

# Light Nuclei Production in RHIC-STAR BES-II Program

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U.S. DEPARTMENT OF  
**ENERGY**

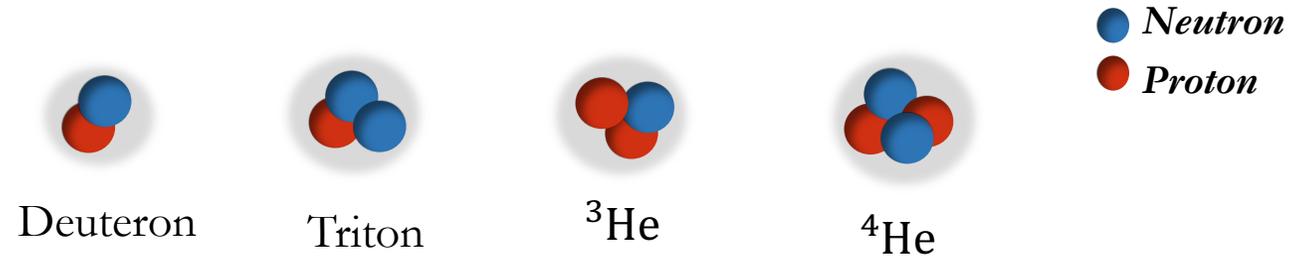
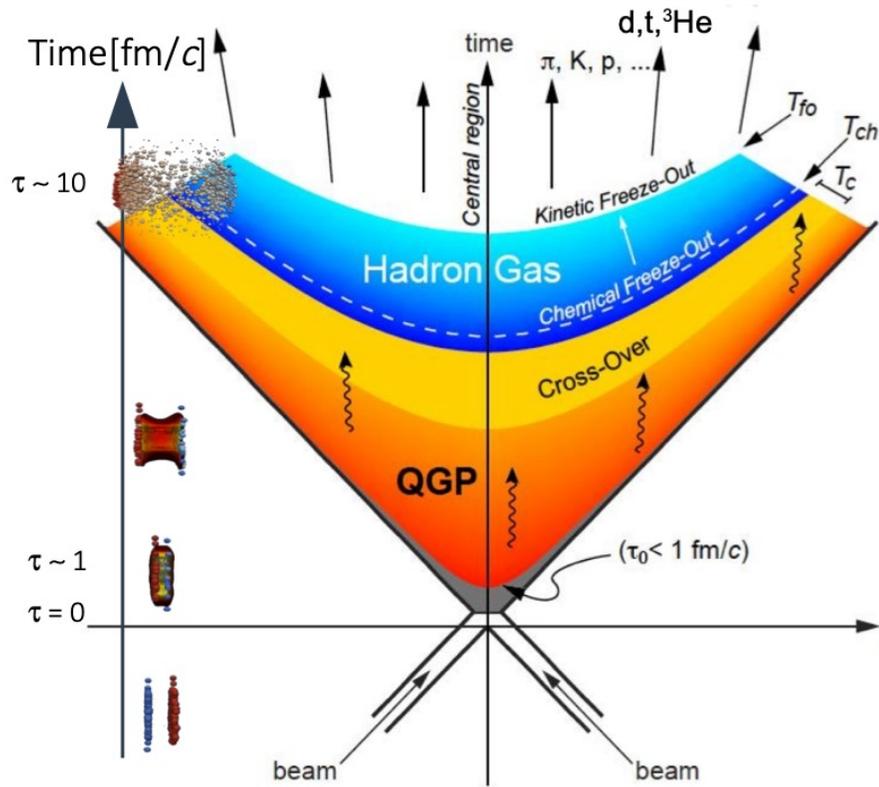
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# Outline

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- Introduction
- Method
- Proton and light nuclei production
  - $p_T$  spectra
  - $dN/dy$  and  $\langle p_T \rangle$
  - Yield ratios
  - $B_A$
  - Kinetic Freeze-out Dynamics
- Summary and Outlook

# Introduction



## ➤ Light Nuclei Production Mechanism

### • Thermal approach

Assumes that all particles, including light nuclei, have the same chemical freeze-out temperature and chemical potential.

### • Coalescence approach

Assumes that light nuclei are produced by the coalescence of nucleons in the late-stage evolution of the hadron gas.

## ➤ Chemical Freeze-out

Hadron composition fixed.

## ➤ Kinetic Freeze-out

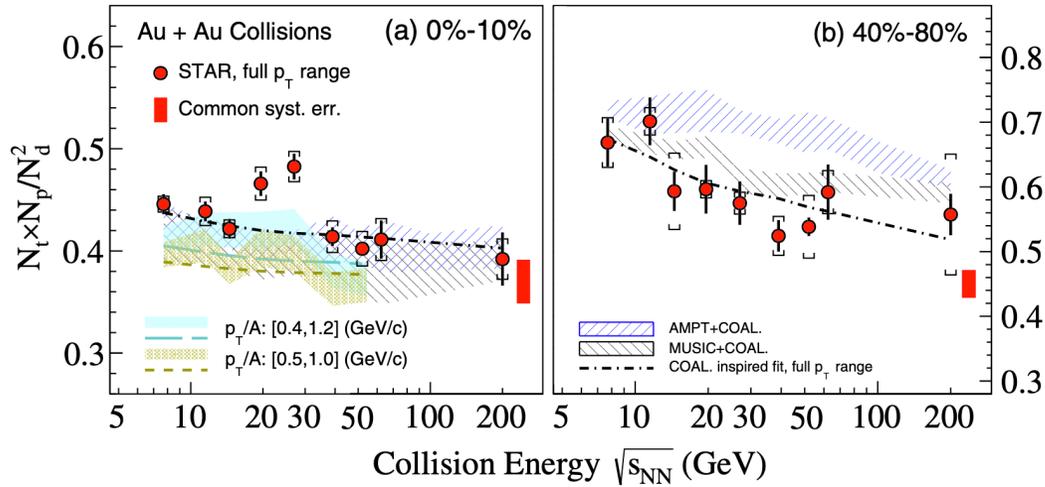
Particle momentum stop changing.

*A. Andronic et al., Nature 561, 321-330 (2018)*

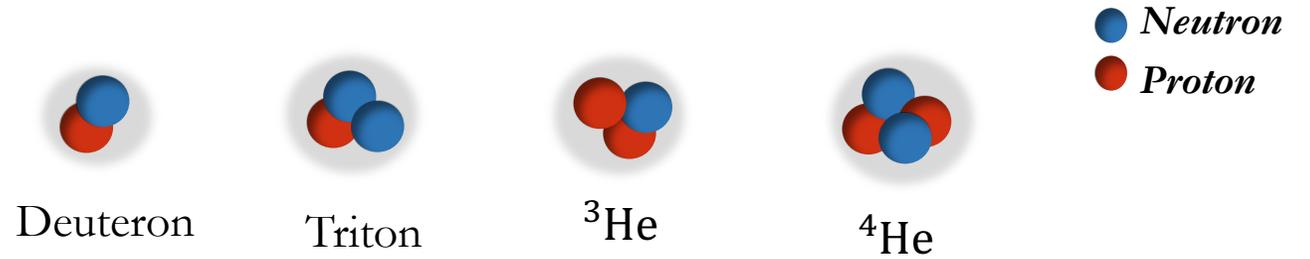
*K. J. Sun et al., Phys. Lett. B 792, 132-137 (2019)*

# Introduction

[STAR Collaboration] Phys.Rev.Lett. 130 (2023) 202301



- Non-monotonic behavior of yield ratio vs. energy observed from 0-10% central Au+Au collisions of STAR experiment, possibly signaling a critical point and/or 1<sup>st</sup> order phase transition.



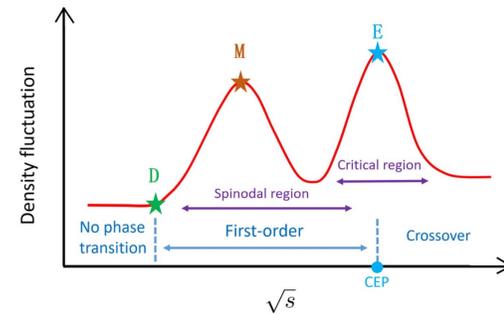
## ➤ Light Nuclei Compound Yield Ratio

Sensitive Observations for Searching Critical Point and 1<sup>st</sup> order boundary

$$\frac{N_t \times N_p}{N_d^2} = \frac{N(p_n) \times N(p)}{N(p_n) \times N(p_n)} \approx \frac{1}{2\sqrt{3}} \left[ 1 + \Delta n + \frac{\lambda}{\sigma} G\left(\frac{\xi}{\sigma}\right) \right]$$

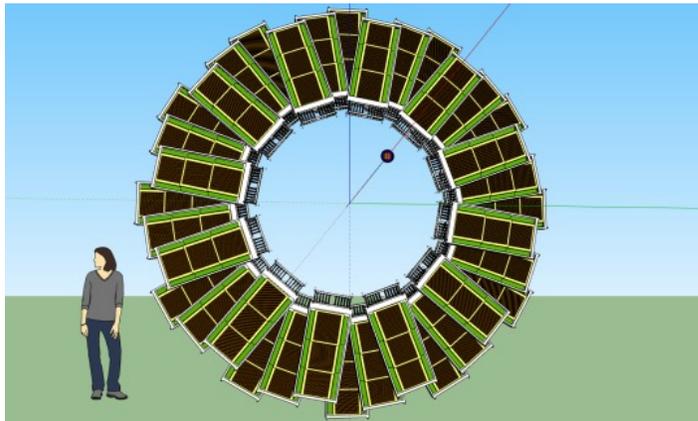
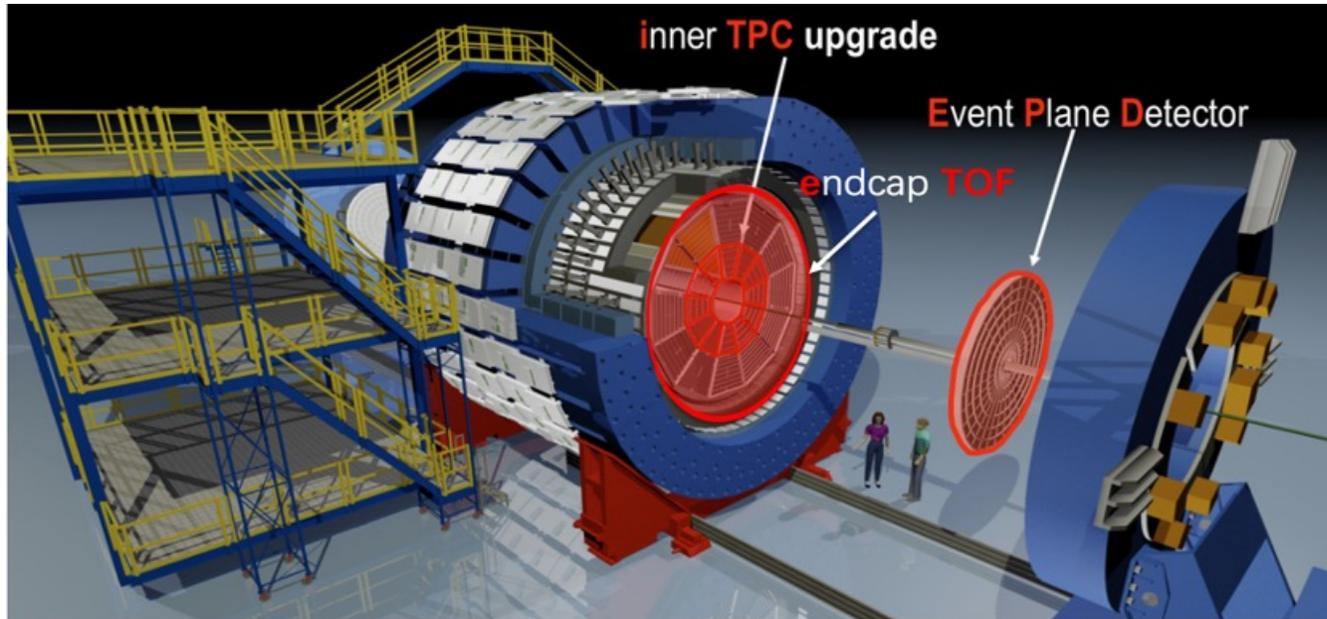
$\Delta n$ : Neutron Density Fluctuation

$G\left(\frac{\xi}{\sigma}\right)$ : Long-range Correlation

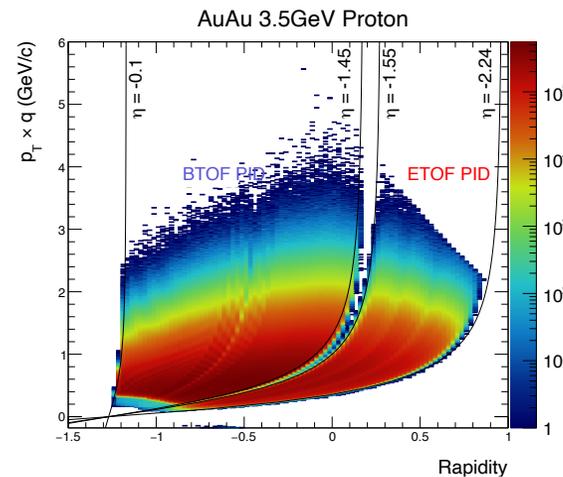


K.J. Sun et al, Phys.Lett.B 781 (2018) 499-504;  
K.J. Sun et al, Phys.Lett.B 816 (2021) 136258

# STAR Detector



Full ETOF wheel



Proton acceptance in BTOF/ETOF

## ➤ BES-II detector upgrade

In Au+Au collisions at  $\sqrt{s_{NN}} = 3.0, 3.2, 3.5, 3.9$  and  $4.5$  GeV

- iTPC

- cover full area,  $-2.4 < \eta < 0$
  - better  $dE/dx$ ,  $p_T > 60$  MeV/c.

- eTOF

- at the east end of STAR,
  - $-2.15 < \eta < -1.55$

## ➤ Fixed target mode

$\sqrt{s_{NN}} = 3.0-13.7$  GeV

$\mu_B$ : 750-280 MeV

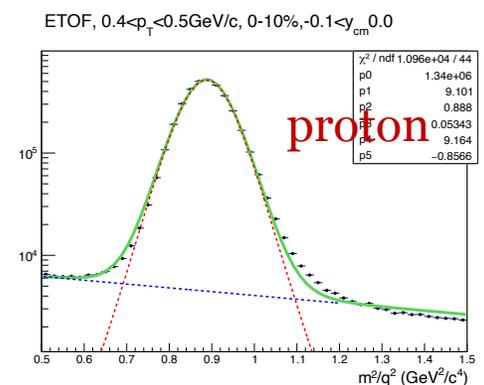
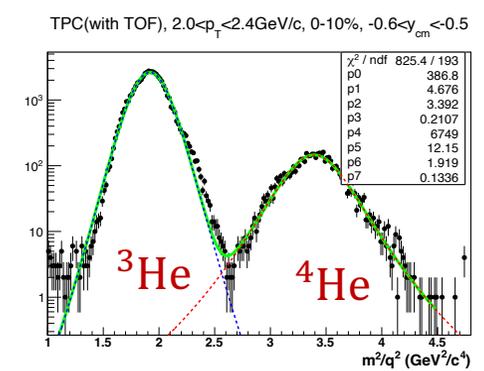
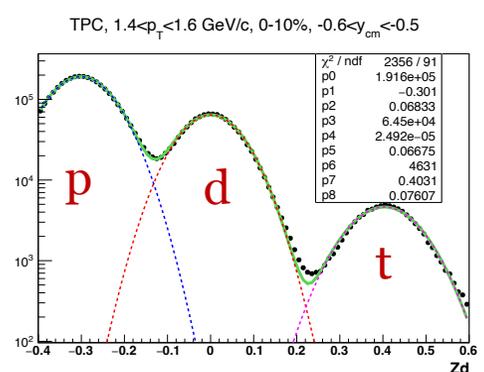
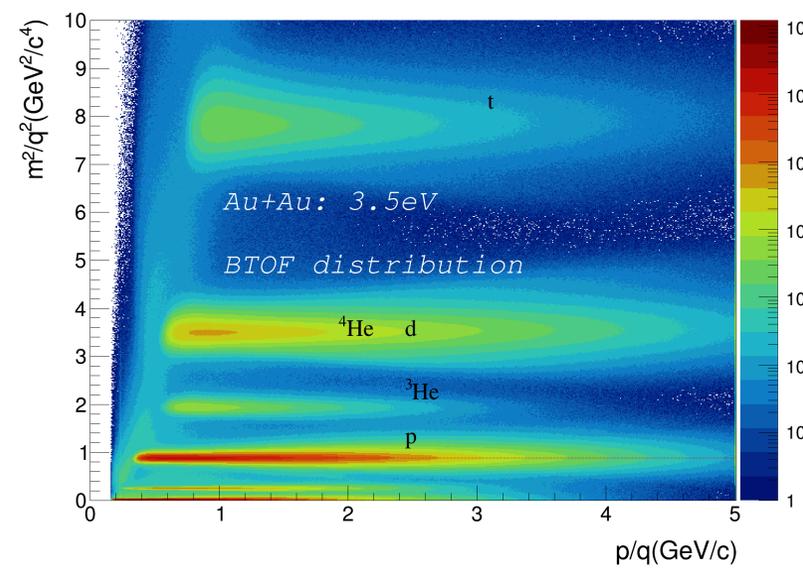
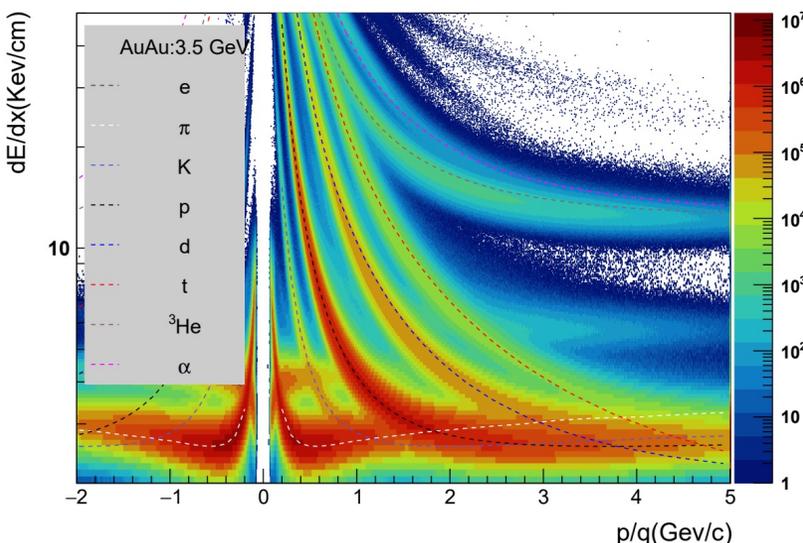
## ➤ Collider mode

$\sqrt{s_{NN}} = 7.7-27$  GeV

$\mu_B$ : 420-156 MeV

➤ ETOF extends more acceptance in mid rapidity, and will be use for light nuclei production in the future.

# Signal Extraction



➤ TPC:

- Signal(**red**):  
Gaussian function  
$$f(p_T) = p_0 e^{-\frac{1}{2} \left( \frac{p_T - p_1}{p_2} \right)^2}$$
- Background(**blue/Magenta**):  
Gaussian

➤ TPC with TOF/ETOF:

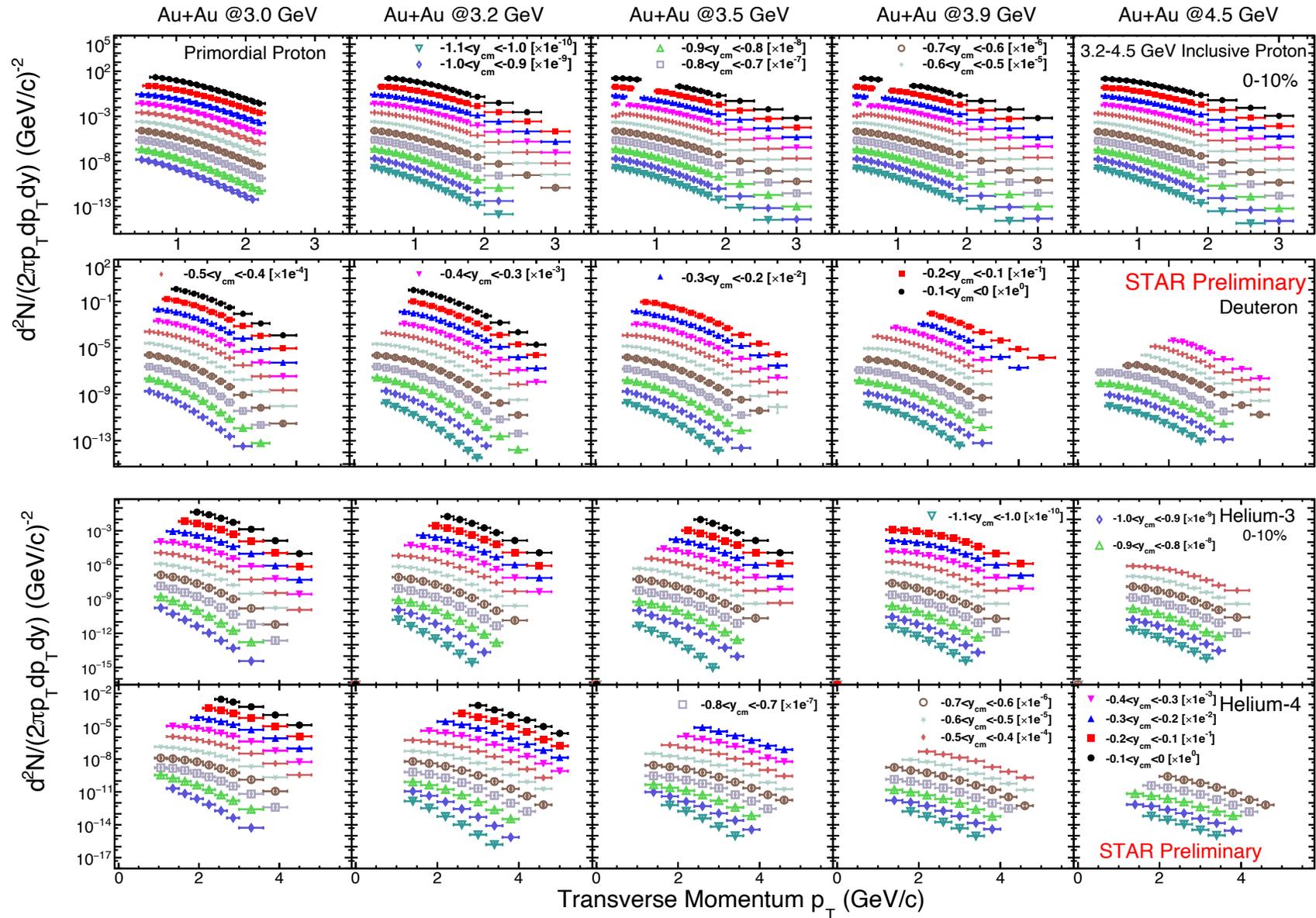
- Signal: **Student-t function**

$$f(x) = p_0 \frac{\Gamma\left(\frac{p_1 + 1}{2}\right)}{p_1 \pi \Gamma\left(\frac{p_1}{2}\right)} \times \left(1 + \frac{t^2}{p_1}\right)^{-\frac{p_1 + 1}{2}}$$

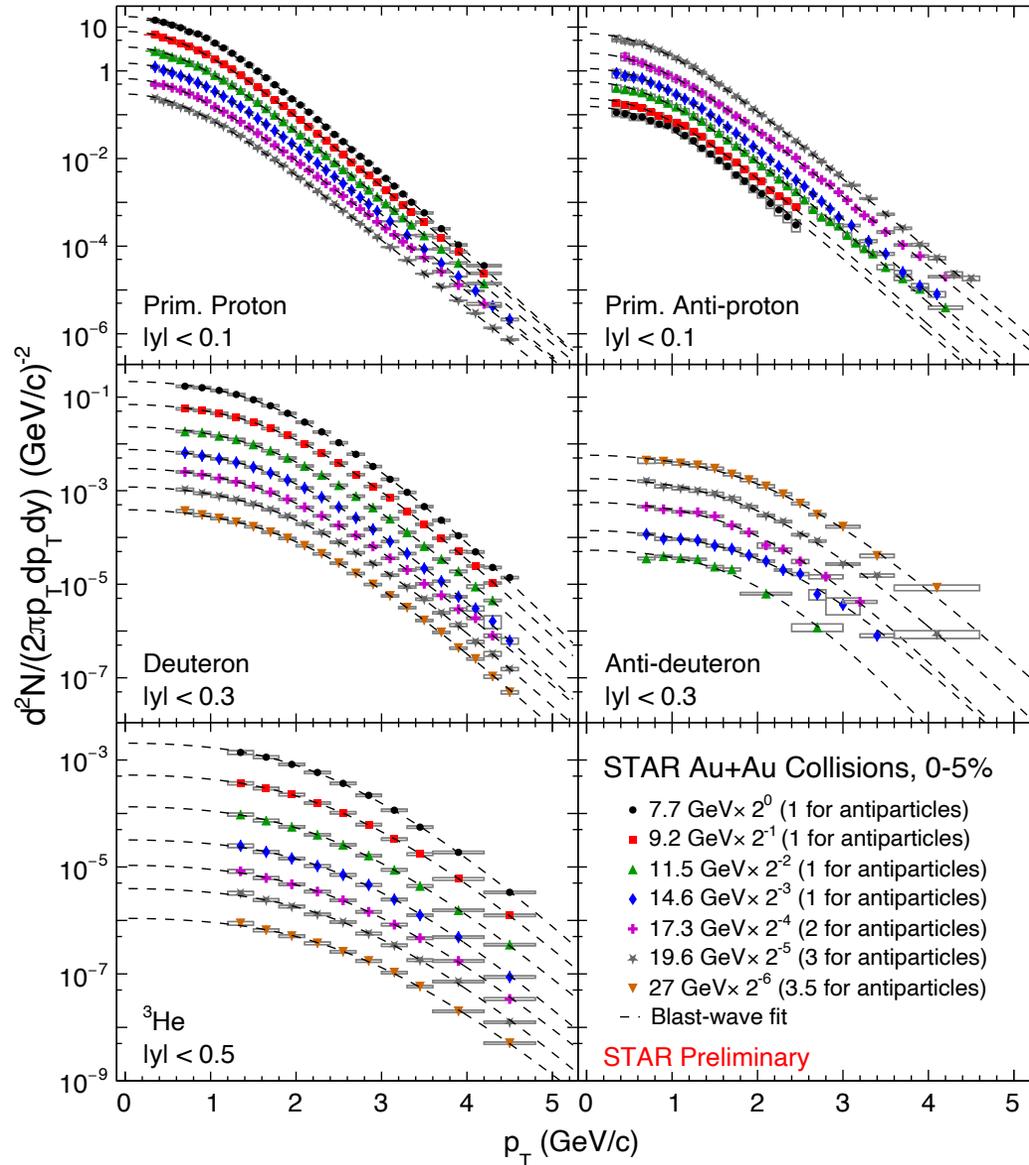
$$t = \left(\frac{m^2 - p_2}{p_3}\right)^2$$

- Background: **exponential, Student-t**( ${}^3\text{He}, {}^4\text{He}$ )

# Transverse Momentum Spectra



# Transverse Momentum Spectra



➤ Default fit function (**Blast-wave**):

$$\frac{d^2N}{2\pi dp_T dy} \propto \int_0^R r dr m_T I_0\left(\frac{p_T \sinh \rho(r)}{T_{kin}}\right) K_1\left(\frac{m_T \sinh \rho(r)}{T_{kin}}\right)$$

$$\rho = \tanh^{-1} \beta_r, \quad \beta_r(r) = \beta_T \left(\frac{r}{R}\right)^n, \quad \text{fix } n=1.$$

**Kinetic Freeze-out Parameters:**

$T_{kin}$  : kinetic freeze-out temperature

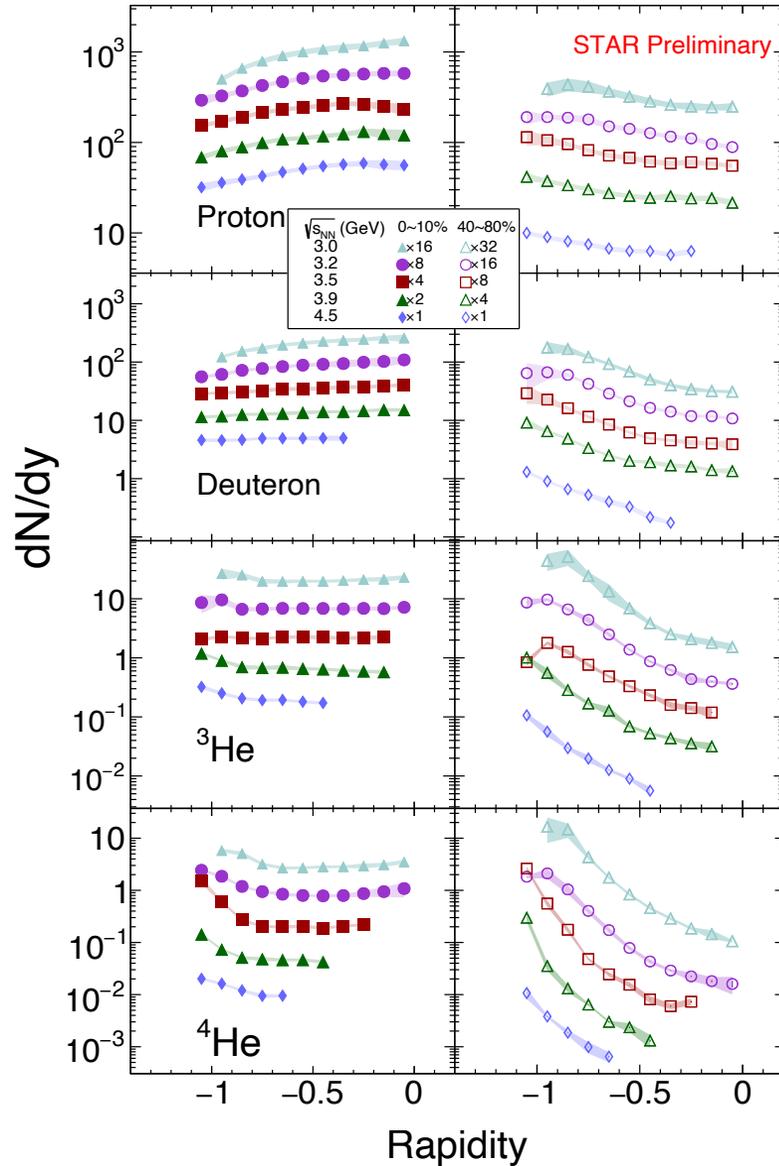
$\langle \beta_T \rangle$  : average radial flow velocity

➤ Alternative fit function (**Double  $p_T^2$  exp.**):

$$\frac{d^2N}{2\pi dp_T dy} \propto p_0 \exp\left(-\frac{p_T^2}{p_1^2}\right) + p_2 \exp\left(-\frac{p_T^2}{p_3^2}\right)$$

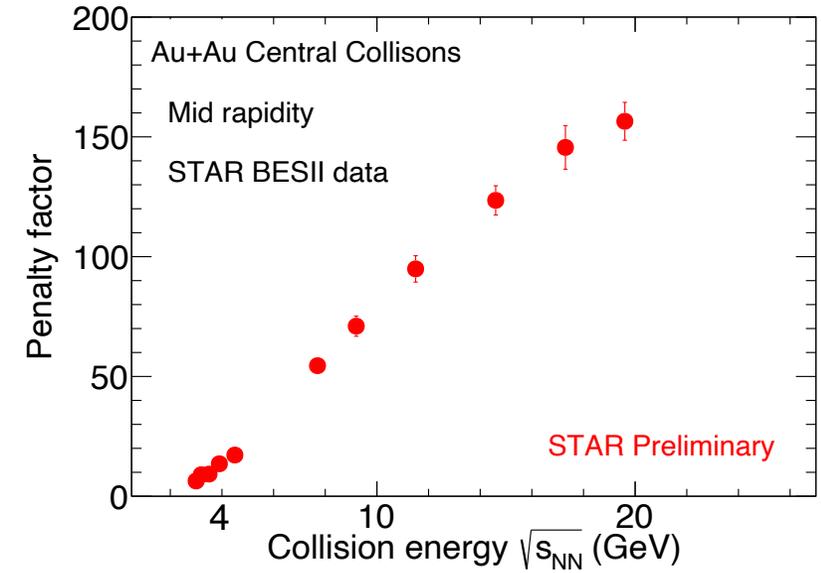
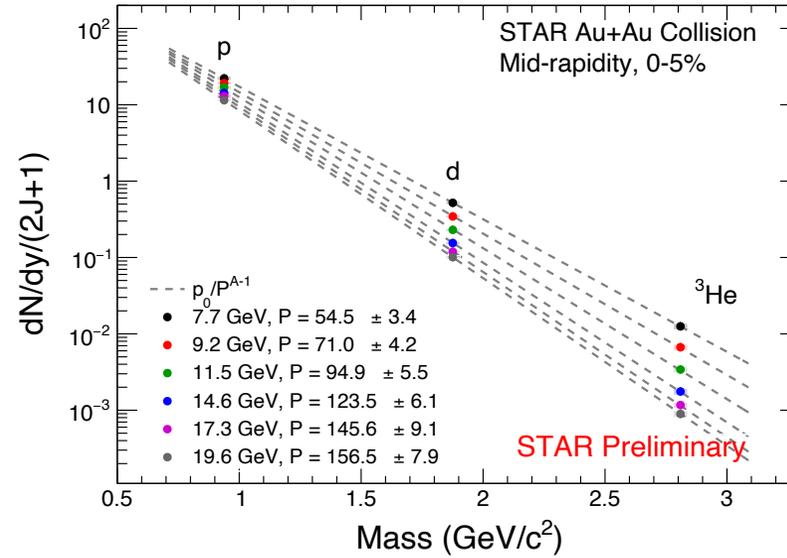
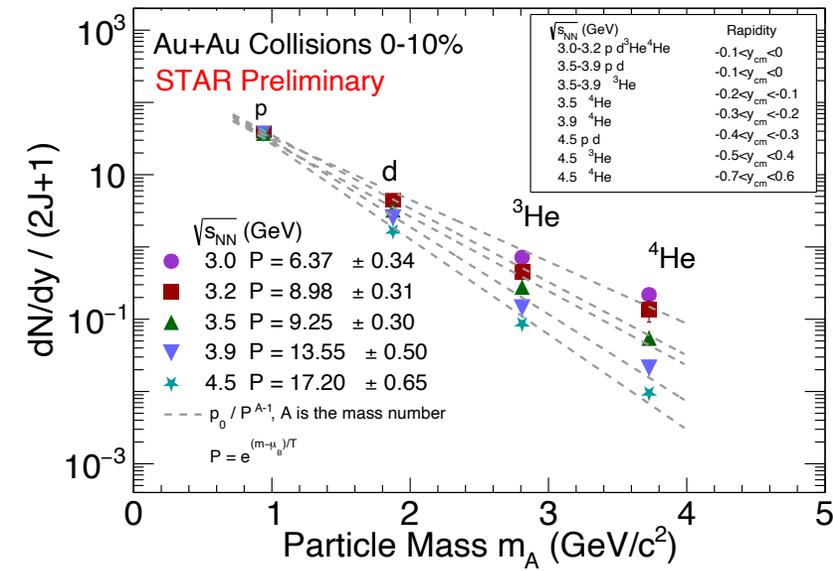
QM2025, Yixuan Jin

# dN/dy distribution



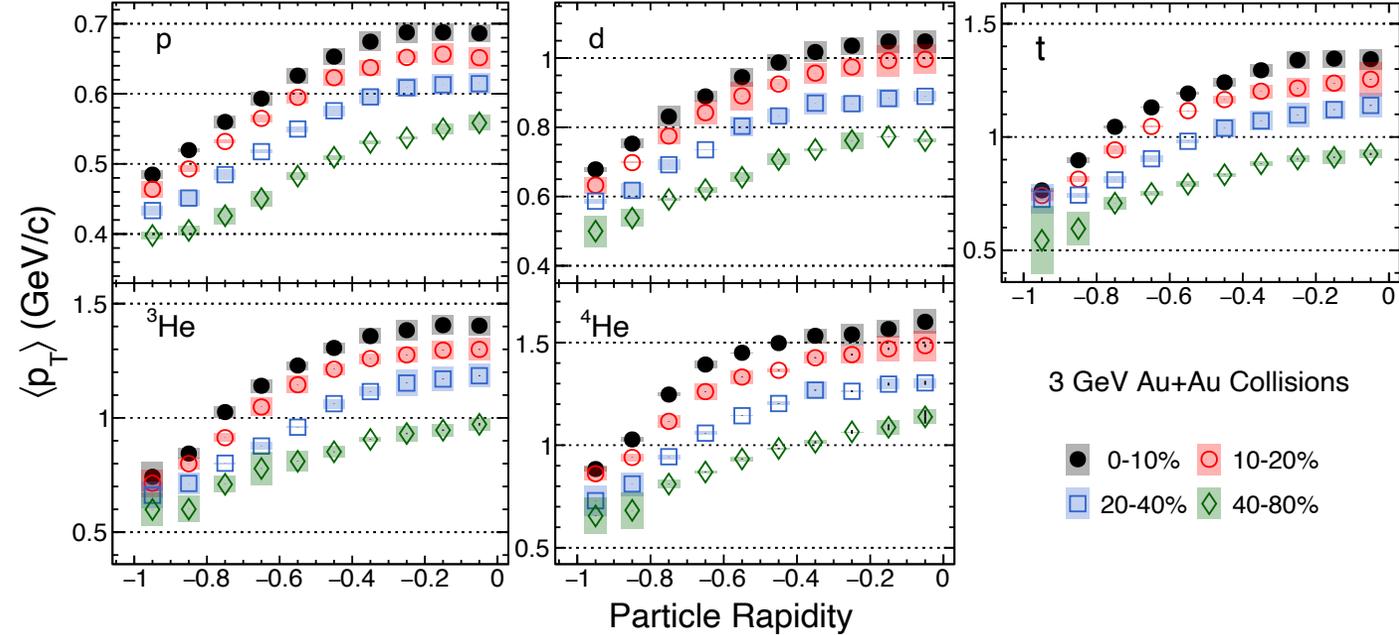
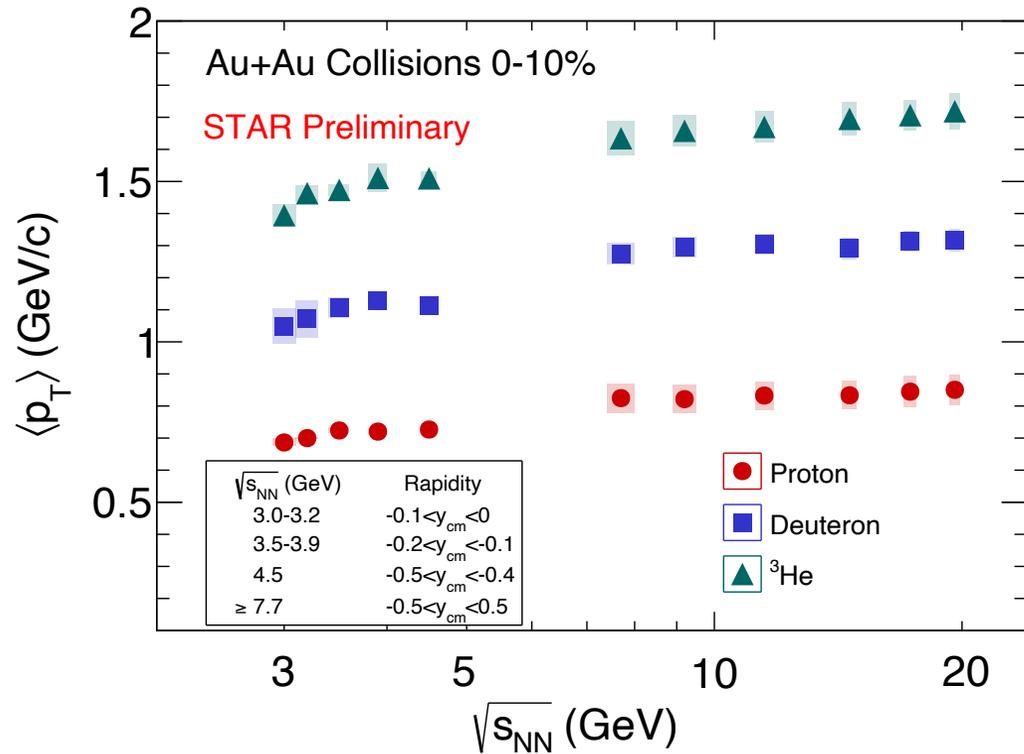
- dN/dy for proton, deuteron, triton,  $^3\text{He}$  and  $^4\text{He}$  in  $\sqrt{s_{NN}} = 3\text{-}4.5$  GeV with centrality and rapidity dependence.
- Light nuclei with larger mass numbers show higher dN/dy from target to mid-rapidity and central to peripheral collisions, suggesting fragment contributions to their production.
- The band indicate systematical uncertainty.

# dN/dy distribution



- $dN/dy/(2J+1)$  was fit with  $\frac{p_0}{P^{A-1}}$ , where P is the penalty factor and determined by Boltzmann factor  $e^{\frac{m-\mu_B}{T}}$ . P value increases with increasing beam energy, indicating suppressed formation of high-mass objects at higher energies.

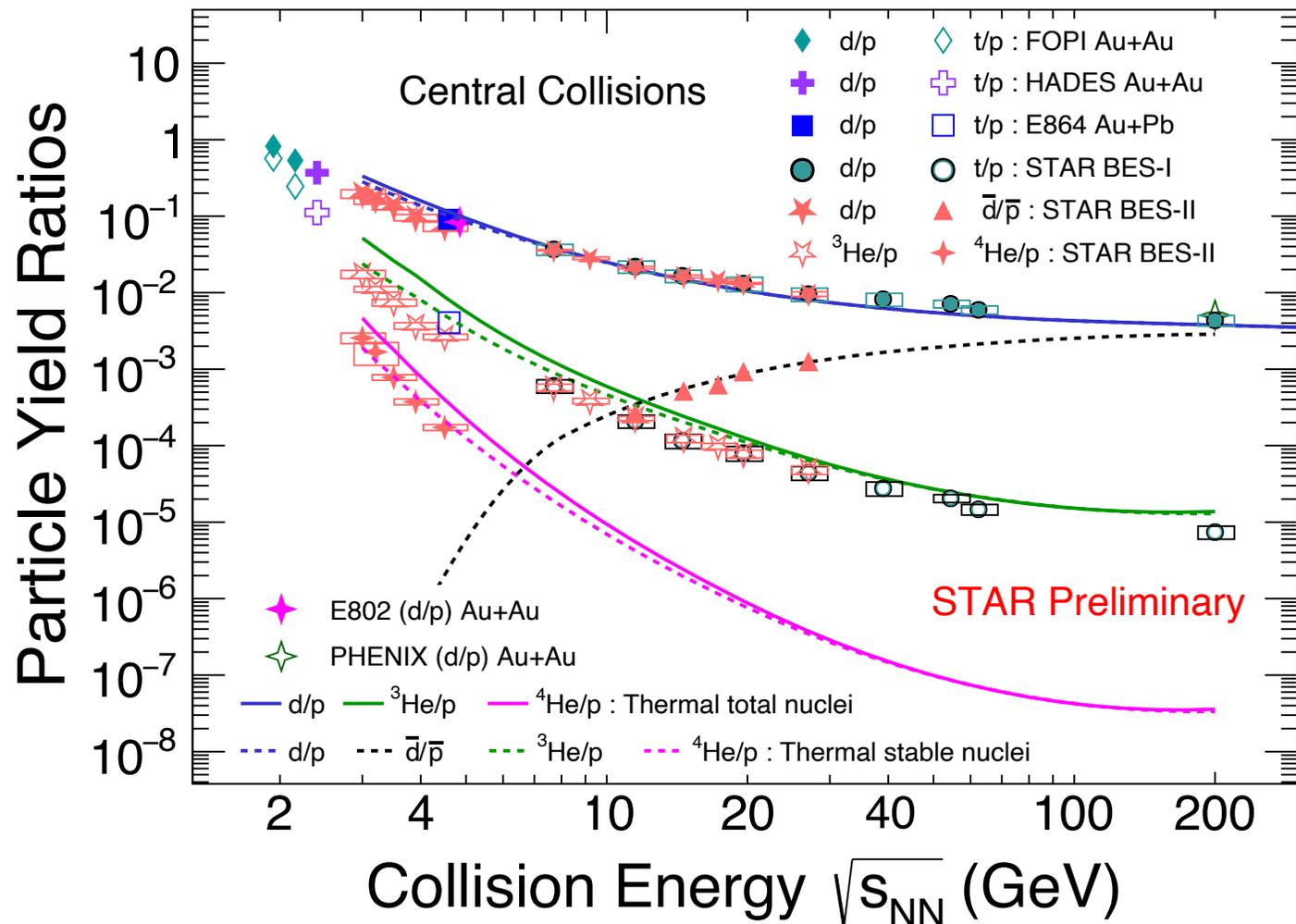
# Averaged Transverse Momentum $\langle p_T \rangle$



STAR Collaboration. *Phys.Rev.C* 110 (2024) 5, 054911

- $\langle p_T \rangle$  of protons and light nuclei as a function of centrality, rapidity, and collision energy.
- Hint of  $\langle p_T \rangle$  increase with energy for 4.5 GeV and below, flat trend between 7.7 and 19.6 GeV. This behavior will be further studied in 4.5 - 7.7 GeV in the future.

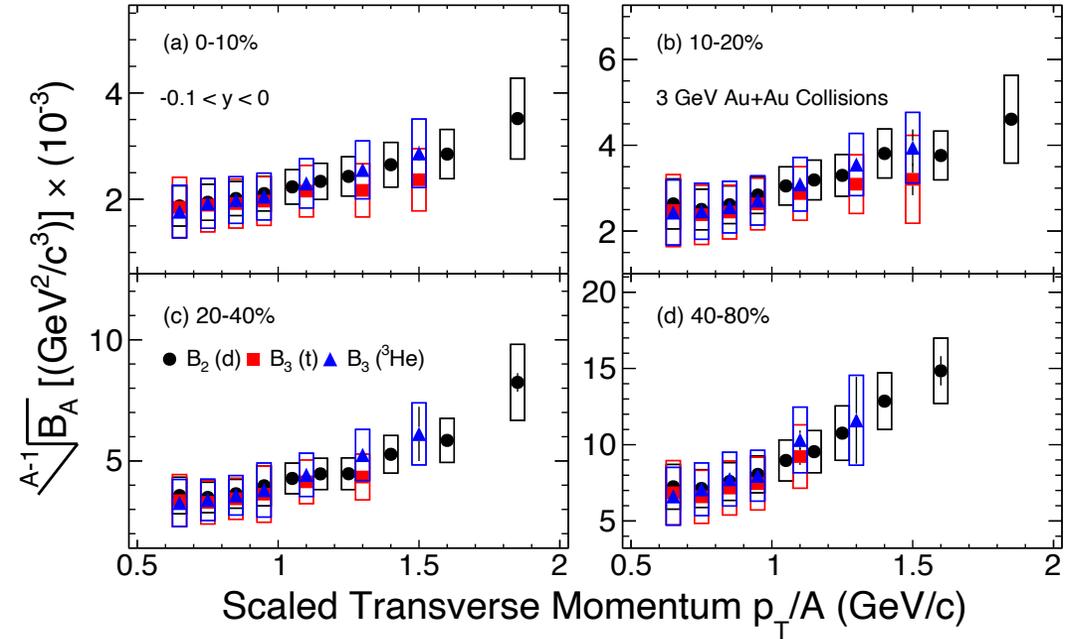
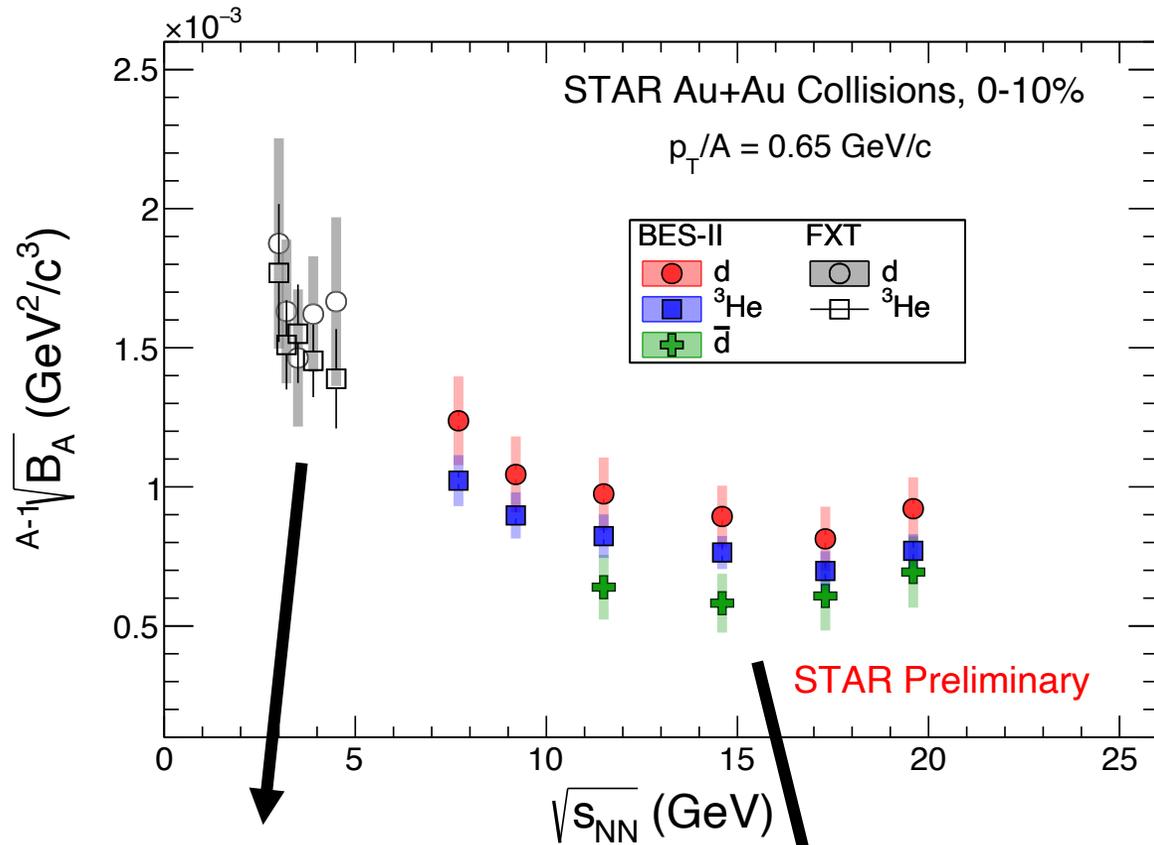
# Particle Ratio



- Clear energy dependence is observed for both  $d/p$ ,  $\bar{d}/\bar{p}$ ,  $t/p$ ,  $^3\text{He}/p$ , and  $^4\text{He}/p$  ratios.
- The trends of ratios can be described qualitatively by the thermal model.
- $^3\text{He}/p$  is overestimated by thermal model, possibly due to the hadronic re-scattering effect.
- Considering only stable nuclei,  $^4\text{He}/p$  from thermal model is consistent with the experiment data.

[STAR Collaboration] *Phys. Rev. C* 96, 044904 (2017); *Phys.Rev.Lett.* 130 (2023) 202301;  
 [E802 Collaboration] *Phys.Rev.C* 60 (1999) 064901; [E864 Collaboration] *Phys.Rev.C* 61 (2000) 064908;  
 [FOPI Collaboration] *Nucl.Phys.A* 848 (2010) 366-427; V. Vovchenko, et al. *Phys. Rev. C* 93(2016) 6, 064906;

# Coalescence parameters $B_A$



$$E_A \frac{d^3 N_A}{d^3 p_A} = B_A \left( E_p \frac{d^3 N_p}{d^3 p_p} \right)^Z \left( E_n \frac{d^3 N_n}{d^3 p_n} \right)^{A-Z} \approx B_A \left( E_p \frac{d^3 N_p}{d^3 p_p} \right)^A, B_A \propto \left( \frac{1}{V_{eff}} \right)^{(A-1)}$$

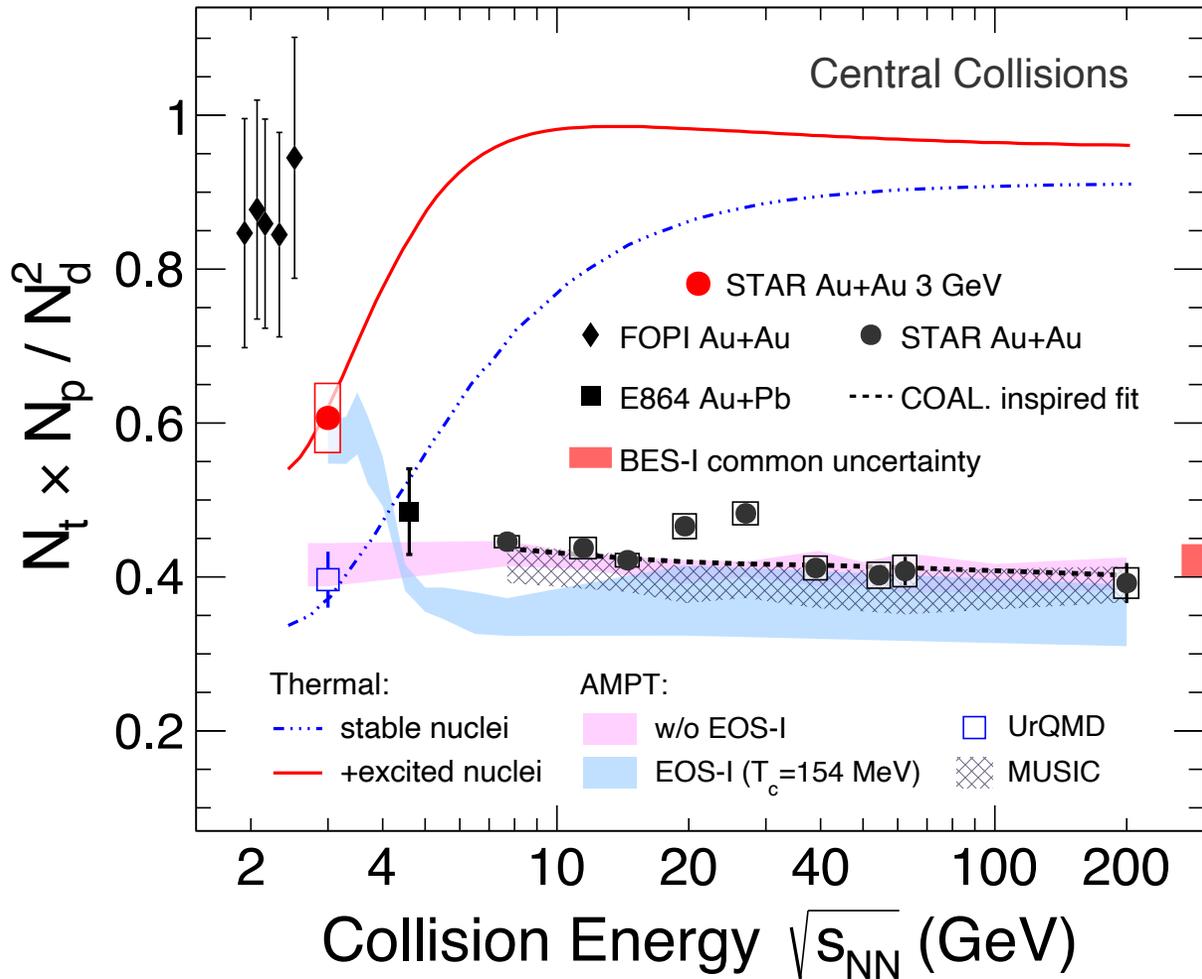
$V_{eff}$  is the effective volume of the nucleon emission source.

- As the energy increases,  $B_A$  becomes smaller, reflecting that the effective volume of the system becomes larger.
- Length of homogeneity becomes smaller in peripheral collisions and at higher  $p_T$  region.

STAR Collaboration. *Phys.Rev.C* 110 (2024) 5, 054911

R. Scheibl and U. Heinz *Phys.Rev.C* 59 (1999) 1585-1602; S. Zhang et al. *Phys.Lett.B* 684 (2010) 224-227

# Compound Ratio



$$\triangleright \frac{N_t \times N_p}{N_d^2} \approx \frac{1}{2\sqrt{3}} \left[ 1 + \Delta n + \frac{\lambda}{\sigma} G\left(\frac{\xi}{\sigma}\right) \right]$$

$\Delta n$ : Neutron Density Fluctuation

$G\left(\frac{\xi}{\sigma}\right)$ : Long-range Correlation

$\triangleright$  Energy dependence of  $\frac{N_t \times N_p}{N_d^2}$  in most central Au+Au collisions.

$\triangleright$  AMPT model employing a first order phase transition with input critical temperature of 154 MeV can reproduced 3 GeV yield ratio.

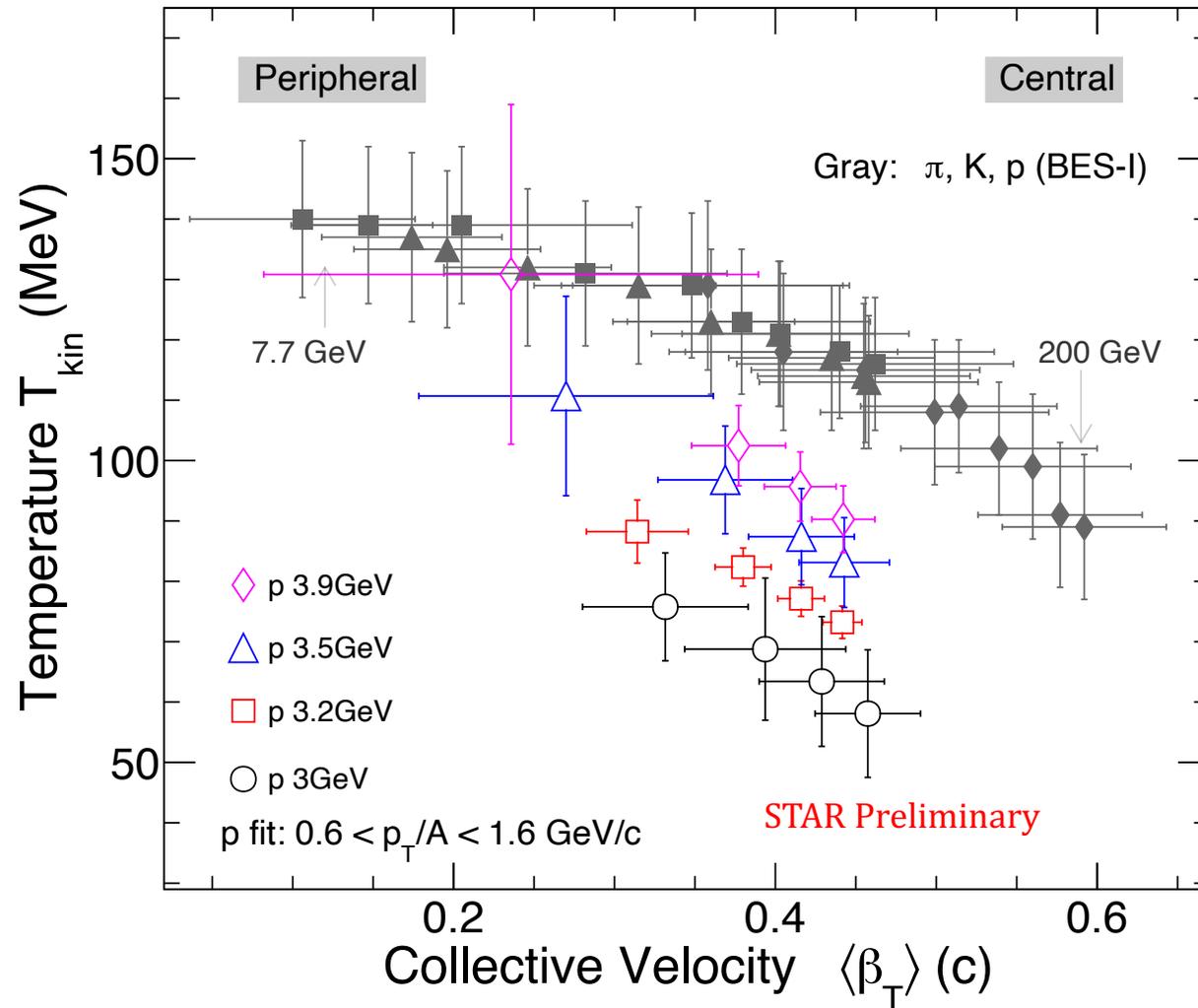
$\triangleright$   $\sqrt{s_{NN}} = 3.2-27$  GeV compound ratio from STAR BESII is in progress.

*K. Sun et al. Phys.Rev.C 103 (2021) 6, 064909*

*STAR Collaboration. Phys.Rev.C 110 (2024) 5, 054911*

# Kinetic Freeze-out Dynamics

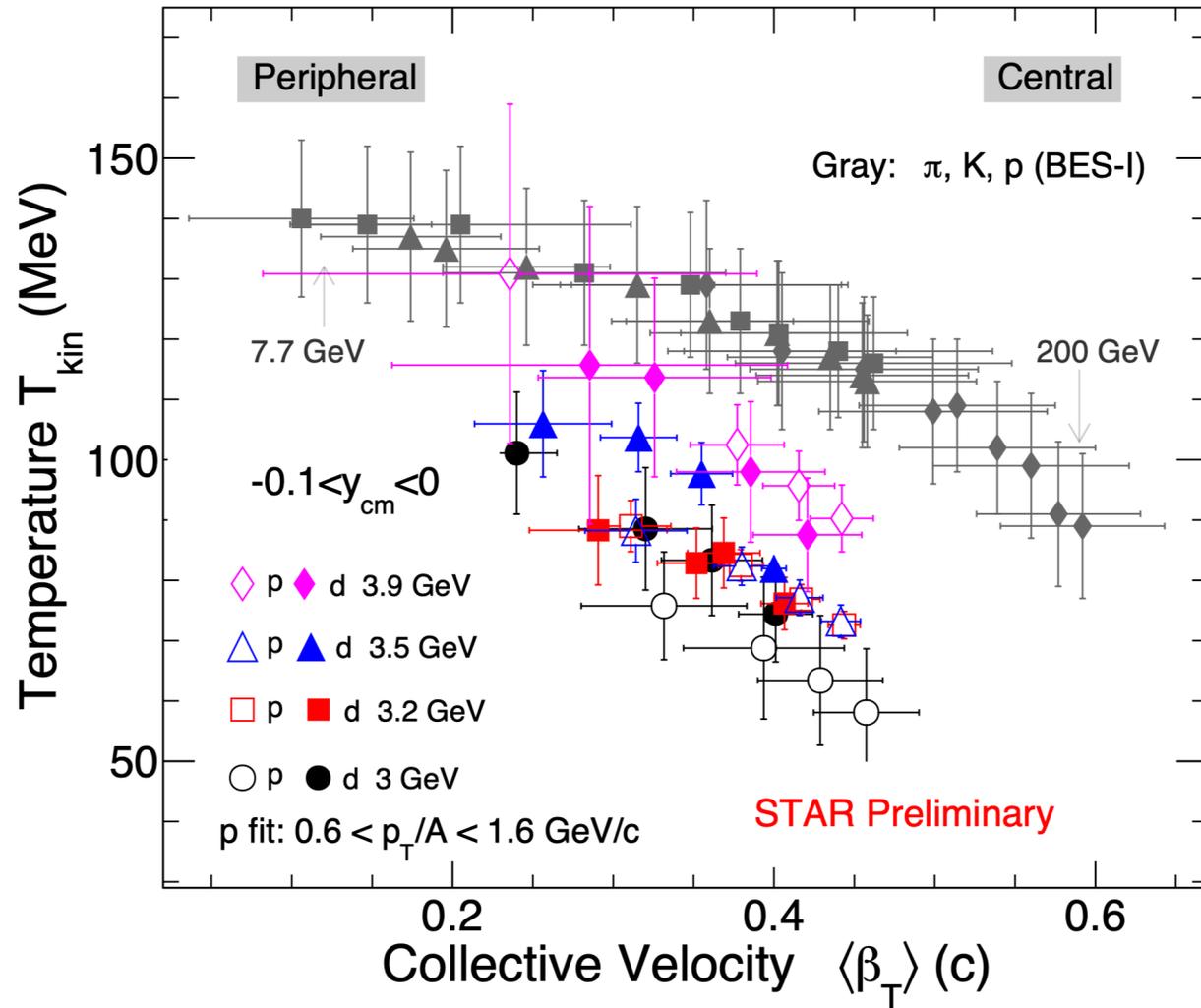
Au + Au Collisions at Mid-rapidity



➤ For  $\sqrt{s_{NN}} = 3.0-3.9$  GeV, proton  $T_{kin}$  increases with energy while  $\langle \beta_T \rangle$  stays approximately constant. This trend is different for  $\sqrt{s_{NN}} \geq 7.7$  GeV, may imply a different medium equation of state (EoS).

# Kinetic Freeze-out Dynamics

Au + Au Collisions at Mid-rapidity



- The differing trends in  $T_{kin}$  and  $\langle\beta_T\rangle$  for protons and deuterons ( $\sqrt{s_{NN}} = 3.0-3.9$  GeV) imply they have distinct kinetic freeze-out surfaces.
- $T_{kin}$  versus  $\langle\beta_T\rangle$  distribution shows a clear gap region between 3 GeV and energies above 7.7 GeV.
- The gap can be filled by collision energies  $\sqrt{s_{NN}} = 3.0 - 3.9$  GeV, may imply a different medium equation of state (EoS).

# Summary and Outlook

## ➤ **Summary:**

- ✓ We presented light nuclei production ( $p_T$  spectra,  $dN/dy$ ,  $\langle p_T \rangle$ , particle ratio, and  $B_A$ ) in Au+Au collisions at  $\sqrt{s_{NN}} = 3.0-27$  GeV by STAR experiment, studying their rapidity and energy dependence.
- ✓ The thermal model overestimates light nuclei ratio  $d/p$  and  ${}^3\text{He}/p$ , but consistent with  ${}^4\text{He}/p$  only considering stable nuclei.
- ✓ The extracted kinetic freeze-out parameters ( $T_{kin}$ ,  $\langle \beta_T \rangle$ ) may imply that the equation of state describing the hot, dense nuclear matter at low collision energies ( $\sqrt{s_{NN}} = 3.0-3.9$  GeV) differs from that observed at higher energies.

## ➤ **Outlook:**

- Continue to calculate the compound ratio ( $N_t \times N_p / N_d^2$ ) in Au+Au collisions at  $\sqrt{s_{NN}} = 3.2-27$  GeV.
- Measure  ${}^4\text{He}$  production in  $\sqrt{s_{NN}} \geq 7.7$  GeV to gain more insight on  $N_{{}^4\text{He}} \times N_p / (N_{{}^3\text{He}} \times N_d)$  and production mechanism.

*Thank you*