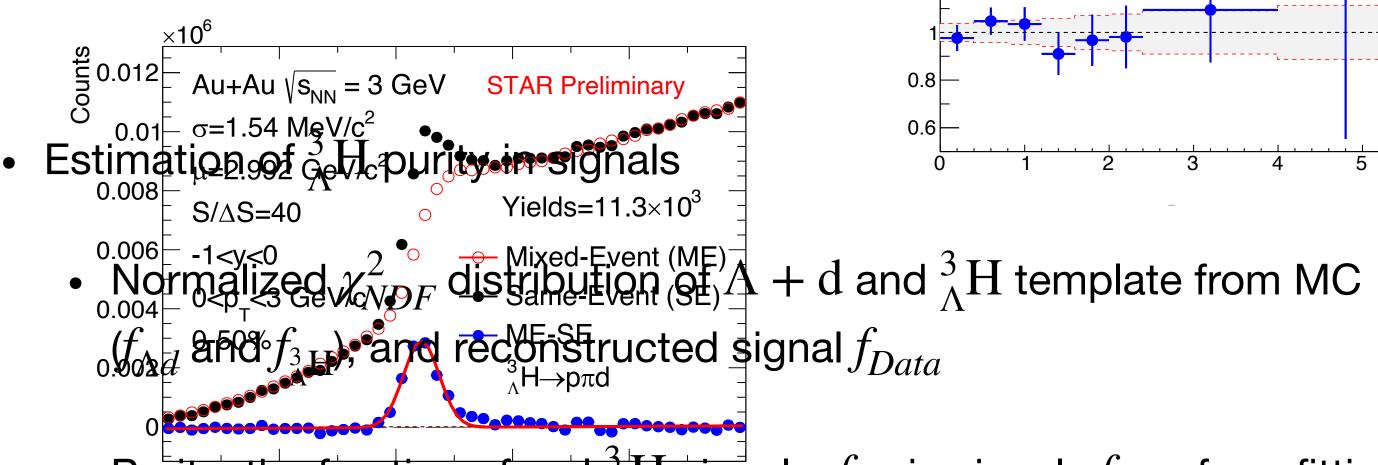
• SE-ME signals contains real signal and kinematically correlated $\Lambda + d(\Lambda \to p\pi^-)$



Real signal: lower χ^2



Backgrounds: higher χ^2



1<p_<3 GeV/c

H MC

⊣₃ta

-d MC

= 0.62± 0.04

 6 3 H χ^{2} $^{\prime}$ NDF

3 GeV/c

y<-0.25

Purity = 0.62 ± 0.04

□ Λ+d MC

 $1 < p_{\tau} < 3 \text{ GeV/c}$

-0.50<y<-0.25

0-50%

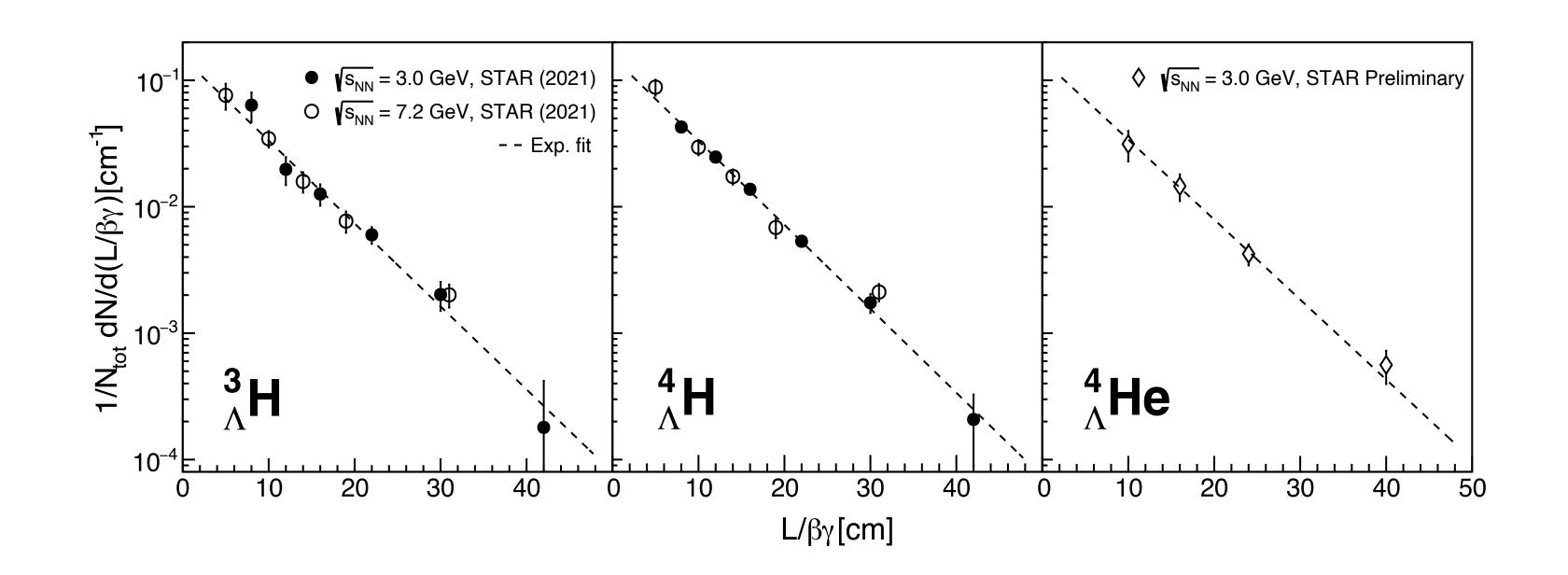
STAR preliminary

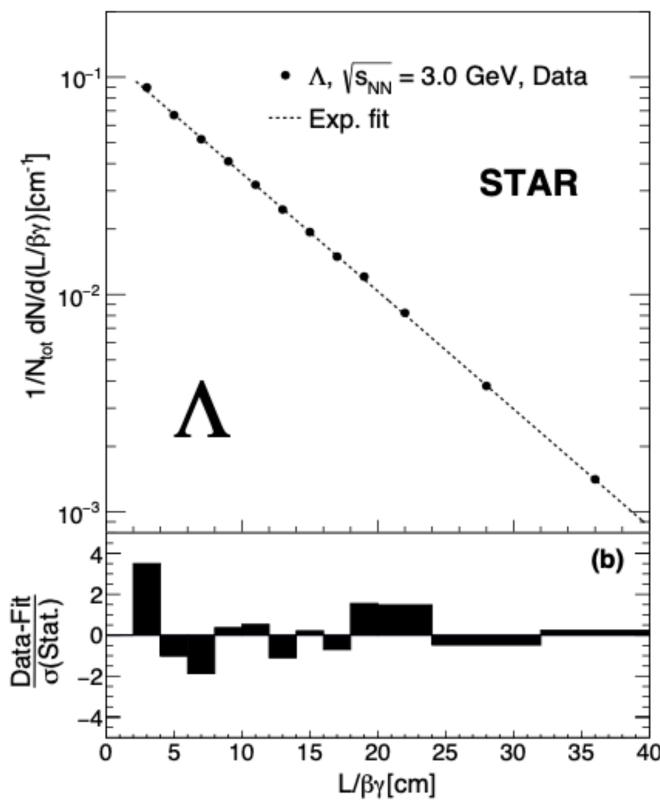
Purity = 0.62 ± 0.04

• Purity: the fraction of real 3H signals f_{3H} in signals f_{Data} from fitting $f_{Data} = p_0 \cdot (f_{\Lambda d} + p_1 \cdot f_{3H})^{\text{Mass}(p\pi d) (GeV/c^2)}$

Lifetime

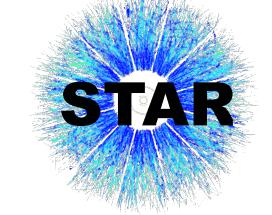


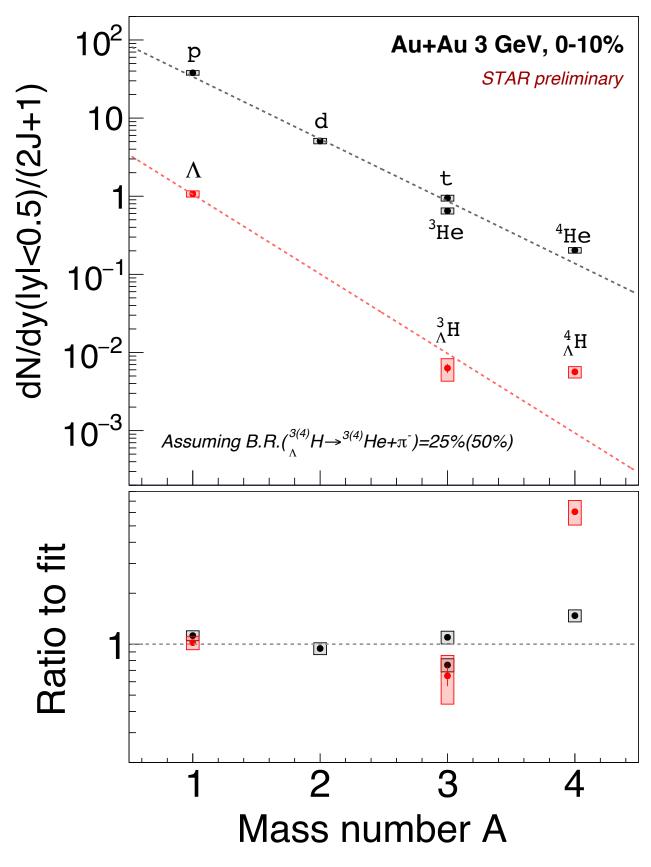




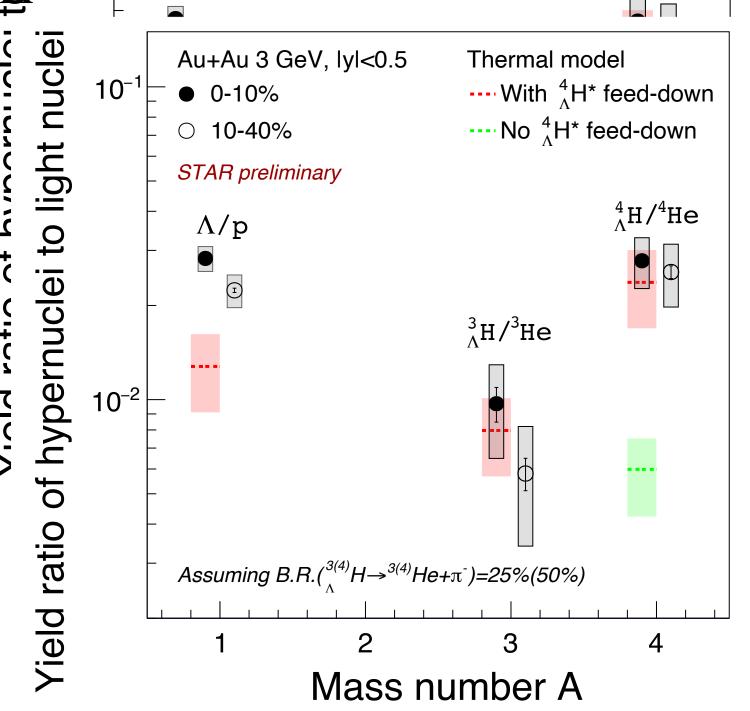
- Lifetime τ extracted via $N(t) = N_0 e^{-L/\beta \gamma c \tau}$
- Λ lifetime cross check : 267±4 ps, consistent with PDG value (263±2 ps)
- $^3_\Lambda H$ and $^4_\Lambda H$ lifetimes from 3.0 GeV consistent with 7.2 GeV results

Hyper-to-light nuclei ratio See poster by: Yingjie Zhou (4/8 T11_2)





- Thermal/coalescence models predict approx. exponential dependence of yields (25 down) vs A
- $\frac{1}{2}H$ les a factor of 6 above exponential fit to $(\Lambda, {}^3_{\Lambda}H, {}^4_{\Lambda}H)$



- Non-mononic behavior in light-tohyper-nuclei ratio vs A observed
 - Thermal model calculations including excited $^4_\Lambda H^*$ feed-down shows a similar trend

A. Andronic et al, PLB 697 (2011) 203 (Thermal model)

- Non-existence of bound $^3_{\Lambda}H^*$ (J⁺=3/2)
 - Data support creation of unstable A = 4 hypernuclei from heavy-ion collisions