

# Estimate of nonflow baseline for the chiral magnetic effect in isobar collisions at STAR



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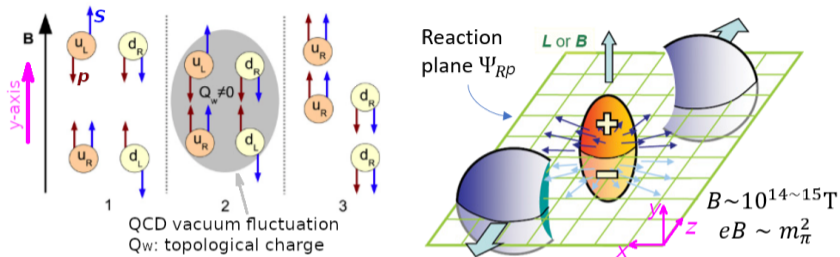
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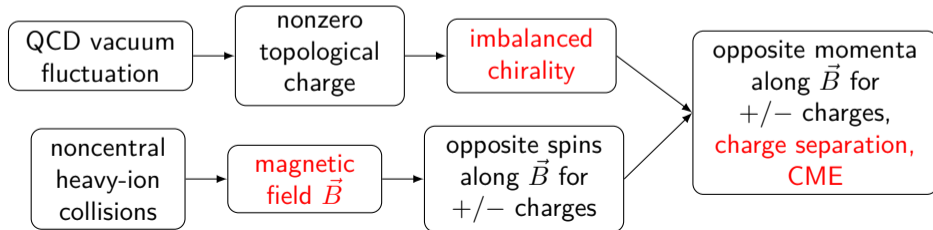
# Outline

- 1 Introduction
- 2 Isobar  $\Delta\gamma$  nonflow baseline
- 3 Summary

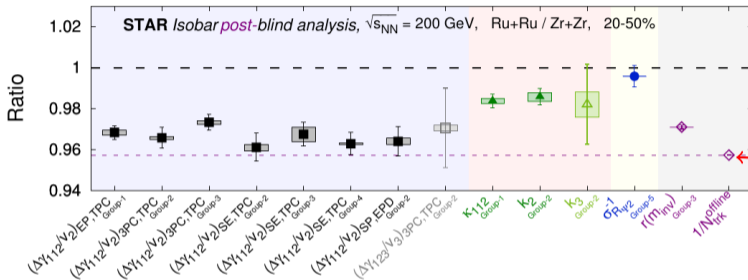
# The Chiral Magnetic Effect (CME)



CME schematics. [Kharzeev, McLerran and Warringa, NPA 803, 227 (2008)]



# Isobar Results



Post-blind results from STAR isobar analysis [STAR, PRC 105, 014901 (2022)].

- ▶ Isobar expectation:  $\Delta\gamma/v_2$  in  $^{96}_{44}\text{Ru} + ^{96}_{44}\text{Ru}$  is larger than in  $^{96}_{40}\text{Zr} + ^{96}_{40}\text{Zr}$ .
- ▶ The main reason that the observed isobar ratio is less than unity is the multiplicity difference.
- ▶ The better quantity is  $N\Delta\gamma/v_2$ . Its naive background baseline is unity.
- ▶ Isobar data are all above this naive baseline. Investigate nonflow effects.

- ▶ The CME-sensitive observable  $\Delta\gamma \equiv C_3/v_2^*$ :

$$C_{3,os} = \langle \cos(\phi_\alpha^\pm + \phi_\beta^\mp - 2\phi_c) \rangle,$$

$$C_{3,ss} = \langle \cos(\phi_\alpha^\pm + \phi_\beta^\pm - 2\phi_c) \rangle,$$

$$C_3 = C_{3,os} - C_{3,ss}$$

OS: opposite-sign pair  
SS: same-sign pair

- ▶ The asterisk (\*) on  $v_2$  indicates it is the measured  $v_2$  containing nonflow
- ▶  $\Delta\gamma$  contains CME and a major background proportional to  $v_2$  (true  $v_2$  flow)

# Nonflow Contribution to Isobar Baseline

- ▶ The naive baseline of unity would be correct if there was no nonflow. Nonflow correlations will cause the baseline to deviate from unity.

Note:

$\epsilon$  is not eccentricity

- ▶ Nonflow in  $v_2^*$ :  $v_2^{*2} = v_2^2 + v_{2,\text{nf}}^2$ ,  $\epsilon_{\text{nf}} \equiv v_{2,\text{nf}}^2/v_2^2$
- ▶  $C_3$  is composed of flow-induced background (major), 3p nonflow correlations (minor), and possible CME (not written out) [Y. Feng, et al., PRC 105, 024913 (2022)]:

$$C_3 = \frac{N_{2p}}{N^2} C_{2p} v_{2,2p} v_2 + \frac{N_{3p}}{2N^3} C_{3p} = \frac{v_2^2 \epsilon_2}{N} + \frac{\epsilon_3}{N^2},$$

$$\frac{N \Delta \gamma}{v_2^*} = \frac{N C_3}{v_2^{*2}} = \frac{\epsilon_2}{1 + \epsilon_{\text{nf}}} + \frac{\epsilon_3}{N v_2^2 (1 + \epsilon_{\text{nf}})} = \frac{\epsilon_2}{1 + \epsilon_{\text{nf}}} \left( 1 + \frac{\epsilon_3 / \epsilon_2}{N v_2^2} \right)$$

- 2-particle (2p) nonflow (e.g., resonance):  $C_{2p} \equiv \langle \cos(\phi_\alpha + \phi_\beta - 2\phi_{2p}) \rangle$ ,  $\epsilon_2 \equiv \frac{N_{2p} v_{2,2p}}{N v_2^2} C_{2p}$
- 3-particle (3p) nonflow (e.g., jets):  $C_{3p} \equiv \langle \cos(\phi_\alpha + \phi_\beta - 2\phi_c) \rangle_{3p}$ ,  $\epsilon_3 \equiv \frac{N_{3p}}{2N} C_{3p}$
- $N \approx N_+ \approx N_-$  is POI (particle of interest) mult.  $N_{2p}$  ( $N_{3p}$ ) is 2p (3p) nonflow pair (triplet) mult.

# Nonflow Contribution to Isobar Baseline

► Isobar ratio:

$$\frac{(N\Delta\gamma/v_2^*)^{\text{Ru}}}{(N\Delta\gamma/v_2^*)^{\text{Zr}}} \equiv \frac{(NC_3/v_2^{*2})^{\text{Ru}}}{(NC_3/v_2^{*2})^{\text{Zr}}} = \frac{\epsilon_2^{\text{Ru}}}{\epsilon_2^{\text{Zr}}} \cdot \frac{(1 + \epsilon_{\text{nf}})^{\text{Zr}}}{(1 + \epsilon_{\text{nf}})^{\text{Ru}}} \cdot \frac{[1 + \epsilon_3/\epsilon_2/(Nv_2^2)]^{\text{Ru}}}{[1 + \epsilon_3/\epsilon_2/(Nv_2^2)]^{\text{Zr}}}$$

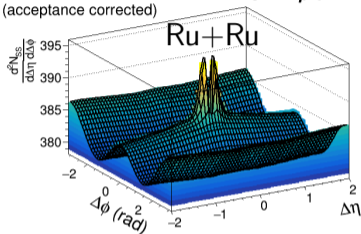
$$\approx 1 + \frac{\Delta\epsilon_2}{\epsilon_2} - \frac{\Delta\epsilon_{\text{nf}}}{1 + \epsilon_{\text{nf}}} + \frac{\epsilon_3/\epsilon_2/(Nv_2^2)}{1 + \epsilon_3/\epsilon_2/(Nv_2^2)} \left( \frac{\Delta\epsilon_3}{\epsilon_3} - \frac{\Delta\epsilon_2}{\epsilon_2} - \frac{\Delta N}{N} - \frac{\Delta v_2^2}{v_2^2} \right)$$

► Need  $\epsilon_{\text{nf}}$ ,  $\epsilon_2$ ,  $\epsilon_3$  for background estimate

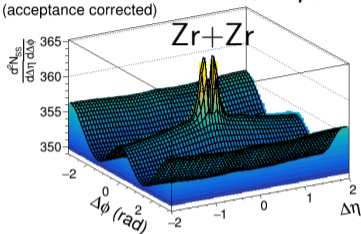
$$\Delta X = X^{\text{Ru}} - X^{\text{Zr}}$$

- $\epsilon_{\text{nf}} \equiv \frac{v_{2,\text{nf}}^2}{v_2^2} = \frac{v_2^{*2} - v_2^2}{v_2^2}$
- $\epsilon_2 \equiv \frac{C_{2p}N_{2p}v_{2,2p}}{Nv_2} = \frac{N_{2p}v_{2,2p}}{Nv_2} \langle \cos(\phi_\alpha + \phi_\beta - 2\phi_{2p}) \rangle$
- $\epsilon_3 \equiv \frac{C_{3p}N_{3p}}{2N} = \frac{N_{3p}}{2N} \langle \cos(\phi_\alpha + \phi_\beta - 2\phi_c) \rangle_{3p}$

Ru+Ru 200GeV, 20–50%, SS pair  
(acceptance corrected) **STAR preliminary**



Zr+Zr 200GeV, 20–50%, SS pair  
(acceptance corrected) **STAR preliminary**



- $0.2 < p_T < 2.0$  GeV
- $|\eta| < 1$
- Centrality 20 – 50% defined by POI multiplicity
- Mixed-event acceptance corrected

Fit function:

$$\begin{aligned}
 f(\Delta\eta, \Delta\phi) = & \underbrace{A_1 G_{NS,W}(\Delta\eta) G_{NS,W}(\Delta\phi)}_{\text{Nearside}} + \underbrace{A_2 G_{NS,N}(\Delta\eta) G_{NS,N}(\Delta\phi)}_{\text{Nearside}} + \underbrace{A_3 G_{NS,D}(\Delta\eta) G_{NS,D}(\Delta\phi)}_{\text{Nearside}} \\
 & + \underbrace{\frac{B}{2-|\Delta\eta|} \text{erf}\left(\frac{2-|\Delta\eta|}{\sqrt{2}\sigma_{\Delta\eta,AS}}\right) G_{AS}(\Delta\phi \pm \pi)}_{\text{Awayside}} + \underbrace{DG_{RG}(\Delta\eta)}_{\text{Ridge}} + \underbrace{C[1 + 2V_1 \cos(\Delta\phi) + 2V_2 \cos(2\Delta\phi) + 2V_3 \cos(3\Delta\phi)]}_{\text{Flow}}
 \end{aligned}$$

$G_s(x)$  Gaussian function,  $V_n = v_n^2$  assumed to be  $\eta$ -independent.  
 NS–nearside, AS–awayside, RG–ridge; W–wide, N–narrow, D–dip.



<i>STAR preliminary</i>		Ru+Ru	Zr+Zr
ss	fit parameter $C$	$381.651 \pm 0.011$	$351.988 \pm 0.009$
	fit parameter $V_2 = v_2^2$	$0.0029716 \pm 0.0000029$	$0.0028668 \pm 0.0000025$
	$\langle \cos(2\Delta\phi) \rangle_{ss} ( \Delta\eta  > 0.05)$	$0.0035968 \pm 0.0000010$	$0.0034930 \pm 0.0000010$
inclusive	$\langle \cos(2\Delta\phi) \rangle = v_2^{*2} ( \Delta\eta  > 0.05)$	$0.0037161 \pm 0.0000007$	$0.0036088 \pm 0.0000007$
	nonflow $U = \langle \cos(2\Delta\phi) \rangle - V_2$	$0.0007446 \pm 0.0000030$	$0.0007420 \pm 0.0000026$
	$\epsilon_{nf} = U/V_2$	$(25.06 \pm 0.10)\%$	$(25.88 \pm 0.09)\%$

► Nonflow in  $v_2^{*2}$  is  $\sim 25\% / (1 + 25\%) = 20\%$ .

► The nearside wide Gaussian ( $A_1$  term) is dominant.

► We take half of it as systematics:

$$\Delta\epsilon_{nf} = (-0.82 \pm 0.13 \mp 0.30)\%, \quad -\Delta\epsilon_{nf}/(1 + \epsilon_{nf}) = (0.65 \pm 0.11 \pm 0.22)\%.$$

$$\Delta v_2^2/v_2^2 = \Delta V_2/V_2 = (3.7 \pm 0.1 \mp 0.3)\%.$$

- ▶  $\epsilon_2$  can be obtained from ZDC measurement (no nonflow, assuming negligible CME) [STAR, PRC 105, 014901 (2022)]

$$\epsilon_2 = \frac{N\Delta\gamma\{\text{ZDC}\}}{v_2\{\text{ZDC}\}} \approx 0.57 \pm 0.04 \pm 0.02 \text{ (tracking efficiency } \sim 80\%)$$

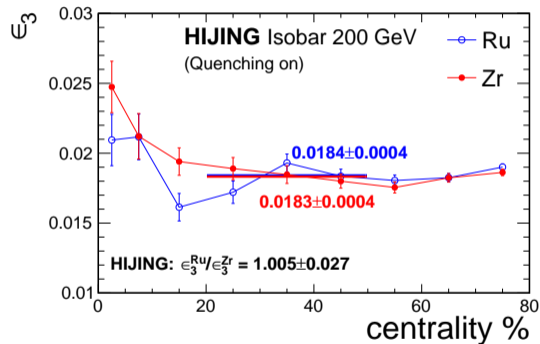
- ▶ The  $\Delta\epsilon_2$  precision from ZDC is too poor:  $\Delta\epsilon_2/\epsilon_2 \approx (2.3 \pm 9.2)\%$ , but we can estimate it as follows:

- Assuming  $C_{2p}^{\text{Ru}} = C_{2p}^{\text{Zr}}$ , then  $\epsilon_2 \propto Nr$ , where the pair multiplicity difference  $r \equiv \frac{N_{\text{os}} - N_{\text{ss}}}{N_{\text{os}}}$  is precisely measured [STAR, PRC 105, 014901 (2022)]

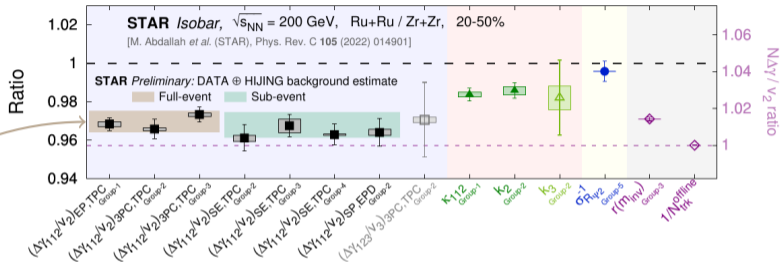
$$\Delta\epsilon_2/\epsilon_2 = \Delta r/r + \Delta N/N = (-2.95 \pm 0.08)\% + 4.4\% = (1.45 \pm 0.08)\%$$

- For a point of reference, AMPT simulation w.r.t. RP gives  $\Delta\epsilon_2/\epsilon_2 \approx (3.5 \pm 1.4)\%$

- ▶ 3p nonflow study in real data is difficult (work ongoing)
- ▶ We use HIJING simulation (which has no flow) to obtain  $\epsilon_3 \approx (1.84 \pm 0.04)\%$ , and  $\Delta\epsilon_3/\epsilon_3 = (0.5 \pm 2.7)\%$  ( $\sim 8.6 \times 10^8$  events for each isobar).
- ▶ HIJING without jet quenching gives  $\epsilon_3 = (2.24 \pm 0.05)\%$ , differing by 22%.
- ▶ We assign 50% systematic uncertainty for  $\epsilon_3$  ( $\pm 0.92\%$ ), and assume  $\Delta\epsilon_3/\epsilon_3$  is presently dominated by statistics.



# Estimated Background Level for Isobar $N\Delta\gamma/v_2$ Ratio



$$\frac{(N\Delta\gamma/v_2^*)^{\text{Ru}}}{(N\Delta\gamma/v_2^*)^{\text{Zr}}} \approx 1 + \frac{\Delta\epsilon_2}{\epsilon_2} - \frac{\Delta\epsilon_{\text{nf}}}{1 + \epsilon_{\text{nf}}} + \frac{\epsilon_3/\epsilon_2/(Nv_2^2)}{1 + \epsilon_3/\epsilon_2/(Nv_2^2)} \left( \frac{\Delta\epsilon_3}{\epsilon_3} - \frac{\Delta\epsilon_2}{\epsilon_2} - \frac{\Delta N}{N} - \frac{\Delta v_2^2}{v_2^2} \right)$$

$$= 1 + (1.45 \pm 0.08)\% + (0.65 \pm 0.11 \pm 0.22)\%$$

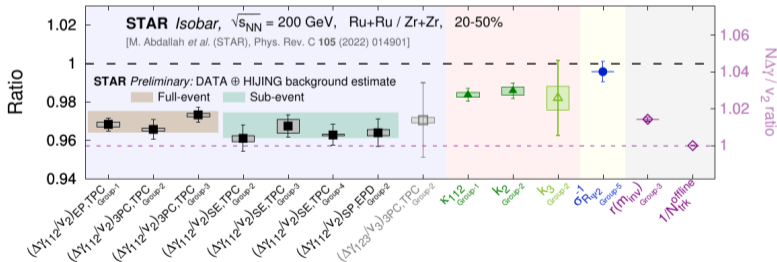
$$+ (0.094 \pm 0.007 \pm 0.048) [(0.5 \pm 2.7)\% - (1.45 \pm 0.08)\% - 4.4\% - (3.7 \pm 0.1 \pm 0.3)\%]$$

$$= 1 + (1.45 \pm 0.08)\% + (0.65 \pm 0.11 \pm 0.22)\% - (0.85 \pm 0.26 \pm 0.44)\%$$

$$= \mathbf{1.013 \pm 0.003 \pm 0.005} \quad (\text{full-event})$$

# Summary

- ▶  $v_2$  nonflow and 2p nonflow in  $C_3$  are measured. 3p nonflow in  $C_3$  is estimated by HIJING. Large degree of cancellation between 2p and 3p nonflow.
- ▶ New preliminary isobar background estimate  $\frac{(N\Delta\gamma/v_2^*)^{Ru}}{(N\Delta\gamma/v_2^*)^{Zr}} \approx (1.013 \pm 0.003 \pm 0.005)$  for full-event,  $(1.011 \pm 0.005 \pm 0.005)$  for sub-event.



# Backup

# Table: fit results on slide 8

<i>STAR preliminary</i>	Ru+Ru	Zr+Zr
$A_1$	$2.967 \pm 0.009$	$2.801 \pm 0.007$
$\sigma_{\Delta\eta,NS,W}$	$0.9878 \pm 0.0030$	$0.9550 \pm 0.0025$
$\sigma_{\Delta\phi,NS,W}$	$0.6329 \pm 0.0009$	$0.6364 \pm 0.0008$
$A_2$	$15.615 \pm 0.011$	$14.515 \pm 0.009$
$\sigma_{\Delta\eta,NS,N}$	$0.12668 \pm 0.00008$	$0.12839 \pm 0.00008$
$\sigma_{\Delta\phi,NS,N}$	$0.12889 \pm 0.00006$	$0.12977 \pm 0.00006$
$A_3$	$-72.522 \pm 0.018$	$-66.943 \pm 0.016$
$\sigma_{\Delta\eta,NS,D}$	$0.022288 \pm 0.000006$	$0.022314 \pm 0.000005$
$\sigma_{\Delta\phi,NS,D}$	$0.102971 \pm 0.000029$	$0.102619 \pm 0.000027$
$B$	$0.2140 \pm 0.0037$	$0.1943 \pm 0.0031$
$\sigma_{\Delta\eta,AS}$	$0.591 \pm 0.005$	$0.589 \pm 0.005$
$\sigma_{\Delta\phi,AS}$	$1.1 \times 10^5 \pm 18.3 \times 10^5$	$1.4 \times 10^5 \pm 11.7 \times 10^5$
$D$	$0.2759 \pm 0.0032$	$0.2660 \pm 0.0026$
$\sigma_{\Delta\eta,RG}$	$0.2600 \pm 0.0018$	$0.2524 \pm 0.0015$
$C$	$381.651 \pm 0.011$	$351.988 \pm 0.009$
$V_1$	$-0.001916 \pm 0.000006$	$-0.001943 \pm 0.000005$
$V_2$	$0.0029716 \pm 0.0000029$	$0.0028668 \pm 0.0000025$
$V_3$	$0.0001766 \pm 0.0000012$	$0.0001842 \pm 0.0000011$
$\chi^2/\text{NDF}$	$1018458.1/159982 = 6.4$	$1136361.1/159982 = 7.1$

## Fit function

$$\begin{aligned}
 & A_1 G_{NS,W}(\Delta\eta) G_{NS,W}(\Delta\phi) \\
 & + A_2 G_{NS,N}(\Delta\eta) G_{NS,N}(\Delta\phi) \\
 & + A_3 G_{NS,D}(\Delta\eta) G_{NS,D}(\Delta\phi) \\
 & + \frac{B}{2 - |\Delta\eta|} \operatorname{erf}\left(\frac{2 - |\Delta\eta|}{\sqrt{2}\sigma_{\Delta\eta,AS}}\right) \\
 & \quad \times G_{AS}(\Delta\phi \pm \pi) \\
 & + DG_{RG}(\Delta\eta) \\
 & + C[1 + 2V_1 \cos(\Delta\phi) \\
 & \quad + 2V_2 \cos(2\Delta\phi) + 2V_3 \cos(3\Delta\phi)]
 \end{aligned}$$

Large  $\sigma_{\Delta\phi,AS}$  turns  $G_{AS}$  into a flat line.

# Awayside $\Delta\eta$ correlation

Suppose two particles (1, 2) correlated in  $\eta$  by momentum conservation or other nonflow effect. We let

$$\begin{cases} \Delta\eta = \eta_1 - \eta_2 \\ \delta = \eta_1 + \eta_2 \end{cases} \Rightarrow \begin{cases} \eta_1 = \frac{\delta + \Delta\eta}{2} \\ \eta_2 = \frac{\delta - \Delta\eta}{2} \end{cases} \quad (1)$$

where  $\eta_1 = -\eta_2 + \delta$ . For momentum conservation, the two particles tend to be back-to-back in  $\eta$  direction ( $\eta_1 \sim -\eta_2$ ).  $\delta$  serves as fluctuations, and the correlation could be a function of  $\delta$ .

Since  $|\eta_1| < 1$  and  $|\eta_2| < 1$ , the range of  $\delta$  is

$$\begin{cases} \left| \frac{\delta + \Delta\eta}{2} \right| < 1 \\ \left| \frac{\delta - \Delta\eta}{2} \right| < 1 \end{cases} \Rightarrow |\delta| < 2 - |\Delta\eta| \quad (2)$$

Suppose the correlation function between two particles is  $f(\eta_1, \eta_2) = g(\Delta\eta, \delta)$ .

$$\begin{aligned} f(\eta_1, \eta_2) d\eta_1 d\eta_2 &= g(\Delta\eta, \delta) \frac{\partial(\eta_1, \eta_2)}{\partial(\Delta\eta, \delta)} d\Delta\eta d\delta \\ &= g(\Delta\eta, \delta) \begin{vmatrix} 1/2 & 1/2 \\ -1/2 & 1/2 \end{vmatrix} d\Delta\eta d\delta = \frac{1}{2} g(\Delta\eta, \delta) d\Delta\eta d\delta \end{aligned} \quad (3)$$

Integral over  $\delta$  to get the marginal distribution of  $\Delta\eta$

$$h(\Delta\eta) = \int_{-2+|\Delta\eta|}^{2-|\Delta\eta|} \frac{1}{2} g(\Delta\eta, \delta) d\delta \quad (4)$$

If there is no correlation, then

$g(\Delta\eta, \delta) = f(\eta_1, \eta_2) = f(\eta_1)f(\eta_2) = \frac{1}{4}$ , and the integral becomes  $h(\Delta\eta) = \frac{1}{4}(2 - |\Delta\eta|)$ , the acceptance triangle.



# Awayside $\Delta\eta$ correlation

An intuitive assumption of the correlation from momentum conservation is  $\delta$  obeys a Gaussian distribution centering at 0, which is  $\delta \sim \mathcal{N}(0, \sigma)$ .

$$g(\Delta\eta, \delta) = \frac{1}{\sigma\sqrt{2\pi}} \exp\left(-\frac{\delta^2}{2\sigma^2}\right) \quad (5)$$

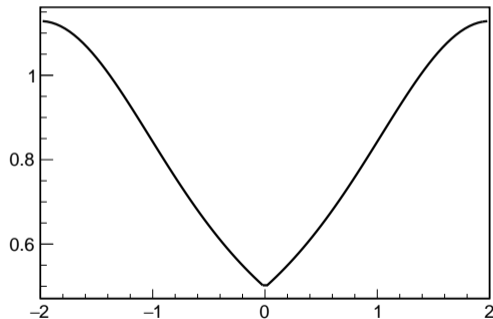
(may differ by a constant factor). And the marginal distribution becomes

$$h(\Delta\eta) = \frac{1}{2} \operatorname{erf}\left(\frac{2 - |\Delta\eta|}{\sqrt{2}\sigma}\right) \quad (6)$$

After the acceptance correction, the function form should be

$$\frac{1}{2 - |\Delta\eta|} \operatorname{erf}\left(\frac{2 - |\Delta\eta|}{\sqrt{2}\sigma}\right) \quad (7)$$

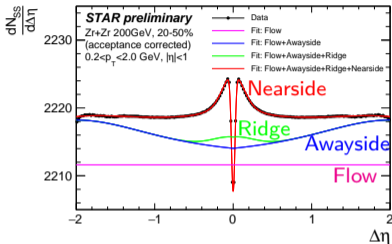
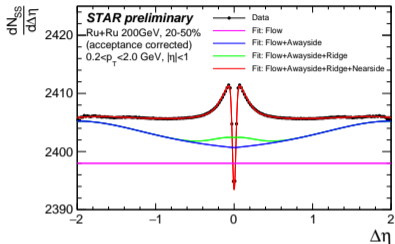
If we set  $\sigma = 1$ , then the function looks like below, which seems similar to the STAR data shape at awayside (large  $|\Delta\phi|$ ).



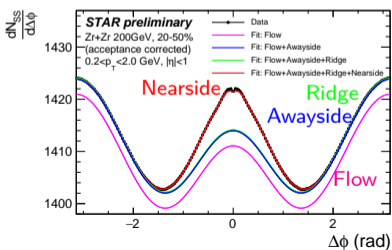
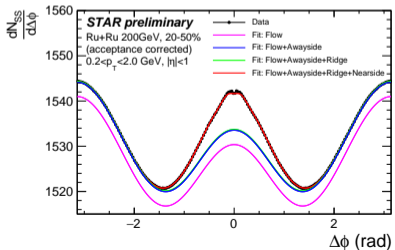
Ru+Ru

Zr+Zr

$\Delta\eta$  projections



$\Delta\phi$  projections



- ▶ Data: markers in black.
- ▶ Flow: flow component in fit.
- ▶ Flow+Awayside
- ▶ Flow+Awayside+Ridge: the  $\Delta\phi$  ridge is a 1D Gaussian centering at  $\Delta\eta = 0$ , which is small and independent from  $\Delta\phi$ .
- ▶ Flow+Awayside+Ridge+Nearside: the total fit function, where the nearside includes 3 2D Gaussians centering at  $\Delta\eta = 0$ ,  $\Delta\phi = 0$ , which are the wide, the narrow, and the dip.

# Estimated Background Components for Isobar $N\Delta\gamma/v_2$ Ratio

Quantity		Method	Systematic uncertainty	Full-event value	Sub-event value
Multiplicity $\Delta N/N$	Measured		Negligible	4.4%	4.4%
Flow $\Delta v_2^2/v_2^2$	Measured	Nonflow subtracted as per below	From nonflow syst.	$\Delta v_2^2/v_2^2 = (3.7 \pm 0.1 \pm 0.3)\%$	$\Delta v_2^2/v_2^2 = (3.7 \pm 0.1 \pm 0.3)\%$
$v_2$ nonflow	Measured	$(\Delta\eta, \Delta\phi)$ correlations, experimentally measured	Nonflow $\sim 25\%$ (full event), dominated by NS wide Gaus; consider $\pm 1/2$ WG as syst. uncertainty	$-\Delta\epsilon_{\text{nf}} = (0.82 \pm 0.13 \pm 0.30)\%$ $\frac{-\Delta\epsilon_{\text{nf}}}{1+\epsilon_{\text{nf}}} = (0.65 \pm 0.11 \pm 0.22)\%$	$-\Delta\epsilon_{\text{nf}} = (0.59 \pm 0.15 \pm 0.27)\%$ $\frac{-\Delta\epsilon_{\text{nf}}}{1+\epsilon_{\text{nf}}} = (0.48 \pm 0.12 \pm 0.22)\%$
$v_2$ -induced bkgd: $\epsilon_2 = N\Delta\gamma/v_2$	Measured	Measured by ZDC (assume negligible CME)	Small	$\epsilon_2 = (0.57 \pm 0.04 \pm 0.02)\%$	$\epsilon_2 = (0.79 \pm 0.05 \pm 0.01)\%$
$v_2$ -induced bkgd difference: $\frac{\Delta\epsilon_2}{\epsilon_2} \sim \frac{\Delta(N_{2p}/N)}{(N_{2p}/N)} = \frac{\Delta(rN)}{rN}$	Measured	$r = (N_{\text{OS}} - N_{\text{SS}})/N_{\text{OS}}$ experimentally measured	Negligible	$\frac{\Delta\epsilon_2}{\epsilon_2} = (1.45 \pm 0.08)\%$	$\frac{\Delta\epsilon_2}{\epsilon_2} = (1.45 \pm 0.08)\%$
3p contribution to $C_3$ : $\epsilon_3 = C_{3p}N_{3p}/(2N)$	Model estimate	HIJING simulations quenching-on	Quenching-on and off difference $\sim 20\%$ . Take $\pm 50\%$ as syst. uncertainty	$\epsilon_3 = (1.84 \pm 0.04 \pm 0.92)\%$	$\epsilon_3 = (1.91 \pm 0.09 \pm 0.95)\%$
3p contribution difference: $\Delta\epsilon_3/\epsilon_3$	Model estimate	HIJING simulation quenching-on	Assumed negligible relative to the large stat. uncertainty	$\frac{\Delta\epsilon_3}{\epsilon_3} = (0.5 \pm 2.7)\%$ $\frac{\epsilon_3/\epsilon_2}{Nv_2^2} = 0.104 \pm 0.008 \pm 0.053$	$\frac{\Delta\epsilon_3}{\epsilon_3} = (-1.8 \pm 6.3)\%$ $\frac{\epsilon_3/\epsilon_2}{Nv_2^2} = 0.079 \pm 0.006 \pm 0.040$
background estimate				$1.013 \pm 0.003 \pm 0.005$	$1.011 \pm 0.005 \pm 0.005$

# Outlook

- ▶ Improve the  $\epsilon_3$  estimate.
- ▶ Improve the  $(\Delta\eta, \Delta\phi)$  2D fittings.
- ▶ Background estimates for each centrality bin separately.