# Spectra and radial flow at RHIC with Tsallis distribution in Blast Wave model

Zebo Tang, Fuqiang Wang, Yichun Xu, Zhangbu Xu Thanks: Lijuan Ruan, Gene van Buren, Aihong Tang, Mike Lisa

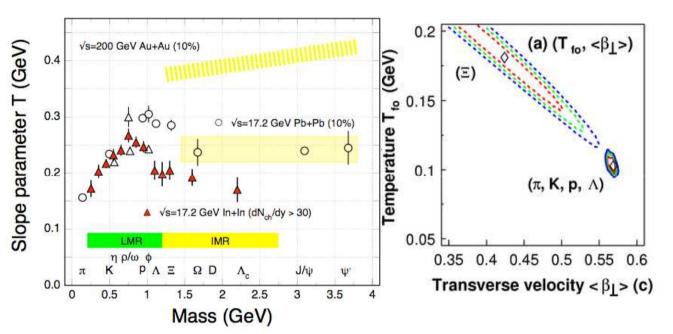
- •What physics can spectra address?
- Why do we need a new BlastWave model
- How to implement Tsallis statistics in BlastWave framework
- •Can spectra tell us about fluctuation and bulk viscosity?
- •Who said p+p spectra are similar to Au+Au?
- Summary and Outlook

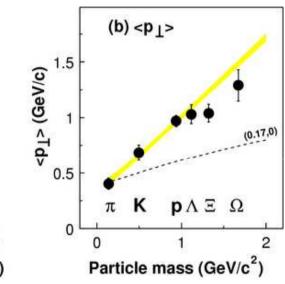
## What physics can Spectra tell us?

- Low p<sub>T</sub>
  - Integrated particle yields (dN/dy)
  - Radial Flow and freeze-out temperature
- Intermediate p<sub>⊤</sub>
  - Coalescence
- High p<sub>T</sub>
  - Jet quenching
- What are the connections among them
  - Bulk medium interaction and pressure gradient drives thermalization and radial flow
  - Thermalization and quark degree of freedom provides quark coalescence
  - Jet quenching dissipates energy into the system
- Bulk Viscosity, Fluctuation?

# m<sub>T</sub> slope vs mass

#### Nu Xu's plot





Nu Xu, QM2008

STAR whitepaper, PRL92(2004)

$$T_{eff} = T+1/2m\beta^2$$

#### Radial flow

#### Spectral shape depends on PID mass Higher mass => larger inverse slope More central => larger inverse slope

	Central	Mid-central	Peripheral
data			
π, K, p spectra [79]	0-5%	15-30%	60-92%
$\Lambda$ spectra [80]	0-5%	20  35%	35 - 75%
pion radii [67]	0-12%	12-32%	32-72%
Elliptic flow [38]	0-11%	11-45%	45-85%
$\chi^2/(\# \text{ data points})$			
$\pi^+ \& \pi^- \text{ spectra}$	7.2/10	26.5/10	13.0/9
$K^+$ & $K^-$ spectra	24.2/22	21.4/22	10.1/10
$p \& \overline{p}$ spectra	10.6/18	23.2/18	28.0/12
$\Lambda \& \overline{\Lambda} \text{ spectra}$	9.5/16	12.8/16	11.0/16
$\pi v_2$	14.6/12	29.8/12	5.2/12
$p v_2$	1.6/3	9.2/6	0.8/3
$\pi r_{out}$	1.9/6	0.4/2	0.4/2
$\pi r_{side}$	2.7/6	0.07/2	0.06/2
$\pi r_{long}$	5.3/6	0.003/2	0.1/2
Total	77.6/99	107.7/90	68.7/68
parameters			
T (MeV)	$106 \pm 3$	$107 \pm 2$	$100 \pm 5$
$\rho_0$	$0.89 \pm 0.02$	$0.85\pm0.01$	$0.79\pm0.02$
$\langle \beta_T \rangle$	$0.52\pm0.01$	$0.50\pm0.01$	$0.47\pm0.01$
$\rho_2$	$0.060\pm0.008$	$0.058\pm0.005$	$0.05\pm0.01$
$R_x(fm)$	$13.2\pm0.3$	$10.4\pm0.4$	$8.00\pm0.4$
$R_y(fm)$	$13.0\pm0.3$	$11.8\pm0.4$	$10.1\pm0.4$
$\tau(fm/c)$	$9.2\pm0.4$	$7.7\pm0.9$	$6.5\pm0.6$
$\Delta t(fm/c)$	$0.003 \pm 1.3$	$0.06 \pm 1.3$	$0.6 \pm 1.8$

TABLE II: Upper section: data used in the fit. Middle section: number of  $\chi^2/$  data points for each measure. Lower section: best fit parameters. Note that  $\langle \beta_T \rangle$  is not a fit parameter, but it is calculated from  $\rho_0$ .

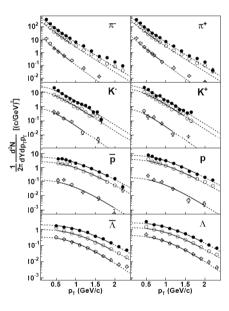
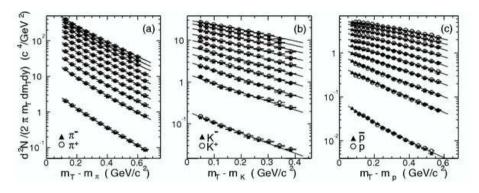
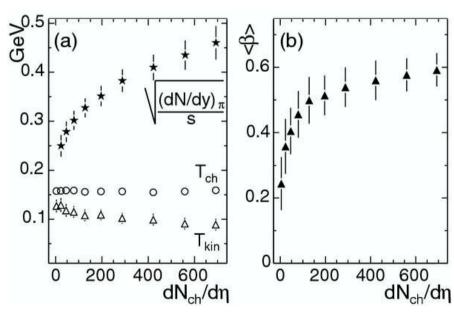


FIG. 49: Comparison of the data with the blast-wave calculations performed with the best fit parameters in three centrality bins. The closed circles are central data, the open circles are mid-central data and the crosses are peripheral data. The plain lines show the blast wave calculation within the fit range while the dash lines show the extrapolation over the whole range.



#### STAR PRL92



#### F. Retiere and M. Lisa PRC70; PHENIX PRL88

#### **Blast Wave**

Because of azimuthal symmetry we can integrate over  $\phi$  making use of the modified Bessel function  $I_0(z) = (2\pi)^{-1} \int_0^{2\pi} e^{z \cos \phi} d\phi$ :

$$E\frac{d^{3}n}{d^{3}p} = \frac{g}{(2\pi)^{2}} \int_{-\mathcal{Z}}^{\mathcal{Z}} d\zeta \left[ m_{T} \cosh y \frac{\partial z}{\partial \zeta} - m_{T} \sinh y \frac{\partial t}{\partial \zeta} \right]$$

$$\times \int_{0}^{R} r dr \exp \left( -\frac{m_{T} \cosh \rho \cosh(y - \eta) - \mu}{T} \right) I_{0} \left( \frac{p_{T} \sinh \rho}{T} \right)$$
(14)

For the transverse mass spectrum we integrate with the help of another modified Bessel function  $K_1(z) = \int_0^\infty \cosh y \, e^{-z \cosh y} \, dy$ :

$$\frac{dn}{m_T dm_T} = \frac{g}{\pi} m_T \int_{-\mathcal{Z}}^{\mathcal{Z}} d\zeta \left[ \cosh \eta \frac{\partial z}{\partial \zeta} - \sinh \eta \frac{\partial t}{\partial \zeta} \right] \int_0^R r dr \, K_1 \left( \frac{m_T \cosh \rho}{T} \right) I_0 \left( \frac{p_T \sinh \rho}{T} \right) 
= \frac{2g}{\pi} m_T Z_{\tilde{t}} \int_0^R r dr \, K_1 \left( \frac{m_T \cosh \rho}{T} \right) I_0 \left( \frac{p_T \sinh \rho}{T} \right) \tag{15}$$

E. Schnedermann, J. Sollfrank, U. Heinz, nucl-th/9307020, PRC48 (cited 312)

#### **Assumptions:**

- 1) Local thermal equilibrium → Boltzmann distribution
- 2) Longitudinal and transverse expansions (1+2)
- 3) Radial flow profile  $\rho(r) \propto A \tanh(\beta_m (r/R)^n)$ , (n=1)
- 4) Temperature and  $\langle \beta \rangle$  are global quantities

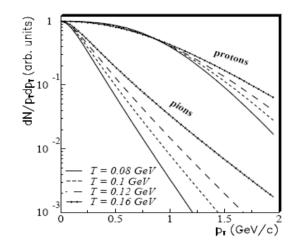


FIG. 4: Transverse momentum spectra for protons (upper curves) and pions (lower curves), as calculated by Equation  $\boxed{19}$  for several values of the temperature parameter T. Other parameters follow the "round" source defaults of Table  $\boxed{1}$  All spectra are arbitrarily normalized to unity at  $p_T=0$ .

#### F. Retiere, M. Lisa, PRC70

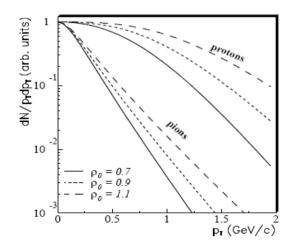


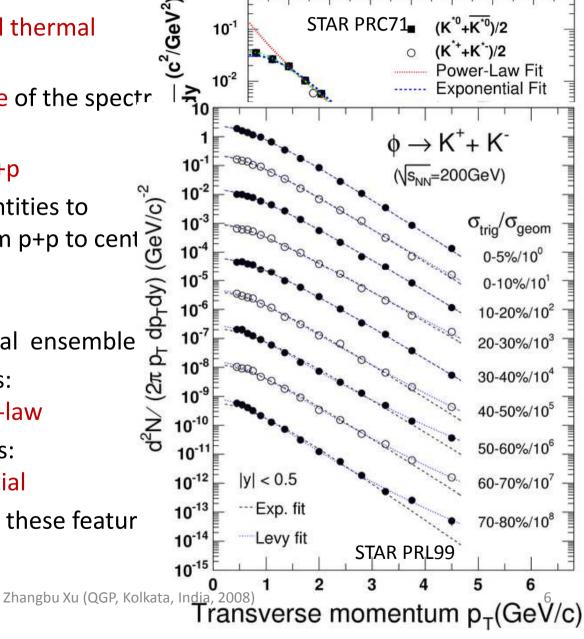
FIG. 5: Transverse momentum spectra for protons (upper curves) and pions (lower curves), as calculated by Equation [9], for several values of the radial flow parameter  $\rho_0$ . Other parameters follow the "round" source defaults of Table []. All spectra are arbitrarily normalized to unity at  $\rho_T = 0$ .

#### Limitations of THE BlastWave

Strong assumption on local thermal equilibrium

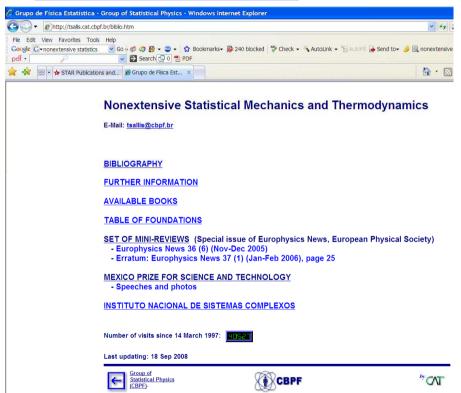
 Arbitrary choice of p<sub>T</sub> range of the spectre (low and high cuts)

- Flow velocity  $<\beta>=0.2$  in p+p
- Lack of non-extensive quantities to describe the evolution from p+p to cent A+A collisions
- example in chemical fits: canonical to grand canonical ensemble
- m<sub>T</sub> spectra in p+p collisions:
   Levy function or m<sub>T</sub> power-law
- m<sub>T</sub> spectra in A+A collisions:
   Boltzmann or m<sub>T</sub> exponential
- What function can capture these featur



#### **Tsallis Statistics**

- Nice web based notebooks: Tsallis Statistics, Statistical Mechar for Non-extensive Systems and Long-Range Interactions http://www.cscs.umich.edu/~crshalizi/notabene/tsallis.html
- http://tsallis.cat.cbpf.br/biblio.htm



Negative Binomial Distribution:  $\kappa=1/(q-1)$ 

Temperature fluctuation:  $\frac{\left\langle 1/T^2\right\rangle - \left\langle 1/T\right\rangle^2}{\left\langle 1/T\right\rangle^2} = 1 - q$ 

G. Wilk: arXiv: 0810.2939

ENTROPY AND STATISTICAL MECHANICS: FOUNDATIONS				
	BG (thermal equilibrium)	q ≠ 1 (thermal metaequilibrium, nonequilibrium)		
Distribution of velocities at equilibrium	Maxwell 1860	R. Silva, A.R. Plastino, J.A.S. Lima Phys Lett A 249, 401 (1998)		
		R.S. Mendes and C. Tsallis Phys Lett A 285, 273 (2001)		
Kinetic equation Molecular chaos hypothesis (Stosszahlansatz)	Boltzmann 1872	J.A.S. Lima, R. Silva, A.R. Plastino Phys Rev Lett 86, 2938 (2001)		
		G. Kaniadakis Physica A 296, 405 (2001)		
Optimization of entropy with	Gibbs 1902	C.Tsallis J Stat Phys 52, 479 (1988)		
constraints		E.M.F. Curado and C.Tsallis J Phys A 24, L69 (1991)		
		C.Tsallis, R.S. Mendes, A.R. Plastino Physica A 261, 534 (1998)		
Steepest descent	Darwin-Fowler 1922	S. Abe and A.K. Rajagopal J Phys A 33, 8733 (2000)		
Microcanonical ensemble counting	Balian-Balazs 1987 Kubo et al 1988	S. Abe and A.K. Rajagopal Phys Lett A 272, 341 (2000) Europhys Lett 55, 6 (2001)		
Conditions of uniqueness of S	Shannon 1948	R. J. V. Santos J Math Phys 38, 4104 (1997)		
Compact conditions of uniqueness of S	Khinchin 1953	S. Abe Phys Lett A 271, 74 (2000)		
Limit theorems	De Moivre 1733 Laplace 1744 Gauss 1809 Lévy 1937	S. Umarov, C. Tsallis and S. Steinberg cond-mat/0603593 S. Umarov, C. Tsallis, M. Gell-Mann and S. Steinberg cond-mat/0606038 and /0606040		
ndia 2009)	Khinchin 1949	S. Abe and A.K. Rajagopal Europhys Lett 52, 610 (2000)		

# It is all about the q-statistics

# $\begin{aligned} & \text{BASIC QUANTITIES} \\ & q\text{-exponential}: \exp_q(x) \equiv [1 + (1-q)\,x]^{\frac{1}{1-q}} \longrightarrow_{q \to 1} e^x \\ & q\text{-logarithm}: \ln_q(x) \equiv \frac{x^{1-q}-1}{1-q} \longrightarrow_{q \to 1} \ln x \\ & \text{Boltzmann-Gibbs entropy}: S_{BG} \equiv -k \sum_{i=1}^W p_i \ln p_i \\ & q\text{-entropy}: S_q \equiv k \frac{1-\sum_{i=1}^W p_i^q}{q-1} = k \sum_{i=1}^W p_i \ln_q(1/p_i) = -k \sum_{i=1}^W p_i^q \ln_q p_i \longrightarrow_{q \to 1} S_{BG} \\ & \text{Escort distribution}: P_i \equiv p_i^q / \sum_{j=1}^W p_j^q \\ & \text{Ensemble } q\text{-average}: \langle A \rangle_q \equiv \sum_{i=1}^W A_i P_i = \sum_{i=1}^W A_i p_i^q / \sum_{j=1}^W p_j^q \end{aligned}$

■ Box: The two basic functions that appear in Nonextensive Statistical Mechanics are the q-exponential and the q- logarithm with  $\ln_q(\exp_q x) = \exp_q(\ln_q x) = x$ . They are simple generalizations of the usual exponential and logarithmic functions which are retrieved by performing a |1-q| << 1 expansion. Similarly the q-entropy generalizes the standard Boltzmann-Gibbs entropy. The escort distribution is a generalization of the usual ensemble averaging function to which it reduces for q = 1.

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- Why is this relevant to us (Heavy-ion physics)?
  - We have dealt with Boltzmann distribution
     But the spectra are clearly non-Boltzmann
  - It is easy to make a change
  - It is easy to compare
  - Change m<sub>T</sub> exponential to m<sub>T</sub> power law

#### Tsallis statistics in Blast Wave model

Because of azimuthal symmetry we can integrate over  $\phi$  making use of the modified Bessel function  $I_0(z) = (2\pi)^{-1} \int_0^{2\pi} e^{z \cos \phi} d\phi$ :

$$E\frac{d^{3}n}{d^{3}p} = \frac{g}{(2\pi)^{2}} \int_{-\mathcal{Z}}^{\mathcal{Z}} d\zeta \left[ m_{T} \cosh y \frac{\partial z}{\partial \zeta} - m_{T} \sinh y \frac{\partial t}{\partial \zeta} \right]$$

$$\times \int_{0}^{R} r dr \exp \left( -\frac{m_{T} \cosh \rho \cosh(y - \eta) - \mu}{T} \right) I_{0} \left( \frac{p_{T} \sinh \rho}{T} \right)$$
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For the transverse mass spectrum we integrate with the help of another modified Bessel function  $K_1(z) = \int_0^\infty \cosh y \, e^{-z \cosh y} dy$ :

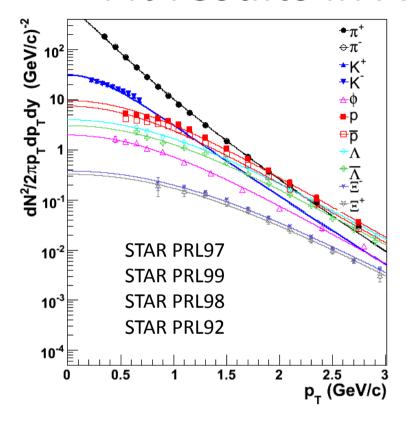
$$\frac{dn}{m_T dm_T} = \frac{g}{\pi} m_T \int_{-\mathcal{Z}}^{\mathcal{Z}} \left[ \cosh \eta \frac{\partial z}{\partial \zeta} - \sinh \eta \frac{\partial t}{\partial \zeta} \right] \int_0^R r dr \, K_1 \left( \frac{m_T \cosh \rho}{T} \right) I_0 \left( \frac{p_T \sinh \rho}{T} \right) 
= \frac{2g}{\pi} m_T Z_{\tilde{t}} \int_0^R r dr \, K_1 \left( \frac{m_T \cosh \rho}{T} \right) I_0 \left( \frac{p_T \sinh \rho}{T} \right) \tag{15}$$

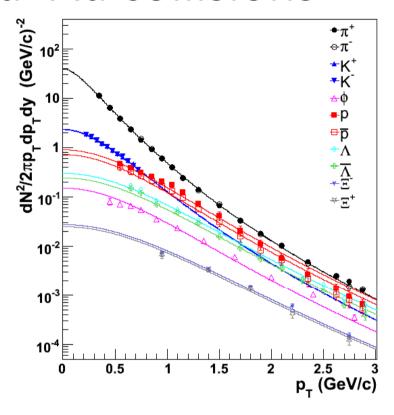
#### With Tsallis distribution, the BlastWave equation is:

$$\frac{dN}{m_T dm_T} \propto m_T \int_{-Y}^{+Y} \cosh(y) dy \int_{-\pi}^{+\pi} d\phi \int_{0}^{R} r dr (1 + \frac{q-1}{T} (m_T \cosh(y) \cosh(\rho) - p_T \sinh(\rho) \cos(\phi)))^{-1/(q-1)}$$

Where  $\rho$ =Atanh( $\beta_m(r/R)^n$ ), n=1; any of the three integrals is HypergeometryF1

#### Fit results in Au+Au collisions



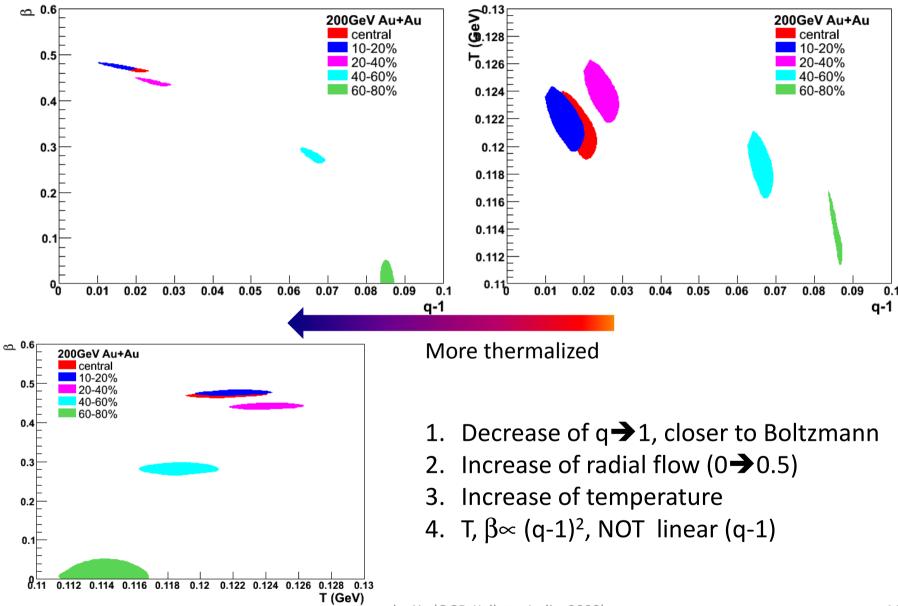


Au+Au 0—10%:  $<\beta>=0.470+-0.009$  T=0.122+-0.002 q=1.018+-0.005 $chi^2/nDof=130/125$ 

Au+Au 60—80%:  

$$<\beta>=0$$
  
T = 0.114 +- 0.003  
q = 1.086 +- 0.002  
chi^2/nDof = 138/123

#### Dissipative energy into flow and heat



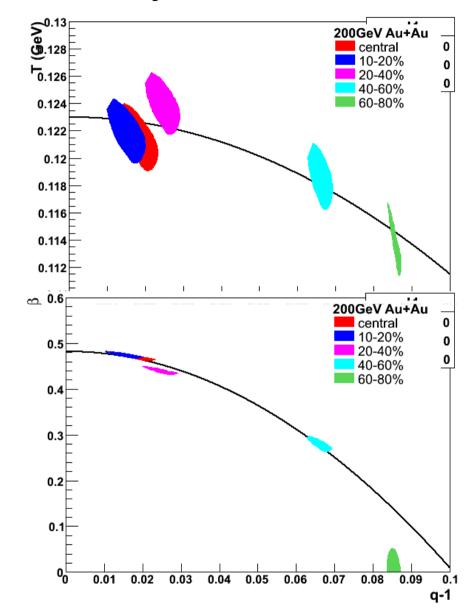
#### Related to bulk viscosity ( $\xi$ )

$$T_{eff} = T_0 + \frac{\xi}{a} f(\beta)$$

$$= T_0 + (q-1) \frac{(\xi/\rho)(c_p \rho/a)}{(c_p/c_V)} f(\beta)$$

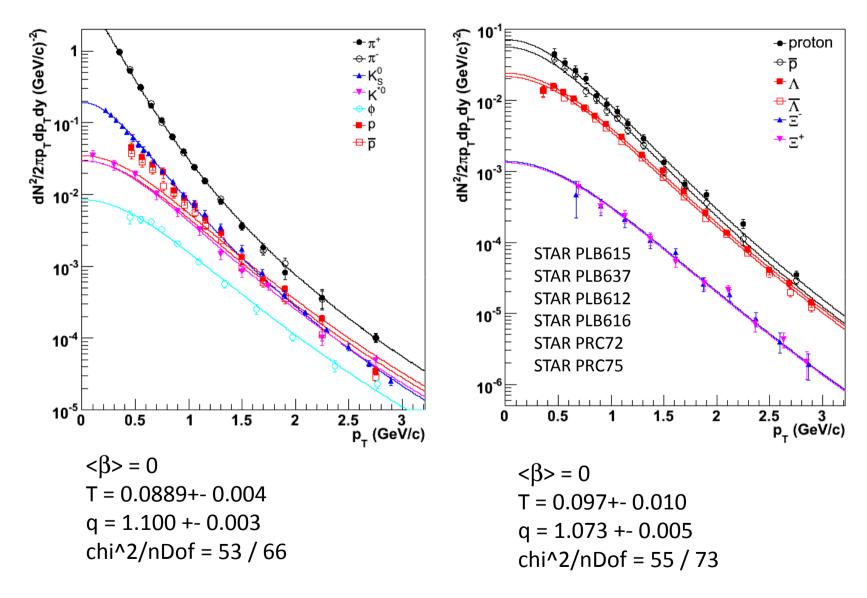
$$= T_0 + (q-1)^2 \frac{(\xi/\rho)}{(c_p/c_V)D} f(\beta)$$

c<sub>p</sub>, ρ and a are, respectively,
 the specific heat under
 constant pressure,
 density and
 the coefficient of external conductance

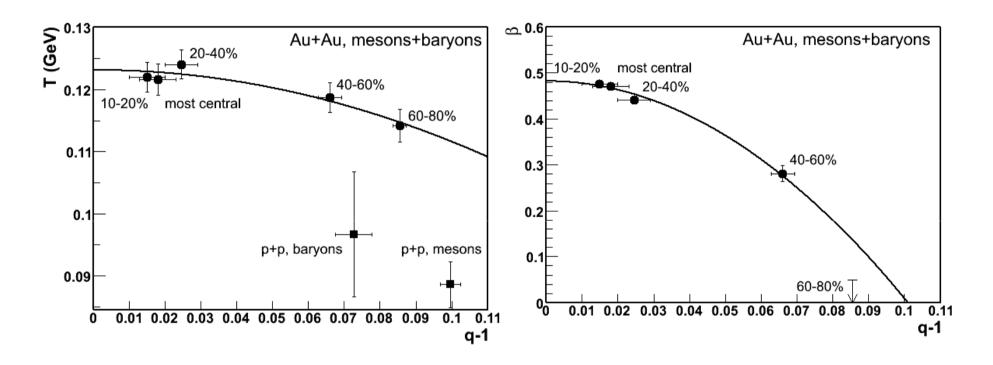


G. Wilk: arXiv: 0810.2939

# Results in p+p collisioins

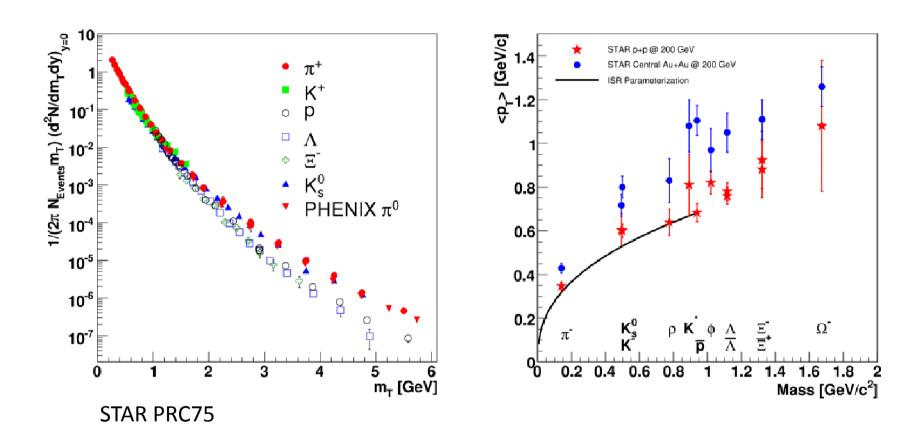


# Evolution from p+p to Au+Au



- •Sharp increase of <T> from p+p to peripheral Au+Au
- •Similar q from p+p to peripheral Au+Au
- •Radial flow is zero at p+p and peripheral Au+Au

### Baryon and meson are different classes



In p+p collisions, the  $m_T$  spectra of baryons and mesons are in two groups However, equilibrated toward more central Au+Au collisions

## Observations from the q-statistics

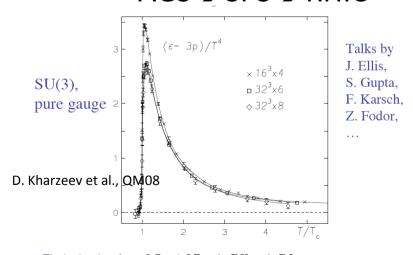
- Fit spectra well for all particles with p<sub>T</sub><~ 3 GeV/c</li>
- Radial flow increases from 0 to 0.5c
- Kinetical freeze-out temperature increases from 90 (110) to 130 MeV
- q-1 decreases from 0.1 to 0.01
- T and  $\beta$  depend on  $(q-1)^2$
- p+p collisions are very different, split between mesons and baryons

- Tsallis statistics describes the data better than Boltzmann-Gibbs statistics
- Radial flow is zero in p+p and peripheral Au+Au collisions
- Evolution from peripheral to central Au+Au collisions: hot spots (temperature fluctuation) are quenched toward a more uniform Boltzmann-like distribution
- dissipative energy into heat and flow, related to bulk viscosity
- Energy conservation is a built-in requirement in any statistical model (that is where you get the temperature)

#### Outlook

#### Search for critical point:

- large bulk viscosity at phase transition
- PID spectra to 3 GeV/c
- Study T, β vs q-1 with centrality and energy
   AGS→SPS→RHIC



#### • Higher energy at LHC:

- Large power-law tail due to semi-hard processes
- Without Tsallis
   distribution, it is likely
   impossible to extract
   radial flow from spectra
- Good (large) nonextensive effect and easy to extract bulk viscosity

# Application of Tsallis statistics has a long history at RHIC

1) Non-extensive thermodynamics, heavy ion collisions and particle production at RHIC energies. Bhaskar De (Maulana Azad Coll.), S. Bhattacharyya (Indian Statistical Inst., Calcutta), Goutam Sau (Unlisted), S. Published in Int.J.Mod.Phys.E16:1687-1700,2007.

<u>LaTeX(US) | LaTeX(EU) | Harvmac | BibTeX | Keywords | Cited 1 time Journal Server</u>
Bookmarkable link to this information

2) Nonextensive hydrodynamics for relativistic heavy-ion collisions.

Bookmarkable link to this information

T. Osada (Musashi Inst. Tech.), G. Wilk (Warsaw, Inst. Nucl. Studies). Feb 2008. 23pp. Published in Phys.Rev.C77:044903,2008. e-Print: arXiv:0710.1905 [nucl-th]

References | LaTeX(US) | LaTeX(EU) | Harvmac | BibTeX | Cited <u>6 times</u>
Abstract and Postscript and PDF from arXiv.org (mirrors: <u>au br cn de es fr il in it jp kr ru tw uk za aps lanl</u>)
Journal Server

3) Signals of non-extensive statistical mechanics in high-energy nuclear collisions.

W.M. Alberico (Turin U. & INFN, Turin), P. Czerski (Cracow, INP), A. Lavagno (Turin Polytechnic & INFN, Turin), N. CERN & INFN, Turin), V. Soma (Turin U.). Oct 2005. 13pp. e-Print: hep-ph/0510271

References | LaTeX(US) | LaTeX(EU) | Harvmac | BibTeX | Keywords | Cited 4 times | Abstract and Postscript and PDF from arXiv.org (mirrors: au br cn de es fr il in it jp kr ru tw uk za aps lanl ) Bookmarkable link to this information

4) A Nonextensive model for quark matter produced in heavy ion collisions.

Tamas S. Biro, Gabor Purcsel (Budapest, RMKI) . Mar 2004. 16pp. e-Print: hep-ph/0403038

References | LaTeX(US) | LaTeX(EU) | Harvmac | BibTeX | Keywords

Abstract and Postscript and PDF from arXiv.org (mirrors: <u>au br cn de es fr il in it jp kr ru tw uk za aps lanl</u>)

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#### m<sub>T</sub>-m<sub>0</sub> power-law

- STAR PRD74 (2006)
- STAR PRC71 (2005)
- STAR PRL99 (2007)
- Energy conservation

Z. Chajecki and M. Lisa arXiv:0807.3569

Soft+Minijets

T. Trainor, arXiv:0710.4504