

International Journal of Modern Physics E
 © World Scientific Publishing Company

Investigating the CME in isobaric ($^{96}\text{Ru}+^{96}\text{Ru}$ and $^{96}\text{Zr}+^{96}\text{Zr}$) collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV using Sliding Dumbbell Method with the STAR detector at RHIC

Jagbir Singh (for the STAR Collaboration) *

*Instituto de Alta Investigación, Universidad de Tarapacá, Casilla 7D, Arica, 1000000, Chile
 jsingh2@bnl.gov*

1 Received Day Month Year
 2 Revised Day Month Year

3 The chiral imbalance, coupled with the presence of a strong magnetic field produced
 4 during heavy-ion collisions, can result in charge separation along the magnetic field
 5 axis, a phenomenon known as the Chiral Magnetic Effect (CME). A novel technique,
 6 the Sliding Dumbbell Method (SDM) has been developed to investigate the CME with
 7 the RHIC's isobar program. The SDM facilitates the selection of events correspond-
 8 ing to various charge separations (f_{D_bCS}) across the dumbbell. A partitioning of the
 9 charge separation distributions for each collision centrality into ten percentile bins is
 10 done in order to find potential CME-like events corresponding to the highest charge
 11 separation across the dumbbell. The study reports the results on CME sensitive γ -
 12 correlator ($\gamma = \langle \cos(\phi_a + \phi_b - 2\Psi_{RP}) \rangle$) and δ -correlator ($\delta = \langle \cos(\phi_a - \phi_b) \rangle$) for each
 13 bin of f_{D_bCS} in each collision centrality for isobaric collisions (Ru+Ru and Zr+Zr) at
 14 $\sqrt{s_{\text{NN}}} = 200$ GeV measured with the STAR detector. Furthermore, the background
 15 scaled ratio ($\Delta\gamma_{\text{Ru/Zr}}/\Delta\gamma_{\text{Bkg}}$) is presented to check for the expected enhancement of
 16 the CME in Ru+Ru collisions as compared to Zr+Zr collisions.

17 *Keywords:* Quark-Gluon plasma; chiral magnetic effect; heavy-ion collisions; multi-
 18 particle correlations; Q-cumulant.

19 **1. Introduction**

20 Quantum Chromodynamics (QCD) predicts the existence of metastable domains
 21 characterized by fluctuating topological charges, which can induce chirality imbal-
 22 ance in quarks under extreme conditions of temperature and/or density, such as
 23 those present during the formation of the quark-gluon plasma (QGP).^{1,2} In non-
 24 central heavy-ion collisions, P-odd meta-stable states and the strong magnetic field
 25 generated by highly energetic spectator protons leads to the separation of oppo-
 26 sitely charged particles along the magnetic field direction and perpendicular to the
 27 reaction plane, known as the Chiral Magnetic Effect (CME).³⁻⁶ The CME is under
 28 active theoretical⁷⁻¹⁰ and experimental¹¹⁻¹⁵ investigation in relativistic heavy-ion

*J. Singh (for the STAR Collaboration)

2 *J. Singh*

29 collisions, particularly at the Relativistic Heavy Ion Collider (RHIC) and the Large
30 Hadron Collider (LHC).

31 Isobaric collisions of ${}^{96}_{44}\text{Ru}+{}^{96}_{44}\text{Ru}$ and ${}^{96}_{40}\text{Zr}+{}^{96}_{40}\text{Zr}$ nuclei were proposed as a
32 promising approach to address the challenges associated with the detection of the
33 CME in heavy-ion collisions.^{16,17} The STAR experiment at RHIC collected data
34 in 2018 on $Ru + Ru$ and $Zr + Zr$ collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV to search for
35 the CME in isobaric collision systems.¹⁷ The larger atomic number of Ruthenium
36 (${}^{96}_{44}\text{Ru}$) compared to Zirconium (${}^{96}_{40}\text{Zr}$) leads to an increase of approximately 15%
37 in the squared magnetic field in Ru+Ru collisions. This enhanced magnetic field is
38 expected to give rise to a proportional increase in the CME contribution in Ru+Ru
39 collisions, while the similarity in mass numbers in both isobars ensures compara-
40 ble flow-driven backgrounds. After rigorous data analysis and examination of vari-
41 ous CME-sensitive observables, no significant enhancement of the CME signal was
42 found in Ru+Ru collisions compared to Zr+Zr collisions, within uncertainties.¹⁸

43 The most commonly used observable in the search for the Chiral Magnetic
44 Effect (CME) is the three particle γ -correlator, initially proposed by Voloshin.³ It
45 is defined as:

$$\begin{aligned} \gamma_{a,b} &= \langle \cos(\phi_a + \phi_b - 2\Psi_{RP}) \rangle \approx \frac{\langle \cos(\phi_a + \phi_b - 2\phi_c) \rangle}{v_{2,c}} \\ &= \langle \cos(\Delta\phi_a) \cos(\Delta\phi_b) \rangle - \langle \sin(\Delta\phi_a) \sin(\Delta\phi_b) \rangle \end{aligned} \quad (1)$$

46 where, ϕ_a and ϕ_b denote the azimuthal angles of particles “a” and “b”, respectively,
47 while Ψ_{RP} represents the reaction plane angle. $v_{2,c}$ is the elliptic flow of third
48 particle “c” having azimuthal angle ϕ_c . $\Delta\phi_a$ and $\Delta\phi_b$ represent the azimuthal
49 angles of particles “a” and “b” measured with respect to the reaction plane. To
50 estimate v_2 and compute 2- and 3-particle correlators, the Q-cumulant technique¹⁹
51 is used.

52 The reaction plane independent 2-particle δ -correlator is also used, which is as
53 follows:

$$\delta_{a,b} = \langle \cos(\phi_a - \phi_b) \rangle = \langle \cos(\Delta\phi_a) \cos(\Delta\phi_b) \rangle + \langle \sin(\Delta\phi_a) \sin(\Delta\phi_b) \rangle \quad (2)$$

54 From equations (1) and (2), one can determine in-plane ($\langle \cos(\Delta\phi_a) \cos(\Delta\phi_b) \rangle$) and
55 out-of-plane ($\langle \sin(\Delta\phi_a) \sin(\Delta\phi_b) \rangle$) correlations to study the preferential emission
56 of charged particles.

57 In this study, the Sliding Dumbbell Method (SDM)^{20–22} is utilized to identify po-
58 tential CME-like events. These events are subsequently analyzed using the γ and
59 δ correlators to verify whether or not they exhibit the expected characteristics of
60 CME events.

61 2. Sliding Dumbbell Method

62 The “Sliding Dumbbell Method”^{20–22} has been designed to identify potential CME-
63 like events in heavy-ion collisions that display higher back-to-back charge separation

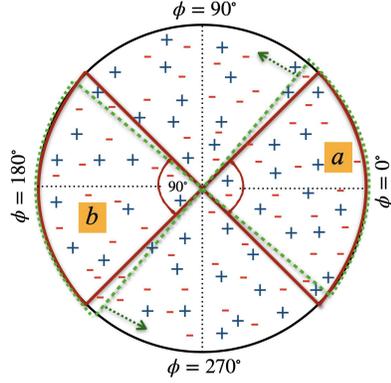


Fig. 1. (Color online) A graphical representation of the transverse (azimuthal) plane showing positive (+) and negative (-) particle hits in an event.

64 on an event-by-event basis. In the SDM, the azimuthal plane of each event is scanned
 65 by sliding a dumbbell-shaped region of $\Delta\phi = 90^\circ$ in steps of $\delta\phi = 1^\circ$ as shown in
 66 Fig. 1. This approach helps in identifying the region exhibiting the highest back-
 67 to-back charge separation. To quantify this separation, the parameter Db_{+-} is
 68 calculated. It represents the sum of the positive charge fraction on one side (“a”) of
 69 the dumbbell and the negative charge fraction on the opposite side (“b”) for each
 70 orientation of the dumbbell across the azimuthal plane, is obtained as:

$$Db_{+-} = \frac{n_+^a}{(n_+^a + n_-^a)} + \frac{n_-^b}{(n_+^b + n_-^b)}, \quad (3)$$

71 where, $n_a^+(n_b^+)$ and $n_a^-(n_b^-)$ represent the number of positive and negative charged
 72 particles on sides “a” and “b” of the dumbbell, respectively. The $Db_{+-} = 2$ corre-
 73 sponds to 100% back-to-back charge separation while $Db_{+-} = 1$ means no back-
 74 to-back charge separation. Furthermore, the charge excess asymmetry across the
 75 dumbbell, Db_{asy} , is defined as:

$$Db_{+-}^{asy} = \frac{(n_a^+ - n_a^-) - (n_b^- - n_b^+)}{(n_a^+ - n_a^-) + (n_b^- - n_b^+)} \quad (4)$$

76 Here, $n_a^+ - n_a^-$ represents the excess positive charge on the “a” side of the dumb-
 77 bell, while $n_b^- - n_b^+$ denotes the excess negative charge on the “b” side. By sliding
 78 the dumbbell in steps of $\delta\phi = 1^\circ$ across the azimuthal plane, 360 values of both
 79 Db_{+-} and Db_{asy} are obtained. The maximum value of Db_{+-} , denoted as Db_{+-}^{max} ,
 80 is then selected under the constraint $|Db_{asy}| < 0.25$, ensuring the identification
 81 of CME-like events with balanced charge excess asymmetry. The charge separation
 82 across the dumbbell, f_{DbCS} , referred to as charge separation, defined as:

83

$$f_{DbCS} = Db_{+-}^{max} - 1 \quad (5)$$

4 *J. Singh*

84 The charge separation (f_{D_bCS}) distributions are obtained for each centrality
 85 class for data as well as charge shuffle background (as discussed in section 2.1).
 86 These distributions are subdivided into ten percentile bins, ranging from 0-10%
 87 (highest charge separation) to 90-100% (lowest charge separation) for each collision
 88 centrality. Two (δ) and three particle (γ) correlators are computed for different
 89 charge combinations and for each f_{D_bCS} bin in each centrality for both data and
 90 background.

91 **2.1. Background Estimation**

92 We estimate the background contributions to the γ -correlator across different f_{D_bCS}
 93 percentile bins using the SDM, we account for contributions that may result in
 94 higher charge separation purely by statistical fluctuations while preserving the in-
 95 trinsic particle correlations. This is achieved by randomly shuffling the charges of
 96 particles within each event while keeping their momenta (i.e., θ and ϕ) unchanged.
 97 The charge-shuffled sample for a given centrality is then analyzed in the same man-
 98 ner as the original dataset.²⁰ The γ -correlator obtained from this charge-shuffled
 99 sample in a particular f_{D_bCS} bin is denoted as γ_{ChS} .

100 Conversely, the charge correlations disrupted by the shuffling process are re-
 101 covered from the original events corresponding to the same f_{D_bCS} bins and the
 102 resulting γ -correlator is referred to as γ_{Corr} . Consequently, the total background
 103 contribution to the γ -correlator is calculated as:

104

$$\gamma_{Bkg} = \gamma_{ChS} + \gamma_{Corr} \quad (6)$$

105 **2.2. Data Analyzed**

106 We used the minimum bias events from the isobar collisions ($^{96}_{44}\text{Ru}+^{96}_{44}\text{Ru}$ and
 107 $^{96}_{40}\text{Zr}+^{96}_{40}\text{Zr}$) at $\sqrt{s_{NN}} = 200$ GeV. Events with $-35 < V_z < 25$ cm and tracks with
 108 $|\eta| < 1$, $0.2 < p_T < 2.0$ GeV/ c and $DCA < 3$ cm are used for the analysis. After
 109 all event selection cuts, we analyze approximately 1.7 billion minimum-bias (MB)
 110 events each for Ru+Ru and Zr+Zr collisions.

111 **3. Results and Discussions**

112 The f_{D_bCS} distributions for Ru+Ru and Zr+Zr collisions are compared in Fig. 2 for
 113 different collision centralities. It is observed that $f_{D_bCS}^{Ru+Ru}$ distributions are almost
 114 similar to $f_{D_bCS}^{Zr+Zr}$ distributions. Additionally, these distributions shift toward higher
 115 values of f_{D_bCS} as the collision centrality decreases.

116 The dependence of γ -correlator (left) and δ -correlator (right) on f_{D_bCS} is dis-
 117 played in Fig. 3 for opposite-sign and same-sign charge pairs for 0-60% collision cen-
 118 trality for Ru+Ru (blue color) and Zr+Zr (red color) collisions at $\sqrt{s_{NN}} = 200$ GeV.
 119 It is observed that $\gamma_{OS} > 0$ and $\gamma_{SS} < 0$ for the highest f_{D_bCS} bins, specifically 0-
 120 20% (0-30%) for 0-40% (40-60%) collision centralities, consistent with expectations

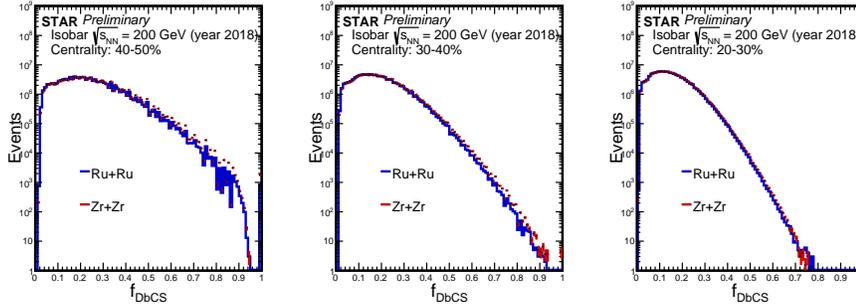


Fig. 2. (Color online) Comparison of f_{DbCS} distributions for Ru+Ru and Zr+Zr collisions at $\sqrt{s_{NN}} = 200$ GeV for different collision centralities.

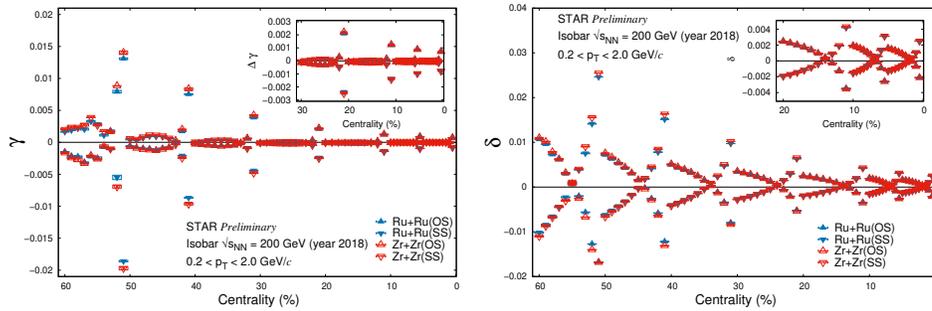


Fig. 3. (Color online) γ (Left) and δ (Right) dependence on f_{DbCS} for 0-60% collision centralities for Ru+Ru and Zr+Zr collisions at $\sqrt{s_{NN}} = 200$ GeV. Boxes represent the systematic errors while the statistical errors (represented by bars) are within the marker sizes.

121 for CME-like events. Also, $\delta_{SS} > 0$ and $\delta_{OS} < 0$ for the top f_{DbCS} bins, while for
 122 lower f_{DbCS} bins, $\delta_{SS} < 0$ and $\delta_{OS} > 0$. Additionally, it is noteworthy that the
 123 magnitude of the γ (δ)-correlator is slightly higher for Zr+Zr collisions compared to
 124 Ru+Ru collisions.

125 Figure 4 shows the dependence of in-plane and out-of-plane correlations on
 126 f_{DbCS} for opposite-sign (left) and same-sign (right) charge pairs in Ru+Ru and
 127 Zr+Zr collisions at $\sqrt{s_{NN}} = 200$ GeV for 0-60% collision centralities. For the top
 128 20% f_{DbCS} bins, out-of-plane correlations are stronger than in-plane correlations
 129 for both same-sign and opposite sign charge pairs, suggesting a out-of-plane charge
 130 separation.

131 Figure 5 illustrates the dependence of $\Delta\gamma$ on f_{DbCS} for Ru+Ru (Left) and
 132 Zr+Zr (Right) collisions, along with their respective backgrounds i.e., $ChS_{Ru/Zr}$
 133 and $Corr_{Ru/Zr}$, for 0-60% collision centralities. The values of $\Delta\gamma$'s in the top f_{DbCS}
 134 bins are enhanced many times than the average values.¹⁸ The data points ($\Delta\gamma_{data}$)
 135 for the top 20% f_{DbCS} bins have larger values than the total background contribu-

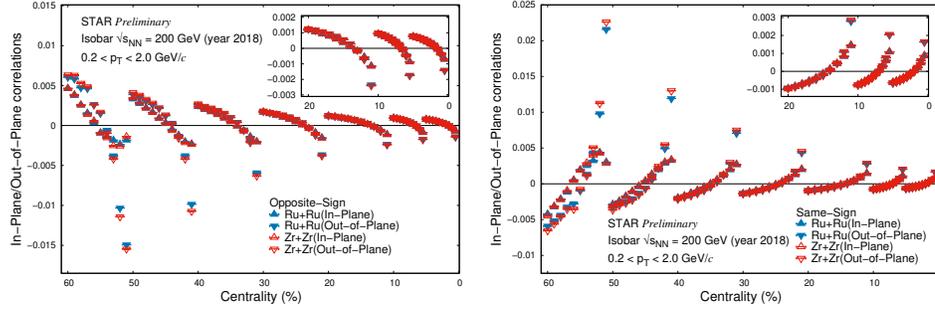
6 *J. Singh*


Fig. 4. (Color online) The dependence of in-plane and out-of-plane correlations on f_{D_bCS} for 0-60% collision centralities in Ru+Ru and Zr+Zr collisions at $\sqrt{s_{NN}} = 200$ GeV is shown for opposite-sign (left) and same-sign (right) charge pairs.

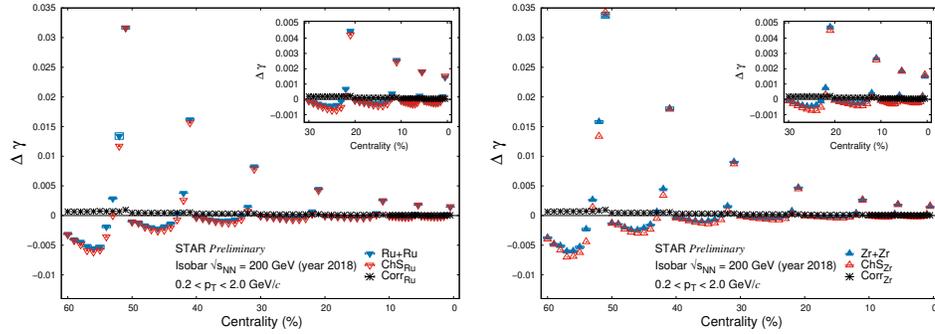


Fig. 5. (Color online) $\Delta\gamma$ dependence on f_{D_bCS} for Ru+Ru (Left) and Zr+Zr (Right) collisions, including $ChS_{Ru(Zr)}$ and $Corr_{Ru(Zr)}$ backgrounds for 0-60% collision centralities. The boxes indicate the systematic errors, while the statistical errors, represented by bars, are within the marker sizes.

136 tion ($\Delta\gamma_{ChS} + \Delta\gamma_{Corr}$) in the 30-50% centrality range.

137 The ratio $\Delta\gamma_{Data}/\Delta\gamma_{Bkg}$ for Ru+Ru and Zr+Zr collisions for top 20% f_{D_bCS}
 138 bins is presented for the first time in Fig. 6. It is observed that this ratio exceeds
 139 unity by approximately 2-6% for collision centralities within the 10-50% range.
 140 The ratios agree within uncertainties for both Ru+Ru and Zr+Zr collisions. No
 141 significant enhancement is observed in the background-scaled $\Delta\gamma$ of Ru+Ru relative
 142 to Zr+Zr for the top 20% f_{D_bCS} bins.

143 Figure 7 presents the “double ratio”, which compares the $\Delta\gamma_{Data}/\Delta\gamma_{Bkg}$ ratios
 144 between Ru+Ru and Zr+Zr collisions for top 20% f_{D_bCS} bins. A “Pol0 fit” (straight
 145 line fit) was applied to the double ratio, resulting in a fitted value of $1.007 \pm$
 146 0.003 for 0-60% centralities. This result suggests no significant enhancement of the
 147 CME signal in Ru+Ru collisions relative to Zr+Zr collisions, contrary to the initial
 148 expectations for isobar collisions.¹⁸

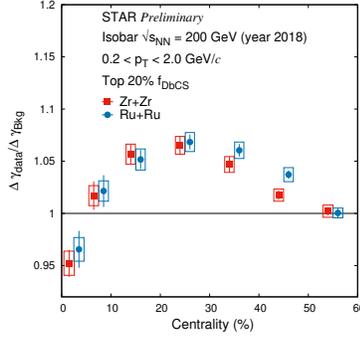


Fig. 6. (Color online) $\Delta\gamma_{Data}/\Delta\gamma_{Bkg}$ for Ru+Ru and Zr+Zr collisions at $\sqrt{s_{NN}} = 200$ GeV as a function of collision centrality. The statistical errors are represented by bars, while the systematic errors are depicted by boxes.

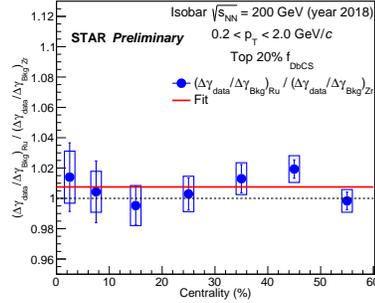


Fig. 7. (Color online) The double ratio $\left(\frac{\Delta\gamma_{Data}/\Delta\gamma_{Bkg}}{\Delta\gamma_{Data}/\Delta\gamma_{Bkg}}\right)_{Ru}$ for 0-60% collision centralities. A straight-line (Pol0) fit is applied to the data, represented by the red line. The statistical errors are represented by bars, while the systematic errors are depicted by boxes.

149 4. Conclusions

150 In this study, isobaric collision data was analyzed using the Sliding Dumbbell
 151 Method (SDM), which identifies back-to-back charge separation on an event-by-
 152 event basis within each collision centrality. The γ and δ correlators exhibited a
 153 significant enhancement in the top f_{DbCS} bins compared to the average values
 154 across centralities for both same-sign (SS) and opposite-sign (OS) charge pairs.
 155 The γ -correlator was found to be positive for OS pairs and negative for SS pairs,
 156 while the δ -correlator displayed the opposite trend in top f_{DbCS} bins. Addition-
 157 ally, both SS and OS charge pairs exhibited stronger out-of-plane correlations in
 158 the top f_{DbCS} bins, consistent with out-of-plane charge separation, indicating that
 159 these bins contain the potential CME-like events. Similar trends were observed in
 160 charge-shuffle background but with reduced magnitudes in both isobars. However,

8 *J. Singh*

161 the double ratio $\left(\frac{\Delta\gamma_{Data}/\Delta\gamma_{Bkg}}{\Delta\gamma_{Data}/\Delta\gamma_{Bkg}}\right)_{Ru}$ for top 20% f_{DvCS} bins suggests no significant
 162 enhancement of the CME signal in Ru+Ru collisions compared to Zr+Zr collisions,
 163 contradicting initial expectations for isobaric collisions. In the various techniques
 164 studied by the STAR experiment, the anticipated enhancement in Ru+Ru, due to a
 165 stronger magnetic field, has not been observed. Further theoretical and experimen-
 166 tal investigations are necessary to refine the interpretation of these findings and to
 167 develop more robust methodologies for isolating the CME signatures in heavy-ion
 168 collisions.

169 References

- 170 1. E. V. Shuryak, *Phys. Lett. B* **78**, 150 (1978).
- 171 2. D. Kharzeev, R.D. Pisarski and M.H.G. Tytgat, *Phys. Rev. Lett.* **81**, 512 (1998).
- 172 3. S. A. Voloshin, *Phys. Rev. C* **70**, 057901 (2004).
- 173 4. D. Kharzeev, *Phys. Lett. B* **633**, 260 (2006).
- 174 5. K. Fukushima, D. E. Kharzeev, and H. J. Warringa, *Phys. Rev. D* **78**, 074033 (2008).
- 175 6. D. E. Kharzeev, L. D. McLerran, and H. J. Warringa, *Nucl. Phys. A* **803**, 227 (2008).
- 176 7. D. Kharzeev, J. Liao, S. Voloshin and G. Wang, *Progress in Particle and Nuclear*
 177 *Physics* **88**, 1 (2016).
- 178 8. V. Koch, S. Schlichting, V. Skokov, P. Sorensen, J. Thomas, S. Voloshin et al., *Chinese*
 179 *Physics C* **41**, 072001 (2017).
- 180 9. B. Schenke, C. Shen and P. Tribedy, *Phys. Rev. C* **99**, 044908 (2019).
- 181 10. J. Zhao and F. Wang, *Progress in Particle and Nuclear Physics* **107**, 200 (2019).
- 182 11. B. I. Abelev et al. [STAR collaboration], *Phys. Rev. Lett.* **103**, 251601 (2009).
- 183 12. L. Adamczyk et al. [STAR collaboration], *Phys. Rev. C* **88**, 064911 (2013).
- 184 13. A. M. Sirunyan et al. [CMS collaboration], *Phys. Rev. C* **97**, 044912 (2018).
- 185 14. S. Acharya et al. [ALICE collaboration], *Phys. Lett. B* **777**, 151 (2018).
- 186 15. M. S. Abdallah et al. [STAR collaboration], *Phys. Rev. Lett.* **128**, 092301 (2022).
- 187 16. S. A. Voloshin, *Phys. Rev. Lett.* **105**, 172301 (2010).
- 188 17. V. Koch, et al., *Chin. Phys. C* **41**, 072001 (2017).
- 189 18. M. S. Abdallah et al. [STAR Collaboration], *Phys. Rev. C* **105**, 014901 (2022).
- 190 19. A. Bilandzic, R. Snellings and S. Voloshin, *Phys. Rev. C* **83**, 044913 (2011).
- 191 20. M.M. Aggarwal, A. Attri, S. Parmar, A. Sharma and J. Singh, *Pramana - J. Phys*
 192 **98**, 117 (2024).
- 193 21. J. Singh (for the STAR Collaboration), *Springer Proc. Phys.* **304**, 464 (2024).
- 194 22. Jagbir Singh, Ph.D. Thesis, Panjab University, Chandigarh, INDIA (2024).