

Dielectron Production in Au+Au Collisions from STAR

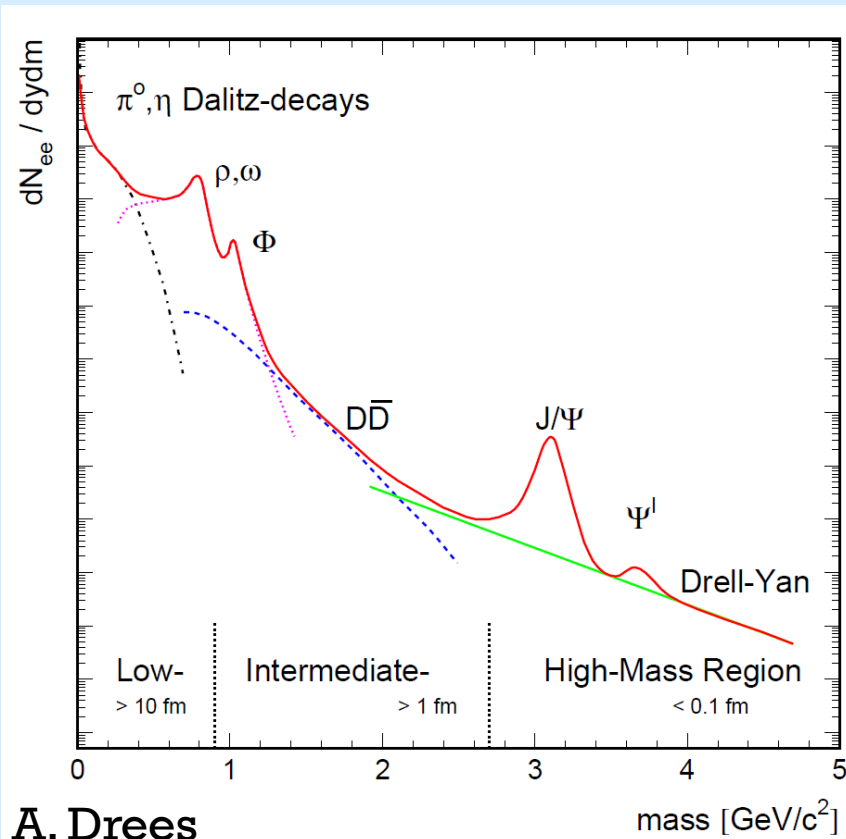
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[For the STAR Collaboration]
Rice University
November 6th, 2013



Outline

- **Motivation**
- **STAR**
- **Analysis Techniques**
- **Data + Observations**
- **Summary & Outlook**

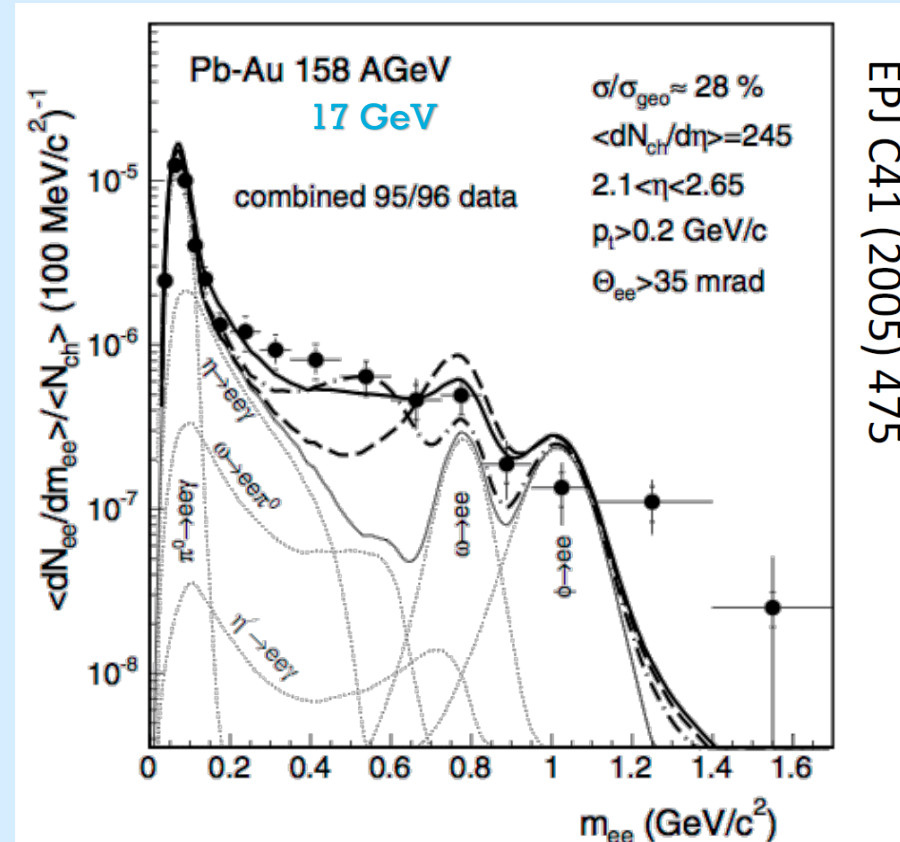
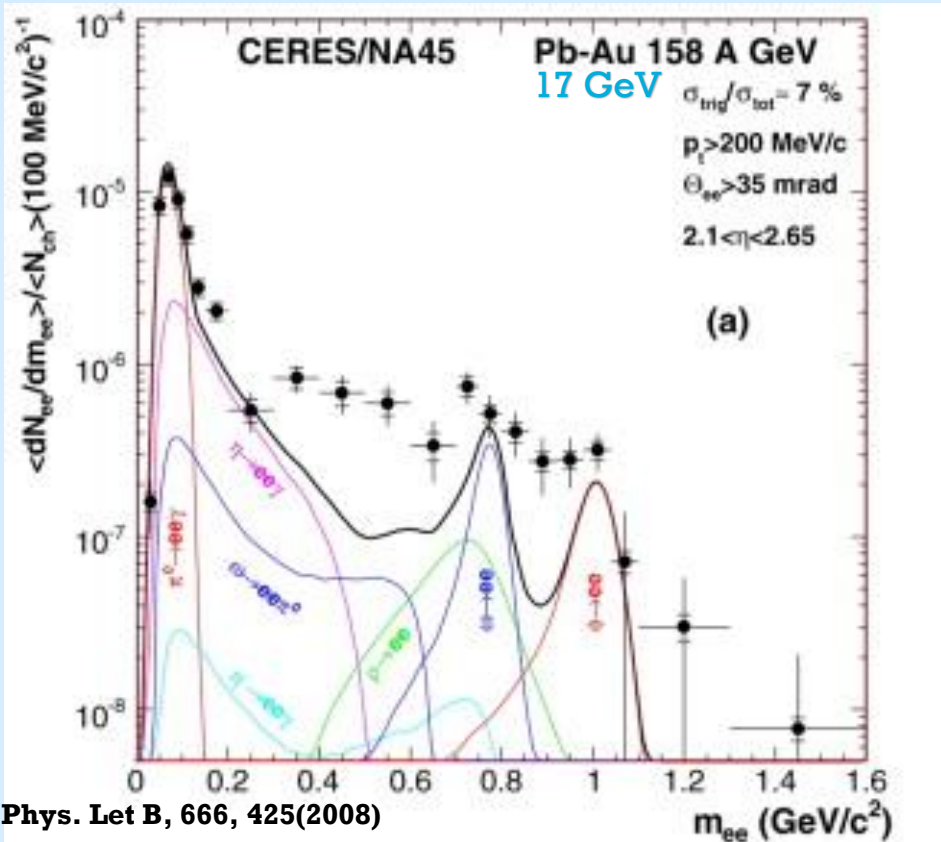
Dielectrons



- **Excellent Probe**
 - Minimal final state interactions
 - Generated at all stages of the collision
- **Chronological Phases [early to latest]**
 - High Mass Region [HMR]
 - Drell-Yan
 - $J/\psi + \Upsilon$ suppression
 - Intermediate Mass Region [IMR]
 - Heavy flavor modification
 - QGP thermal radiation
 - Low Mass Region [LMR]
 - Vector meson modification
 - Possible link to chiral symmetry restoration

Modification of ρ -meson

- CERES Measurements
 - Vacuum ρ not sufficient
- Possible explanations
 - Vacuum ρ [dash]
 - Mass dropping [dot-dash]
 - Broadening of spectral function [solid]



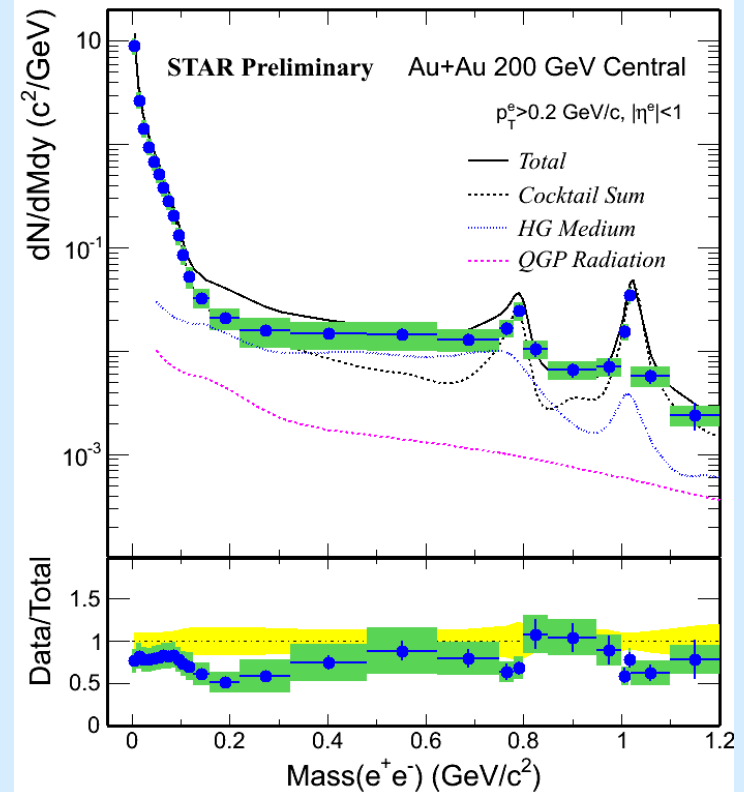
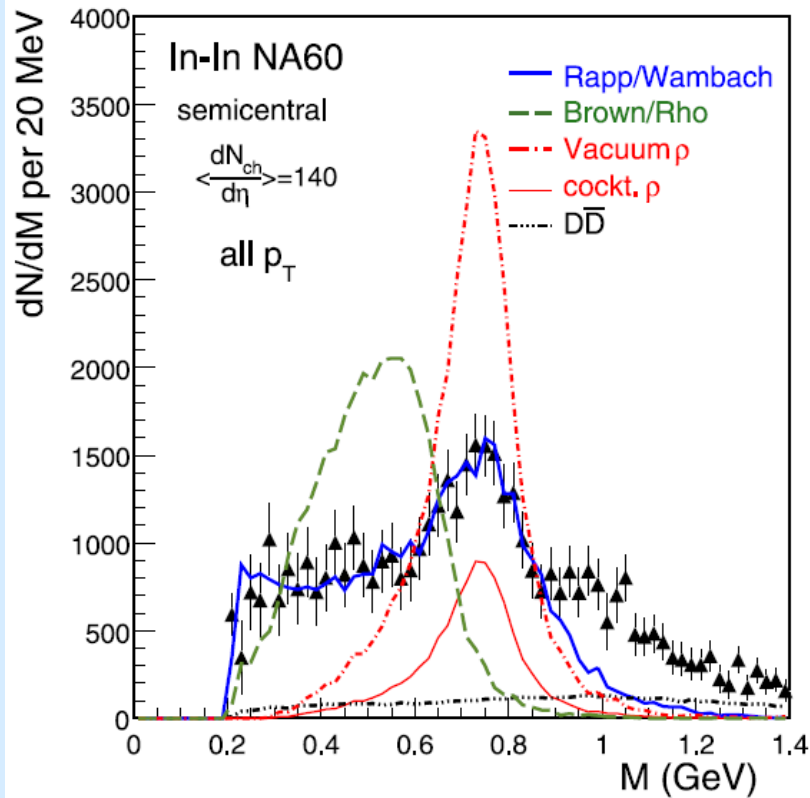
Modification of ρ -meson

- NA 60

- Vacuum ρ is inadequate [**dash-dot**]
- Excludes mass-dropping [**dash**]
- Supports broadening of ρ spectral function [**solid**]

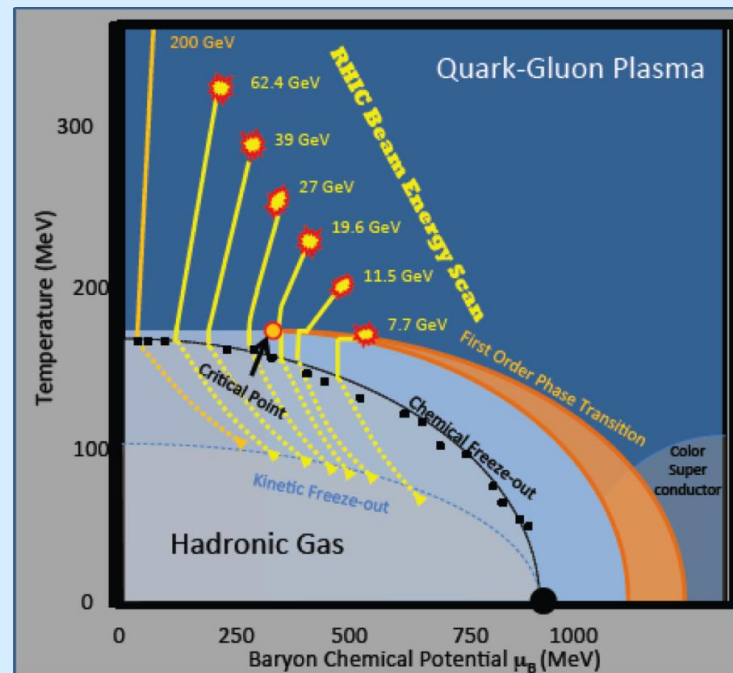
- RHIC AuAu 200

- Vacuum ρ inadequate
- Supports ρ broadening



Beam Energy Scan

- RHIC Beam Energy Scan Program [2010-2021]
 - Au+Au @19.6, 27, 39, & 62.4 GeV
 - Same colliding species & detector
 - Opportunity to extensively study ρ spectral function
 - Connect between SPS & RHIC Au+Au 200 GeV
 - Dependence on \sqrt{s} ?
 - Compare to models



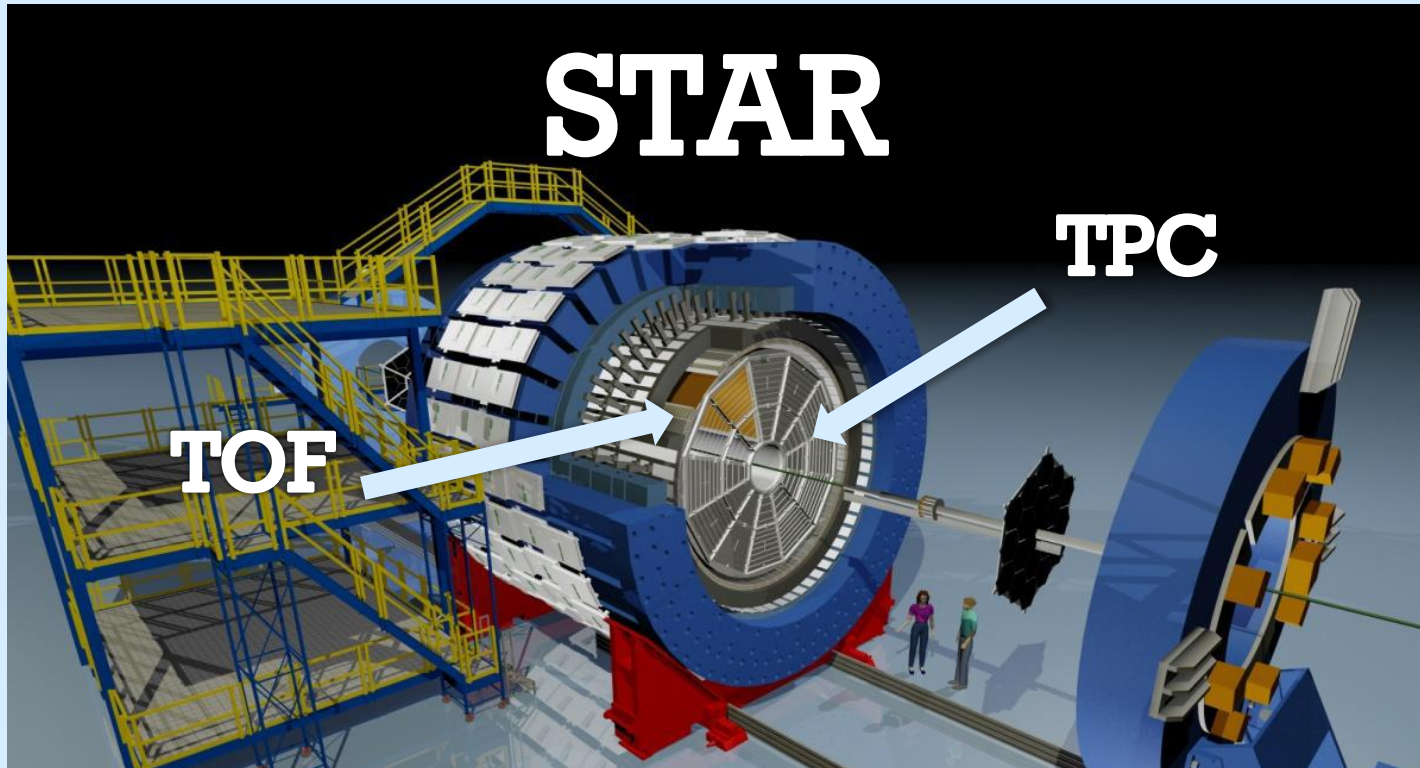
STAR Detector

Time Projection Chamber [TPC]

- Tracking
- Ionization energy loss
- Full azimuthal coverage

Time of Flight [TOF]

- Precise timing ($<90\text{ps}$)
- Improves TPC's purity
- Full azimuthal coverage



Electron Identification

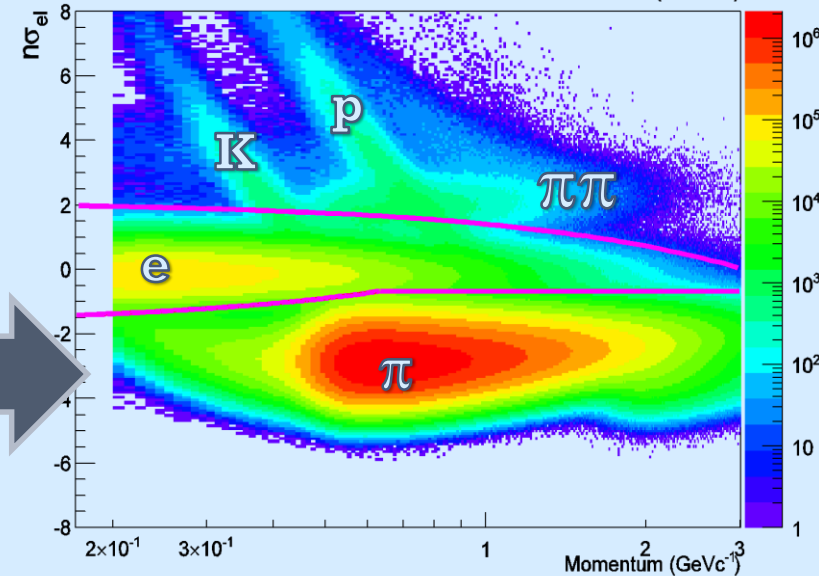
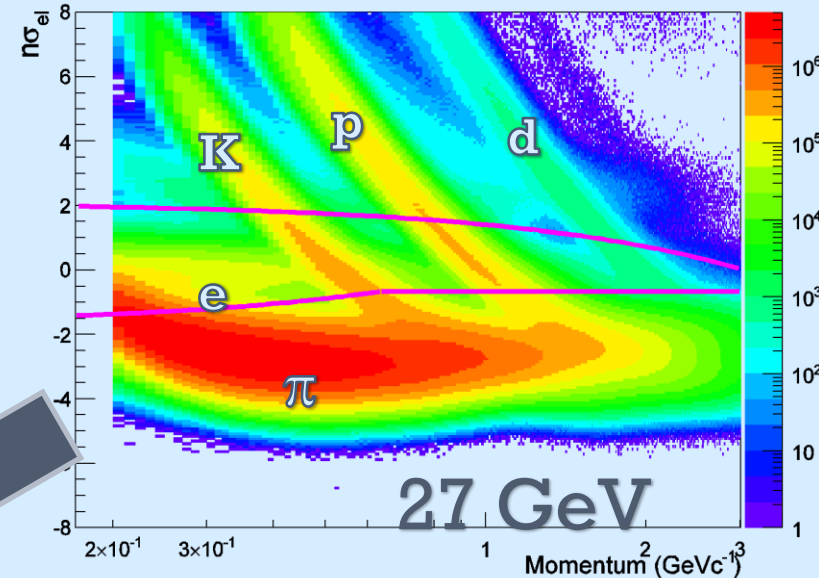
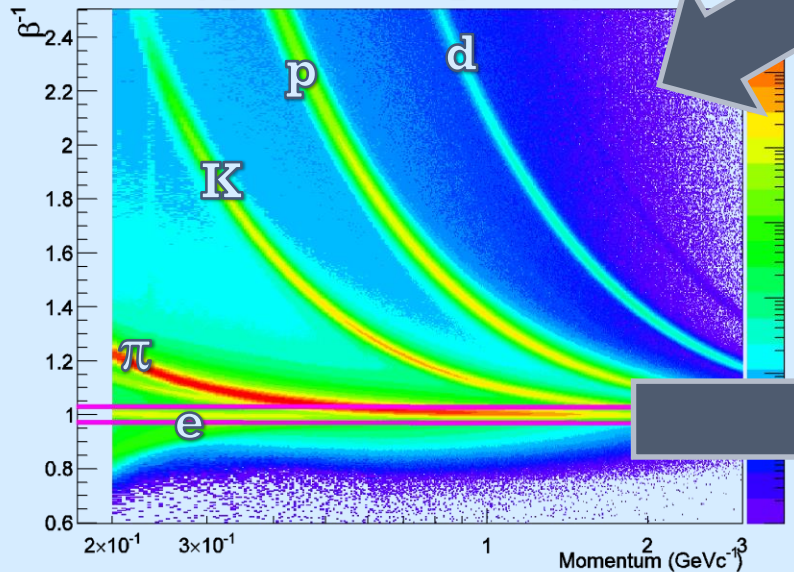
- **Combine the TPC+TOF**

- **TPC provides:**

$$n\sigma_{el} = \left(R_{dE/dx}\right)^{-1} \ln \frac{\langle dE/dx \rangle^{\text{exp}}}{\langle dE/dx \rangle^{\text{th}}}$$

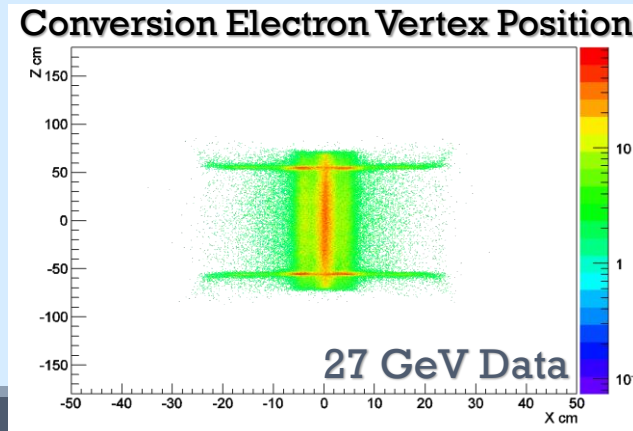
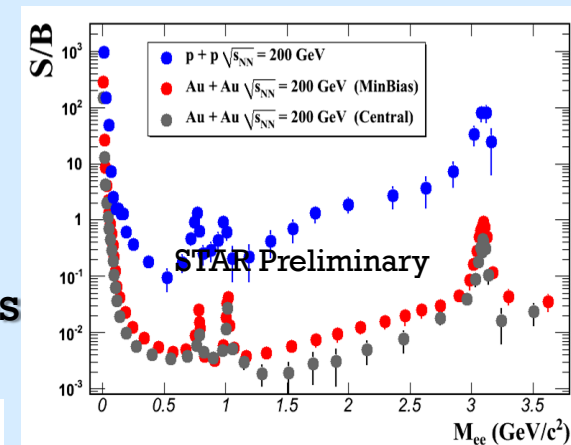
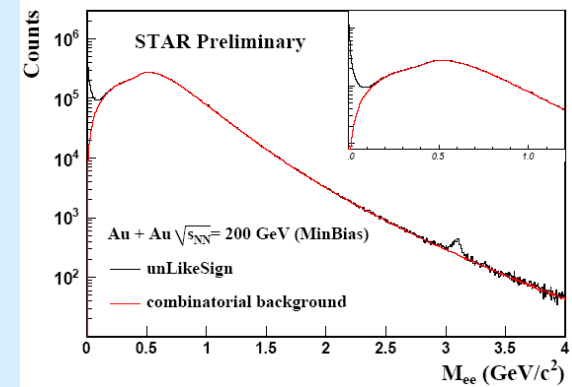
- **Use TOF to remove slow hadrons**

- **Improves purity**
 - **Typical eID purity ~95%**



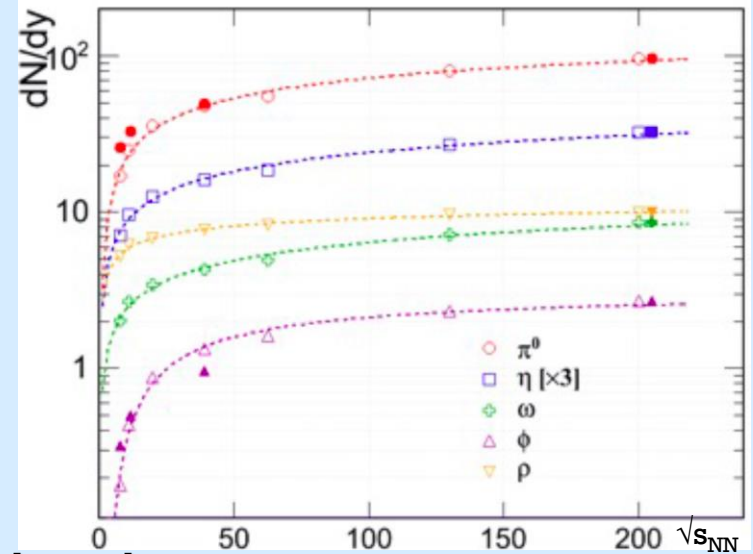
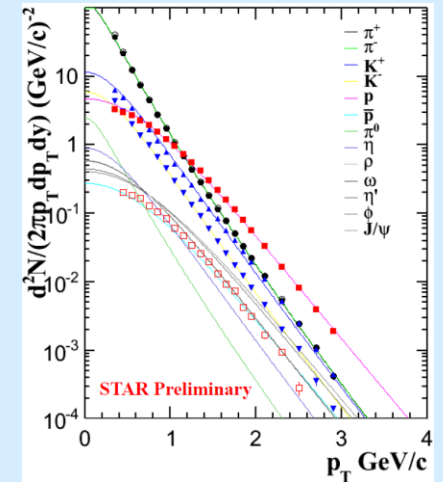
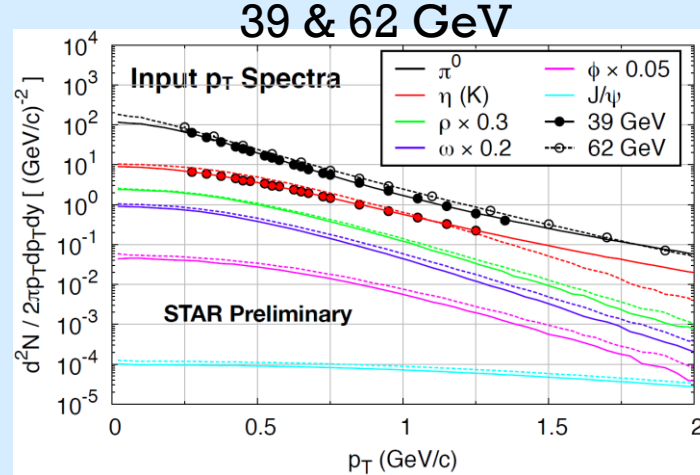
Background Removal

- Types of background: 3 C's
 - Combinatorial, Correlated, & Conversion
- Techniques
 - Same event like-sign pairs [LS]
 - Combine all like-sign pairs & average
 - Removes combinatorial & correlated
 - Like-sign/unlike-sign acceptance difference
 - Corrected with mixed events
 - Mixed event pairs [ME]
 - Pair e^+/e^- from different events w/ similar properties
 - Removes combinatorial
 - Pair cuts [PC]
 - Removes conversions



Cocktail Determination

19.6 GeV



- **Simulation Inputs**

- Flat ϕ & η

- p_T spectra

- 19.6 GeV

- Tsallis fits to NA49

- 39 & 62 GeV

- π^0 from $\pi^{-/+}$ & η from K spectra @ STAR

- Everything else: AMPT

- **Decay**

- Kroll-Wada Formula ($VM \rightarrow \gamma^* \rightarrow \ell^+ \ell^-$)

- **Yields**

- 19.6 GeV

- Meson/ π^0 ratio from NA49 & π from STAR

- 39 & 62 GeV

- Extrapolated from STAR 200 GeV using AMPT's \sqrt{s} -dependence

- **Correlated charm decays**

- Simulated with PYTHIA and scaled by N_{binary}

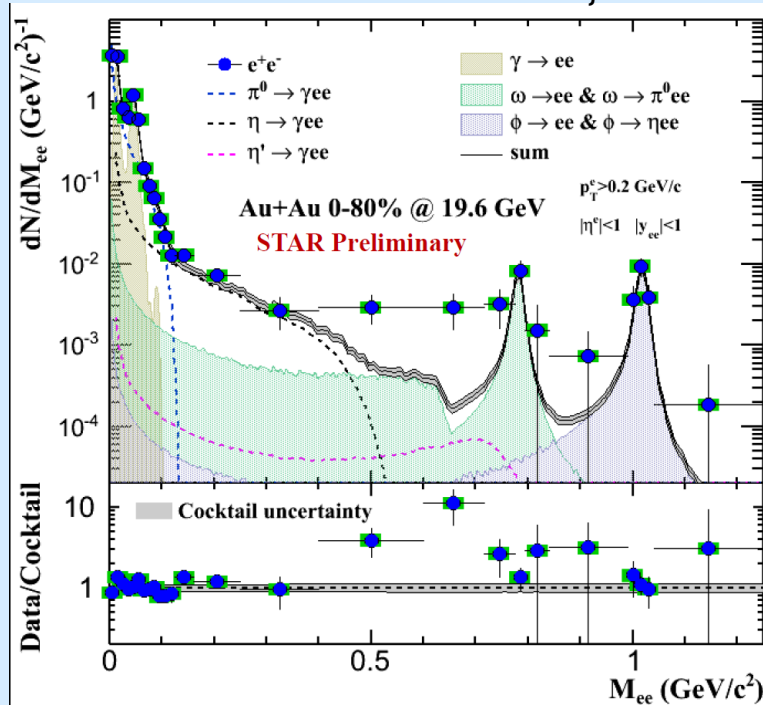
Beam Energy Scan Inv. Mass

$\sqrt{s_{NN}}$ [GeV]	19.6	27	39	62	200(Y10)
Min. Bias Events	28M	70M	99M	55M	240M

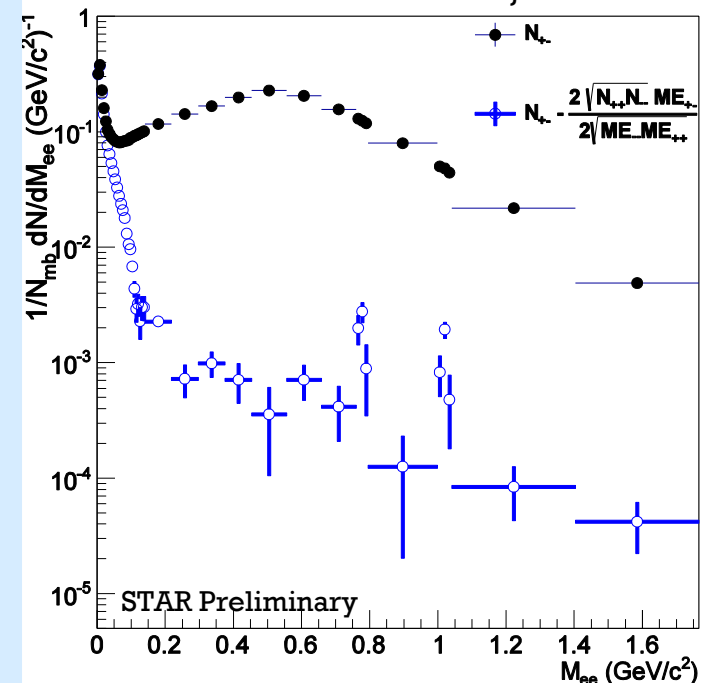
STAR Acceptance cuts:

$p_{Te} > 0.2$ GeV/c, $|\eta_e| < 1.$, and $|Y_{ee}| < 1$
 [Y_{ee} not used in 27GeV]

Corrected Invariant Mass, 19GeV



Raw Invariant Mass, 27 GeV



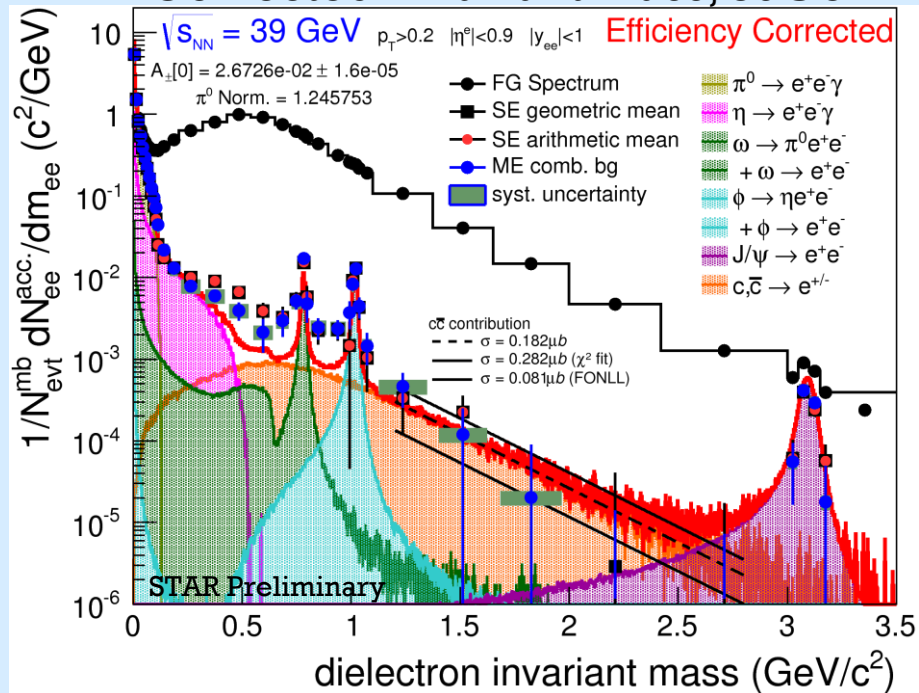
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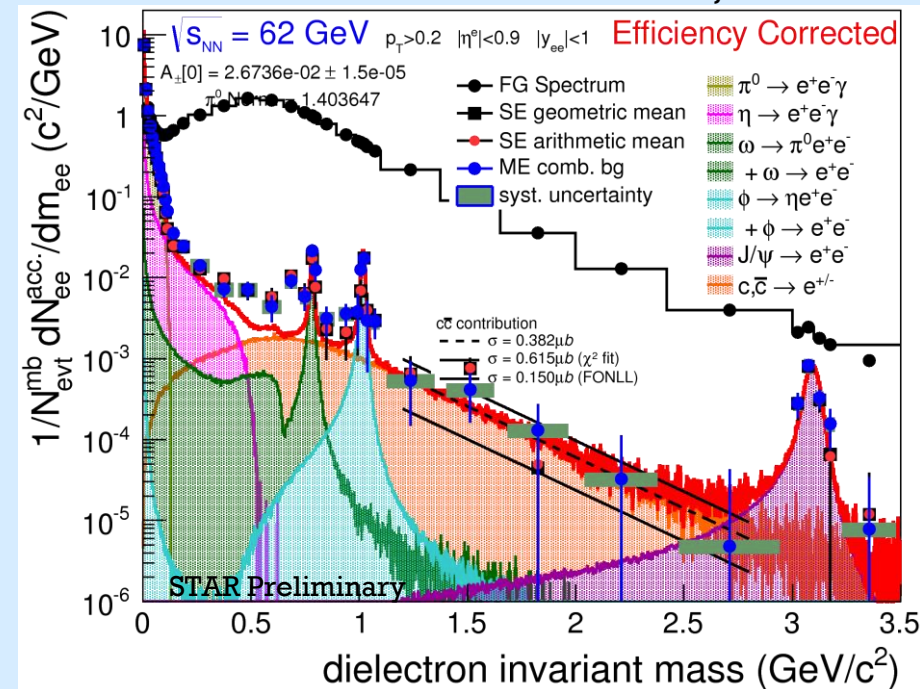
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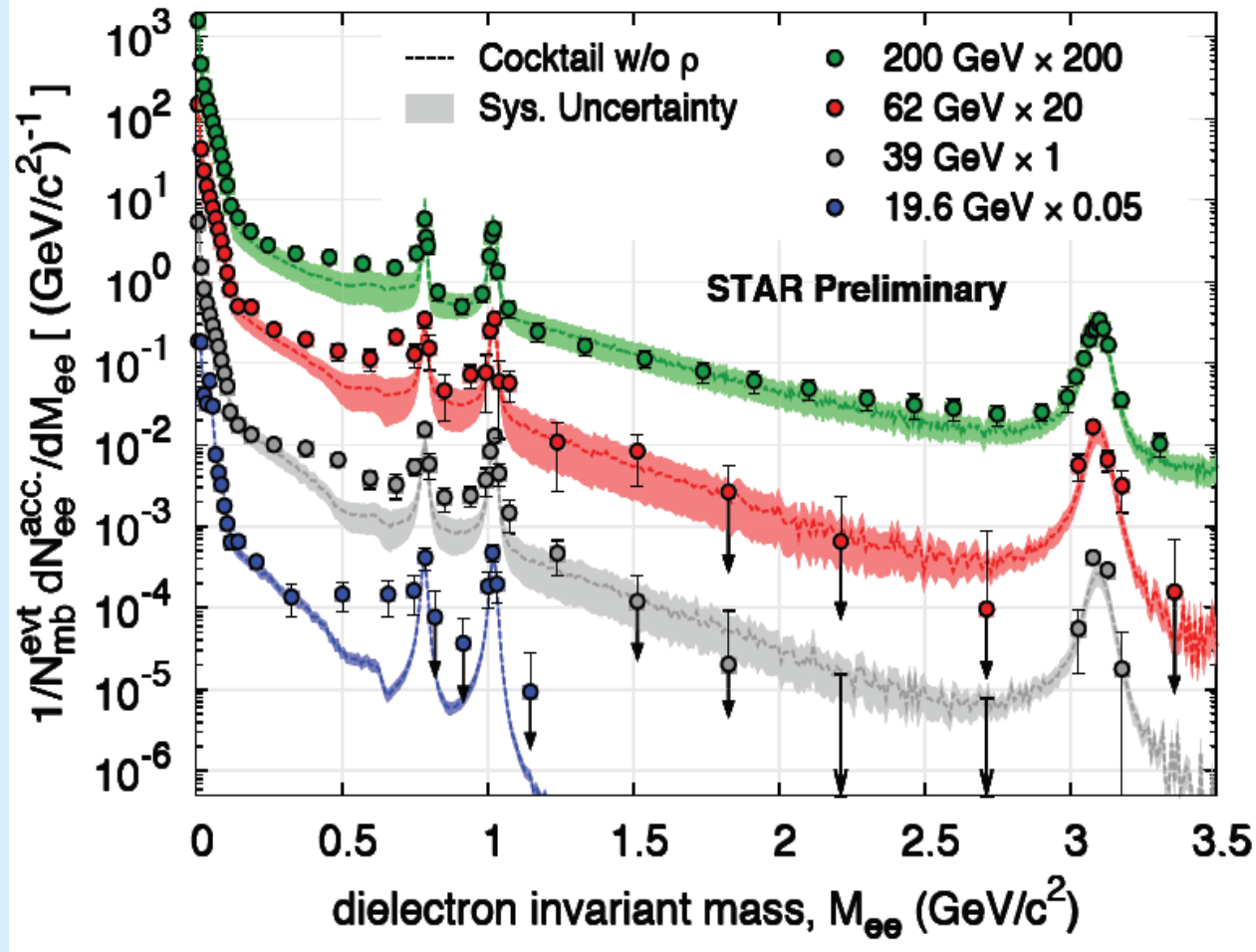
Corrected Invariant Mass, 39GeV



Corrected Invariant Mass, 62GeV

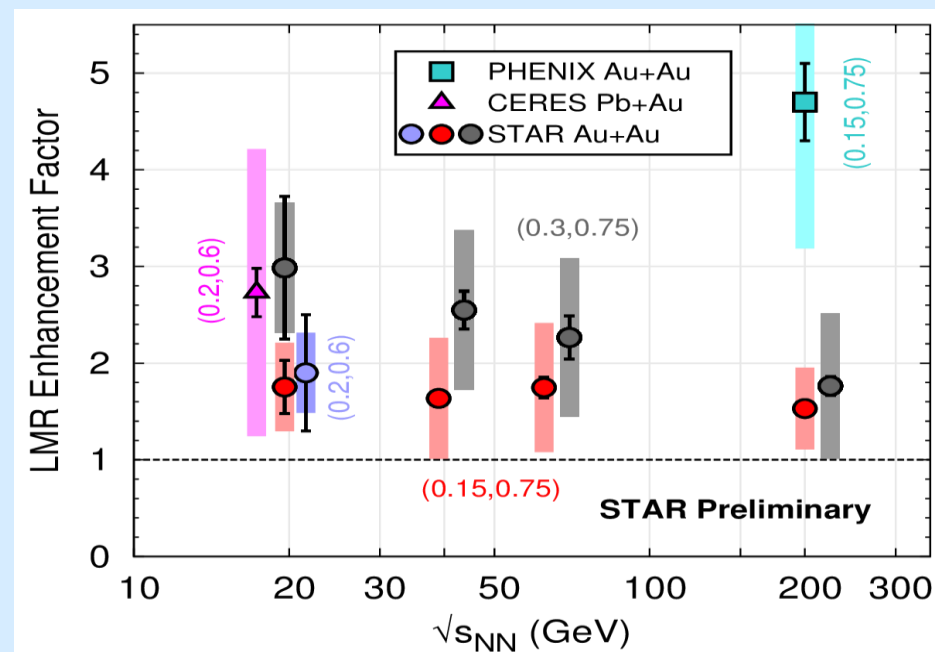
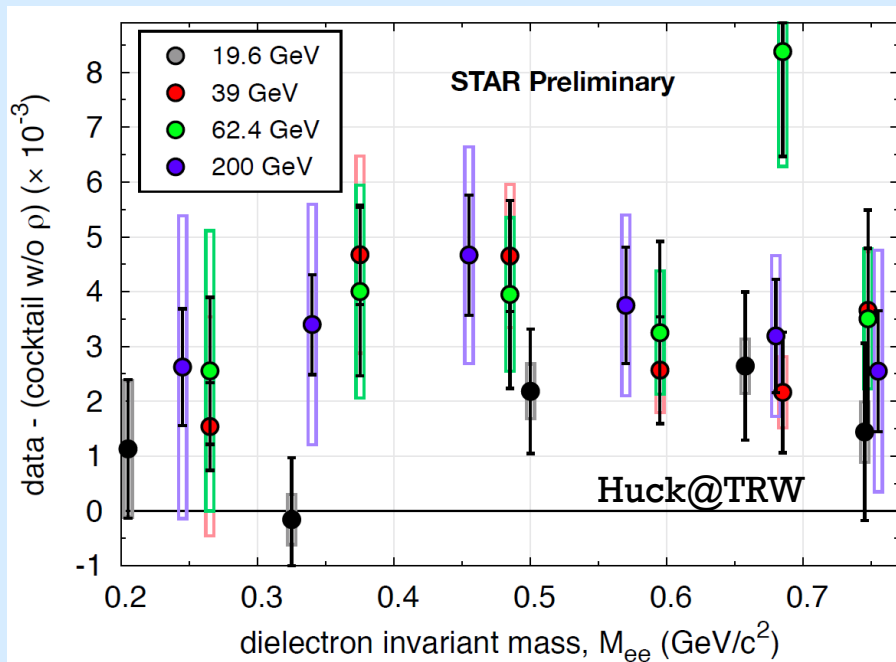


Observations



- Cocktail does not include ρ -meson
- LMR excess at all $\sqrt{s_{\text{NN}}}$

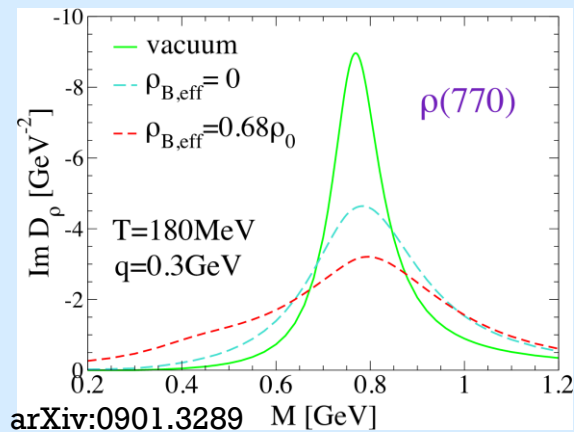
Observations II



- CERES + PHENIX have different acceptances than STAR

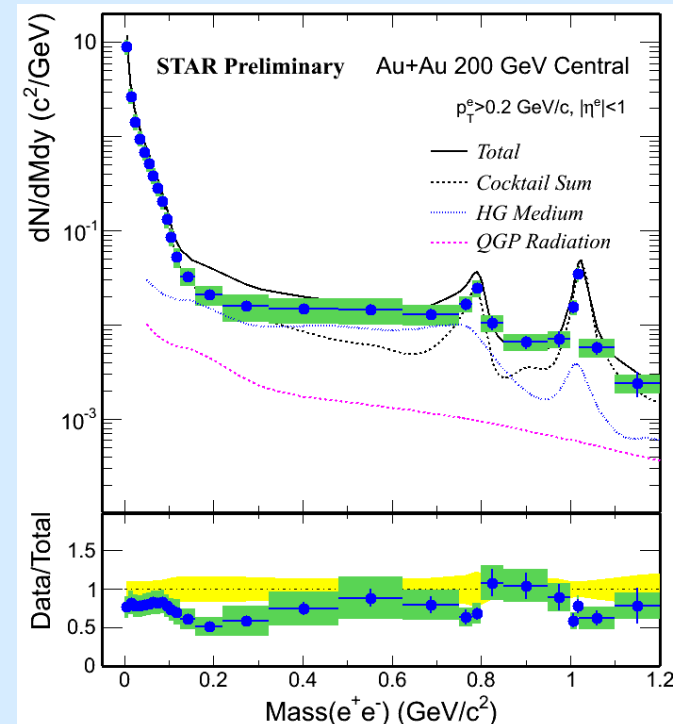
- ρ -meson modifications at all $\sqrt{s_{NN}}$
- The absolute excess yield (data - cocktail) shows no significant $\sqrt{s_{NN}}$ dependence given uncertainties
- The Enhancement Factor is sensitive to the mass range selected
 - The charm component in the cocktail baseline increases with $\sqrt{s_{NN}}$

Theory: Rapp, Wambach, van Hees [RWvH]



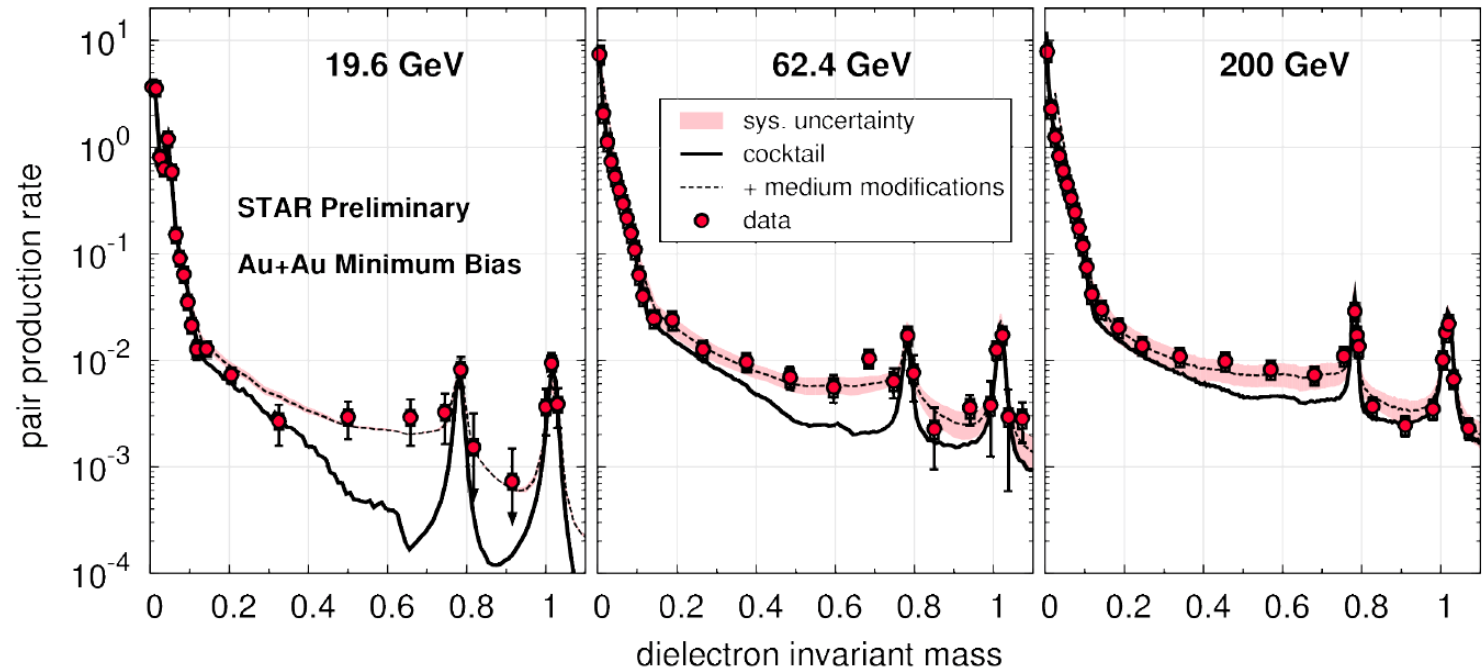
- Hadronic phase: ρ -meson “melts” when extrapolated to phase transition boundary
 - Total baryon density dependent
- Top-down extrapolated QGP rate coincides with bottom-up extrapolated hadronic rates

- STAR’s Run 10 AuAu Central 200GeV
- Model curves provided by Rapp
- Complete evolution (HG + QGP)
- Agrees within uncertainties



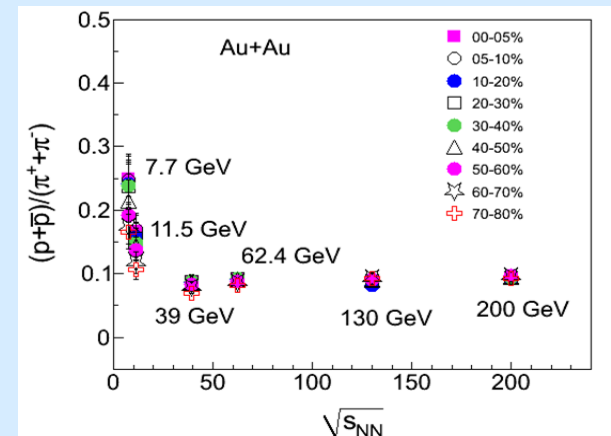
R.Rapp, PRC 63 (2001) 054907. Rapp & Wambach, EPJ A 6 (1999) 415. Calculations via Priv. Comm. w/ Rapp

BES Comparisons to RWvH

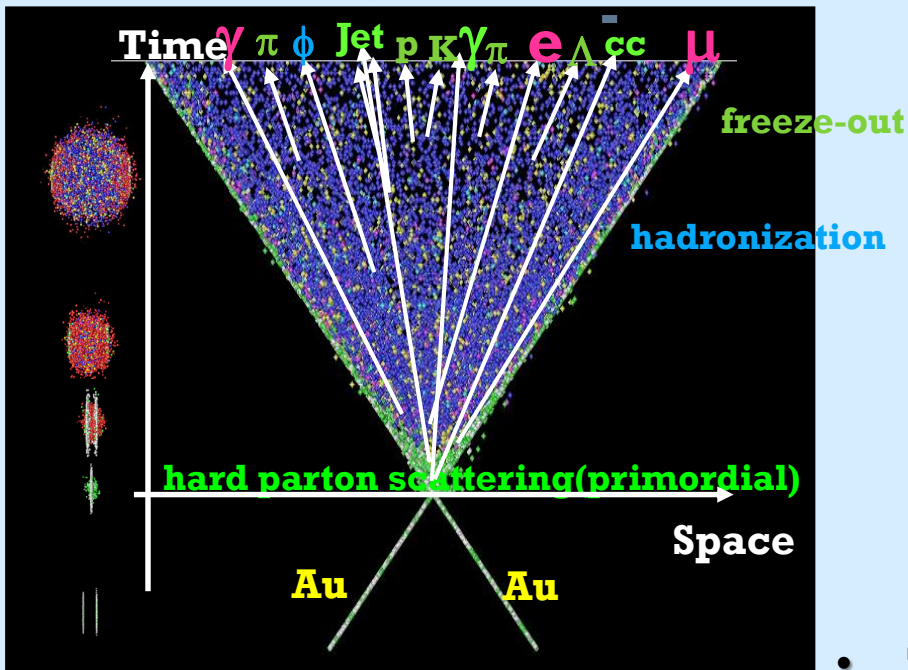


Rapp + Wambach, Adv. Nucl. Phys. 25, 1 (2000). Phys. Rev. 363, 85 (2002). Calculations via Priv. Comm. w/ Rapp

- **Cocktail w/out ρ contributions [solid curve]**
- **Cocktail w/ medium modified ρ [dashed line]**
- **Data consistent with ρ -meson broadening**
 - Model dependent on total baryon density
- **Tests extensive function of the ρ -meson's spectral function**



Study on Direct Virtual Photon



- Relation between real photon production and the associated ee pair production

$$\frac{d^2 N_{ee}}{dm_{ee} dp_T} = \frac{2\alpha}{3\pi} \frac{1}{m_{ee}} L(m_{ee}) S(m_{ee}, p_T) \frac{dN_\gamma}{dp_T}$$

$$L(m_{ee}) = \sqrt{1 - \frac{4m_e^2}{m_{ee}^2} \left(1 + \frac{2m_e^2}{m_{ee}^2}\right)}$$

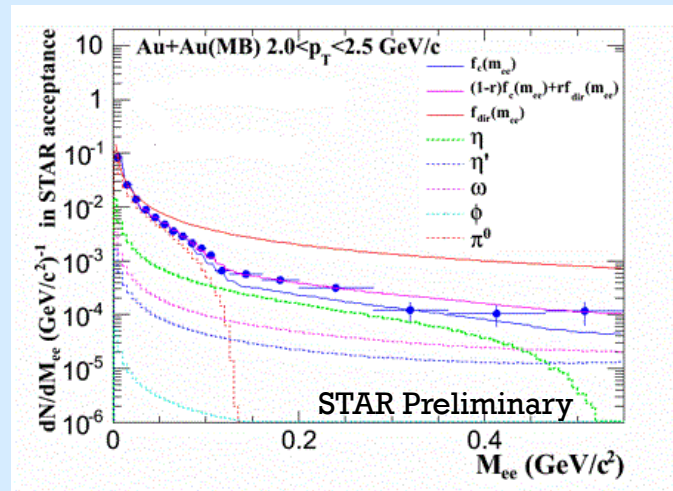
for $m_{ee} \gg m_e$ && $p_T \gg m_{ee}$

$$S \sim 1 \quad L \sim 1$$

$$\frac{d^2 N_{ee}}{dm_{ee}} \approx \frac{2\alpha}{3\pi} \frac{1}{m_{ee}} dN_\gamma$$

Normalize f_{dir}

- Two component fit to quantify the excess



$$f = (1-r)f_c + rf_{dir}$$

cocktail component

direct virtual photon component

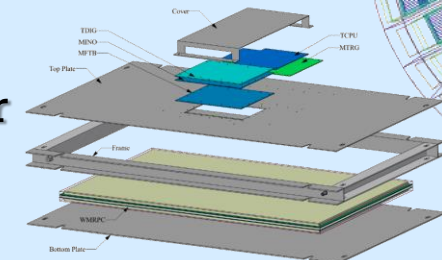
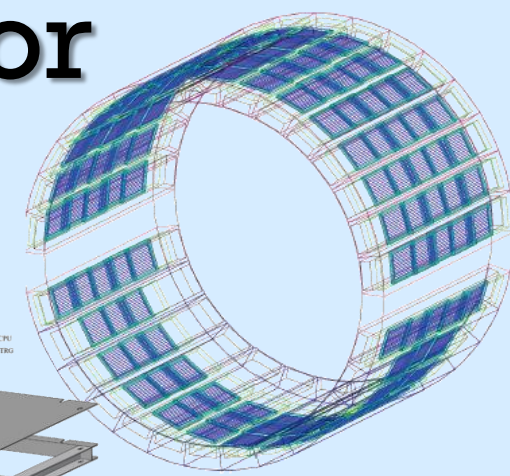
- Ideal probe to study the evolution of the medium by selecting different kinematics
- 1~4 GeV/c
 - Study the properties of QGP
- High $p_T (>6 \text{ GeV}/c)$:
 - Study the photon produced in the primordial step, distinguish initial- and final-state suppression

Beam Energy Scan

- **AuAu 15 GeV [Feb. 2014]**
 - Continue extensive study
 - Test baryon density dependence

- **Future: BES II [\sim 2018+2019]**
 - iTPC, Enhanced Statistics, Dimuons

Muon Telescope Detector



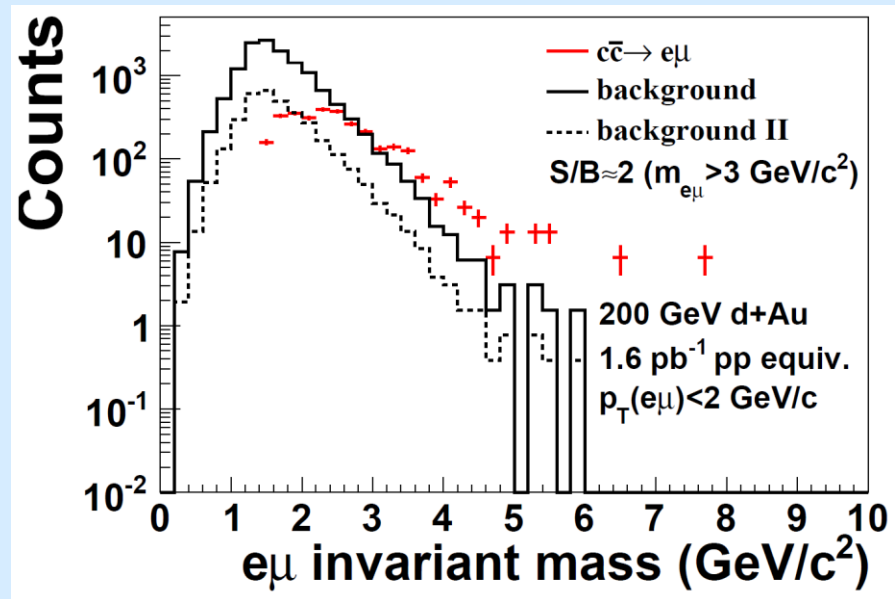
- Multi-gap Resistive Plate Chamber detector
 - Similar to STAR's TOF
 - Located outside the magnet
 - Fully installed for AuAu 200GeV next year

- Dimuon

- Pros vs dielectrons: Less background, triggering possible, less Dalitz decay
- Measure thermal radiation
 - Determine excess dimuon
 - Fit slope m_T -M slope to find T_{eff}

- Electron-muon

- Better handle on charm contribution
 - To better understand background in IMR
 - Aides dielectron thermal radiation study



Summary

- Observed excess w.r.t. the hadronic cocktail and the excess is consistent with model calculations involving broadened- ρ spectra
- No strong energy dependence of the LMR excess from 19.6 – 200 GeV

Outlook

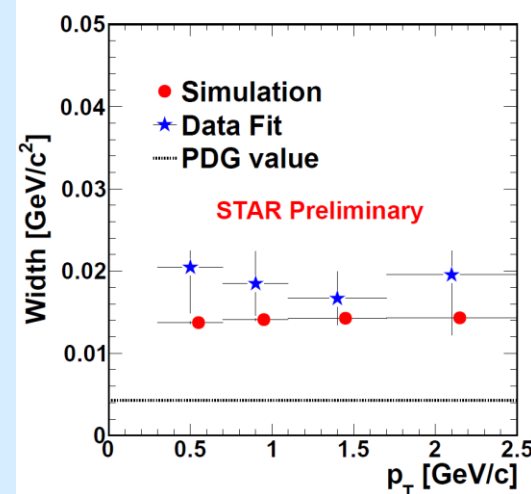
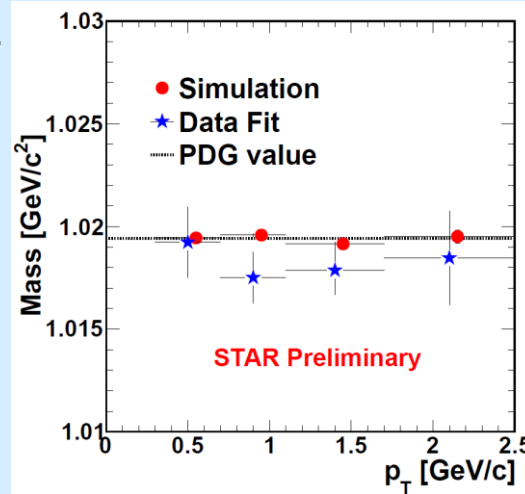
- Direct Virtual Photon
- Beam Energy Scans [I & II]
- Muon Telescope Detector

Thank you!

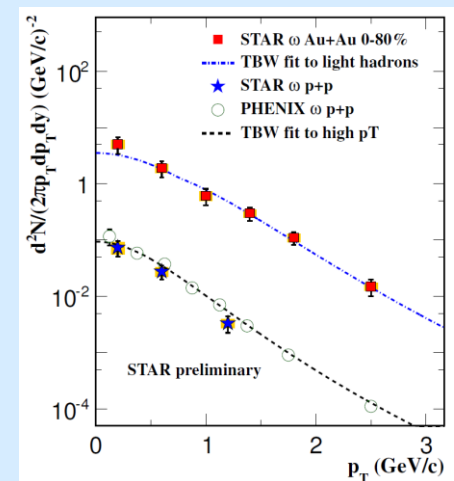
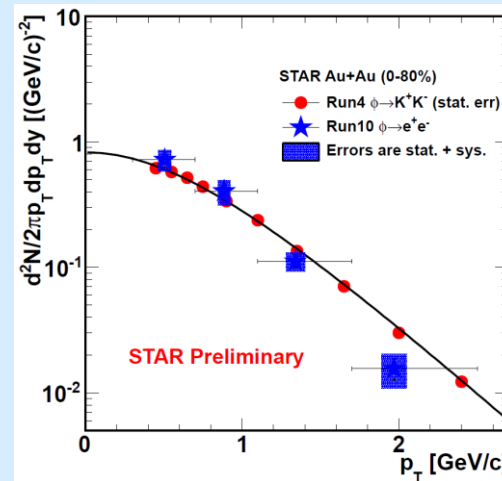
Extras

ω & ϕ Spectra in AuAu @ 200GeV

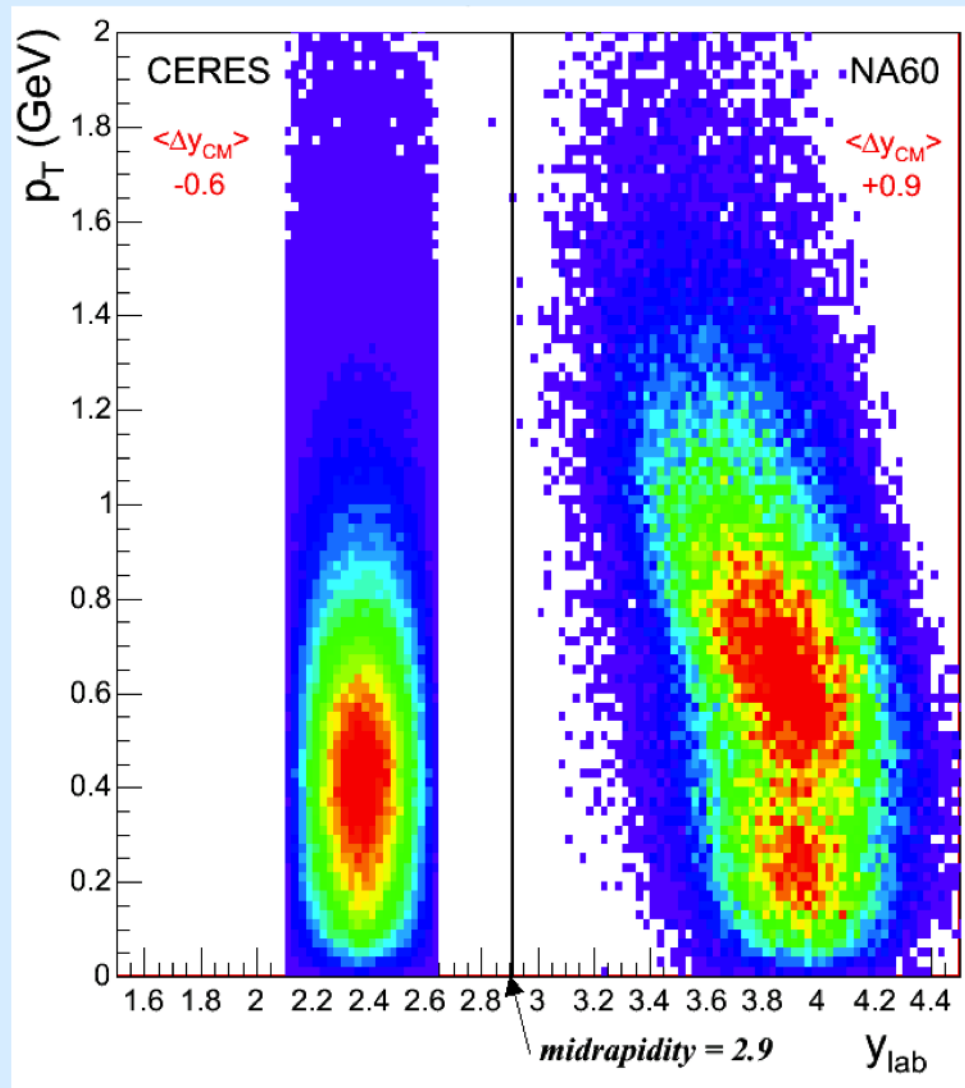
- ϕ consistent with no broadening/mass shift given uncertainties
 - Detector resolution limiting factor
 - ω being studied



- ϕ decay yields are consistent between K^+K^- and e^+e^- channels
- ω p_T spectra matches that of PHENIX and light hadrons

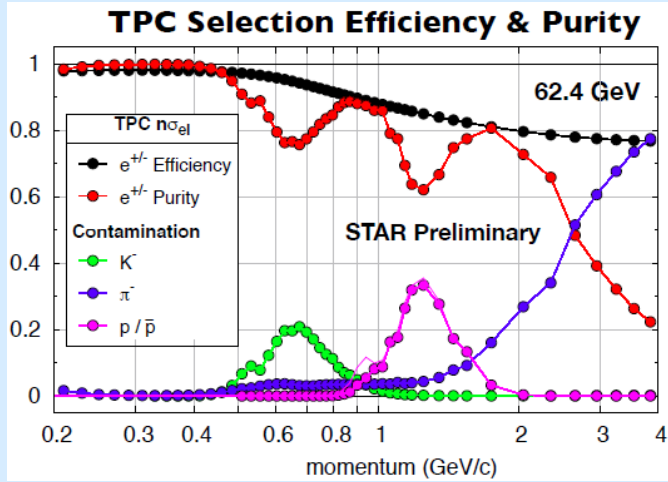


CERES + NA60 Acceptance



Efficiency Corrections

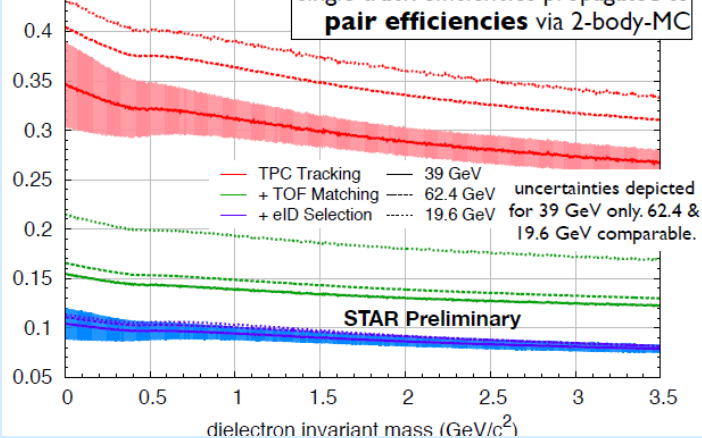
62 GeV



- Pair efficiencies based on single track eff.
 - Pair Methods
 - ToyMC: $\gamma^* \rightarrow e^+e^-$
 - Apply STAR acceptance
 - Single Track
 - Tracking [Embedding]
 - Matching tracks to TOF detector [Data]
 - Electron Identification [Data]

Huck, Thermal Radiations Workshop

39 GeV



Huck, Thermal Radiations Workshop

Pair acceptance correction

