

Charge-dependent directed flow in Cu+Au collisions

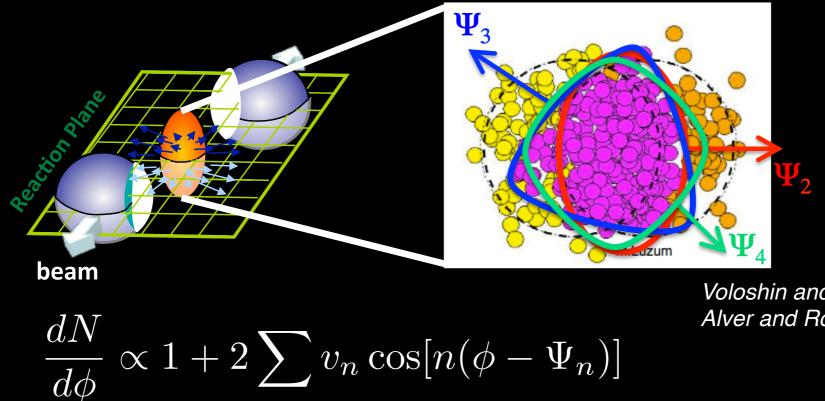
Takafumi Niida for the STAR Collaboration Wayne State University



Initial Stages 2016 @Instituto Superior Técnico, Lisbon

STAR Azimuthal anisotropy

Anisotropies in momentum-space originate from anisotropies in initial geometry (including fluctuations)

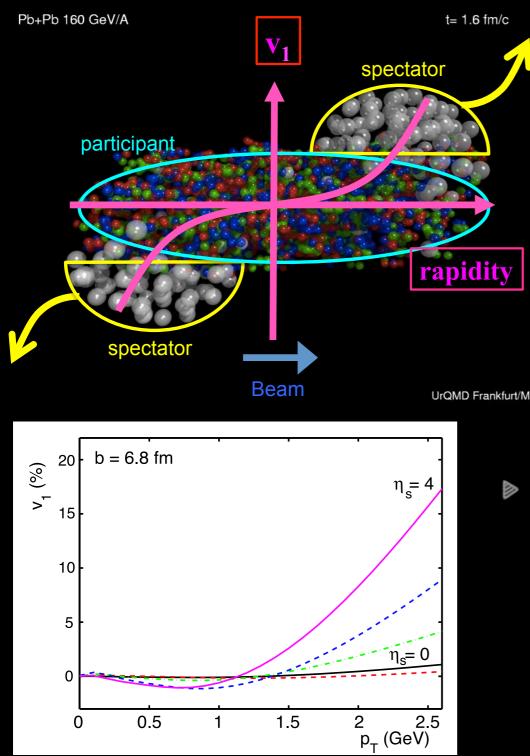


Voloshin and Zhang, Z.Phys.C70, 665 Alver and Roland, PRC81, 054905

Directed flow (v₁): sensitive to EoS and phase transition Elliptic(v₂), Triangular(v₃), \cdots : sensitive to η /s and initial fluctuations

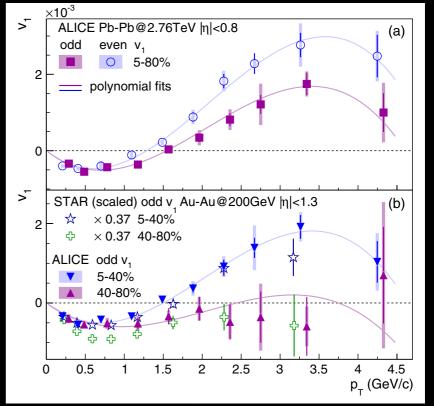
> Csernai and Rohrich, PLB458, 454 (1999) Gale et al., PRL110, 012302 (2013)

STAR Directed flow in A+A



U. Heinz and P. Kolb, J.Phys.G30 (2004) S1229

v₁ in Au+Au vs Pb+Pb ALICE, PRL111.23202

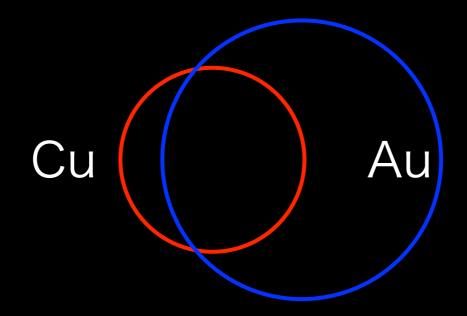


v₁ in A+A collisions

v1 is caused by the initial density asymmetry
 v1 at η=0 is zero due to symmetric density
 non zero v1(pT) comes from the density fluctuation
 Note: <px>=0 if no kick from spectators

How about in asymmetric collisions?

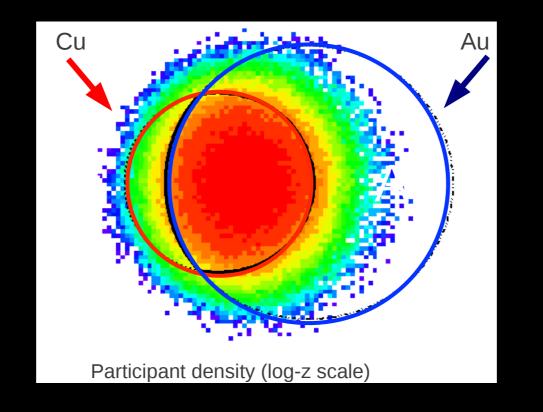




Intrinsic asymmetric density

- Iarger directed flow compared to A+A collisions?
- Sizable initial electric field
 - pointing from Au to Cu, due to the charge difference (# of protons) in both spectators

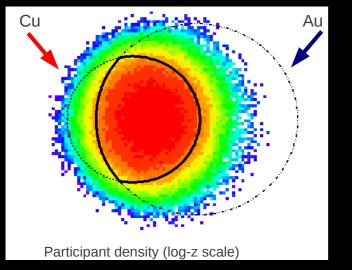
STAR Cu+Au collisions

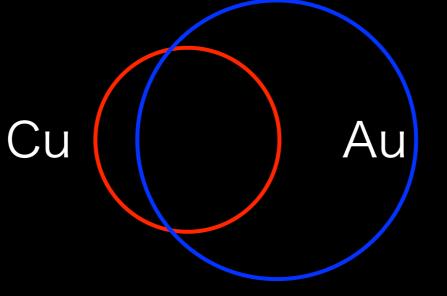


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plot from A. lordanova, @RHIC&AGS users meeting 2013



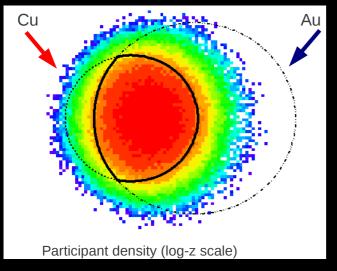


Asymmetric density profile Asymmetric pressure gradient

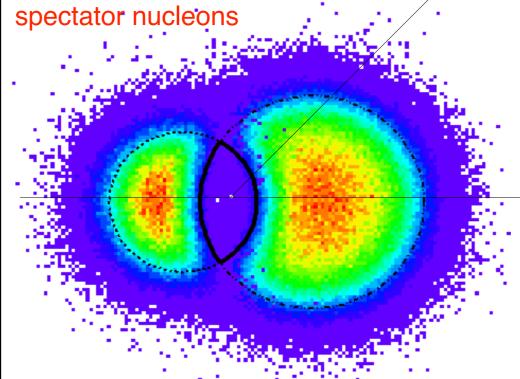
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STAR Cu+Au collisions

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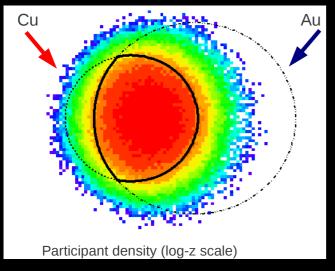
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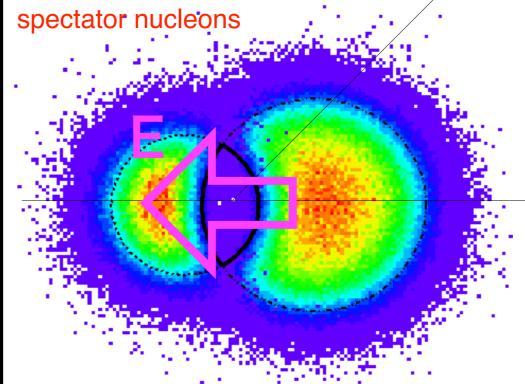
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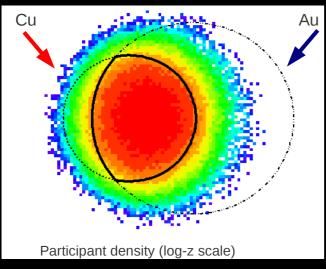
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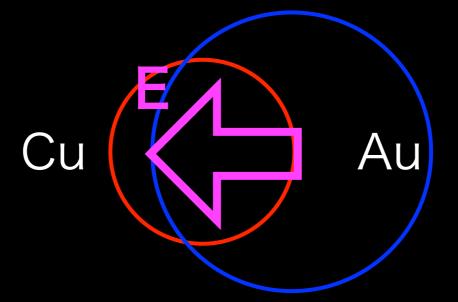


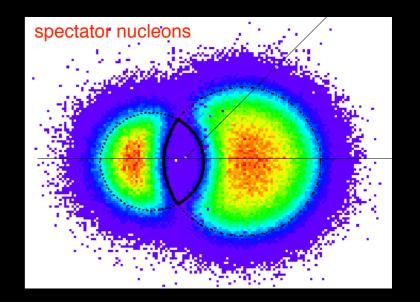
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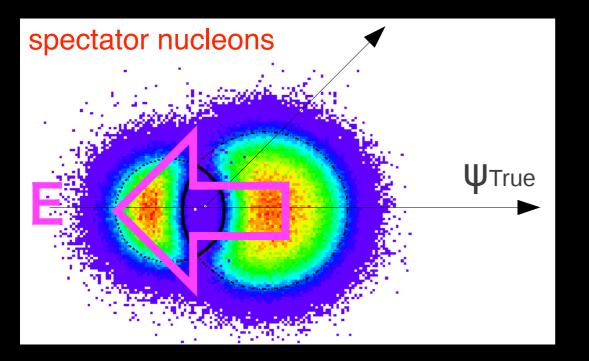


Asymmetric density profile Asymmetric pressure gradient

Dipole-like charge distribution by spectators

- Intrinsic asymmetric density
 - Iarger directed flow compared to A+A collisions?
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STAR Effect of the electric field



If we have the electric field, azimuthal distribution of particles can be written:

$$\frac{dN^{\pm}}{d\phi} \propto 1 + 2v_1 \cos(\phi - \Psi_1) \pm d_E \cos(\phi - \psi_E)$$

 d_E : strength of dipole deformation induced by E-field (proportional to the electric conductivity) ψ_E : azimuthal angle of E-field

Positive particles move to the direction along E-field, and negative particles go to the opposite, which appears as charge dependence of v1

• Y. Hirono et al., Phys. Rev. C90, 021903 (2014), sensitive to the electric conductivity

Note: This idea was first reported at IS2013 conference by Y. Hirono

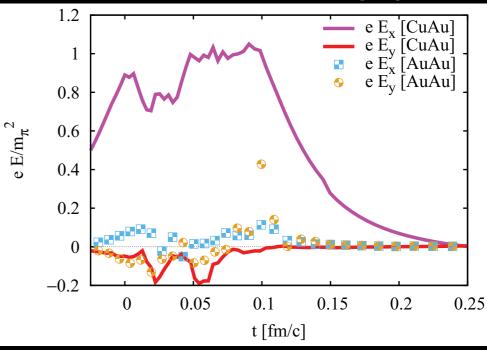
STAR Probe into quark creation time?

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Life time of E-field would be very short

- No signal if there are no quarks (charges) when E-field is strong
- In other words, sensitive to the number of quark & anti-quark at very early stage (V. Voronyuk et al., PRC90.064903)



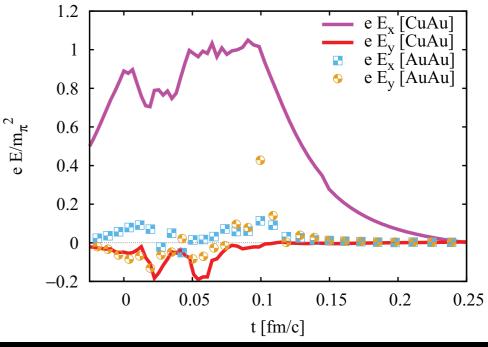


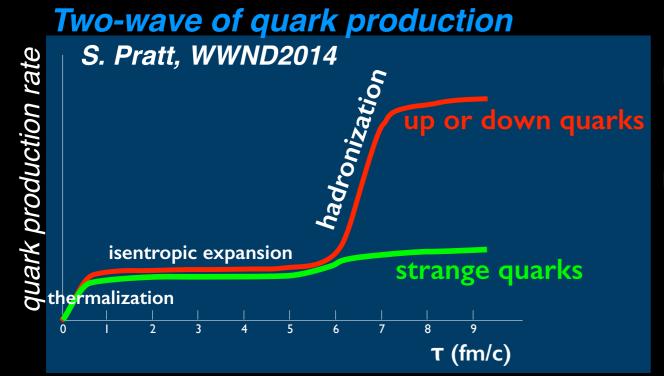
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- Sensitive to the time evolution of quark production
- Also important input for theoretical prediction of CME/CMW

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Solenoidal Tracker At RHIC (STAR)

ER

EEMC

TPC

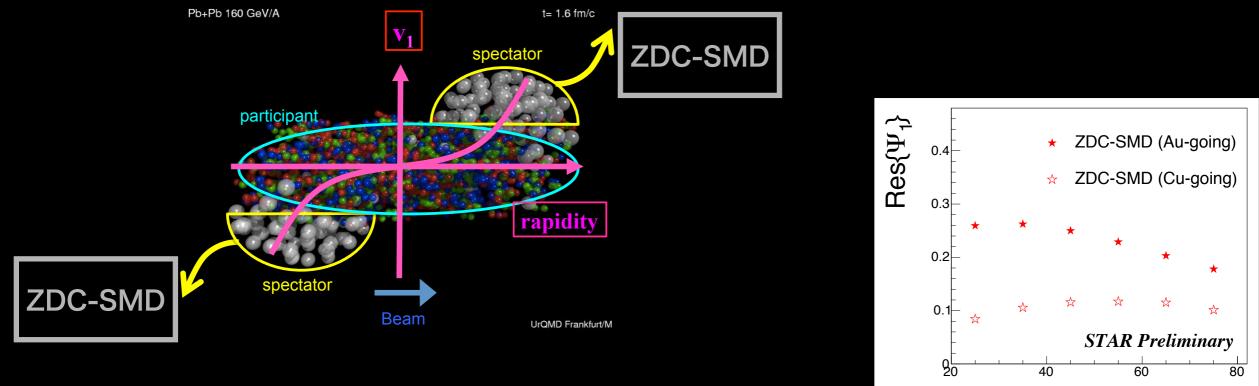
TOF

VPD

BBC

Trigger detectors: VPD, ZDC (detecting spectator neutrons)
 Tracking of charged particles: TPC (|n|<1)
 Event planes: ZDC-SMD

STAR Directed flow measurement



Centrality [%]

$$v_1 = \langle \cos(\phi - \Psi_1) \rangle / \operatorname{Res} \Psi_1$$

Ψ₁ determined by Zero Degree Calorimeter (ZDC) and Shower Max Detector (SMD)

n measure the energy and position of spectator neutrons

Spectator deflects "outward" from the center of collisions (not "inward")
 S. Voloshin and TN, arXiv:1604.04597

provides the direction of E-field

STAR Charge-dependent directed flow

Ψ_1 {Au-spectator} Cu-going direction: $\eta > 0$ 50-60% 20-30% 30-40% 40-50% STAR Preliminary Cu+Au 200 GeV syst. uncert. Au positive ZDC-SMD Au-going |n| < 1> negative |ml<1, Ψ ★ positive syst. uncert. from EP 🛧 negative 0.02 E-Field CU 30 30 2 30 p_ [GeV/c] p_ [GeV/c] p_{_} [GeV/c] p_[GeV/c]

$|v_1 = \langle \cos(\phi - \Psi_1) \rangle$

Sizable v₁ measured relative to Ψ_1 {ZDC-SMD} in Au-going side (Ψ_1^{Au} <0)

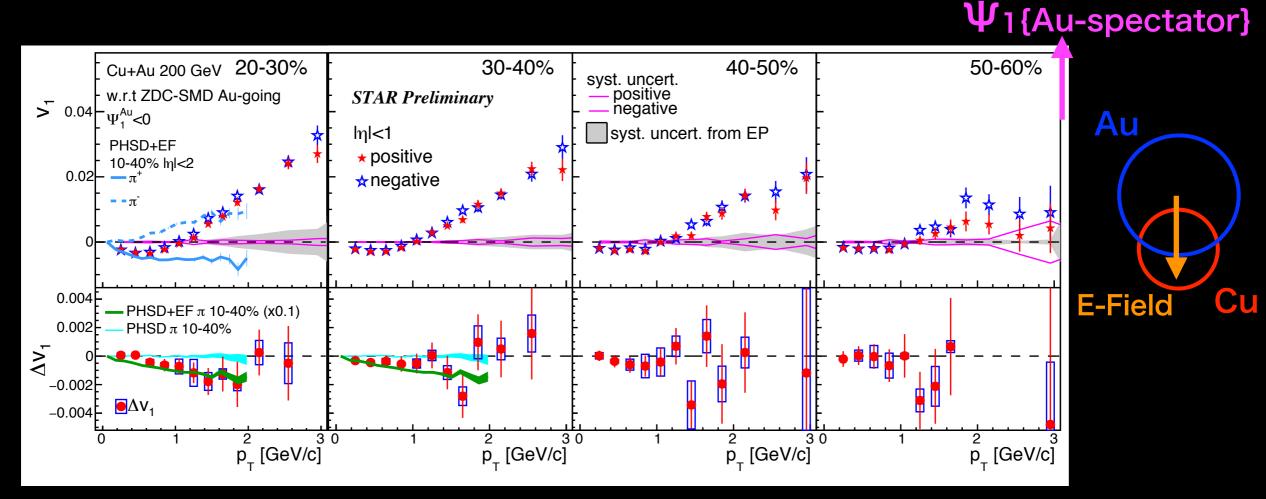
- v1 becomes smaller in more peripheral collisions
- Sign change of v1 around pT=1GeV/c to balance the momentum (more low pT particles in Cu-side, more high pT particles in Au-side)

▶ Larger v₁ compared to A+A collisions

Note: v_1^{even} in A+A is only due to density fluctuations

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STAR Charge-dependent directed flow



$\blacktriangleright \Delta v_1 = v_1(h^+) - v_1(h^-)$, and $v_1 \sim 1\%$, $\Delta v_1 < 0.2\%$

- Δv_1 looks to be negative in p_T<2 GeV/c,
- similar p_T dependence to PHSD model with the electric field (PHSD+EF) (PRC90.064903), but smaller by a factor of 10

Finite but small Δv_1 indicates:

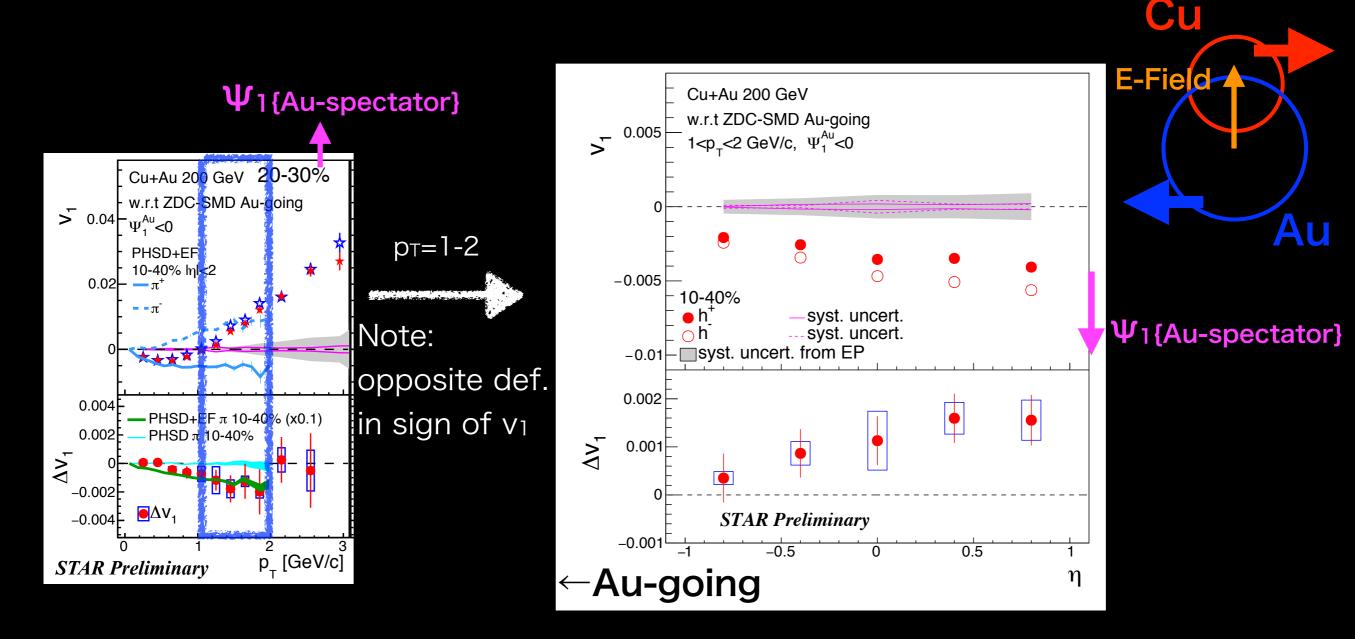
existence of E-field

very small number of quarks at times earlier than the E-field life time(~0.25 fm/c)

 \square PHSD assumes all partons are present at t~0 and affected by the E-field

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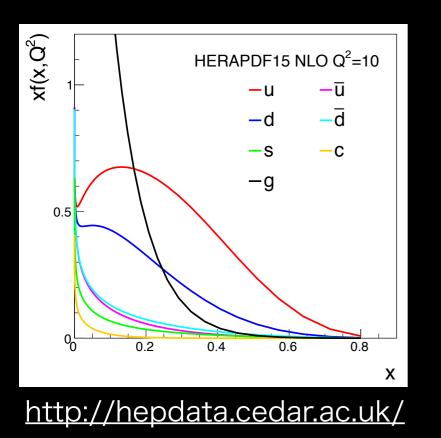
star η dependence of v_1

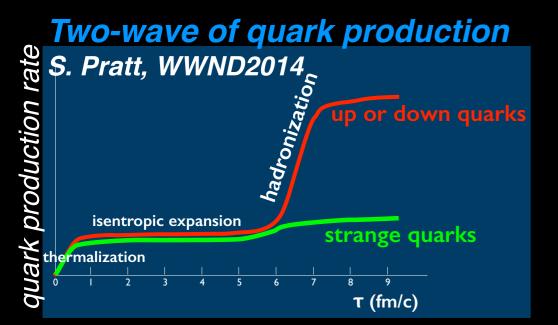


▶ Charge-difference can be seen in $-1 < \eta < 1$ and $1 < p_T < 2$ GeV/c

- Difference appears to be larger in Cu-going direction
- Opposite trend to the PHSD model

STAR How many quarks at initial state?





Rough estimate from PDF

- Quark density in PDF \rightarrow Quarks at initial state
- Quarks + Gluons in PDF \rightarrow All quarks created
 - Assuming gluons are converted to 2 quarks at final state

$$x \sim \frac{p_T}{\sqrt{s}} e^{\eta}$$

- Ø 0.2-4</sup> < x < 0.01
 </p>
- Initial quarks/All quarks created ~15%, which is close to 10% obtained from Δv_1 +PHSD model

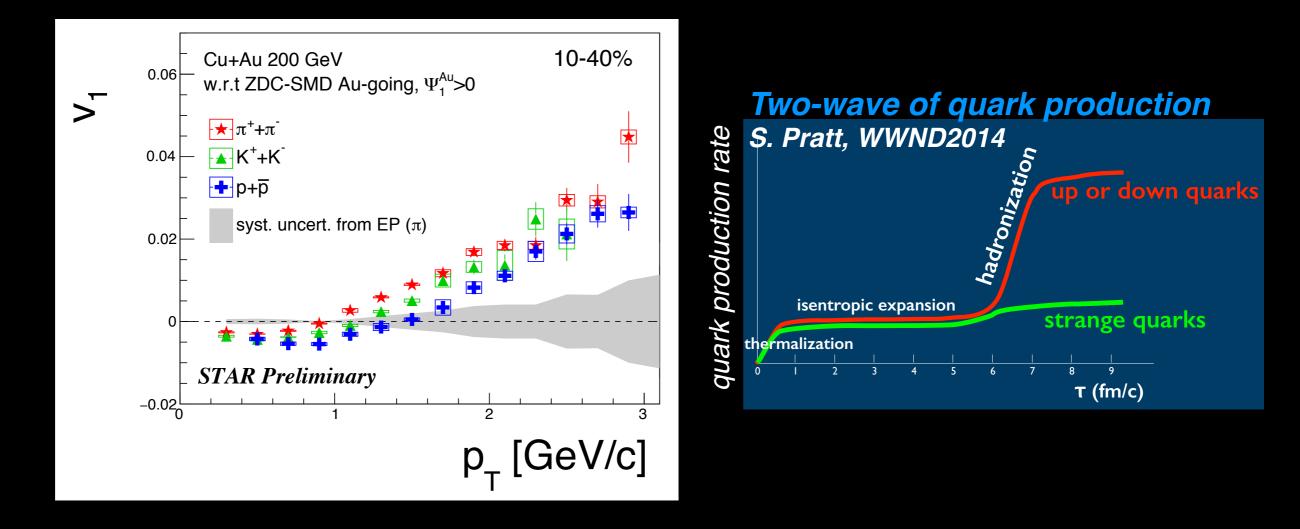
<u>Small fraction of initial quarks to all quarks</u> produced in the collision!

Supporting "two-wave scenario"?

Two waves of light quark production, where small fraction of quarks are created at early time

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STAR Identified Particle v₁



▶ Mass ordering at low p⊤

Can be explained by the radial flow (S. Voloshin, PRC55.R1630(1997))

- Would be interesting to look at charge-dependent kaons
 - To test the two-wave scenario, where s and u quark productions would be expected to be different

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Charge-dependent directed flow in Cu+Au collisions has been measured at the STAR experiment

- Charge dependence of v₁, consistent with an existence of the initial electric field, has been observed
- The magnitude of the difference, Δv₁, is much smaller than the PHSD model prediction, likely indicating that the number of initial (anti-)quarks are very small when the E-Field is strong (t<0.25 fm/c)
- Simple estimate based on the parton distribution functions is consistent with the above interpretation

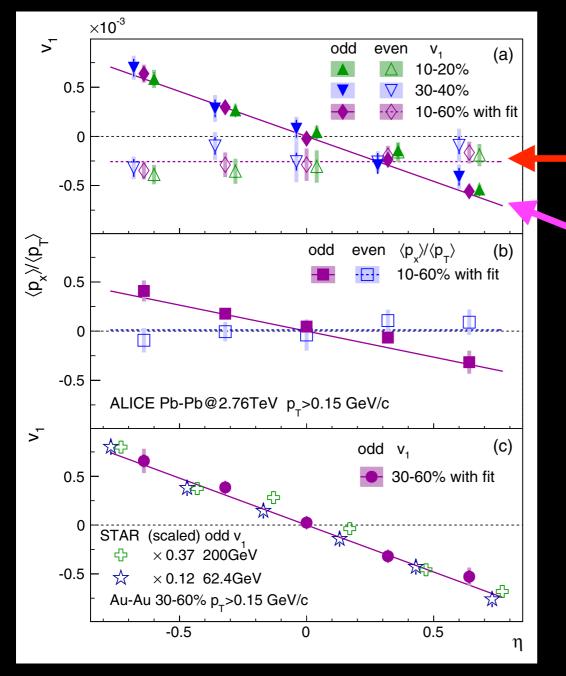
Thank you for your attention!





v_1^{even} and v_1^{odd} in Pb+Pb 2.76 TeV

v1 in Au+Au vs Pb+Pb ALICE, PRL111.23202



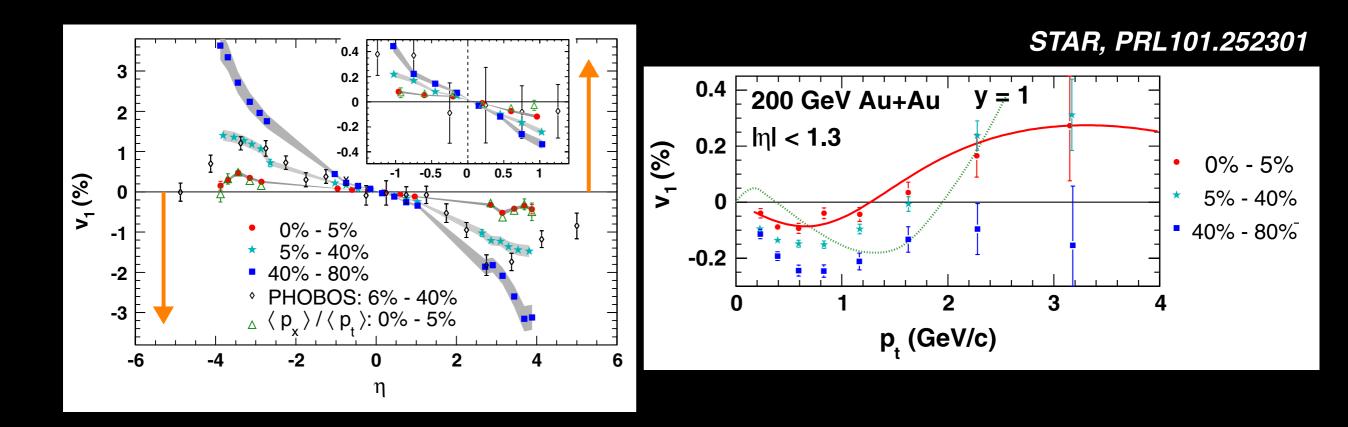
even component independent of η

odd component

$$v_1^{\text{odd}} \{ \Psi_{\text{SP}} \} = [v_1 \{ \Psi_{\text{SP}}^p \} + v_1 \{ \Psi_{\text{SP}}^t \}]/2$$

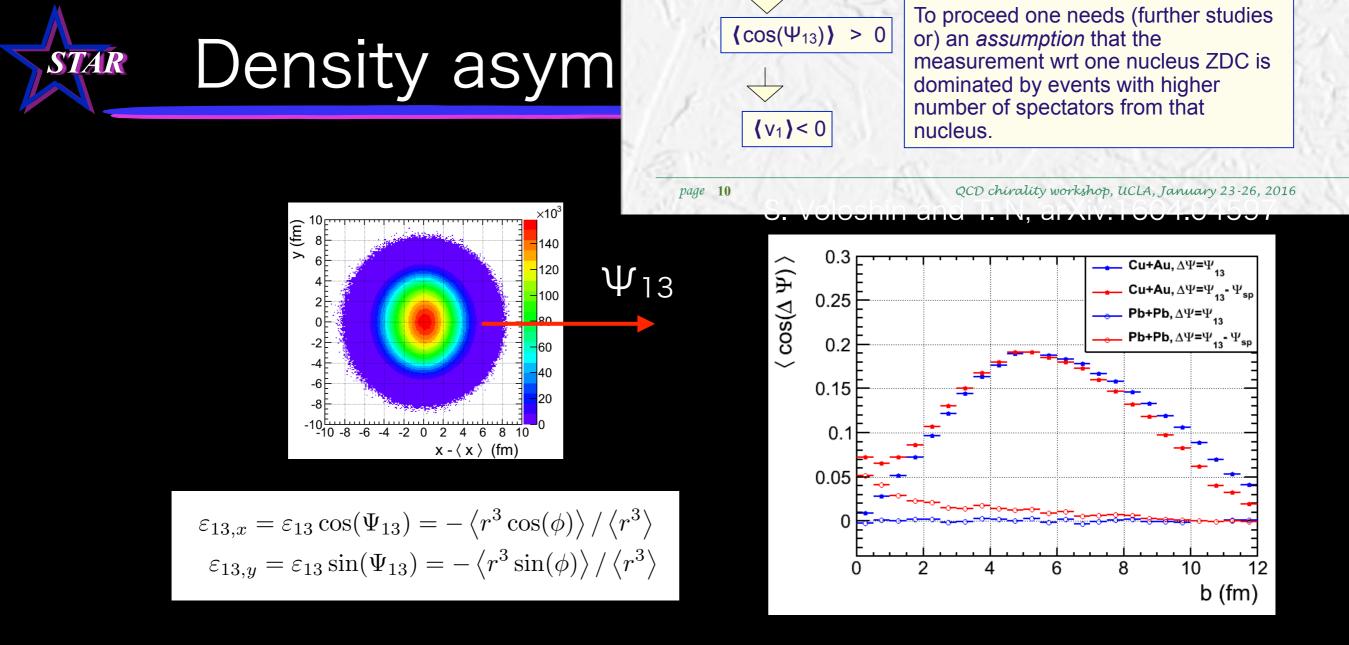
$$v_1^{\text{even}}{\{\Psi_{\text{SP}}\}} = [v_1 \{\Psi_{\text{SP}}^p\} - v_1 \{\Psi_{\text{SP}}^t\}]/2.$$

star v1^{odd} in Au+Au 200GeV



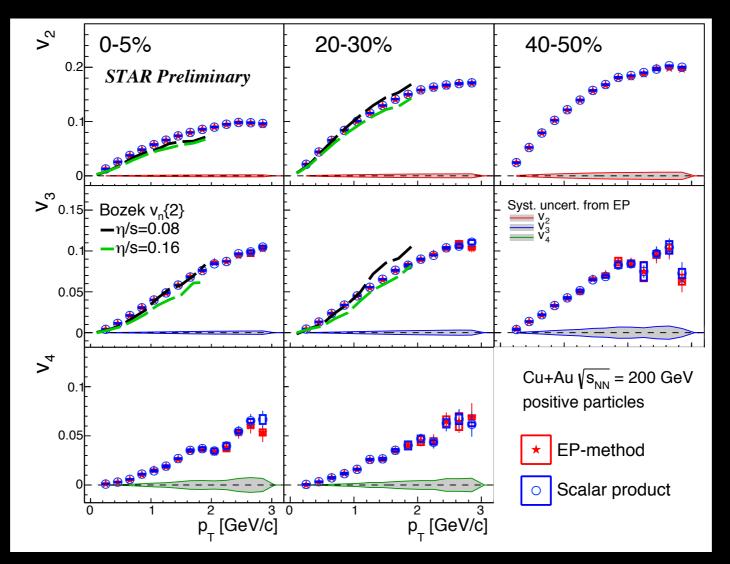
Small signal of v1 at mid-rapidity in Au+Au collisions

$$v_1^{\text{odd}} = \langle sgn(\eta)\cos(\phi - \Psi_1) \rangle$$



- * Ψ_{13} points to the direction where the density gradient is steeper = direction to which more high pT particles are emitted
- ★ Significantly larger $<\cos(\Delta \Psi_{13})>$ in Cu+Au →Larger density asymmetry
- * In Au+Au, $\langle \cos(\Delta \Psi_{13}) \rangle = 0$ but weak correlation between Ψ_{13} and spectator plane Ψ_{SP}

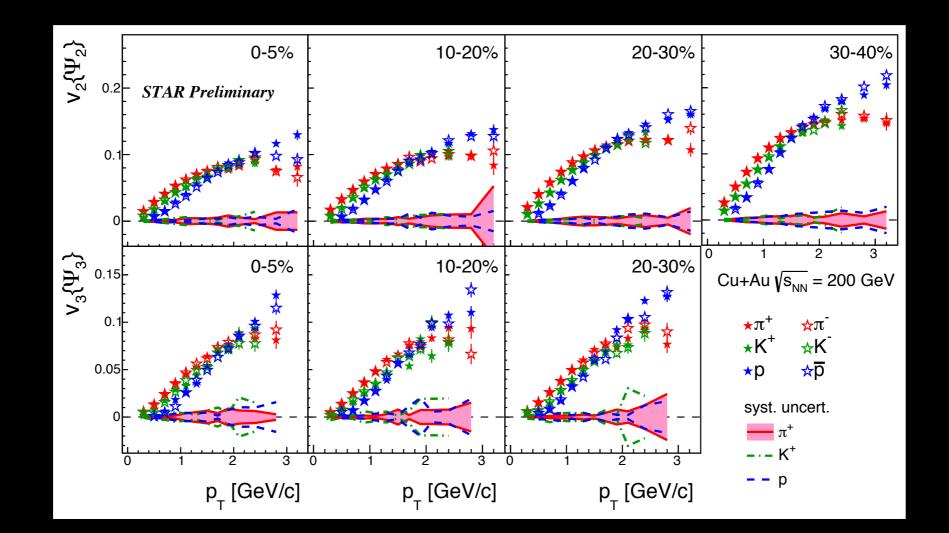
STAR Comparison with Hydro-model



vn{EP} is in good agreement with vn{SP}

- ▶ v₂ and v₃ are described well by e-b-e viscous hydrodynamic model
 - Bozek, PLB.717(2012)287
 - ${\it I}$ The data are close to the model calculations with η /s=0.08 and 0.16

STAR Identified Particle Vn



- ▶ π /K/p identification by TPC + TOF
- Similar trends observed in A+A collisions
 Mass ordering at low p_T (effect of radial flow)
 Baryon/meson splitting at intermediate p_T (partonic flow)

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STAR Measurements of azimuthal anisotropies

Event plane method

• Ψ_n (n>1) determined by TPC(η -sub) and EEMC

$$U_n = \langle \cos[\pi(\phi - \Psi_n)] / \operatorname{Res}\{\Psi_n\} \quad Q_{n,x}$$

 $Q_{n,y}$

Scalar product method

- STAR, PRC66.034904 (2002)
- in forward and backward region $v_n = \frac{\langle \vec{Q}_n^F(B) \cdot \vec{u} \rangle}{\sqrt{\langle \vec{Q}_n^F \cdot \vec{Q}_n^B \rangle}}$

- variation of track selection
- \odot For v₁, EP resolutions from different 3-sub events
- \odot For v_n, difference between TPC η -sub and EEMC

$$\mathbf{Fertice} = \mathbf{F} \begin{bmatrix} \mathbf{TPC} & \mathbf{TPC} & \mathbf{EEMC} \\ (-1 < \eta < -0.4) & (0.4 < \eta < 1) \\ * \Psi_2 & * \Psi_2 & 0 \Psi_2 \\ 0.6 & \Psi_3 & \nabla \Psi_3 & \Delta \Psi_3 \\ + \Psi_4 & \Phi \Psi_4 & \Phi \Psi_4 \\ 0.4 & \mathbf{F} \Psi_4 & \Phi \Psi_4 \\ 0.4 & \mathbf{F}$$

 $\Psi_n = \frac{1}{n} \tan^{-1}(Q_{n,y}/Q_{n,x})$

 $=\Sigma w_i \cos(n\phi)$

 $= \Sigma w_i \sin(n\phi)$

