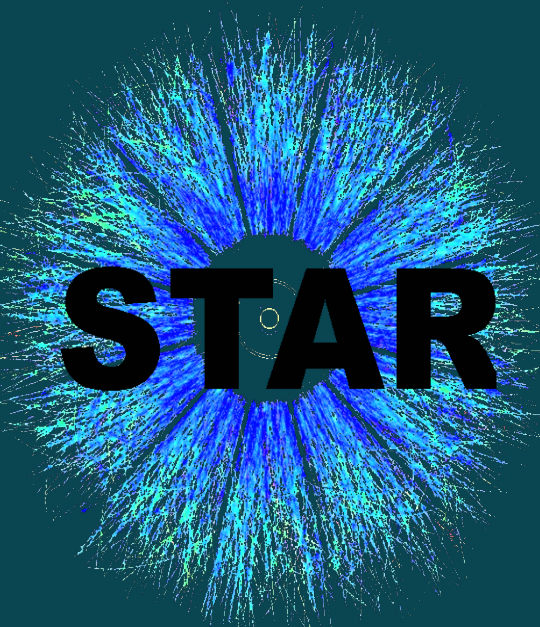




# Search for the Chiral Magnetic Effect using Sliding Dumbbell Method in Isobar Collisions ( $^{44}_{96}\text{Ru} + ^{44}_{96}\text{Ru}$ and $^{40}_{96}\text{Zr} + ^{40}_{96}\text{Zr}$ ) at RHIC



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## Introduction

In the hot dense de-confined medium in non-central heavy-ion collisions, the strong magnetic field created by the fast-moving spectator protons causes the charge separation perpendicular to the reaction plane, a phenomenon known as the Chiral Magnetic Effect (CME) [1] as shown in Fig.1. The charge separation effect has been investigated both at RHIC and LHC using the CME sensitive  $\gamma$ -correlator ( $\langle \cos(\phi_a + \phi_b - 2\Psi_{RP}) \rangle$ ) [2]. It was realized that four more protons in  $^{44}_{96}\text{Ru}$  than those in  $^{40}_{96}\text{Zr}$  will result in an increased magnetic field ( $\sim 10\%$ ) in  $^{96}_{44}\text{Ru} + ^{96}_{44}\text{Ru}$  collisions than those in  $^{96}_{40}\text{Zr} + ^{96}_{40}\text{Zr}$  collisions [3].

This led to the expectation of an enhanced CME effect in Ru+Ru collisions than in Zr+Zr collisions both having similar backgrounds. A new technique, Sliding Dumbbell Method [4] is designed to search for potential CME-like events corresponding to the highest back-to-back charge separation.

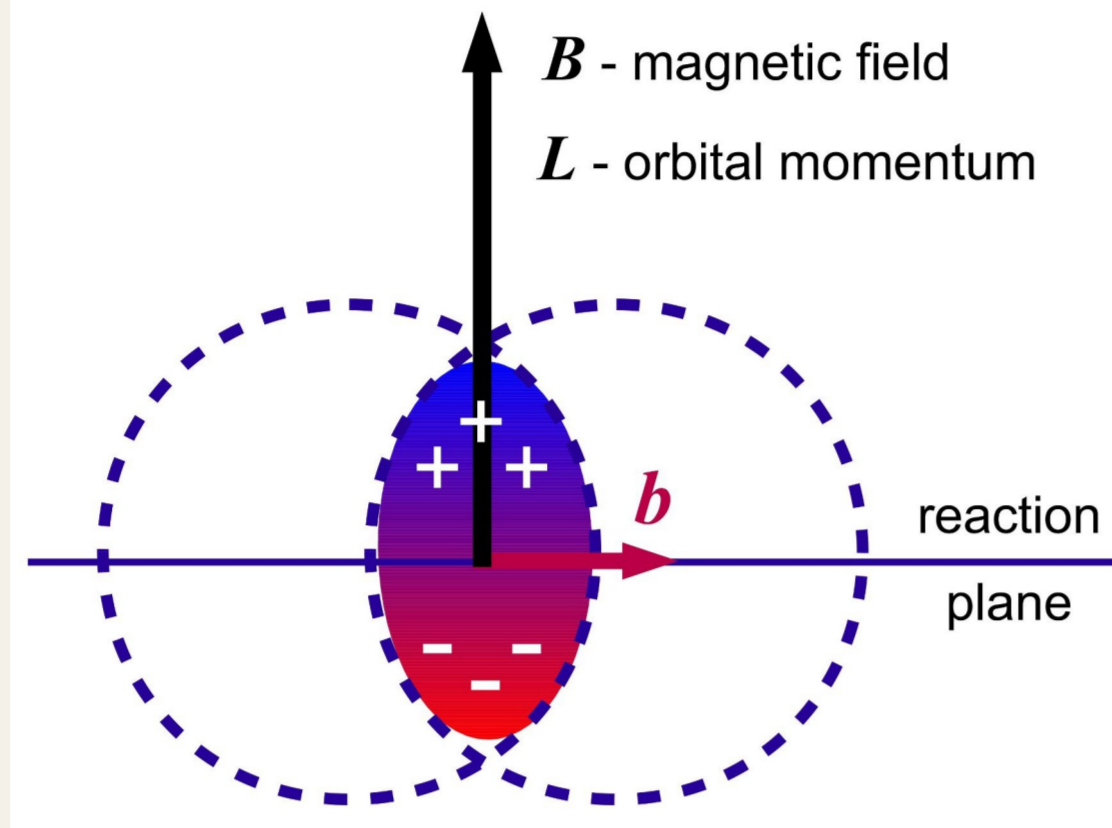


Fig.1: A schematic illustration of charge separation in non-central heavy-ion collision.

## Analysis Method: Sliding Dumbbell Method

The Sliding Dumbbell Method (SDM) [3] is designed to search minutely for the back-to-back charge separation on an event-by-event basis (Fig.2). Azimuthal plane in each event is scanned by sliding the dumbbell of  $90^\circ$  in steps of  $\delta\phi = 1^\circ$  while calculating,  $Db_{+-}$  for each region to obtain the maximum value of  $Db_{+-}$  i.e.,  $Db_{+-}^{max}$  in each event.

$$Db_{+-} = \frac{n_+^a}{(n_+^a + n_-^a)} + \frac{n_-^b}{(n_+^b + n_-^b)}$$

Where,  $n_+^a$  and  $n_-^a$  ( $n_+^b$  and  $n_-^b$ ), are the number of positive and negative charged particles on the "a" ("b") side of the dumbbell.

• Fractional Charge separation ( $f_{DbCS}$ ) across the dumbbell in each event is defined as :

$$f_{DbCS} = Db_{+-}^{max} - 1$$

Fig.2: Pictorial depiction of the transverse (azimuthal) plane with hits of positive (+) and negative (-) charge particles in an event.

## Experimental Setup

The STAR (Solenoidal Tracker at RHIC) detector shown in Fig.3 at RHIC consists of various sub-detectors. Time Projection Chamber (TPC) and Time of Flight (TOF) are the main sub-detectors, with complementary tracking and particle identification capabilities.

Data Set: Isobaric Collisions (Ru+Ru & Zr+Zr) at 200 GeV ( $\sim 50\%$  of the available data).

Event and Track selection cuts are as follows:

- **Vertex Cut:**  $-35 < V_z < 25$  cm
- **Pseudo-rapidity Cut:**  $|\eta| < 1$
- **Transverse momentum Cut:**  $0.2 < p_t < 2.0 \text{ GeV}/c$
- **Distance of Closest Approach:**  $\text{DCA} < 3$  cm

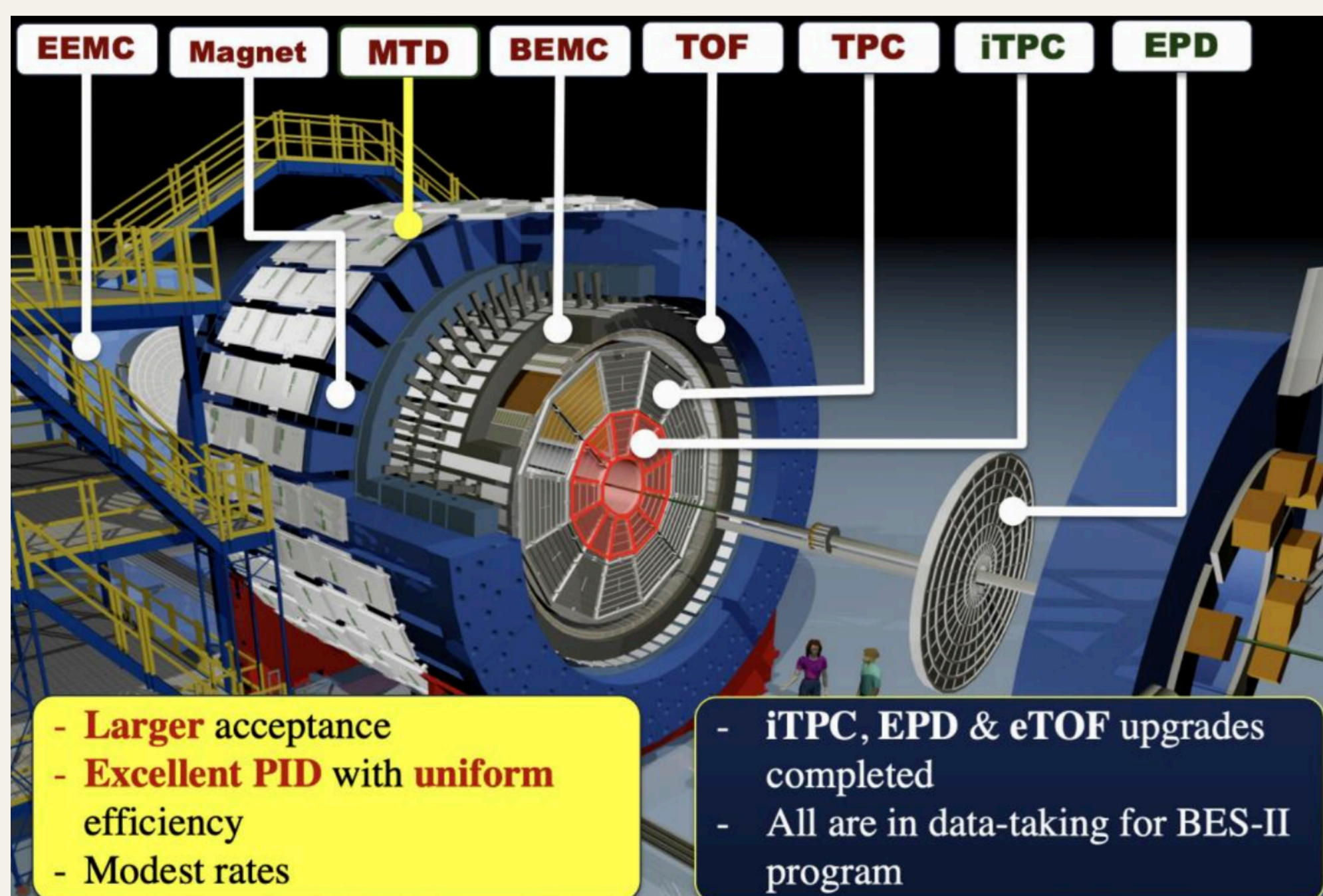


Fig.3: Layout of sub-detectors of the STAR experiment.

## Background Estimation

- **Charge Shuffle (ChS):** The charges of particles in each event are shuffled randomly to destroy the charge-dependent correlations amongst charged particles but keeping  $\theta$  and  $\phi$  of each particle unchanged in an event.
- **Correlated (Corr.) Background:** The shuffling of charges of particles in an event keeping the flow in, kills not only the CME-like correlations but also correlations amongst produced particles in an event. Correlations amongst particles that were destroyed during charge shuffling were recovered from the corresponding original events in a particular  $f_{DbCS}$  bin. This is termed as the correlated background.

## Results

Fig.4: It is seen that the charge separation ( $f_{DbCS}$ ) distributions extend towards higher  $f_{DbCS}$  values with decreasing collision centrality. Different  $f_{DbCS}$  distributions are divided into ten percentile bins for each centrality.

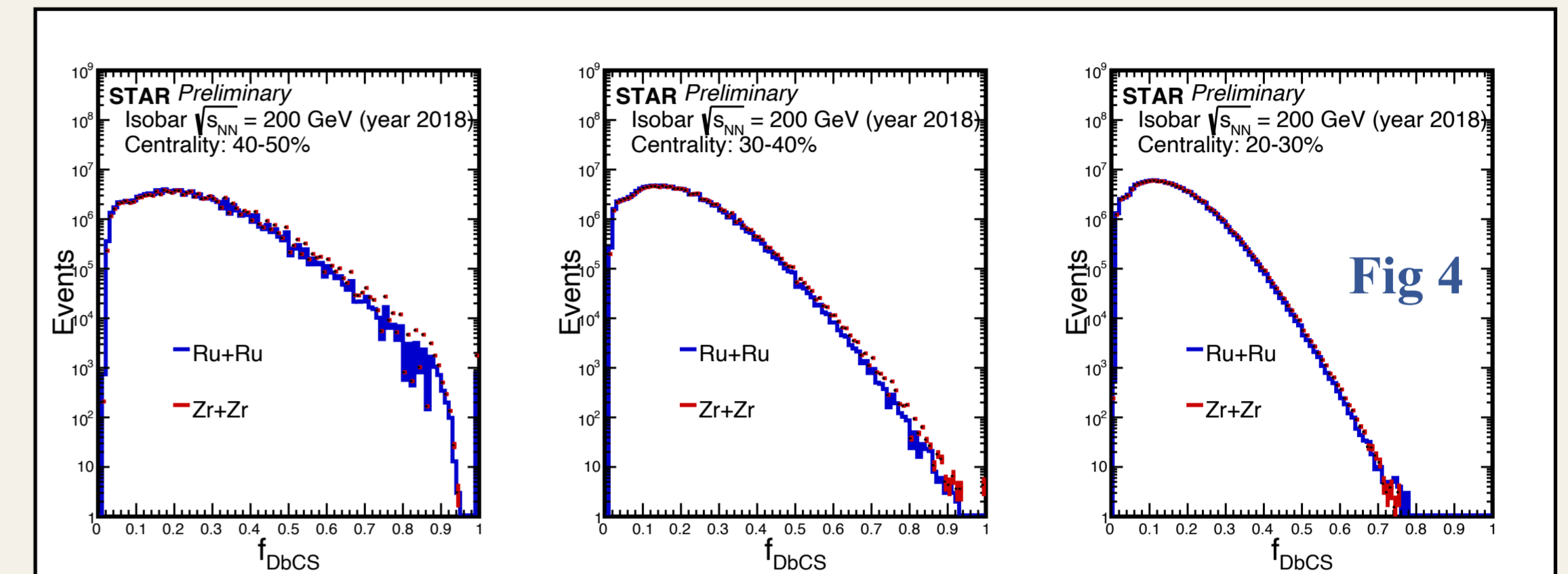


Fig.5:  $\gamma_{OS}$  and  $\gamma_{SS}$  are plotted as a function of centrality for Ru+Ru and Zr+Zr collisions along with different charge separation ( $f_{DbCS}$ ) bins. It is seen that for the highest charge separations,  $\gamma_{OS} > 0$  and  $\gamma_{SS} < 0$  for top 20%(30%)  $f_{DbCS}$  bins for 0-40%(40-60%) centralities. Boxes represent systematic uncertainties on the data points whereas statistical errors are smaller than the symbol sizes.

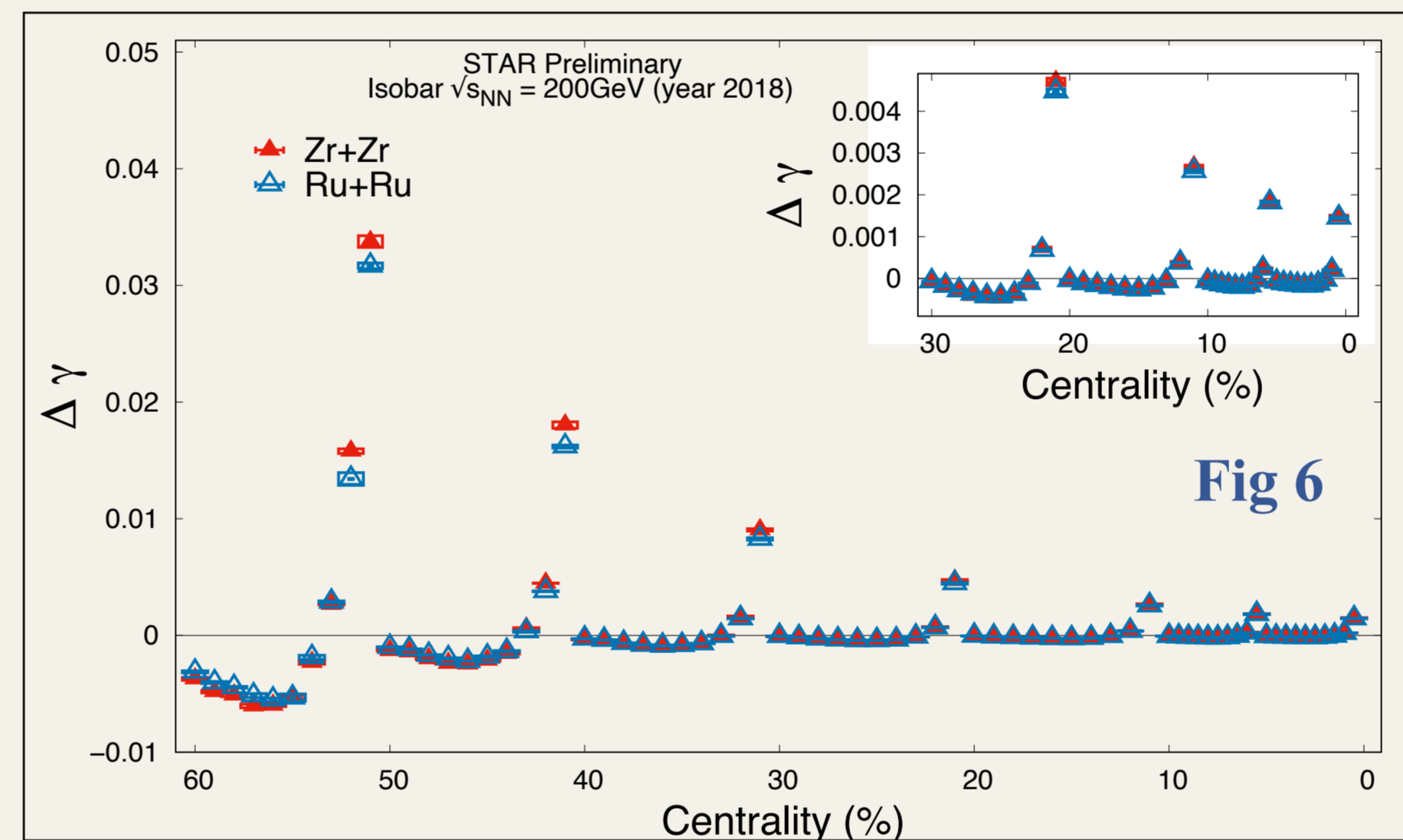
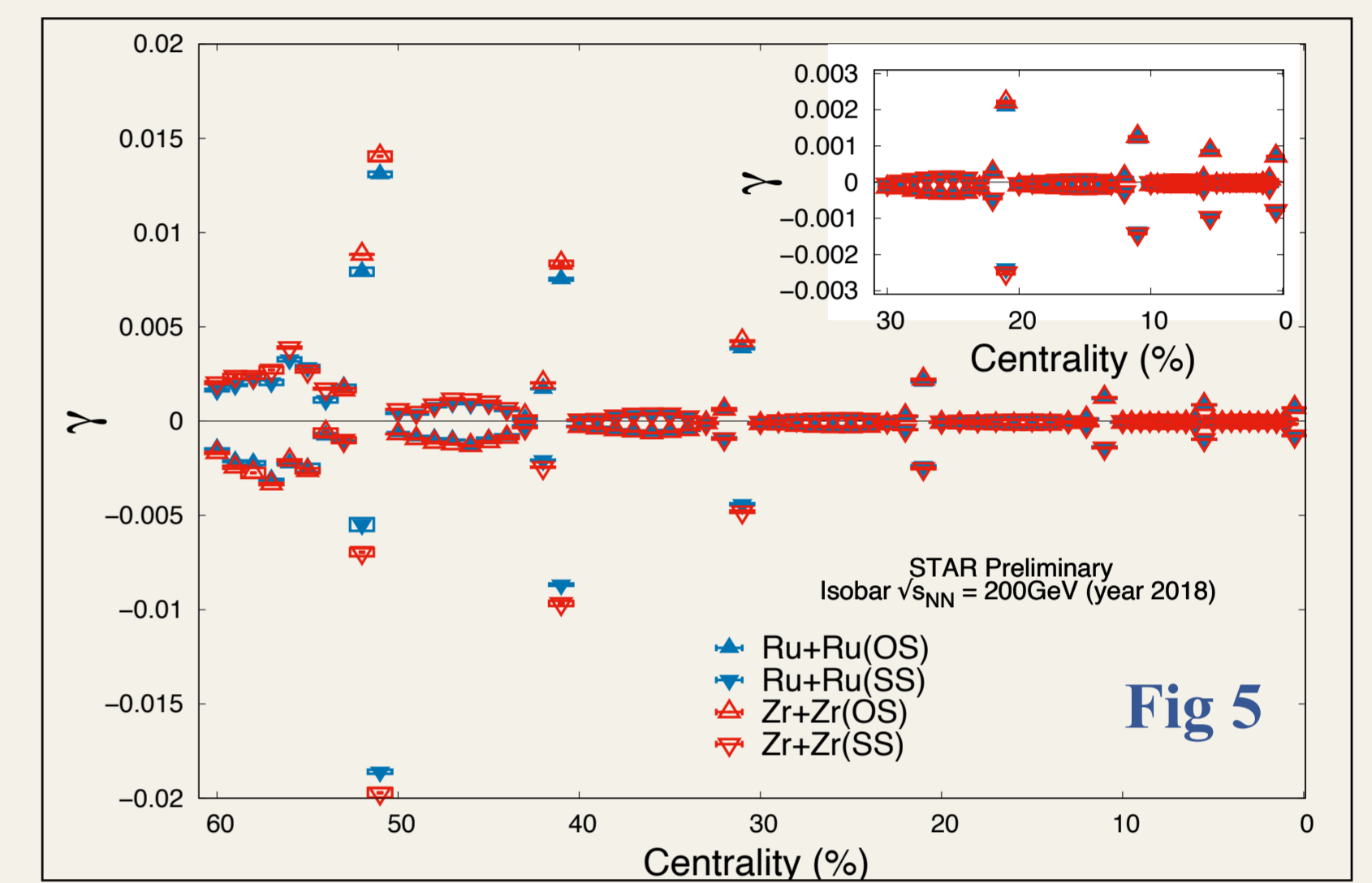


Fig.6:  $\Delta\gamma$  is compared for Ru+Ru and Zr+Zr collisions.  $\Delta\gamma$  is positive for the top 20%(30%)  $f_{DbCS}$  bins in 0-40%(40-60%) centralities.  $\Delta\gamma$  is smaller for Ru than those of Zr for the top 10% (top 20%)  $f_{DbCS}$  bins for 20-40% (40-60%) centralities.

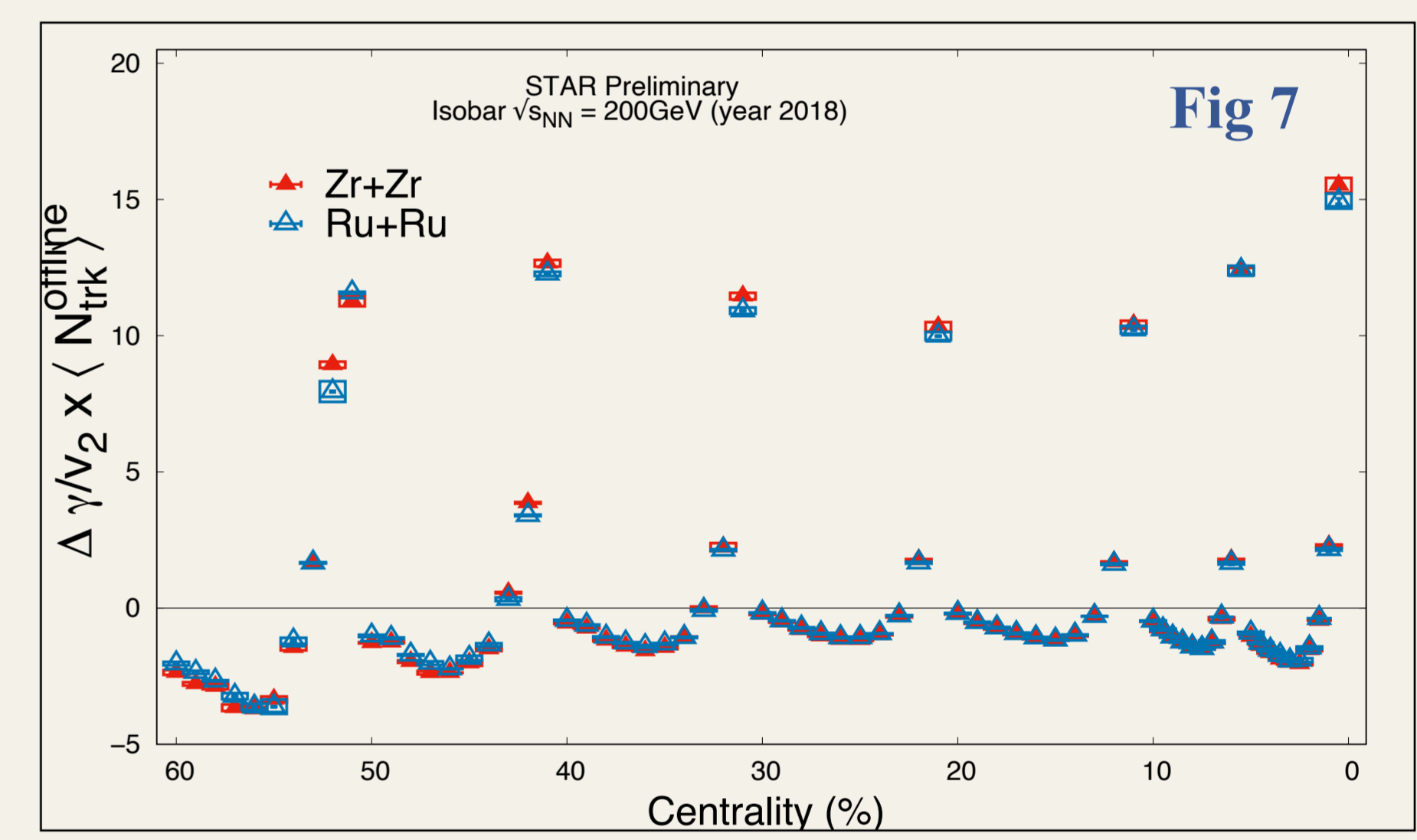


Fig.7:  $\Delta\gamma$  is scaled with  $\langle N_{trk}^{offline} \rangle / v_2$  (where  $\langle N_{trk}^{offline} \rangle$  is multiplicity uncorrected for tracking efficiency in  $|\eta| < 0.5$ ).

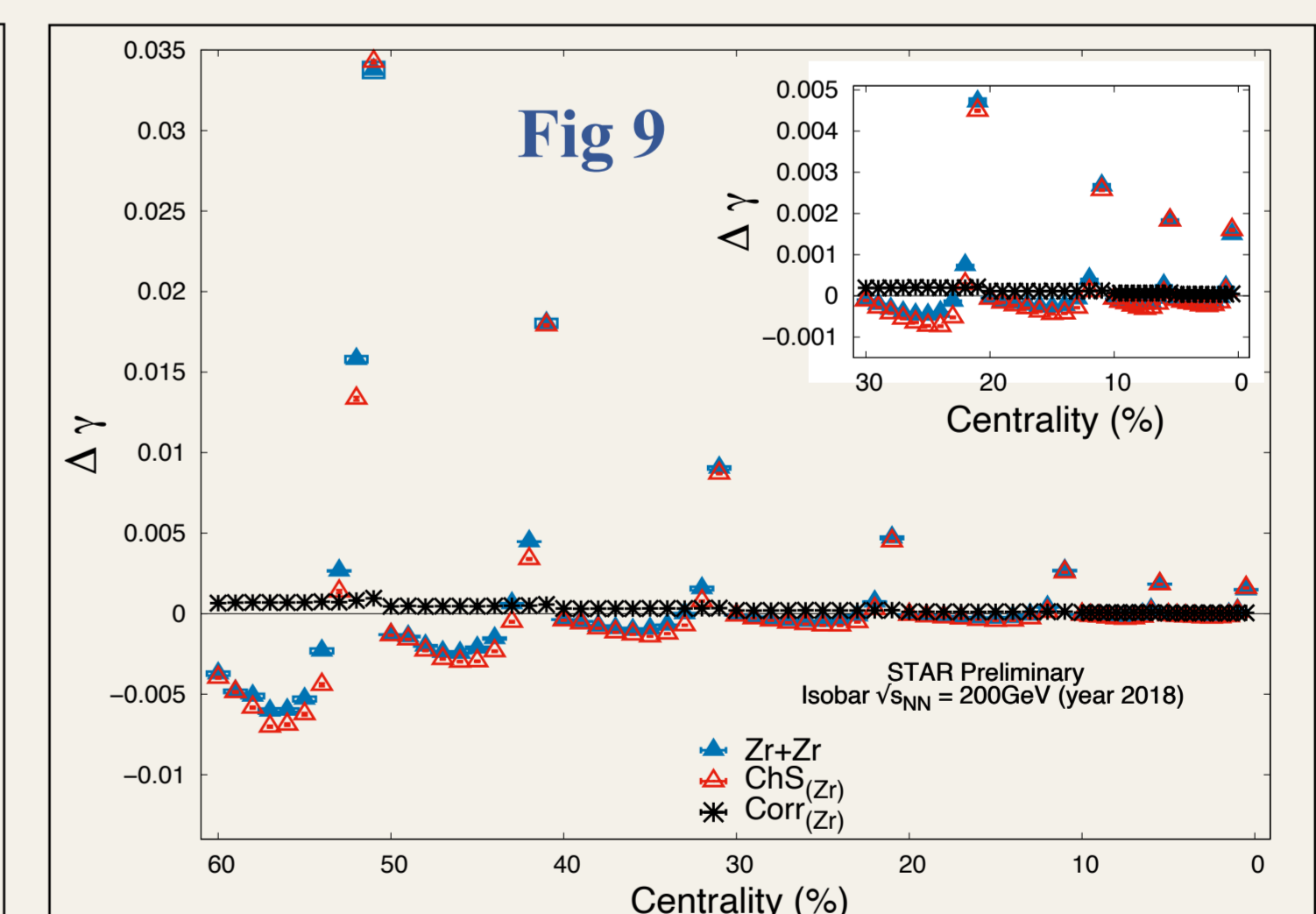
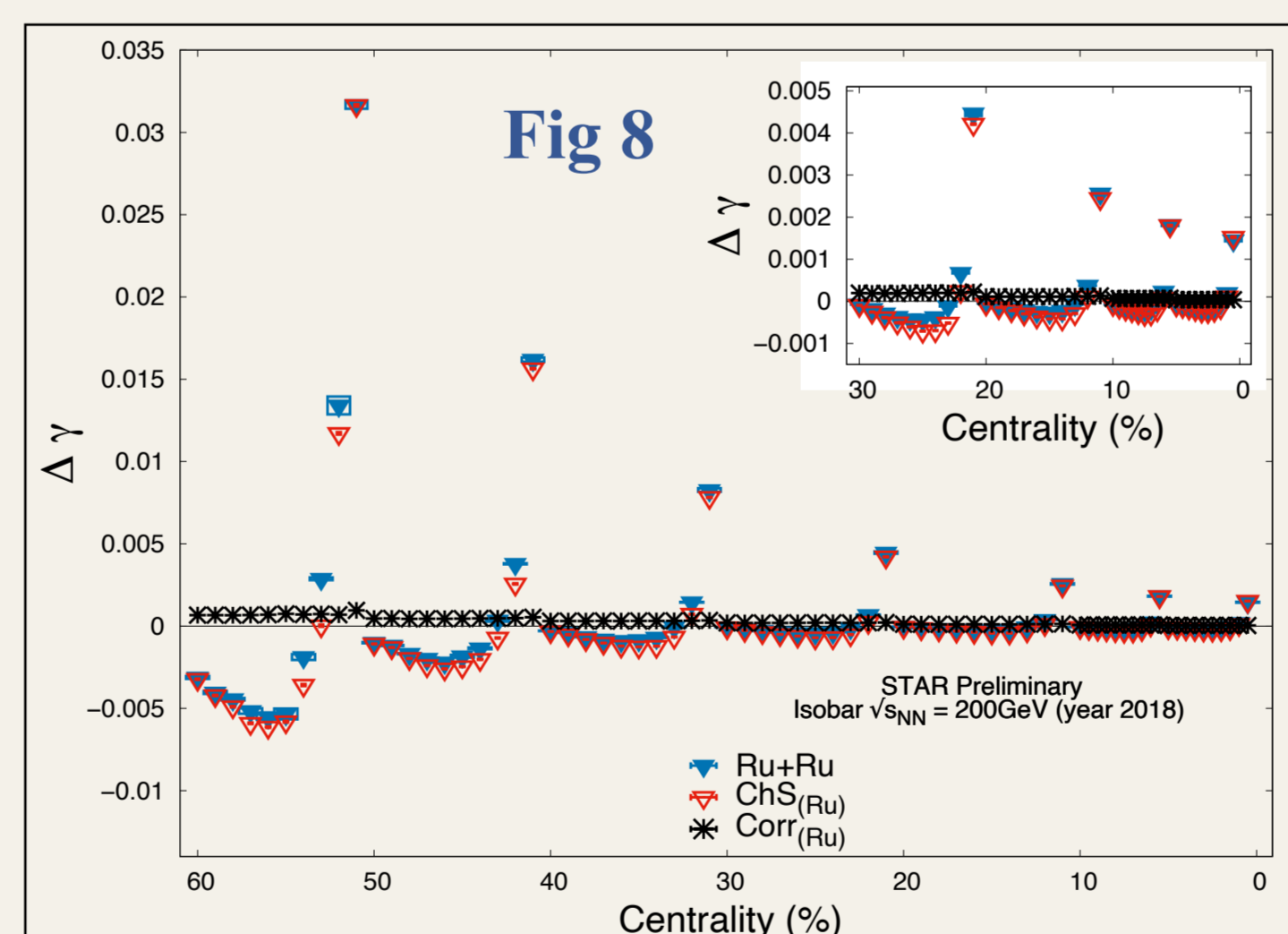


Fig.8 & Fig.9:  $\Delta\gamma$  for Ru+Ru and Zr+Zr are compared with their respective backgrounds (i.e., Charge shuffled (ChS) and Correlated (Corr.)) in each  $f_{DbCS}$  bin for 0-60% collision centralities.

## Summary

- The charge separation ( $f_{DbCS}$ ) distribution extends towards higher  $f_{DbCS}$  values with decreasing collision centrality.
- It is seen that  $\gamma_{OS} > 0$  and  $\gamma_{SS} < 0$  for top 20%(30%)  $f_{DbCS}$  bins for 0-40%(40-60%) centralities as expected in CME like events.
- It can be seen that  $\Delta\gamma$  are smaller for Ru than those of Zr for the top 10% (top 20%)  $f_{DbCS}$  bins for 20-40% (40-60%) centralities. However, the difference between them decreases if  $\Delta\gamma$  is scaled with  $\langle N_{trk}^{offline} \rangle / v_2$ .
- $\Delta\gamma$  for Ru+Ru and Zr+Zr are compared with their respective backgrounds (i.e., Charge shuffled (ChS) and Correlated (Corr.)) for 0-60% collision centralities. We are analyzing the full available data set to get a detailed comparison.

## References:

- [1] I. Selyuzhenkov, arXiv: 0910.0464, (2003).
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- [3] M. S. Abdallah *et al.*, STAR Collaboration, Phys. Rev. C 105, 014901 (2022).
- [4] Jagbir Singh, Anjali Attri and Madan M. Aggarwal, DAE Symp.Nucl.Phys. 64, 830, (2019).