# STAR <br> Dynamics of particle emission probed by femtoscopic correlations in the STAR experiment 

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## Abstract

One of methods to study the properties of hot and dense nuclear matter created in high-energy nuclear collisions is femtoscopic measurements. This method provides information about space-time characteristics of the particle emission region, which has a size and lifetime of the order of $10^{-15} \mathrm{~m}$ and $10^{-23} \mathrm{~s}$, respectively. From non-identical particle correlations, one can obtain information about asymmetry in the emission process between those two particles' species [1]. Such an emission asymmetry gives us knowledge of which type of particles, on average, are emitted earlier and from which region of the source. Using different combinations of pion, kaon, and proton pairs, one can obtain comprehensive knowledge on geometric and dynamic (emission time) properties of the particle emitting source. Such investigation could provide information about differences among the emissions of light mesons (pions), strange mesons (kaons), and baryons (protons).

## Particle identification

Time Projection Chamber (TPC) measures particle momentum and ionization energy loss ( $\mathrm{dE} / \mathrm{dx}$ )
Time Of Flight (TOF) - measures time of flight from the collision point to the detector


|  | T | K | p |
| :---: | :---: | :---: | :---: |
| $\mathrm{p}_{\mathrm{T}}[\mathrm{GeV} / \mathrm{c}]$ | [0.1-1.2] | [0.1-1.2] | [0.4-2.5] |
| $\mathrm{p}[\mathrm{GeV} / \mathrm{c}]$ | [0.1-1.2] | [0.1-1.2] | [0.4-3.0] |
| Momentum threshold [GeV/c] | 0.2 | 0.41 | 0.8 |
| Mass squared window $\left[\mathrm{GeV}^{2} / \mathrm{c}^{4}\right]$ | [0.01-0.03] | [0.21-0.28] | [0.76-1.03] |
| \| $\mathrm{N} \sigma$ \| |  | $\leq 3.0$ |  |
| Pseudorapidity $\|\boldsymbol{\eta}\|$ |  | $\leq 0.5$ |  |
| DCA [cm] |  | $\leq 3.0$ |  |

## Femtoscopy

Theoretical correlation function can be described by the Koonin-Pratt formula [3, 4]:

$$
C\left(\overrightarrow{k^{*}}\right)=\int d \vec{r}\left|\psi\left(\overrightarrow{k^{*}}, \vec{r}\right)\right|^{2} S(\vec{r})
$$

where $\overrightarrow{k^{*}}$ is particle momentum in pair rest frame, $\vec{r}$ is relative distance between two particles, $\psi\left(\overrightarrow{k^{*}}, \vec{r}\right)$ is a pair wave function that expresses interactions between particles, and $S(\vec{r})$ is the source function distribution of relative positions of particles.
The source of non-identical particles is assumed as a 3 -dimensional Gauss distribution with sizes R in out, side and long directions, and the mean $\mu$ value corresponding to the emission asymmetry:

$$
S(\vec{r}) \propto \exp \left(-\frac{\left(r_{\text {out }}-\mu\right)^{2}}{2 R_{\text {out }}^{2}}-\frac{r_{\text {side }}^{2}}{2 R_{\text {side }}^{2}}-\frac{r_{\text {long }}^{2}}{2 R_{\text {long }}^{2}}\right)
$$

Bertsch-Pratt parametrization
Long: determined by the beam axis Out: determined by the direction of the pair momentum in transverse plane Side: perpendicular to long and out

## Emission asymmetry

Asymmetries in the emission process may arise from long-lived resonances, bulk collective effects, or differences in the freeze-out scenario for different particle species [5].
The separation comes from:

- space asymmetry (flow)
- emission time difference
Time asymmetry

| $t_{1} \neq t_{2}$ |
| :--- |
| $\Delta r=0$ |

Fig. 3 Two types of asymmetry

$t_{1}=$| $t_{1}=t_{2}$ |
| :--- |
| $\Delta r \neq 0$ |

$\Delta r \neq 0$

- Moving away scenario - faster particle is emitted earlier (or closer to the edge of the source)
Catching up scenario - faster particle is emitted later (or closer to the center of the source


Fig. 7 Source size (left) and asymmetry (right) for different types of pairs.
*Influence of particles from resonance decays not considered,
**color bands represent systematic uncertainties.

## Summary

- Asymmetry is visible in each kind of analyzed pair
- Lighter particles are emitted closer to the center and/or later (indicated by negative value of $\mu$ )
- Pairs from lambda resonance have a negligible impact on the correlation effect
Only kaon-proton $\mathrm{C}\left(k^{*}\right)$ has visible and significant strong interaction


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