

# Investigating the CME in isobaric ( $^{96}_{44}\text{Ru}+^{96}_{44}\text{Ru}$ and $^{96}_{40}\text{Zr}+^{96}_{40}\text{Zr}$ ) collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV using Sliding Dumbbell Method with the STAR detector at RHIC

*Jagbir Singh*

*(for the STAR Collaboration)*

*Instituto de Alta Investigación, Universidad de Tarapacá, Arica, Chile*

email: jagbir@rcf.rhic.bnl.gov

## Abstract

1 The chiral imbalance, coupled with the presence of a strong magnetic  
2 field produced during heavy-ion collisions, results in charge separation along  
3 the magnetic field axis, a phenomenon known as the Chiral Magnetic Effect  
4 (CME). A novel technique, the Sliding Dumbbell Method (SDM) [1, 2] has  
5 been developed to investigate the CME with the RHIC's isobar program.  
6 The SDM facilitates the selection of events corresponding to various charge  
7 separations ( $f_{D\text{bCS}}$ ) across the dumbbell. A partitioning of the charge sep-  
8 aration distributions for each collision centrality into ten percentile bins is  
9 done in order to find potential CME-like events corresponding to the high-  
10 est charge separation across the dumbbell. The study reports the results on  
11 CME sensitive  $\gamma$ -correlator ( $\gamma = \langle \cos(\phi_a + \phi_b - 2\Psi_{RP}) \rangle$ ) and  $\delta$ -correlator  
12 ( $\delta = \langle \cos(\phi_a - \phi_b) \rangle$ ) for each bin of  $f_{D\text{bCS}}$  in each collision centrality for iso-  
13 baric collisions (Ru+Ru and Zr+Zr) at  $\sqrt{s_{\text{NN}}} = 200$  GeV measured with the  
14 STAR detector. Furthermore, the background scaled ratio ( $\Delta\gamma_{\text{Ru/Zr}}/\Delta\gamma_{\text{Bkg}}$ )  
15 will be presented to check for the expected enhancement of the CME in  
16 Ru+Ru collisions as compared to Zr+Zr collisions. Overall, this research  
17 aims to understand and detect the CME through an innovative experimental  
18 method.

## 19 References

- 20 [1] J. Singh, A. Attri, and M. M. Aggarwal, Proceedings of the DAE Symp.  
21 on Nucl. Phys. **64**, 830 (2019).
- 22 [2] J. Singh (for STAR Collaboration), Springer Proc. Phys. **304**, 464  
23 (2024).