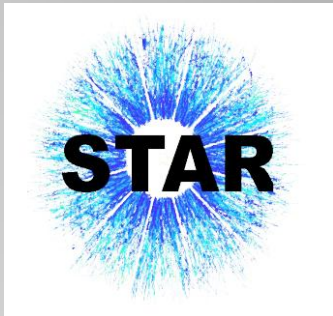


RAPIDITY DEPENDENT FIXED-TARGET PION SPECTRA AT $\sqrt{s_{NN}} = 3.0, 3.5, \text{ and } 4.5 \text{ GeV}$ FROM STAR



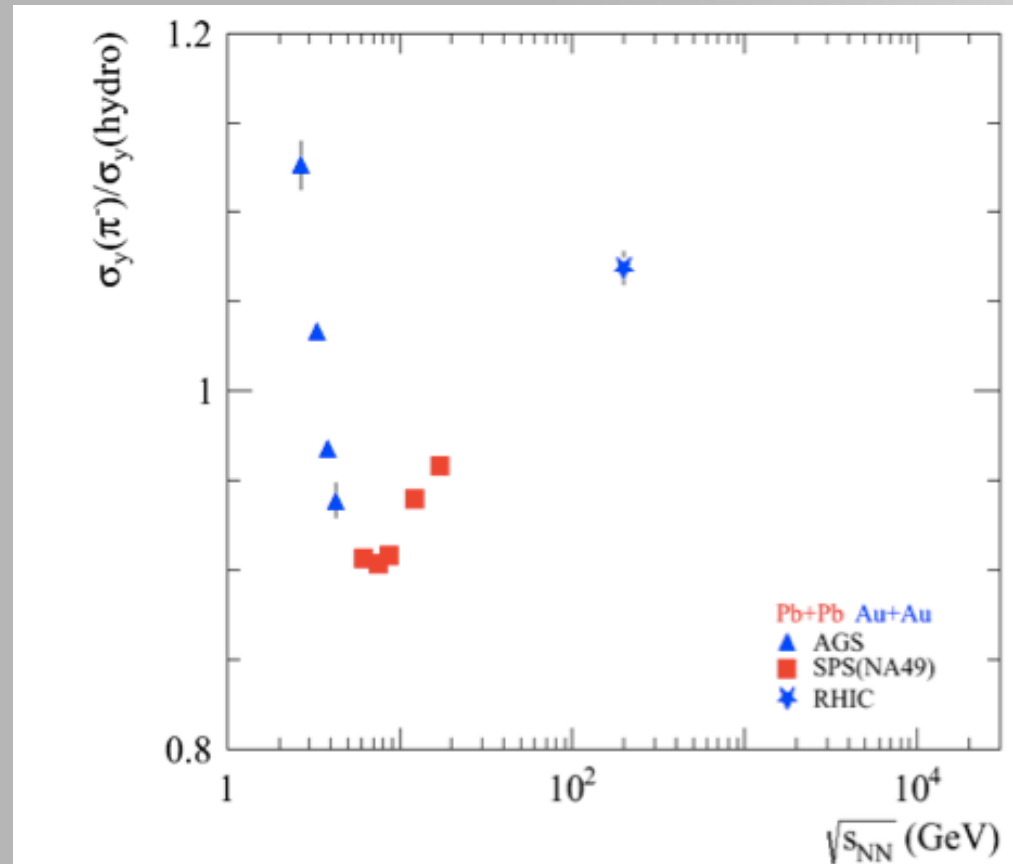
*Brooke Haag- LBNL/ UC Davis
For the STAR Collaboration
(Presented by Daniel Cebra)*

Presented at the Annual Meeting of the Division of Nuclear Physics
October 23-26, 2013
Newport News, VA



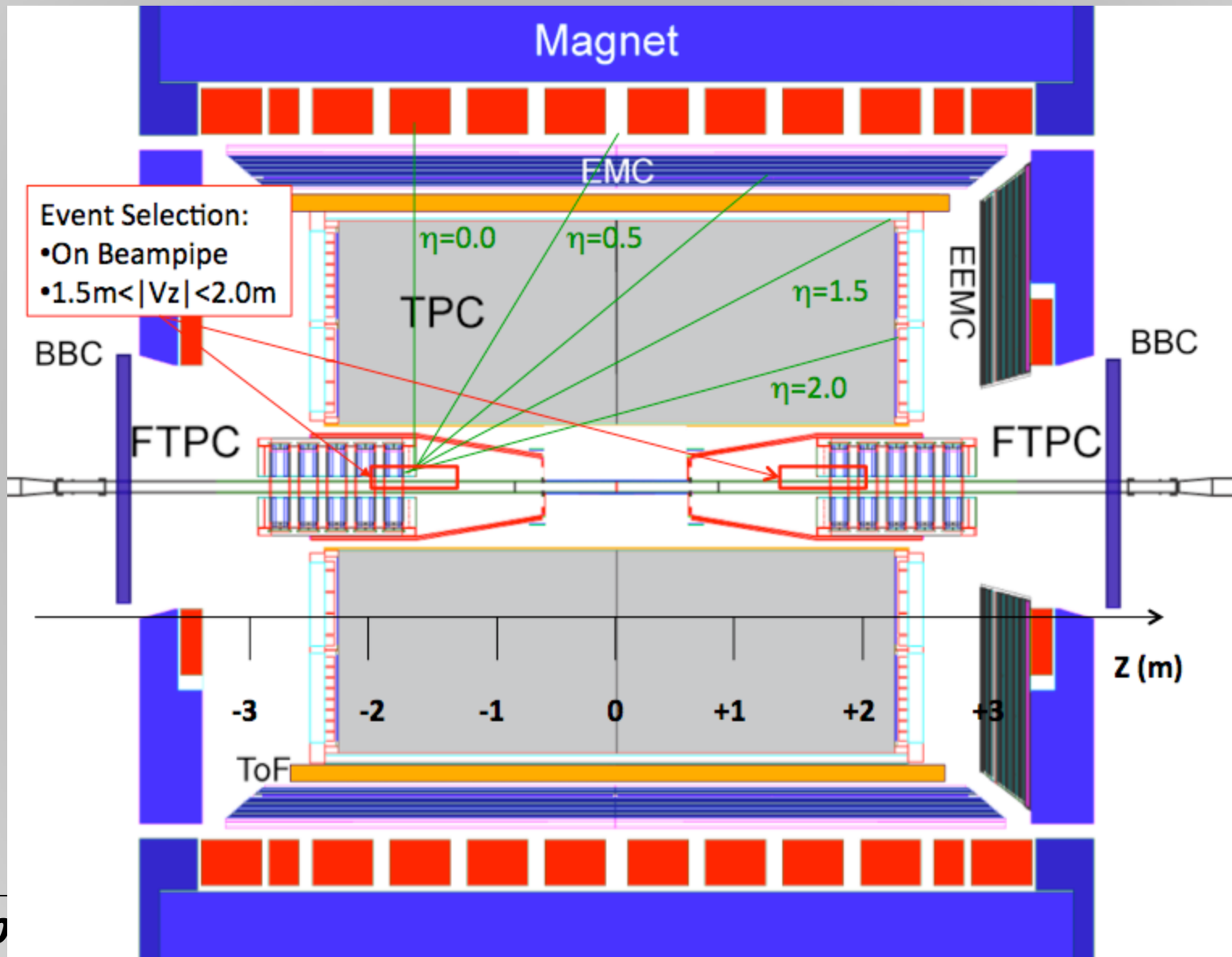
MOTIVATION

- Minimum in the width of the pion rapidity distribution, indicative of a softening of the EOS expected during first order phase transition
- The goals are then to cover a broad rapidity range at low colliding energies



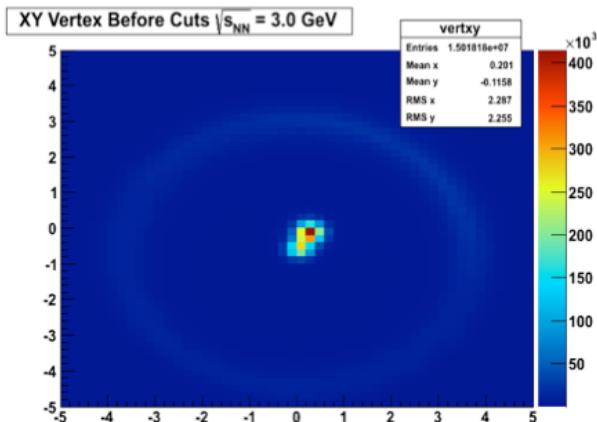
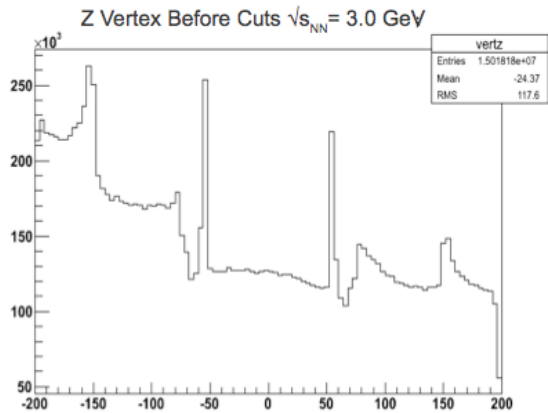
M. Bleicher hep-ph/0509314, H. Petersen nucl-th/0611001, A. Rustamov arXiv:1201:4520

FIXED-TARGET EVENTS



EVENT SELECTION

Before Cuts



Event Cuts

- Vertex Requirements -

$$-200 \leq V_z \leq -150 \text{ cm}$$

$$2 \leq V_R \leq 5 \text{ cm}$$

- Momentum -

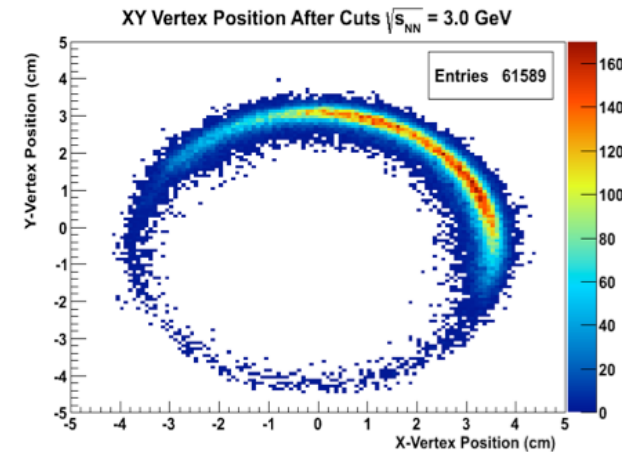
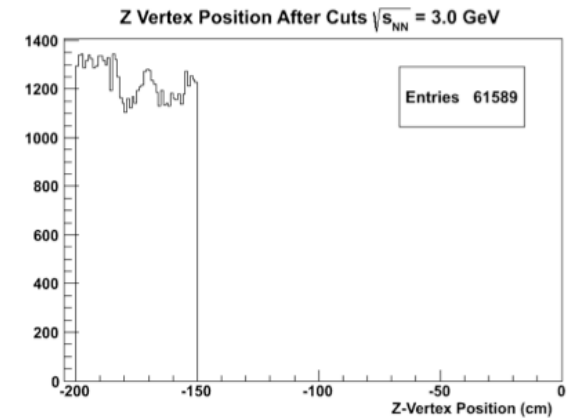
$$\sum_{\text{Tracks}} p_z > 0$$

- Centrality Cut -

Top 10 %

Apply Cuts

After Cuts



KINEMATIC CALCULATIONS

Collision Energy (GeV)	Single Beam Energy	Single Beam P_z (GeV/c)	Fixed Target \sqrt{s}	Single Beam Rapidity	Center of Mass Rapidity
19.6 Au+Au	9.8	9.76	4.47 Au+Al	3.04	1.52
11.5 Au+Au	5.75	5.67	3.53 Au+Al	2.51	1.25
7.7 Au+Au	3.85	3.74	2.99 Au+Al	2.10	1.05

$$\sqrt{(s_{NN})} = \sqrt{(2m^2 + 2Em)}$$

$$m = 0.9315 \text{ GeV}/c^2; E = 9.8 \text{ GeV}$$

$$\sqrt{(s_{NN})} = 4.47 \text{ GeV}$$

$$p_z = \sqrt{(E^2 - m^2)} = 9.76 \text{ GeV}/c$$

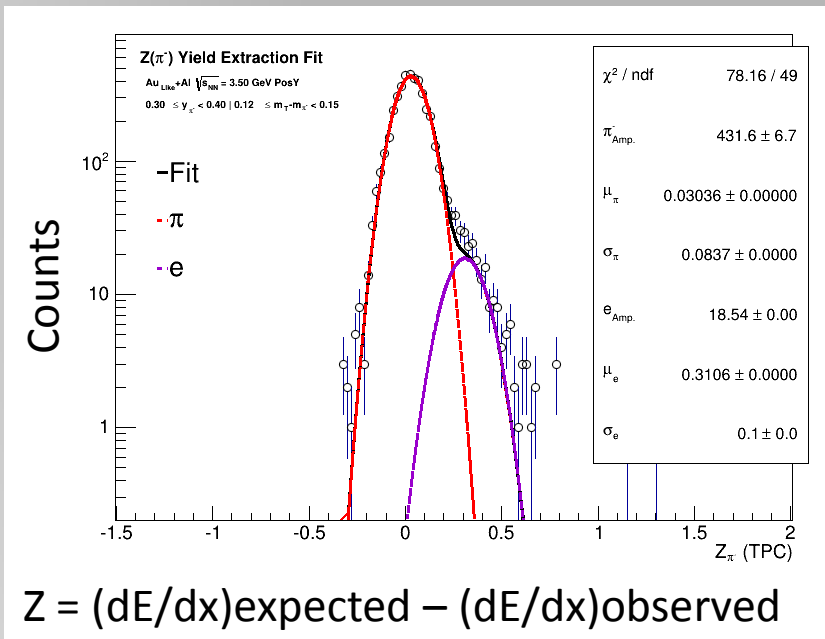
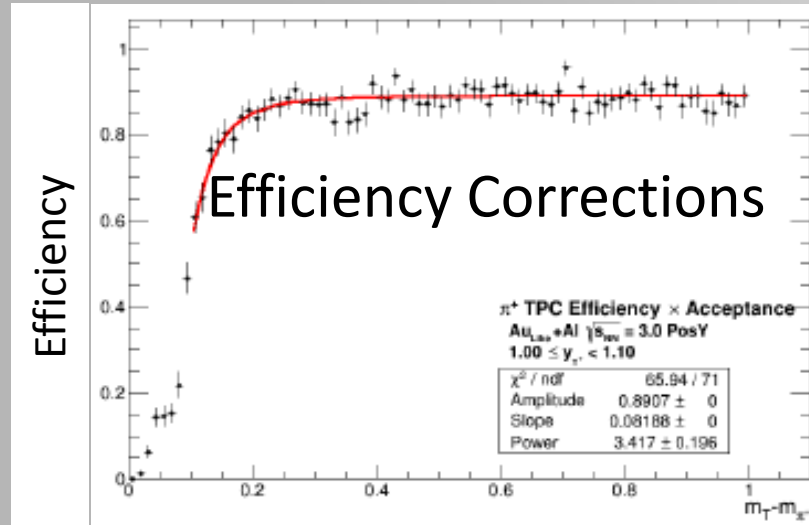
$$y_{\text{beam}} = 0.5 * [\ln(E + p_z) / (E - p_z)]$$

$$y_{\text{beam}} = 3.0$$

$$y_{\text{cm}} = 1.5$$

RAW YIELD EXTRACTION AND CORRECTIONS

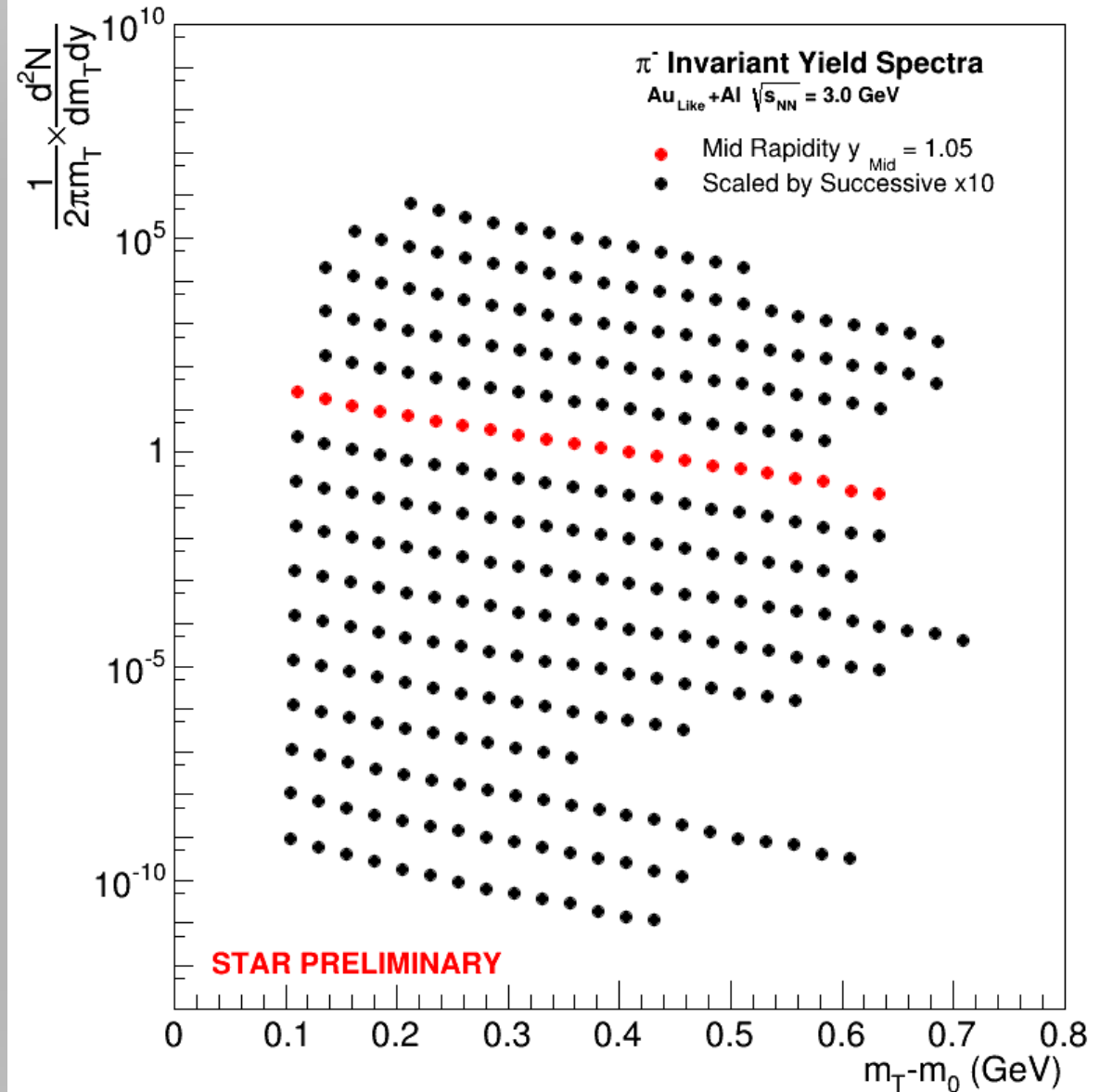
The dE/dx distributions for negative particles were fit for defined rapidity and $m_t - m_0$ bins. Efficiency and other detector effects.



- Electrons could be easily be separated from the π^- s.
- Kaons would overlap in the forward rapidity slides, however the mid-rapidity K^-/π^- ratio is less than 2%
- even fewer anti-protons

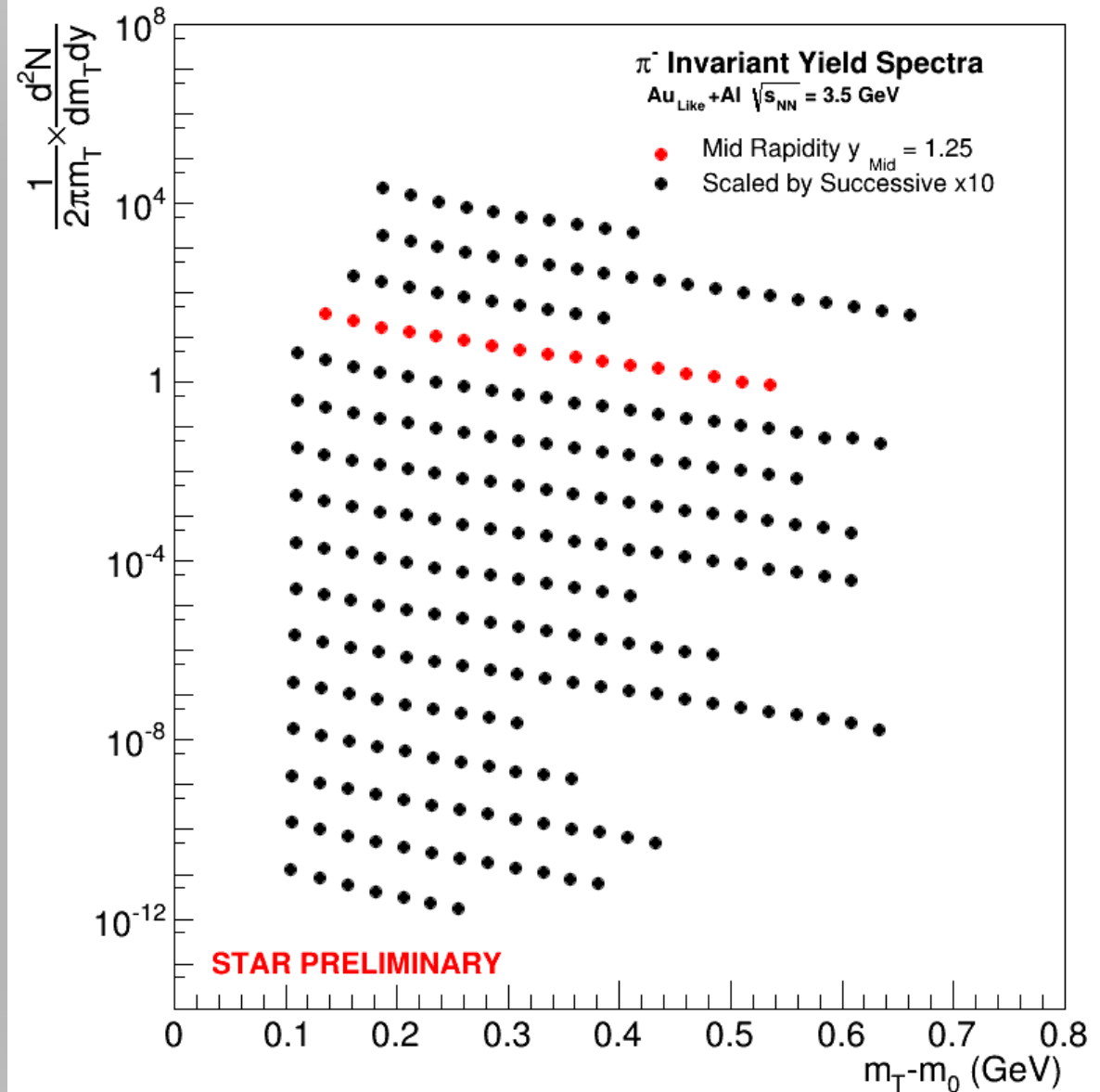
$\sqrt{s_{NN}} = 3.0 \text{ GeV}$

- Each set of points is a different rapidity slice; bottom to top: $0.0 < y < 0.1$ to $1.5 < y < 1.6$
- Curves are offset by powers of ten
- The nucleon-nucleon center-of-mass rapidity slice is indicated in red



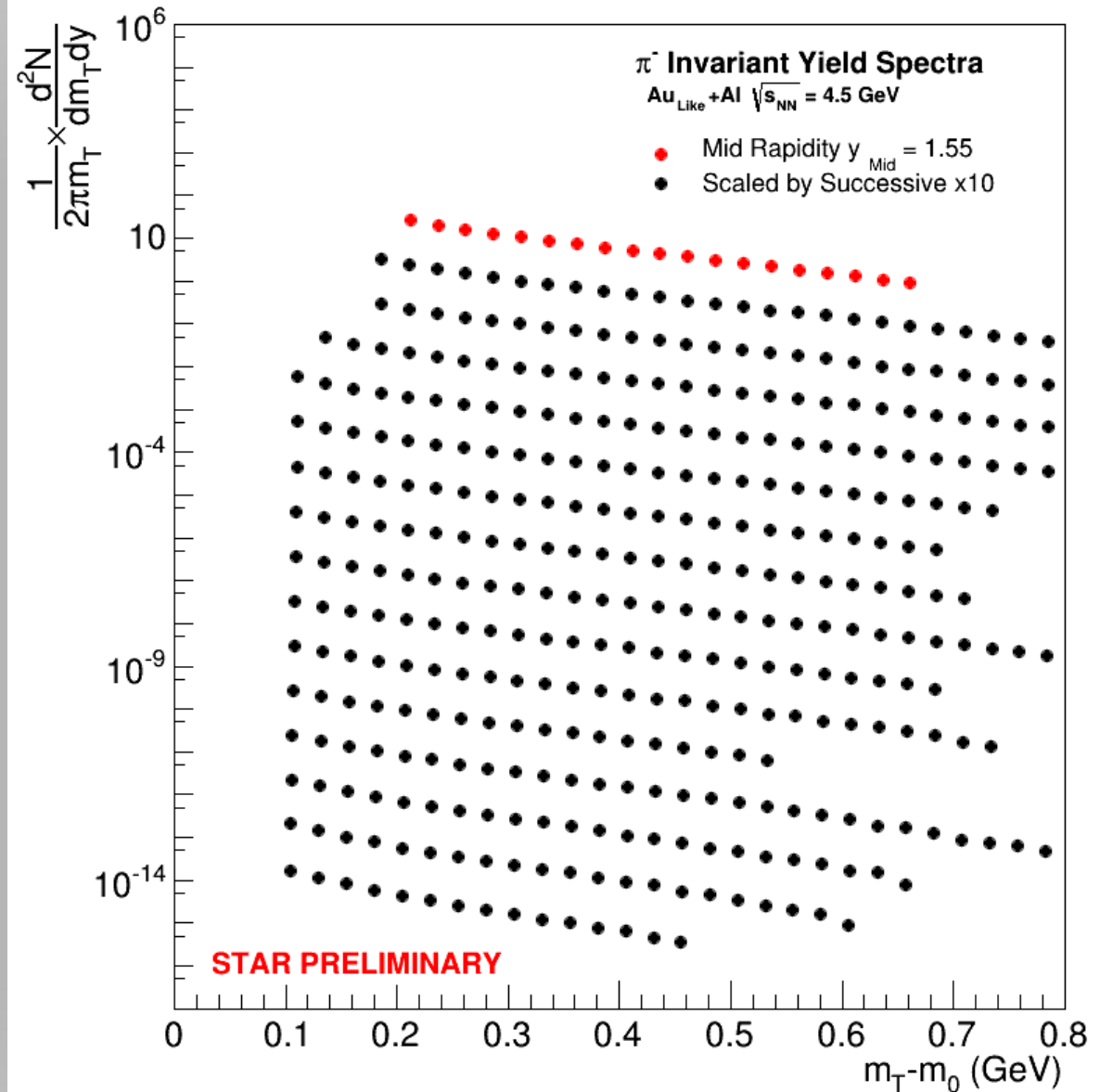
$\sqrt{s_{NN}} = 3.5 \text{ GeV}$

- Each set of points is a different rapidity slice; bottom to top: $0.0 < y < 0.1$ to $1.5 < y < 1.6$
- Curves are offset by powers of ten
- The nucleon-nucleon center-of-mass rapidity slice is indicated in red



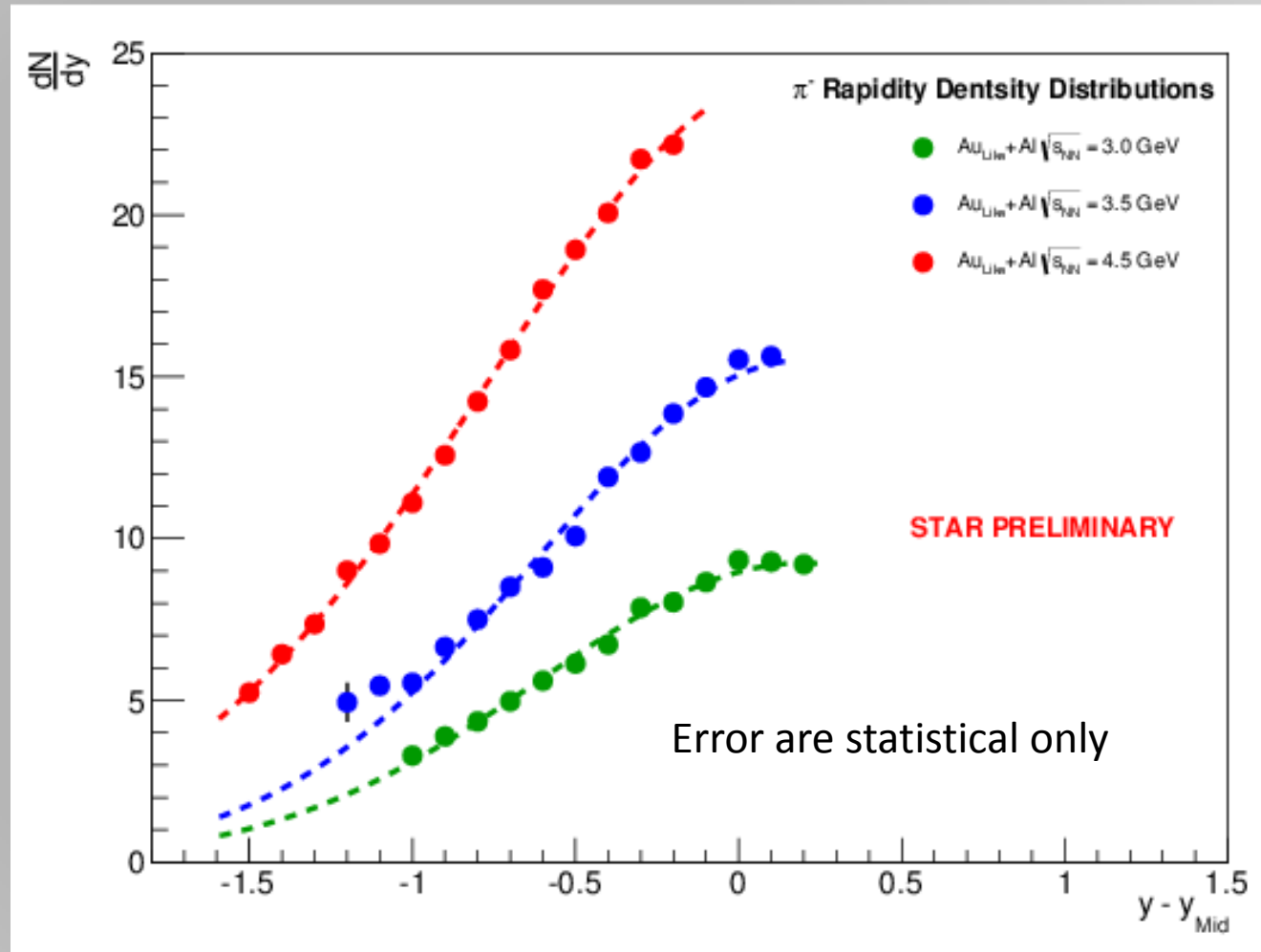
$$\sqrt{s_{NN}} = 4.5 \text{ GeV}$$

- Each set of points is a different rapidity slice; bottom to top: $0.0 < y < 0.1$ to $1.5 < y < 1.6$
- Curves are offset by powers of ten
- The nucleon-nucleon center-of-mass rapidity slice is indicated in red



Rapidity Density Distributions

- The rapidity density distributions have been measured from target rapidity to the nucleon-nucleon center-of-mass rapidity.
- The distributions are fit with Gaussians.
- dN/dy distributions for $Au_{\text{Like}}+Al$ may not peak at zero.



CONCLUSIONS AND OUTLOOK

- We have made the first measurement of broad rapidity dependence of spectra in STAR
- This proof of principle study demonstrates we *can make the dale* measurement, which is an important measurement indicative of the softening of the equation of state and a first-order phase transition
- We are looking forward to fixed target run as we are primed to make the dale measurement with Au+Au collisions and dedicated trigger (currently in development)
- STAR is an excellent detector, still doing important physics as we push its capabilities beyond expectations