# Azimuthal distributions of Righ-pt direct $\gamma$ and $\pi^{0}$ w.r.t reaction plane at STAR 

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## eleiptic flow at low pt

$>$ The azimuthal distributions of the produced particles in heavyion collisions are considered to be sensitive to the initial geometric overlap of the colliding nuclei.

$$
\varepsilon=\frac{\left\langle y^{2}-x^{2}\right\rangle}{\left\langle y^{2}+x^{2}\right\rangle}
$$



$d N / d \phi \propto 1+\sum_{n} 2 v_{n} \cos n\left(\phi-\Psi_{R P}\right)$
$>$ Eccentricity in spatial coordinate is preserved and mapped into the momentum coordinates if the produced particles are not freely streaming.
$>$ At low $\mathrm{p}_{\mathrm{T}}$ the produced particles are not freely streaming and the collectivity is built even before hadronization.
Non-flow

$$
\left.\cos 2\left(\phi_{p_{t}}-\phi_{i}\right)\right\rangle=M v_{2}\left(p_{t}\right) \bar{v}_{2}+\{\text { non-flow }\}
$$

## ECliptic flow at high pt

$\mathrm{V}_{2}$ at high $\mathrm{p}_{\mathrm{T}}$ is finite positive!
Jet quenching : energy loss dependence of path length

$>$ Azimuthal anisotropy at large $\mathrm{p}_{\mathrm{T}}$ seem to be too large for a pure "jet quenching" Phys. Rev. C. 66, 027902 (2002)
$>$ Surface emission is not consistent with the measured value of elliptic flow at high $\mathrm{p}_{\mathrm{T}}$ Phys. Rev. Lett. 93 (2004) 252301

## Why ecliptic flow of direct photons?

$>\mathrm{v}_{2}$ of electromagnetically interacting particles.
$>\mathrm{v}_{2}$ measurements at higher $\mathrm{p}_{\mathrm{T}}$.
$>$ Production mechanisms of photons


## $>$ Path length dependence of parton energy loss.

$\mathrm{V}_{2}<0$ : Particles preferred to traverse through the longer path "out-of-plane"
$\mathrm{V}_{2}=0$ : No preferred direction w.r.t. reaction plane
$\mathrm{V}_{2}>0$ : Particles preferred to traverse through the shorter path "in-plane"

STAR LO production of direct $\gamma$ and fragmentation $\gamma$
Direct photons



Annihilation

Very challenging measurements due to the $\mathrm{S} / \mathrm{B}$ ratio, $\pi^{0}$ is the major source of bg .
$>$ The Compton-scattering process

$$
\rightarrow \gamma / \pi^{0}>\alpha_{\mathrm{em}}
$$

Eragmentation photons
Fragmentation photons $\gamma^{\text {frag }}$


Example of Bremsstrahlung diagramsThe sub-process of $\gamma^{\text {frag }}$ is of order of $\mathrm{O}\left(\alpha_{s}{ }^{2}\right)$ but its yield is comparable to $\gamma^{\mathrm{dir}} \mathrm{LO}$ process $\mathrm{O}\left(\alpha_{\mathrm{s}} \alpha_{\mathrm{em}}\right)$.The $\gamma^{\text {frag }}$ contribution is expected to fall off more rapidly in $\mathrm{x}_{\mathrm{T}}$ than the other lowest order of $\gamma^{\text {dir }}$. (G. Sterman et al. Rev. Mod. Phys. 67, 157 (1995))$\gamma^{\text {frag }} / \gamma^{\text {dir }} \sim 30-40 \%$ at $\mathrm{p}^{\gamma}>8 \mathrm{GeV} / \mathrm{c}$ at mid-rapidity at RHIC energy. D. De Florian and w , Vogelsang, Phys. Rev. D72, 014014 (2005)

## Methods of $\gamma^{\text {dir }}$ measurements

$\checkmark$ Standard Method

1. Measure inclusive photons.
2. Reconstruct other sources of photons "hadrons"!
3. Subtract photons from decay of $\pi^{0}, \eta$ etc.

PHENIX is well-adapted for this method due to the calorimeter granularity and the distance between the calorimeter to the interaction point $\rightarrow \pi^{0}$ reconstruction in central $\mathrm{Au}+\mathrm{Au}$ up to $\mathrm{p}_{\mathrm{T}} \sim 20 \mathrm{GeV} / \mathrm{c}$
$>$ Limited at very high $\mathrm{p}_{\mathrm{T}}$, effective method for both symmetric and asymmetric hadron decays

## $\checkmark$ Transverse Shower Profile Method

STAR is well-suited for the transverse shower shape analysis due
to the Shower Maximum Detector $\rightarrow \gamma / \pi^{0}$ discrimination up to $\mathrm{p}_{\mathrm{T}} \sim 26 \mathrm{GeV} / \mathrm{c}$. M. Beddo et al., Nucl. Instrum. Meth. A499, 725 (2003)
$>$ Effective at very high $\mathrm{p}_{\mathrm{T}}$, but limited only for the symmetric hadron decays

1. Electromagnetic neutral cluster $\left(\pi^{0}, \eta, \rho^{0}, \omega, \ldots, \gamma^{\mathrm{frag}}, \gamma^{\mathrm{dir}}\right)$
2. Reaction plane measurements

3. Transverse shower profile to obtain sample rich/free of $\gamma^{\text {dir }}$
4. Obtain $\mathrm{V}_{2}$ of $\gamma^{\mathrm{dir}}$

# STAR STAR detector and on-line $\gamma-r i c h ~ e v e n t ~ s e l e c t i o n s ~$ 


$1 \gamma$-triggered event each 5k minbias event $\rightarrow \sim 500 \mu \mathrm{~b}^{-1}$ of AuAu 2007 @ 200 GeV 6 k events of minimum bias trigger

# $\wedge$ STAR detector and off-line nentral cluster selections 

$\checkmark$ Select neutral clusters "triggers" (BEMC-BSMD) using charged-particle veto (TPC)

vertex within $\pm 55 \mathrm{~cm}$ of the center of TPC.
At least one cluster with $\mathrm{E}_{\mathrm{T}}>8 \mathrm{GeV}$, Esmd $\eta>0.5 \mathrm{GeV}$, Esmd $\phi>0.5 \mathrm{GeV}$, and no track with $\mathrm{p}>3 \mathrm{GeV} / \mathrm{c}$ pointing to that cluster.

In $\mathrm{Au}+\mathrm{Au}: 28 \%$ of the integrated luminosity
has $\mathrm{E}_{\mathrm{T}}>8 \mathrm{GeV}$ of which $96.5 \%$ left at least 0.5 GeV on each planes of SMD of which $93 \%$ has no track with $\mathrm{p}>3 \mathrm{GeV} / \mathrm{c}$ pointing to it.

Event plane from TPC

$$
\psi=\frac{1}{2} \tan ^{-1}\left(\frac{\sum_{i} \sin \left(2 \delta_{i}\right)}{\sum_{i} \cos \left(2 \delta_{i}\right)}\right)
$$

Shift method for event plane flattening

$$
\Psi^{\prime}=\Psi+\sum_{n} \frac{1}{n}[-\langle\sin (2 n \Psi)\rangle \cos (2 n \Psi)+\langle\cos (2 n \Psi)\rangle \sin (2 n \Psi)]
$$

Sub-event method for reaction plane resolution

$$
\sigma_{\mathrm{RP}}=C \sqrt{<} \cos \left[2\left(\psi^{A}-\psi^{B}\right)\right]>, C=\sqrt{2}
$$

$\underline{\mathbf{v}}_{2}$ of charged and neutral particles

$$
\begin{gathered}
v_{2}^{\text {track,obs }=}<\cos \left(2 \phi_{\text {track }}-2 \psi\right)> \\
v_{2}^{\text {neutral,obs }=}<\cos \left(2 \phi_{\text {tower }}-2 \psi\right)> \\
v_{2}=\frac{v_{2}^{o b s}}{\sigma_{\mathrm{RP}}}
\end{gathered}
$$ 10-40\% AuAu @ $\sqrt{\mathbf{s}}=200 \mathrm{GeV}$


$v_{2}(E P$ off $-\eta)$ reproduces the $v 2\{4\}$ quite well 10-40\% AuAu @ $\sqrt{\mathbf{s}}=200 \mathrm{GeV}$

$\mathrm{v}_{2}$ of charged particles is $\sim 15 \%$ in $10-40 \%(\mathrm{AuAu} @ 200 \mathrm{GeV})$ and constant in pt ( $8-16 \mathrm{GeV} / \mathrm{c}$ ) 10-40\% AuAu @ $\sqrt{\mathbf{s}}=200 \mathrm{GeV}$

$\mathrm{v}_{2}$ of neutral particles is $\sim 10 \%$ in $10-40 \%(\mathrm{AuAu} @ 200 \mathrm{GeV})$ and constant in pt ( $8-16 \mathrm{GeV} / \mathrm{c}$ )

## STAR ECCiptic flow of nentral/charged particles at high pt 10-40\% AuAu @ $\sqrt{\mathbf{s}}=200 \mathrm{GeV}$


$\mathrm{v}_{2}$ of neutral particles is less than $\mathrm{v}_{2}$ of charged particles due to direct photons contributions

## How to separate $\gamma^{\text {dir }}$ from nentral ©g.

$\sim 10 \%$ of all $\pi^{0}(8-16 \mathrm{GeV} / \mathrm{c})$ decay asymmetrically with one gamma has $\mathrm{p}_{\mathrm{T}}>8 \mathrm{GeV} / \mathrm{c}$ within STAR-BEMC acceptance.
$\eta$ causes similar level of background as asymmetric $\pi^{0}$.


Either to reconstruct $\pi^{0}$ or to use the transverse shower shape analysis to distinguish between $\pi^{0}$ and $\gamma^{\mathrm{dir}}$

## STAR BEMC and BSMD



The two photons originated from $\pi^{0}$ hit the same tower at $\mathrm{p}_{\mathrm{T}}>8 \mathrm{GeV} / \mathrm{c}$
The shower shape is quantified with the cluster energy, measured by the BEMC, Normalized by the position-dependent energy moment, measured by the BSMD strips.

## Shower Profile of single $\gamma v$ s. two close $\gamma s$




The probability distribution is peaked at smaller value in AuAu than in pp due to the larger relative fraction of $\gamma^{\text {dir. }}$

The rejection power of direct photons is $\sim 90 \%$
 10-40\% AuAu @ $\sqrt{\text { s }}=200 \mathrm{GeV}$

$\mathrm{v}_{2}$ of $\gamma^{\text {rich }}<\mathrm{v}_{2}$ of $\pi^{0}$ as expected

## Obtain v2 of direct photons

$\oplus$ Select EM neutral clusters
$\oplus$ Use the transverse shower shape to select $\gamma^{\text {dir }}$ free ( $\pi^{0}$-rich) sample and $\gamma^{\text {rich }}$ sample from the neutral clusters.

$$
\begin{gathered}
v_{2}^{\gamma_{\text {rich }}} N^{\gamma_{\text {rich }}}=v_{2}^{b g} N^{b g}+v_{2}^{\gamma_{d i r}} N^{\gamma_{d i r}} \\
\mathcal{R}=\frac{N^{b g}}{N^{\gamma_{r i c h}}} \simeq \frac{N^{\pi^{0}}}{N^{\gamma_{r i c h}}} \\
v_{2}^{\gamma_{\text {direct }}}=\frac{v_{2}^{\gamma_{r i c h}}-v_{2}^{b g} \mathcal{R}}{1-\mathcal{R}} \\
v_{2}^{\gamma_{\text {direct }}}=\frac{v_{2}^{\gamma_{\text {rich }}}-v_{2}^{\pi^{0}} \mathcal{R}}{1-\mathcal{R}}
\end{gathered}
$$

## STAR ECCiptic flow of $\pi^{0}, \gamma^{d i r}$, charged particles at high pt 10-40\% AuAu @ $\sqrt{\mathbf{s}}=200 \mathrm{GeV}$


$v_{2}$ of non decay $\gamma$ is $\sim 1 / 3$ of $v_{2}$ of $\pi^{0}$ and charged particles
$\mathrm{v}_{2}$ of non decay $\gamma$ is not zero and not negative

## STAR ECCiptic flow of $\pi^{0}, \gamma^{d i r}$, charged particles at high pt 10-40\% AuAu @ $\sqrt{\mathbf{s}}=200 \mathrm{GeV}$

$\boldsymbol{> c} 0.5$
$v_{2}$ of non decay $\gamma$ is $\sim 1 / 3$ of $v_{2}$ of $\pi^{0}$ and charged particles
$\mathrm{v}_{2}$ of non decay $\gamma$ is not zero and not negative

## Summary

- The geometrical effect of the medium can be probed by the elliptic flow measurement
- Finite and +ve value of $\mathrm{v}_{2}$ persist up to $\mathrm{pt}=16 \mathrm{GeV} / \mathrm{c}$ for charged and neutral particles
- STAR has reported the first "preliminary" results of non decay photons elliptic flow at high pt at RHIC
- No sign of negative $\mathrm{v}_{2}$ of non decay photons
- Statistically significant value of $+\mathrm{ve} \mathrm{v}_{2}$ of non decay photons
- The $\mathrm{v}_{2}$ at high pt can not be interpreted as path length dependence of energy loss


## Backup slides

## Reaction plane



Only shift method is used to flatten the RP up to the $20^{\text {th }}$ harmonic:

$$
\Psi^{\prime}=\Psi+\sum_{n} \frac{1}{n}[-\langle\sin (2 n \Psi)\rangle \cos (2 n \Psi)+\langle\cos (2 n \Psi)\rangle \sin (2 n \Psi)]
$$

Reaction plane is flat "cos and $\sin \sim 0$ "

## Previous measurements of v2( $\gamma^{\text {dir })}$ at RHJC Phys. Rev. Lett96 (2006) 032302

$$
\begin{aligned}
& v_{2}^{\text {inclusive } \gamma}=\frac{v_{2}^{\text {direct } \gamma} N_{\text {direct } \gamma}+v_{2}^{b . g .} N_{\text {b.g. }}}{N_{\text {direct } \gamma}+N_{\text {b.g. }}} \quad \sigma_{\mathrm{RP}}=\left\langle\cos \left(2\left(\Phi_{\text {measured }}-\Phi_{\mathrm{RP}}\right)\right)\right\rangle \text { is } 0.3 \\
& \text { PHENIX BBC: } 3.1<|\eta|<3.9
\end{aligned}
$$

$$
R=\left(N_{\text {direct } \gamma}+N_{b . g .}\right) / N_{b . g .} .
$$

$$
v_{2}^{\text {direct } \gamma}=\frac{R v_{2}^{\text {inclusive } \gamma}-v_{2}^{\text {b.g. }}}{R-1}
$$

This measurements implies that $\mathbf{v}_{2}$ of direct photons is $\sim 0$


## Path cength dependence of energy loss



$\mathrm{Au}+\mathrm{Au} \sqrt{\mathbf{s}_{\mathrm{NN}}}=\mathbf{2 0 0} \mathrm{GeV}$ - Cent 20-60\%


## STAR measurement does not show path-length dependence.

PHENIX measurement show path-length dependence.

## Path Cength dependence of energy loss








$\pi^{0} \mathbf{v}_{\mathbf{2}}(\mathbf{p t})$ and $\mathbf{R}_{\mathrm{AA}}(\Delta \phi)$ show statistically significant dependence on the path length particularly at $\mathrm{pt}<6 \mathrm{GeV}$

