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# DIPLOMOVÁ PRÁCE



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## Studium proton–protonových srážek na RHIC

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# DIPLOMA THESIS



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## Study of proton–proton collisions at RHIC

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**Abstrakt:**

Byly studovány proton–protonové srážky měřené detektorem STAR na urychlovači RHIC při energii  $\sqrt{s} = 200$  GeV. Minimum bias srážky byly rozděleny do dvou tříd, podle přítomnosti shluku částic s vysokou příčnou hybností (“jetu”). Vlastnosti obou výsledných tříd, měkkých a tvrdých eventů, byly porovnány pomocí jejich  $p_T$  spekter a rozdělení multiplicit nabitých částic. Před tím byly rozdělení opraveny na efekty detektoru (efektivita registrace drah částic a hledání primárního vertexu). Byla provedena analýza rozdělení příčných hybností a multiplicit, včetně jejich fitu pomocí “power-law” ( $p_T$  spektra) a Negativního Binomického rozdělení (multiplicita). Rozdělení byla porovnána s výsledky z antiproton–protonových urychlovačů.

**Klíčová slova:** proton–protonové srážky,  $p_T$  spektra, rozdělení multiplicit nabitých částic, měkké a tvrdé eventy

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**Abstract:**

Proton–proton collisions at  $\sqrt{s} = 200$  GeV measured by the STAR detector at RHIC collider were studied. Minimum bias events were divided into two classes, based on the presence or absence of high- $p_T$  particle cluster (“jet”). Properties of the two resulting event classes, soft and hard, were compared in terms of their transverse momentum spectra and charged particle multiplicities. Prior to this, measured distributions were corrected for detector effects (tracking and primary vertex finding efficiency). Analysis of the transverse momentum and charged multiplicity distributions was performed, including “power-law” ( $p_T$  spectra) and Negative Binomial Distribution (multiplicity) fits. Spectra were compared to the results from antiproton–proton colliders.

**Keywords:** proton–proton collisions, transverse momentum spectra, charged multiplicity distributions, hard and soft events



# Chapter 1

## Introduction

The Relativistic Heavy Ion Collider (RHIC) [1] was built mainly to study the hot and dense strongly interacting matter, created in heavy ion collisions. Under sufficiently extreme conditions, a transition to quark-gluon plasma (QGP), a new form of matter in which quarks and gluons act as free particles, is expected [2, 3].

As a reference a smaller system - proton-proton (p+p) collisions are studied, at the same center of mass energy and using the same detector. The characteristic observables of a possibly created QGP include the enhancement of strange hadron and low-mass dilepton yields and measured  $J/\Psi$ -suppression [3], when compared to the sum of binary nucleon-nucleon collisions.

A large data set of p+p interactions has been taken since RHIC start-up in the year 2000. Besides comparison to heavy ion collisions, it's highly desirable to study their inclusive properties like for example  $p_T$  spectra and also event-wise variables like charged particle multiplicity.

The aim of this thesis is to obtain and analyze  $p_T$  and charged multiplicity distributions (including necessary corrections to the data) and explore the differences between soft and hard events. The latter contain mini-jets from hard parton-parton scattering.

In the past, hadron collisions have been studied at the ISR (p+p) and at the S $\bar{p}$ pS ( $\bar{p}$ +p) colliders at CERN. At Tevatron,  $\bar{p}$ +p collisions are studied. Compared to ISR and S $\bar{p}$ pS, experiments at RHIC and Tevatron are able to record much higher amount of events. More information about RHIC operations, delivered luminosity and numbers of events recorded can be found in appendix A.

For this analysis, I have used minimum bias data from RHIC's first p+p run in 2001-2002, collected by the STAR detector at  $\sqrt{s} = 200$  GeV, the first p+p data available at this center-of-mass energy.

After a brief introduction into p+p collisions in chapter 2, the RHIC and STAR detector setup will be described in chapter 3. In chapter 4 the methods used in data analysis are explained. My results are presented in chapter 5, together with a comparison to other experiments and Hijing Monte Carlo model.

# Chapter 2

## Theoretical concepts

In the following, I will describe high-energy hadron interactions, i.e. hadron collisions with centre of mass energy ( $\sqrt{s}$ ) greater than  $\approx 10$  GeV.

Hadrons are composite objects, the constituents of which are quarks and gluons. These are strongly interacting, presumably structureless and point-like objects. There are three families of quarks, (u,d), (c,s), (t,b). Quarks have non-integer charge and are fermions. A nice review of hadron interactions can be found in [4].

The theory describing the *strong force* is *quantum chromodynamics* (QCD) [5]. The QCD started to emerge in early nineteen sixties, when Gell-Mann [6] and Ne'eman [7] introduced the Eightfold way, arranging mesons and baryons into multiplets of SU(3) group. For prediction of  $\Omega^-$ , discovered in 1964 in Brookhaven, Gell-Mann was awarded a Nobel price in 1969. The quark model was established in 1964 by Gell-Mann [8] and Zweig [9]. It was yet not known, whether quarks were real constituents of hadrons, or a pure mathematical concept. In 1965, Nambu [10] laid the foundation to QCD, giving quarks *color* (quarks formed SU(3)<sub>C</sub> triplet) and suggesting that the strong force is mediated by 8 gauge bosons.

Deep inelastic scattering (DIS) experiments were used to probe inner structure of hadrons. In DIS, leptons scatter off hadrons<sup>1</sup> and the kinematics is reconstructed by measuring the scattered lepton. In 1968, Feynman developed basic ideas of *parton model*<sup>2</sup> - describing DIS as scattering of leptons at quasi-free point-like constituents of hadrons. Partons' quantum numbers (spin  $\frac{1}{2}$ , charge) turned out to be identical with constituent quarks of the quark model, giving rise to quark-parton model (QPM), where partons were identified with quarks.

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<sup>1</sup>mainly protons

<sup>2</sup>parton = PART of protON

But although quarks are real objects (point-like fermions), a free quark was never observed (one of the signatures would be its non-integer electric charge). This is due to the *confinement*. As two quarks move apart, the energy of colour field (a string [11]) between them increases and when their distance  $\approx 1$  fm, it's energetically advantageous to create a quark-antiquark pair.

## 2.1 Perturbative QCD

In QPM, hadrons are described by *parton distribution function* (PDF)  $q(x)$ . That is a probability density of finding a parton inside the hadron, carrying fraction  $x$  of its momenta. PDFs cannot be predicted from first principles, and can be measured in DIS experiments. Quarks turned out to carry only half of the hadron momenta, the remaining half is carried by gluons. These can't be "seen" by leptons in DIS, as they only interact strongly (which leptons don't).

In DIS processes with large 4-momentum ( $Q^2$ ) transfer, hadrons behave as if they consist of *free* partons. This feature of QCD is called *asymptotic freedom*. This is caused by color charge screening effects of QCD vacuum. The effective QCD coupling  $\alpha_s$  decreases with  $Q^2$ :

$$\alpha_s(Q^2) = \frac{4\pi}{(11 - \frac{2}{3}n_f) \log(Q^2/\Lambda_{QCD}^2)} \quad (2.1)$$

Here  $n_f$  is the number of quark flavors (assumed to be 6) and  $\Lambda_{QCD}$  is an important perturbative scale variable, which is determined experimentally:  $\Lambda_{QCD} \approx 0.2$  GeV. The theoretical discovery of asymptotic freedom in 1973 by Gross, Politzer and Wilczek [12, 13] was awarded with Nobel price in 2004.

A hadron-hadron interaction with a sufficiently large  $Q^2$  can be factorized, dividing the interaction into three parts, separated in time: parton distributions, hard processes and hadronization. The inclusive invariant cross section for process  $A + B \rightarrow C + X$  is given in the first order by

$$\frac{d\sigma}{dy} = \sum_{a,b} \int_0^1 dx_a dx_b dz_c f_a(x_a, Q^2) f_b(x_b, Q^2) \frac{D_{C/c}(z_c, Q^2)}{z_c} \frac{d\sigma_{ab \rightarrow cd}}{dy} \quad (2.2)$$

where  $f_a(x_a, Q^2)$  is a PDF of hadron A (containing parton a) and  $D_{C/c}(z_c, Q^2)$  is a *fragmentation function*, a probability of parton c to produce hadron C, carrying fraction  $z$  of its momenta. This comes from the

*model of independent fragmentation* [5]. The  $Q^2$  dependence is given by QCD evolution equations (see [5] for details).

The cross section for process  $ab \rightarrow cd$  is calculated using the Feynman rules of QCD and experimentally determined QCD coupling constant  $\alpha_s$ . Hadronization is a non-perturbative process and can be described by independent fragmentation or by *string model* (to be discussed later).

## 2.2 Jets and Mini-Jets

Events with hard scattering are very rare at  $\sqrt{s} < 100$  GeV. They give rise to *jets*, narrow clusters of particles with high transverse energy, first observed by UA2 experiment at CERN SppS collider [14]. In hard scattering, partons interact at small impact parameter and scatter at wide angles. Hence high transverse energy.

More evidence for jets in high energy p+p collisions comes from the mean number of produced charged particles ( $\langle N_{ch} \rangle$ ). This number is proportional to the cross-section and at lower energies, i.e. less than 100 GeV, it has logarithmic dependence on  $\sqrt{s}$  as shown by Thome et al [15] in the nineteen seventies. SppS experiments at energies 200 – 900 GeV measured cross-sections which are clearly above the values by Thome’s ‘law’ and thus signal new phenomena creating additional particles [16]. Also violation of the so called KNO scaling (to be discussed later) in multiplicity distributions has been attributed to the creation of jets in hadronic collisions.

A mini-jet is a jet whose transverse momentum  $p_T$  is large if compared with the nucleon mass but much less than  $\sqrt{s}$ . According to some authors [17] perturbative calculations are possible down to  $p_T$  of the order of  $\Lambda_{QCD}$ . In this semi-hard momentum regime it would be possible to compute the inclusive jet cross section down to  $p_T > p_{Tmin} \simeq 1$  GeV/c. However, when  $p_T$  is not very large, there are theoretical and experimental difficulties in distinguishing the particles in the mini-jet from the other particles in the reaction.

## 2.3 Soft processes

The overwhelming part of p+p cross section is due to “soft” processes with small transverse momentum transfer. They are associated with coherence effects among (already) interacting partons. Even in soft p+p interactions at high enough energies, tens to hundreds of particles in final state can be produced. As no short distance (high  $Q^2$ ) interaction is involved and the typical

length scale of the dynamics is  $\approx 1$  fm, the corresponding strong coupling constant is of the order of unity. Alternative non perturbative description must be adopted. Unlike perturbative QCD, we are not able to describe a soft process from the first principles. Instead, phenomenological models are used, most of which are QCD-inspired.

### 2.3.1 Exchange forces

A theory of exchange forces [4], inspired by the old one-pion-exchange (OPE)<sup>3</sup> approximation to the nuclear force, describes hadron-hadron scattering by means of Reggeons and Pomerons exchanges. To preserve color neutrality of hadrons, the object exchanged must be colorless. Hence, any hadron, meeting the required quantum numbers (spin, charge, parity,...) can be exchanged. These particles lie on the so called *Regge trajectories*, where their spin is a linear function of the square of their mass.

The Pomeron is an object with vacuum quantum numbers, which can be described as a multiple gluon exchange. Pomeron exchange is the dominant particle production mechanism, with particles being produced from gluons radiated either from the exchanged gluons or from the valence quarks, between which the gluons (Pomeron) were exchanged [4].

### 2.3.2 String model

String model [11] is an effective description of strong interaction phenomena in large-distance processes, such as low-momentum-transfer interactions and hadron mass spectra. A string is a relativistic object, a color flux tube stretched between two quarks. The string tension, that holds the two quarks together, is  $\approx 1$  GeV/fm.

Mesons are formed of a mass-less quark-antiquark pair, connected with a string. Regge trajectories can be described in the model of spinning string. The mass then arises as the kinetic energy of the string, whereas spin is determined by string's angular momentum.

There are four different models to form a baryon [18]: a) quark and diquark connected with one string (q–qq); b) the “three-string” model or Y configuration with three strings from three quarks joined in the fourth mass-less point; c) the “triangle” model or  $\Delta$ -configuration with pair-wise connection of three quarks by three strings; d) the linear configuration (q–q–q) with quarks connected in series.

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<sup>3</sup>In 1934, Yukawa found relationship between the range of nuclear force and mass of its carrier and predicted its mass to be about 100 MeV. Pions were only discovered experimentally in the late nineteen forties.



String theory describes particle production by means of string breaking. One of the models of string breaking mechanism is The Symmetric Lund Model [19]. Another model is Artru-Mennessier Off-shell Resonance model (AMOR) [20]. Unlike Lund model, AMOR permits off-shell resonances to be produced, avoiding otherwise necessary correction procedures to force particles on their mass shells.

## 2.4 Models of high-energy hadron interactions

In most models particle production is assumed to have two components. The hard interaction is described by perturbative QCD. The soft component is usually described as separation of color charges resulting, through different mechanisms, in low- $p_T$  production. At higher energies many of the changes in the event characteristics can be described by increasing the number of strings.

### 2.4.1 Pythia

Pythia [21] is a Monte Carlo generator developed by the Lund University group. The parton cross section is approximated by  $2 \rightarrow 2$  processes and the associated leading order Feynman diagrams. PDFs and fragmentation functions are user-defined from a list of currently available parametrizations. The default hadronization mechanism in Pythia is the Symmetric Lund Model.

In Pythia the simplest possible way to produce a soft event is to have an exchange of a very soft gluon between the two colliding hadrons. Without initially affecting momentum distributions of partons, the hadrons become color octet objects. If only valence quarks are considered, the color octet state of a baryon can be decomposed into a color triplet quark and an antitriplet diquark. Strings are stretched between the quarks and diquarks and as they move apart, the strings fragment.

Pythia assumes that different pairwise interactions can take place independently so that the number of scatterings in the collision, for a fixed hadron-hadron impact parameter, is given by a Poisson distribution. Radial hadronic matter distribution in the hadron is assumed to be double gaussian.

### 2.4.2 Hijing

HIJING (Heavy Ion Jet INteraction Generator) [22] Monte Carlo model uses the successful implementation of pQCD processes of Pythia and is in many

ways similar to this model. Compared to Pythia, it enhances mini-jet production. It's currently the only model describing hard scattering and at the same time able to describe nuclear effects, like nuclear shadowing and jet quenching. Therefore it has been widely used by RHIC experiments, to describe and compare p+p, d+Au and Au+Au collisions.

### 2.4.3 Multiple scattering formalism (VENUS/NeXus)

The VENUS model by Klaus Werner [20] is based on Gribov-Regge theory (GRT) [23]. This theory allows multiple interactions between partons to happen 'in parallel' - opposed to Pythia, where the parton interactions are mainly  $2 \rightarrow 2$  processes. Elementary interactions are described by Pomeron exchanges.

VENUS model was further extended and renamed into NeXus model [24]. It uses a parton-based GRT framework where parton-parton scattering processes are described in a consistent way as 'soft', 'hard' or 'semi-hard'.

Hadronization is achieved by string fragmentation, using a model similar to AMOR, rather than the Lund framework. String breaking points are defined in a covariant and gauge invariant way by an area law [25].

## 2.5 Cross sections definitions

Hadronic processes can be divided into elastic and inelastic collisions. In inelastic collisions, the incident particles transform a part of their energy to create new particles. Inelastic collisions can be subdivided into diffractive and non-diffractive processes.

In diffractive processes one nucleon acts as a region of absorption and the interference of scattering amplitudes gives rise to diffraction pattern in the forward and backward regions. A nucleon suffering a diffractive scattering will become excited and lose a small amount of energy when breaking up into a few particles at a small emission angle. This can occur in one (singly diffractive) or both of the nucleons (doubly diffractive).

In non-diffractive processes the nucleons hit 'head-on' and both disintegrate creating large particle multiplicities at mid-rapidity. For experimental reasons, non-singly diffractive (NSD) cross section is often measured. This consists of doubly-diffractive and non-diffractive inelastic collisions. According to STAR measurement at  $\sqrt{s} = 200$  GeV,  $\sigma_{NSD,proton-proton} = 31.3 \pm 3.7$  mb [26] and represents about 70 % of inelastic p+p cross section.

A frequent object of study are reactions  $p+p \rightarrow C+X$ , commonly named *one particle inclusive processes*. The invariant cross section of this type of

interaction is defined by:

$$\sigma_{\text{pp}}^{\text{C}} = E_{\text{C}} \frac{d^3\sigma}{dp_{\text{C}}^3} \quad (2.3)$$

where  $\frac{d^3\sigma}{dp_{\text{C}}^3}$  is the differential cross section, i.e. the probability per unit incident flux, of detecting particle C within the phase space element  $dp_{\text{C}}^3$ ;  $E_{\text{C}}$  is included to ensure the Lorentz invariance. In a general case invariant cross section is a function of three independent variables (i.e. the components of momentum vector).

This expression can be simplified using *rapidity*  $y$ , definitions and details can be found in appendix B. Rapidity can be approximated by *pseudorapidity*  $\eta$ :

$$\eta = -\ln \tan \frac{\theta}{2} \quad (2.4)$$

where  $\theta$  is the angle between particle's momentum and beam axis. In case of azimuthal symmetry the invariant cross section 2.3 is a function of two independent variables - transverse momentum  $p_{\text{T}}$  and pseudorapidity  $\eta$ .

## 2.6 Event charged multiplicity

One of the basic global event observables is its charged multiplicity ( $N_{\text{ch}}$ ). Sometimes called simply 'multiplicity', it's defined as the number of charged secondary particles plus resonance decay products in the final state, on the other hand weak decay products (of  $\text{K}^0$ ,  $\Lambda$ , ...) are not counted.

Multiplicity distributions in p+p collisions were measured at SPS and ISR facilities at CERN, up to energies of  $\sqrt{s} = 62$  GeV. The distribution shape showed independence of collision energy, so called KNO scaling [27]. If  $P_N$  is the probability to have multiplicity  $N$  and  $\langle N \rangle$  is the mean multiplicity, KNO scaling states that  $\langle N \rangle P_N = \psi(z)$  is a function of the scaled multiplicity  $z = \frac{N}{\langle N \rangle}$  with no further dependence other than that implied from the relation between  $\langle N \rangle$  and  $\sqrt{s}$ .

At higher energies, however, KNO scaling is broken, as discovered by the UA5 collaboration at SppS [28]. Instead, multiplicity distributions were shown [28] to follow a two-parameter Negative Binomial Distribution (NBD):

$$P(N; \langle N \rangle, k) = \binom{N+k-1}{k-1} \left( \frac{\langle N \rangle / k}{1 + \langle N \rangle / k} \right)^N \frac{1}{(1 + \langle N \rangle / k)^k} \quad (2.5)$$

where  $\langle N \rangle$  and  $k$  ( $k > 0$ ) are energy-dependent. NBD is valid at SPS to ISR to S $\bar{p}p$ S energies, in both full phase space and its subspaces (experimentally given cuts on  $|\eta|$ ,  $p_T$  etc.).

There are two special cases for  $k$  value: for  $k = 1$  NBD gives exponential distribution; for  $k \rightarrow \infty$  NBD goes to Poisson distribution. The dispersion of NBD is:

$$D = \frac{\langle N \rangle + k}{k} \langle N \rangle \quad (2.6)$$

which is higher than Poisson distribution's  $\langle N \rangle$ .

It can be easily shown, that for  $\bar{N} \gg k$  NBD goes to  $\Gamma$  distribution in  $z$ :

$$\langle N \rangle P(N; \langle N \rangle, k) \rightarrow \frac{k^k}{\Gamma(k)} z^{k-1} e^{-kz} \quad (2.7)$$

One could get KNO scaling, if  $k$  is independent of  $\sqrt{s}$ , but  $k$  falls with energy [28], so although one gets  $\Gamma$  distribution in  $z$  at high  $\sqrt{s}$ , KNO scaling is broken.

## 2.7 Transverse momentum spectra

For unidentified particles,  $p_T$  spectrum is well described with QCD-inspired 'power-law' [29] functional form:

$$\frac{1}{p_T} \frac{dN}{dp_T} = A \left( 1 + \frac{p_T}{p_0} \right)^{-n} \quad (2.8)$$

that behaves like a power of  $p_T$  at high  $p_T$ , coming from jets fragmentation. At low  $p_T$  it has an exponential form, suggesting thermal behaviour. The factor  $\frac{1}{p_T}$  comes from equation B.9.

The behaviour of the  $p_T$  distribution in relation to event multiplicity has been one of the most important subjects in minimum bias physics. An increase of  $\langle p_T \rangle$  with charged particle density in the central region was first observed by UA1 collaboration [29]. Due to large systematic errors coming from calculation of  $\langle p_T \rangle$  and statistical errors on true multiplicities of high energy events, the shape of the correlation is an open question, but the increasing trend is clear, which is the opposite of what one would expect from kinematics under the condition of particle density saturation. Many interpretations have been proposed in terms of semi-hard effects [30] in the central region or possible hadronic phase transition [31], but the correct explanation is still unknown.

# Chapter 3

## Experimental setup

### 3.1 The accelerator

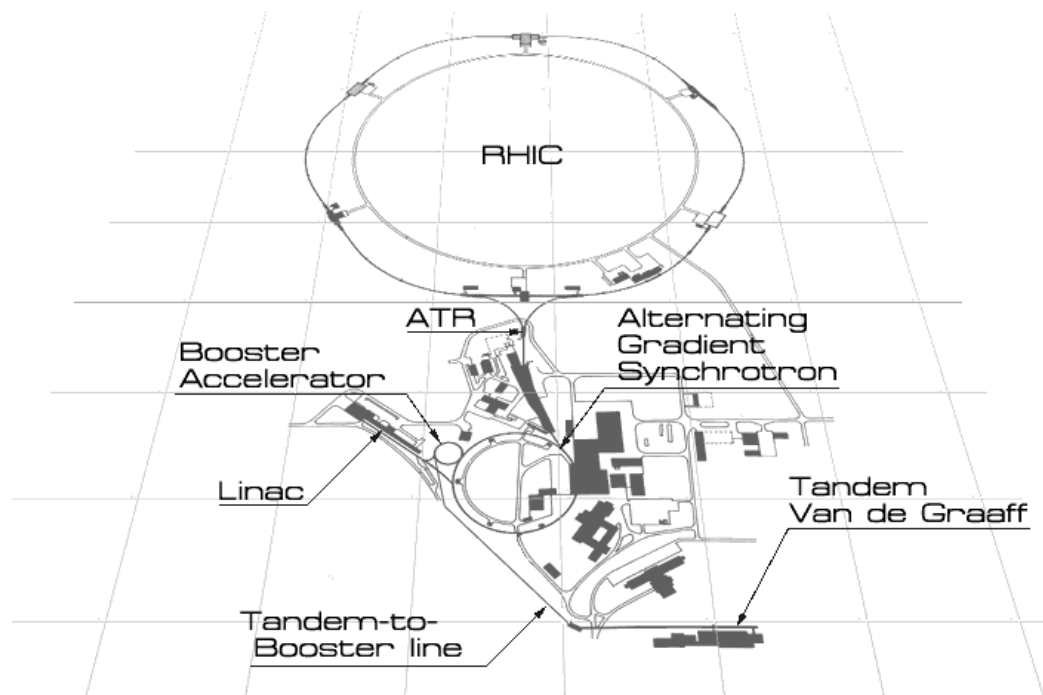


Figure 3.1: The RHIC accelerator complex.

The Relativistic Heavy Ion Collider [1] (RHIC) is currently the world's highest energy accelerator of heavy nuclei and the world's first polarized proton collider. It's situated at Brookhaven National Laboratory (BNL), New York, USA.

The collider consists of two independent concentric acceleration and storage rings with a circumference of 3.8 km, which are made of 1740 superconducting magnets. There are six intersection points, four of which are currently instrumented with experiments. The layout of the RHIC accelerator complex is shown in Fig. 3.1.

RHIC can store and collide nuclei with masses ranging from protons to Au and due to the independence of the storage rings it can collide beams of unequal masses, such as deuterons with Au nuclei. The top energy for Au+Au collisions is  $\sqrt{s_{NN}} = 200$  GeV per nucleon, for p+p it's  $\sqrt{s} = 500$  GeV.

Gold ion beams originate from a pulsed sputter-source that provides negatively charged gold ions to the Tandem Van de Graaf accelerator. They are partially stripped of their electrons by passing through a carbon foil and further accelerated to 1 MeV/u. Now with a charge  $Q = +32$  they're delivered to Booster synchrotron and then accelerated to 95 MeV/u (95 MeV per nucleon). Next the beam is injected into the AGS (Alternating Gradient Synchrotron) with a charge of  $Q = +77$ . The ions are stripped of the two remaining electrons and brought up to the RHIC injection energy (10.8 GeV/u). Once in the RHIC ring, they are accelerated to 100 GeV/u and stored for up to 10 hours.

For the p+p running the injection procedure is different. Protons originate from the 200 MeV LINAC, pass through the booster to the AGS. Accelerated here to 24.3 GeV, they're injected into the RHIC.

In the RHIC rings the protons are stored in so-called 'bunches' with  $100 \times 10^9$  protons per bunch. The main reason for storing the ions in discrete bunches is to allow for optimized acceleration in the RF cavities. Furthermore, by tracking the location of the bunches, the experiments can be informed of their location and triggering is facilitated. Up to 120 bunches can be stored, resulting in the beam crossing rate of 10 MHz.

The six intersection regions of the RHIC storage rings allow the beams to be steered into head-on collisions. Currently, four intersection regions are instrumented with two major detectors, STAR and PHENIX, and two smaller ones, BRAHMS and PHOBOS.

## 3.2 STAR detector

The Solenoidal Tracker at RHIC (STAR) [32] is a detector located at the 6 o'clock position at the RHIC collider ring. STAR is designed to study the behavior of strongly interacting matter at high energy density and to search for signatures of Quark Gluon Plasma (QGP) formation. Besides, it's well equipped to study spin-dependent asymmetries in polarised proton collisions,

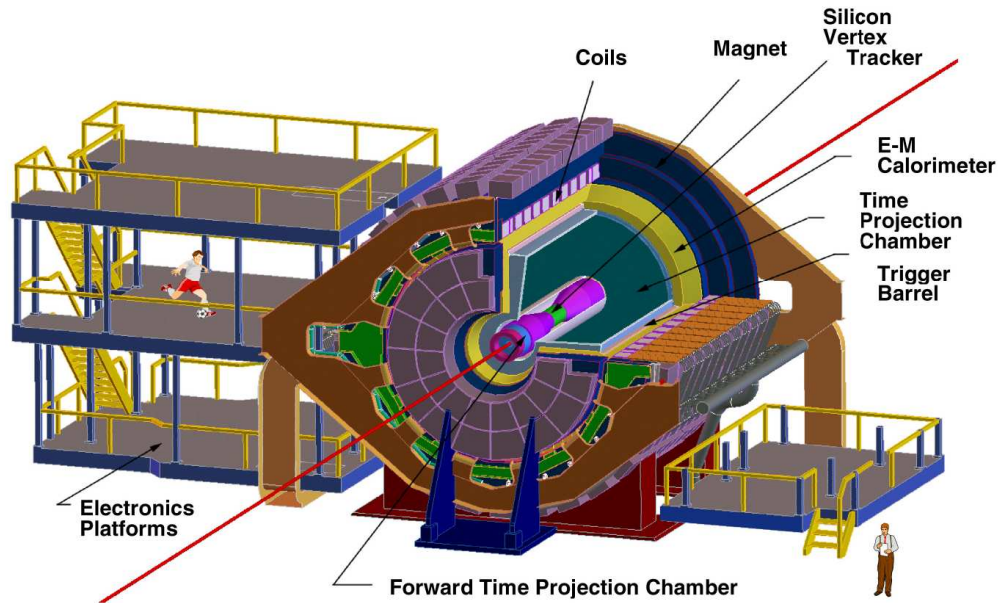


Figure 3.2: A perspective view of the STAR detector.

due to its Endcap Electromagnetic Calorimeter (EMC) (not present in year 2001-2002 running).

STAR is an azimuthally symmetric, solenoidal detector, shown in Fig. 3.2. Its unique feature is the large Time Projection Chamber (TPC) with a full azimuthal and a large pseudorapidity coverage. The large acceptance of the detector enables measurement of a large fraction of the thousands of charged hadrons produced in a heavy ion collision. Thus, very high multiplicity  $p+p$  events can be detected and reconstructed with high tracking efficiency, avoiding track merging. Figure 3.3 shows the subsystems the detector is built of.

A solenoidal magnet with a uniform magnetic field of 0.5 Tesla surrounds the whole detector and provides for charged particle momentum analysis. A brief description of the detector subsystems, starting from the beamline:

- Beam Beam Counters (BBC) (not shown): Trigger detector sensitive to forward charged tracks used in  $p+p$  triggering.
- Zero Degree Calorimeters (ZDC): Trigger detector sensitive to very forward spectator neutrons (neutral, so are not steered into the beam pipe). Used for Au+Au triggering.
- Forward Time Projection Chambers (FTPC): Tracking detector at for-

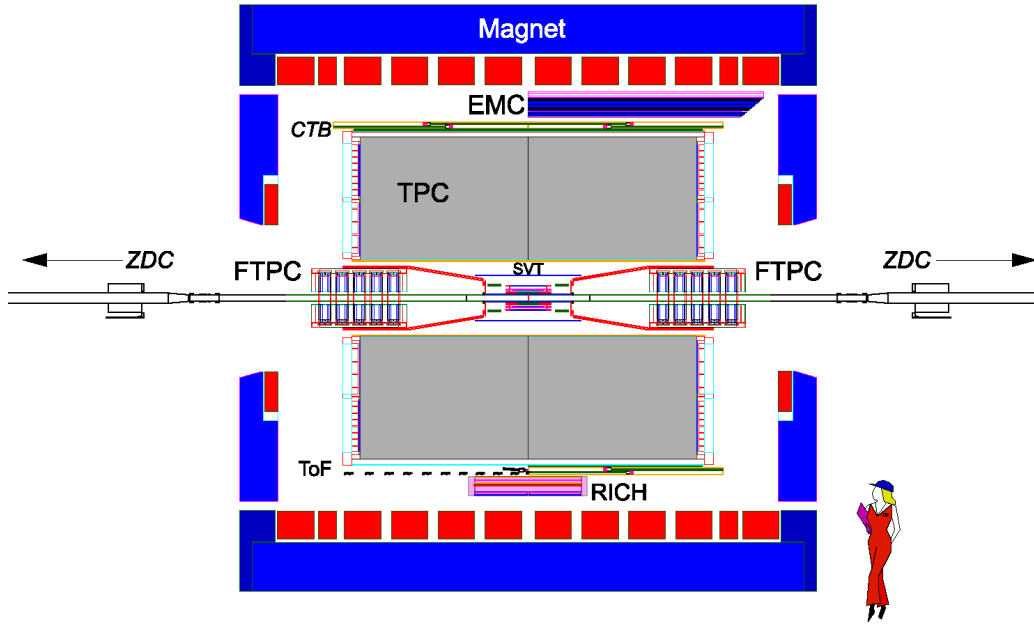


Figure 3.3: Cutaway side view of the STAR detector as configured in 2001-2002 running.

ward pseudorapidities ( $2.7 < |\eta| < 3.9$ ). Still in commissioning during the 2001-2002 running.

- Silicon Vertex Tracker (SVT): High precision measurement of tracks close to the interaction vertex.
- Time Projection Chamber (TPC): Measures charged tracks and specific energy loss ( $dE/dx$ ), allowing for identification of  $\pi^\pm$ ,  $K^\pm$ ,  $p$ ,  $\bar{p}$  and  $e^\pm$ .
- Time Of Flight (TOF): Measures identified charged particles in the intermediate momentum range.
- Central Trigger Barrel (CTB): Fast scintillator detector to estimate charged particle multiplicity (proportional to ADC sum). Used in Au+Au triggering (centrality selection).
- Ring Imaging CHerenkov counter (RICH): Particle identification at high transverse momenta.
- Barrel Electromagnetic Calorimeter (EMC or BEMC): Measures energy of electrons, positrons and gammas, partly of (charged as well as neutral) hadrons.



STAR detector has been further upgraded, including the EEMC (covering  $1 < \eta < 2$  with full azimuth angle), Silicon Strip Detector (SSD), Photon Multiplicity Detector (PMD) and Forward Pion Detector (FPD), on the other hand the RICH is no longer part of STAR detector.

There's a major upgrade ongoing, a large-acceptance ( $|\eta| < 1.0$  and full azimuth) Time Of Flight (TOF) detector, that will allow for PID at intermediate momenta, together with separating pions from electrons at low momenta. It's currently in mass-production and testing stage. A significant R&D is being done for STAR, including the HFT, a hybrid silicon pixel detector, to replace the SVT.

### 3.2.1 Time Projection Chamber (TPC)

STAR TPC [33] is a large cylindrical gas chamber. With a length of 4.2 m it's the largest TPC in the world. Its inner and outer radii are 50 cm and 200 cm, respectively. TPC setup is shown in figure 3.4.

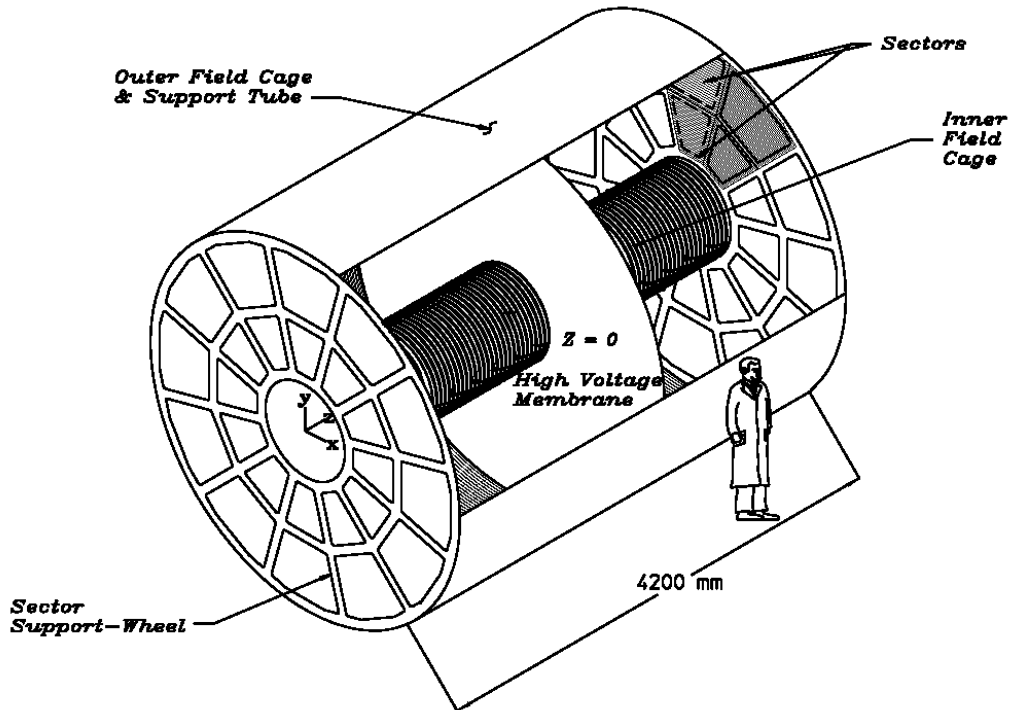


Figure 3.4: STAR Time Projection Chamber.

The TPC is filled with P10 gas (10% methane, 90% argon) regulated at 2 mbar above atmospheric pressure. Its primary attribute is a fast drift

velocity (typically  $5.45 \text{ cm} \cdot \mu\text{s}^{-1}$ ) which peaks at a low electric field. As charged particles pass through the TPC gas, they ionize its molecules and secondary electrons are released.

A uniform electric field of 28 kV between the central membrane and TPC endcaps drives these electrons to anode pad planes at the TPC endcaps. The total drift time is approximately  $40 \mu\text{s}$  and depends on the actual drift velocity. This has to be calibrated during the run with lasers, disposed around the outer radius of the TPC.

The anode pad plane where the electrons are read out is divided into 12 sectors in azimuth. Each sector is equipped with 5692 read-out pads in 45 padrows. A view of a TPC sector with description is shown in figure 3.5.

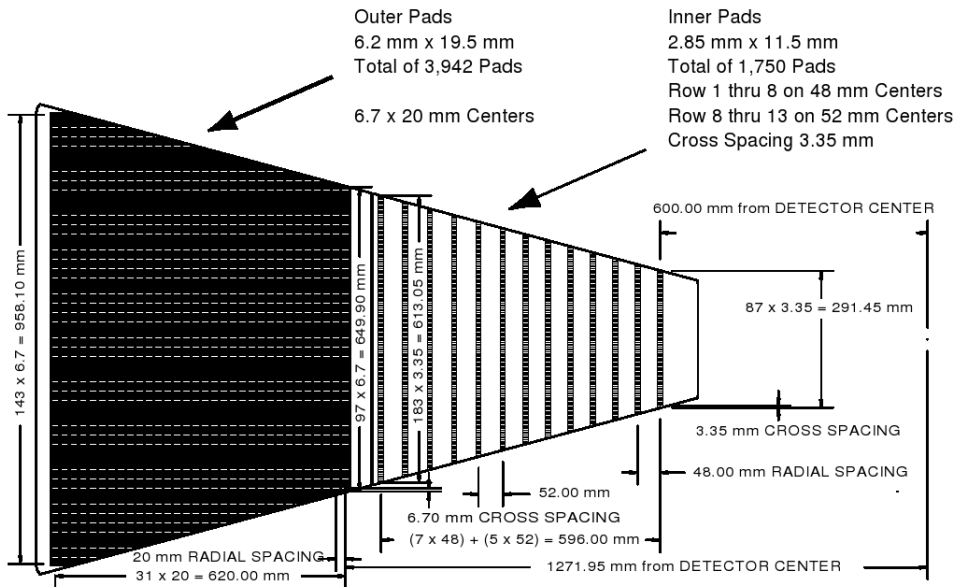


Figure 3.5: TPC sector (in one of the two anode planes) with read-out pads.

Multi-wire proportional chambers (MWPC) are used for the electron signal amplification. The gating grid is used to prevent electrons from untriggered event from reaching the read-out pads and to prevent positively charged ions from drifting into the TPC volume. When the charge is collected, the pad location in  $(r, \phi)$  plane plus the drift time information is used to reconstruct the 3D position of the hit (the coordinate along the beam line is  $z$  in STAR).

The trigger + gating grid opening cause a  $2\mu\text{s}$  delay, which means a drift path of 11 cm at the typical drift velocity at both sides of the TPC, so the effective TPC length is  $420 - 2 \cdot 11 = 398 \text{ cm}$ .

The TPC is capable of measuring charged tracks within  $|\eta| < 1.8$  and with transverse momenta greater than  $0.1 \text{ GeV}/c$ . Particle identification is achieved by measuring  $dE/dx$ . The dependence of  $dE/dx$  on particle's momentum is displayed in figure 3.6. PID based on  $dE/dx$  in the TPC can be used for momenta up to  $1 \text{ GeV}/c$ . In the future, STAR will be instrumented with a new Time Of Flight (TOF) detector, which will increase the PID capability of the TPC into higher momenta.

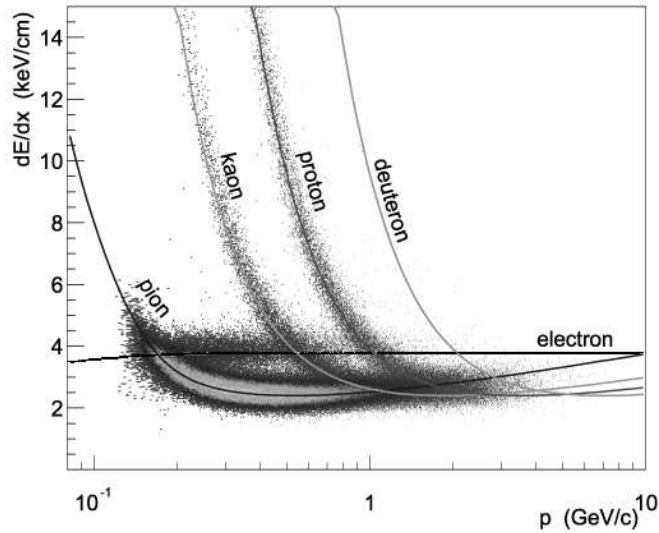


Figure 3.6: Ionization energy loss in the TPC ( $dE/dx$ ) for different particle species. The shaded area shows identified pions.

The acceptance of the TPC is 96 % for high momentum tracks travelling perpendicular the beamline. The 4 % inefficiency is caused by the spaces between the sectors which are required to mount the wires on the sectors. Also, any space points that fall on the last two pads of a pad row are ignored. This fiducial cut is applied to avoid position errors that result from tracks not having symmetric pad coverage on both sides of the track. This fiducial cut reduces the total acceptance to 94 %.

The TPC hit resolution for vertical tracks normal to a padrow is between  $0.4 - 0.7 \text{ mm}$  in the transverse direction and between  $0.7 - 1.2 \text{ mm}$  in the drift direction. Another important quantity is the  $p_T$  resolution. It increases roughly linearly from 2% for  $p_T = 0.5 \text{ GeV}/c$  to 7% at  $p_T = 9.25 \text{ GeV}/c$  [26].

### 3.2.2 Trigger detectors

Since the crossing rate of 10 MHz is much higher than the maximum read-out rate of the slow detectors (for example TPC:  $\approx 100$  Hz, given by its front end electronics), trigger detectors have to be used. They have to tell whether there was an interaction within given beam crossing and whether this interaction is of interest and the event should be read out. As a trigger, the calorimeters (EEMC, BEMC, ZDC) are often used. During the p+p running of 2001/2002 trigger detectors used were the BBC and the CTB.

The Beam Beam Counters (BBC) are used to provide minimum bias trigger for p+p running. They are installed  $\approx 4$  m away from the interaction point and cover the pseudorapidity region  $3.3 < |\eta| < 5.0$ . To reject background interactions (such as interactions with beam pipe gas), the signals are required in coincidence. This dual-arm topology makes the BBC coincidence trigger sensitive to Non Single Diffractive (NSD) part of inelastic cross-section. Its efficiency has been estimated 86% during the 2001/2002 run [26]. BBC coincidence rate is the primary luminosity measure in STAR. In the 2001/2002 run, it ranged from 1 kHz to 20 kHz with average value  $\approx 6$  kHz.

The Central Trigger Barrel (CTB) consist of 240 scintillator slats arranged cylindrically around the outer TPC radius and covering  $|\eta| < 1.0$ . The CTB is read out every beam-crossing, so it can be used to identify TPC tracks that have matching hits in the CTB. This is useful for identifying tracks from pile-up events which will be discussed further in detail.

### 3.2.3 STAR operations

The last running of RHIC so far was in the period of 20 weeks from February to June 2006. Polarised protons were collided at  $\sqrt{s} = 200$  GeV and for the last 3 weeks, at  $\sqrt{s} = 62$  GeV.

During this run, I've had an opportunity to participate in data taking. As a shift crew, I served one week of daily shifts in STAR control room. The main task of the shift crew person is so called fast-offline Quality Assurance (QA), observing the main detector-related quantities and assuring data quality.

I was also responsible for monitoring of TPC gas system and water cooling system. Being a run-time operator, I took part in starting and stopping runs, both for physics data taking and for detector system calibrations (pedestal and laser runs).

# Chapter 4

## Data analysis

### 4.1 Event reconstruction

After the event is triggered, all detectors are read out and signals are converted into ADC-counts by the read-out electronics. Digitized data are collected in a standard DAQ (Data Acquisition) format and stored into the HPSS (High Performance Storage System), a tape library.

When the run is finished, raw data is passed through an offline reconstruction chain, and stored on disks in the so called MuDst compressed format, making it accessible for physics analyses. MuDst files are ROOT [34] files, containing TTree objects with branches corresponding to various objects, like tracks, kinks, vertices, calorimeter cells, etc.

ROOT4STAR [35] framework is used in both reconstruction chain and physics analyses. It's basically ROOT with STAR-specific classes (objects). Classes called makers are used to run an analysis, like for example StMuDstMaker for reading MuDst files and accessing their content.

The reconstruction process and most of the analyses in STAR are performed at RHIC Computing Facility (RCF), a computer farm located at BNL. Other important (besides ROOT4STAR) parts of STAR analysis framework are Scheduler (a program used to make analysis jobs and run them at RCF) and FileCatalog, a database of all files stored at RCF, which is described in more detail in appendix A.

#### 4.1.1 TPC tracking

During the TPC drift time the 5690 read-out pads of its each sector are sampled 512 times, thus give a total of  $\approx 70$  million of pixels in 3D. Due to diffusion of secondary electron in TPC gas, charge from one hit is distributed between multiple pads and time bins, so a clustering algorithm has to be

applied to determine the center and the amplitude of the hit. There are 45 padrows in each TPC sector, so a track can have a maximum of 45 hits.

Once the hits are found, the TPT (TPC Tracker) algorithm is used (starting from the outermost padrow) to construct segments of tracks. In the second step, a helical extrapolation is used to add more points to track segments and extend them inwards and outwards. Once all hits have been associated to tracks, a procedure of helix-merging is performed to identify low  $p_T$  tracks, that have been wrongly identified as two separate tracks. TPC tracking efficiency will be discussed in the next section.

Tracks reconstructed in the TPC are called global tracks. They include tracks from the triggered event, tracks from pile-up events and background tracks, such as particles from cosmic ray showers. The data volume from TPC of one p+p event is  $\approx 10$  kB. Global tracks having  $\geq 15$  hits are called 'good global tracks' and are used for further analysis.

### 4.1.2 Primary vertex finding and event pile-up

As the probability of two interactions in a single beam crossing is very low, it can be neglected. However since the TPC read-out rate is 25 kHz, there is a high probability of more interactions happening during the read out of triggered event. Details on pile-up rates and luminosities can be found in appendix C. Tracks from the pile-up event will be shifted in  $z$  (due to the time difference) and some of them will be split into two parts by the central membrane. There's no usable method based on these tracks' topology to distinguish between the triggered and the pile-up event.

However, the CTB detector contains the time information. So only those global tracks, having a matching hit in the CTB, are guaranteed to come from the triggered event. As the average multiplicity of charged tracks in a p+p event is low, pile-up tracks could easily smear the primary vertex ('vertex') finding. Therefore only good global tracks with a matching hit in the CTB are used to find the vertex.

The position of the beam in the  $(x, y)$  plane is known with high precision and the  $z$  position of the primary vertex is determined by minimizing the  $\chi^2$  value, using distance of closest approach ( $DCA$ ) of the tracks to the vertex. In some special cases, when there's no good global track with a CTB-matching hit, the vertex can still be found, but the probability of finding the vertex in such cases is much lower than with the CTB-matching algorithm.

Once the primary interaction vertex is found, all global tracks having a  $DCA < 3$  cm are refitted to pass through it and unless the 'refit' fails, they are marked primary tracks. For my analysis I used good primary tracks. Note: primary tracks are *primary* from the detector's point of view, from the

p+p collision point of view, they are tracks of *secondary* particles (primary particles are protons that collide).

## 4.2 Corrections to the Data

Minimum bias sample (events triggered by the BBC coincidence trigger) was used for this analysis. To better understand any measured quantity and to be able to compare to other experiments it's necessary to correct the data for detector effects, i.e. to estimate, what an ideal detector would measure.

I have used a simulation, made by my STAR colleagues Eric Hjort and Jon Gans. Particles from Hijing [22] generated p+p events were filtered through a GEANT model of the STAR detector and the electron drift and read-out pads were further simulated. The primary vertices of Hijing events were randomly generated to meet Gaussian shape in  $z$  coordinate with  $\sigma \approx 40$  cm, which is shown in the left panel of figure 4.1. Details on Hijing setting can be found in appendix D.

These events were embedded into zerobias<sup>1</sup> events to simulate background. Afterwards the same reconstruction software was used, that is used for the data reconstruction.

### 4.2.1 Vertex finding efficiency

If the vertex finding algorithm (i.e.  $\chi^2$  minimization) fails, vertex is not found. Although such an event is still recorded for physics analysis, it has no primary tracks. The distribution of  $z$  position of generated and reconstructed vertices (*'VertexZ'*) is plotted in the left panel of figure 4.1. As the right panel of figure 4.1 shows, vertex finding efficiency is independent of its original position, so this doesn't constrain the *VertexZ* range we'll be able to use.

Vertex finding efficiency depends on the number of CTB-matched good global tracks (*'goodglCTB'*), that were used to find it. This information can be obtained directly from the data, because events without found vertex have global tracks too. We denote the vertex finding efficiency  *vtx\_found*. See the left panel of figure 4.2. To correct for events with lost vertices, each event is weighed by  $1/vtx\_found[goodglCTB]$ . Within statistical error the weight is 100% for  $goodglCTB > 11$ .

The right panel of figure 4.2 shows distribution of *goodglCTB*. The average vertex finding efficiency in the data is  $86.8 \pm 0.1$  %, close to the MC value. The discrepancy is caused by slightly different multiplicity distributions (to be discussed in the next chapter).

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<sup>1</sup>randomly fired trigger, not synchronized to beam crossing

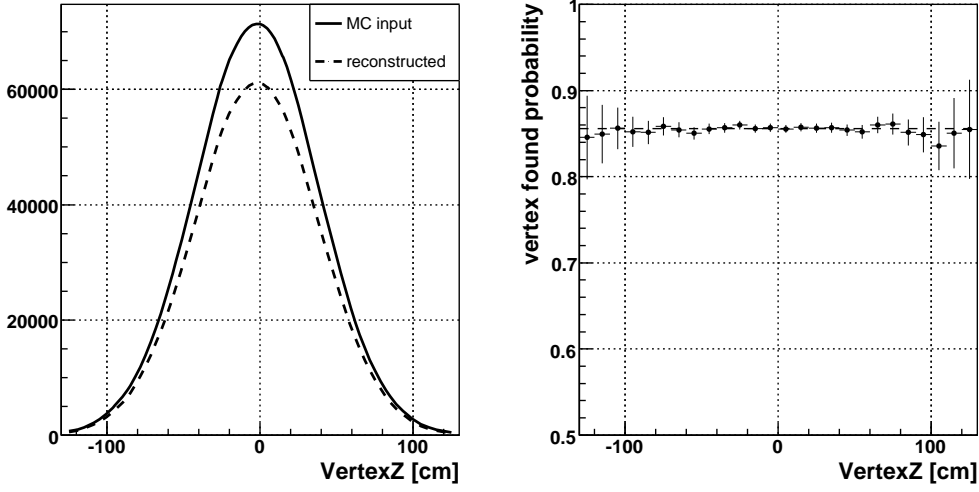


Figure 4.1: Left: Vertex  $z$  position from MC simulation and reconstructed by vertex finder. Right: vertex finding probability versus its  $z$  position. The dashed line shows the average  $85.6 \pm 0.4$  %.

## 4.2.2 Tracking efficiency

The efficiency of detecting given charged particle (its track) within the TPC acceptance depends mainly on its  $p_T$ . Given the magnetic field, TPC inner radius and the minimum of 15 hits required, the lowest measurable  $p_T \approx 0.1$  GeV/ $c$ , which is around what one would expect. Tracking is further affected with multiple scattering), causing the efficiency to drop at low  $p_T$ . In the mid-rapidity ( $\eta \approx 0$ )  $p \approx p_T$ , so the  $\beta\gamma$  factor is approximately proportional to  $p_T$ . This is shown in equation 4.1, where  $\tau$  is particle's mean life time and  $l$  its mean path before it decays.

$$l = \beta\gamma \cdot c\tau \quad (4.1)$$

$$\beta\gamma = \frac{p}{m} \quad (4.2)$$

$$l \approx \frac{p_T}{m} \cdot c\tau \quad (4.3)$$

Therefore, as  $\langle p_T \rangle \approx 0.4$  GeV/ $c$ , a considerable number of weakly decaying particles (mainly  $K^\pm$ ) decays before they can be detected.

Single-particle tracking efficiency turns out to depend on event multiplicity, in the following  $N_{ch}$  is MC multiplicity in the acceptance of  $|\eta| < 0.5$  and  $p_T > 0.15$  GeV/ $c$ . The left panel of figure 4.3 shows an obvious rise



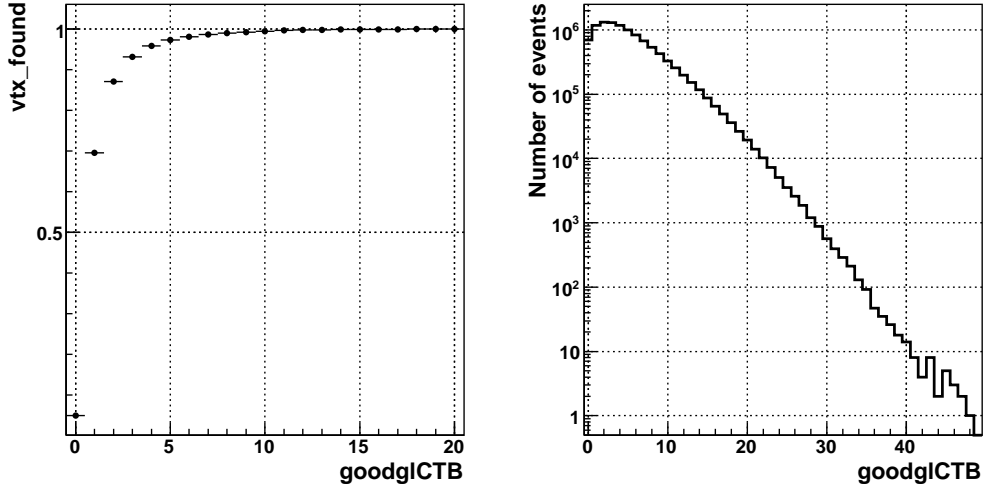


Figure 4.2: Left: vertex finding efficiency. Right: distribution of CTB-matched good global tracks.

of tracking efficiency with  $p_T$ . For  $p_T > 1$  GeV/ $c$  it's constant within statistical errors. We observe lower efficiency of tracking in events with low multiplicity, caused by the vertex finding efficiency (in a non-trivial manner) and possibly other effects, which are beyond the scope of this analysis. In the right panel of figure 4.3 the actual (MC generated)  $p_T$  distribution is shown. Given these two plots, the  $N_{ch}$  and  $p_T$  averaged tracking efficiency is 78 % for  $p_T > 0.1$  GeV/ $c$  and 86 % for  $p_T > 0.5$  GeV/ $c$ .

The dependence of tracking efficiency on particle's  $p_T$  changes the shape of the  $p_T$  spectrum, that is observed. Therefore, using this  $p_T$ -dependent tracking efficiency, the  $p_T$  spectrum can be corrected. Bin-to-bin correction was used to correct the  $p_T$  spectra. To correct the measured  $p_T$  spectrum, it's divided by a correction histogram (efficiency), similar to that in figure 4.3. Given the  $p_T$  resolution of the TPC and choosing reasonable binning, inter-bin migration is negligible, so this method can be used [36]. As the simulation used only has limited statistics, in the end we decided to use only  $p_T$  (not  $N_{ch}$ ) dependent correction factor.

Single-particle tracking efficiency depends also on its pseudorapidity and event (primary) vertex position. The left panel of figure 4.4 shows  $\eta$  dependence of tracking efficiency. In *VertexZ* we decided to use 25 cm-wide bins. Displacement of primary vertex in  $z$  direction causes a sort of shift of the plot in  $\eta$ . For this plot, only tracks with  $p_T > 0.15$  GeV/ $c$  were included (at both MC input and reconstructed levels).

Since the  $\eta$  dependence is approximately flat, we decided to average all

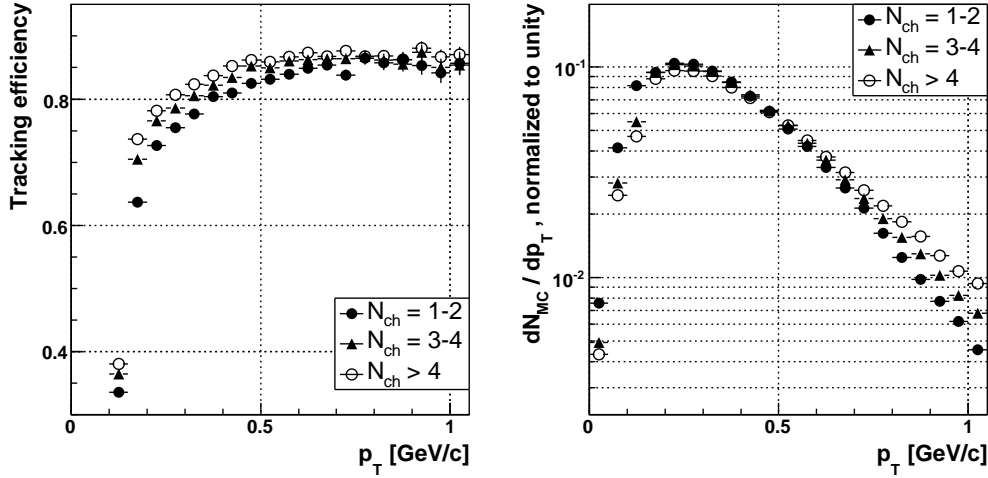


Figure 4.3: Left: single-particle tracking efficiency dependence on  $p_T$  and  $N_{ch}$ ,  $|\eta| < 0.5$ ,  $|VertexZ| < 25$  cm. Right: inclusive  $p_T$  spectra for charged primaries, taken from MC.

corrections over the  $\eta$  range used, so for  $p_T$  spectra we only bin in  $p_T$  and  $VertexZ$  and then use  $\eta$  averaged correction factor. This can only be done, if the  $p_T$  spectra don't depend very much on  $\eta$  range used. The ratios of  $p_T$  spectra for various  $\eta$  ranges are displayed in the right panel of figure 4.4. Independence of  $p_T$  spectra of  $\eta$  also means, that we can correct the pseudorapidity spectrum the same way as  $p_T$  spectrum - bin the correction histogram in  $\eta$  and  $VertexZ$  and average over  $p_T$ .

### 4.2.3 Charged multiplicity correction

In p+p collisions the multiplicities are small enough, that one should bin the distributions with single particle widths. Therefore, inter-bin migration is significant and one can't neglect adjacent bin correlations either. So a more complex correction method has to be used.

As before, I'm using 25 cm-wide bins in  $VertexZ$  and given the symmetry the absolute value is used. Details are presented in table 4.1.

We denote pseudorapidity cut used  $\eta_c$  and transverse momentum cut  $p_{Tc}$ . So  $|\eta| < \eta_c$  and  $p_T > p_{Tc}$ . These cuts we use at both MC and reconstructed level in simulation and also for the measured data. In the following examples,  $\eta_c = 0.5$  and  $p_{Tc} = 0.15$  GeV/c.

We denote  $N_{ch}$  or  $N$  the number of generated MC primary tracks (true multiplicity) and  $M$  number of detected (measured) good primary tracks

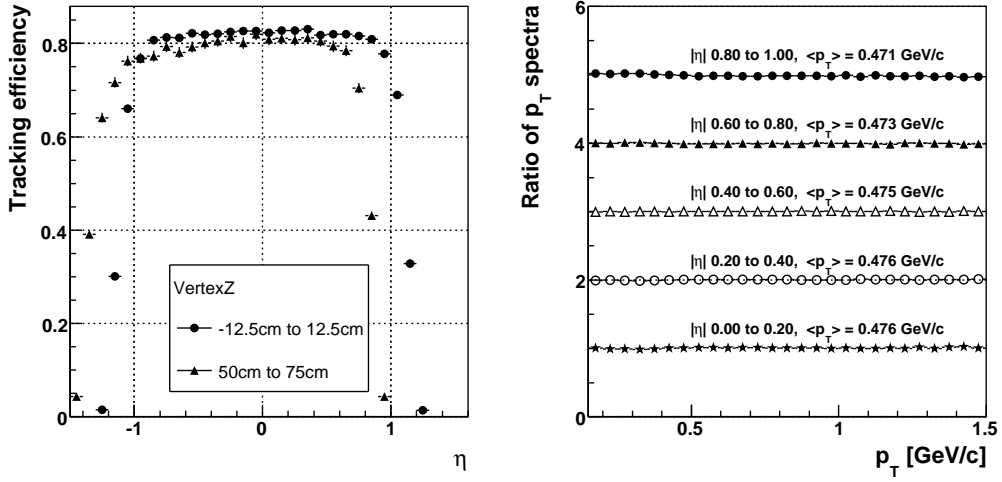


Figure 4.4: Left: Tracking efficiency dependence on  $\eta$ , for  $p_T > 0.15$  GeV/c. TPC can only detect good primary tracks with  $|\eta| < 1.3$ . Right:  $p_T$  spectra ratios - certain  $|\eta|$  range / full eta range ( $|\eta| < 1.0$ ), shifted successively by 0, 1, 2, 3, 4. For  $|VertexZ| < 25$  cm.

<i>VertexZ</i> bin number	$ z $ from	$ z $ to
1	0 cm	25 cm
2	25 cm	50 cm
3	50 cm	75 cm

Table 4.1: *VertexZ* bins

(observed multiplicity).  $M$  distributions are obtained for the three vertex bins as well as the correction factors, each of the three is corrected separately and the results are summed.

### Input for corrections

Using simulations, one must first obtain the conditional probability of detecting  $M$  particles, when  $N$  were generated (passed through the detector):  $P(M|N)$ . This is shown in the left panel of figure 4.5 by the means of a two-dimensional histogram filled with  $(N, M)$ . In the right panel of figure 4.5, the average tracking efficiency ( $\langle M \rangle / N$ ) is plotted as a function of  $N$ .

As the single-particle tracking efficiency depends strongly on particle's  $p_T$  (as shown in figure 4.3),  $P(M|N)$  is determined mainly by the  $p_T$  spectrum in events with multiplicity  $N$ . This may differ between simulation and reality

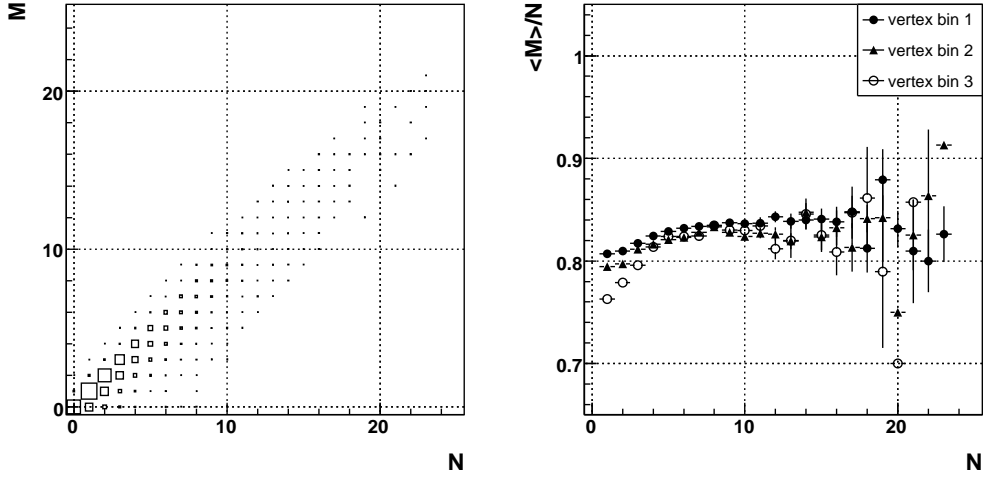


Figure 4.5: Left: Distribution of  $(N, M)$  for vertex bin 1. Right: Average tracking efficiency dependence on  $N$ . Both for  $\eta_c = 0.5$ ,  $p_{Tc} = 0.15$  GeV/ $c$ .

(experiment) , as will be shown in the following chapter. So we had to check, using  $p_T$  spectra and corrections, whether the average tracking efficiency for events with multiplicity  $N$  is the same in simulation and experiment.

The integral of the raw  $p_T$  spectrum gives average multiplicity times number of events. Therefore, the ratio of integrals of uncorrected and corrected spectra gives the efficiency. For each measured multiplicity  $M_{STAR}$  we obtain uncorrected  $p_T$  spectrum, which is then corrected using a bin-to-bin correction. The ratio of their intervals is the efficiency, i.e.  $\frac{M_{STAR}}{\langle N_{STAR} \rangle}$ . The same procedure is applied to  $p_T$  spectra from the simulation (MC), yielding  $\frac{M_{MC}}{\langle N_{MC} \rangle}$ . We found, that these two efficiencies equal within the statistical errors for all multiplicities. So, although the  $p_T$  spectra are different between the simulation and STAR data, this doesn't influence  $P(M|N)$ .

In the previous charged multiplicity study in STAR [26],  $P(M|N)$  was taken binomial, with the probability independent of  $N$ , which implies constant  $\langle M \rangle / N$ . As we can see (right panel of figure 4.5), this is not correct, especially for vertex bin 3. I also studied  $M$  distribution for fixed  $N$  and even this is not binomial ( $p_{binom} = \langle M \rangle / N$ ; tested by  $\chi^2$  test). One of the reasons is that it can be  $M > N$ , due to  $p_T$  and  $\eta$  resolution and some contamination. We decided to use  $P(M|N)$  obtained from simulation, as it's the best correction input available.

### Correction method

We would like to obtain the conditional probability  $P(N|M)$ , that would, applied to the distribution of  $M$ , yield the true multiplicity  $N$ . The simplest method is to use matrix inversion, when both (measured and true) multiplicity distributions are taken as vectors. Applying  $(P(M|N))^{-1}$  to measured distribution yields the true one. However, this method, used for example in [26], can't handle large statistical fluctuations. Moreover, one can get into trouble when  $P(M|N)$  is a singular matrix. As can be seen in [26], the results of this method are highly unstable.

Therefore, another method, "Unfolding method based on Bayes' theorem" [36], was used. The method consists in the following: let us have several true multiplicities  $(N_i, i = 1, 2, \dots, n_N)$  that can *cause* multiplicity  $M_j$  ( $j$  fixed) to be measured. Let us assume we know the *initial probability*  $P_0(N_i)$ . Then, the Bayes formula reads

$$P(N_i|M_j) = \frac{P(M_j|N_i) \cdot P_0(N_i)}{\sum_{l=1}^{n_N} P(M_j|N_l) \cdot P_0(N_l)} \quad (4.4)$$

The measured multiplicity  $M$  can have values  $M_j, j = 1, 2, \dots, n_M$ . These symbols  $(N_i, M_j, n_N, n_M)$  were chosen to keep my equations consistent with [36] - in the correction program itself,  $N_1$  means  $N = 0$ ,  $N_2$  means  $N = 1$  etc and the same for  $M$  too. Then,  $n_N$  and  $n_M$  have to be higher than the maximum multiplicity.

We recall that:

- $\sum_{i=1}^{n_N} P_0(N_i) = 1$  as usual
- $\sum_{i=1}^{n_N} P(N_i|M_j) = 1$ , which comes directly from equation 4.4
- $\sum_{j=1}^{n_M} P(M_j|N_i) = 1$ , meaning that each true multiplicity produces one measured multiplicity.

Using equation 4.4 and the definition of conditional probability, we get the final probability for  $N$ :

$$P(N_i) = \sum_{j=1}^{n_M} P(N_i|M_j) \cdot P(M_j) \quad (4.5)$$

If the initial distribution  $P_0(N)$  is not consistent with the data, it will not agree with the final distribution  $P(N)$ . The closer the initial distribution is to the true distribution, the better the agreement is. This suggest to proceed iteratively. First, the  $P_0(N)$  is chosen, I used a uniform distribution:  $P_0(N_i) = 1/n_N$ . Then, the process goes as:

1. calculate  $P(N_i|M_j)$  and  $P(N)$
2. make a  $\chi^2$  comparison between  $P(N)$  and  $P_0(N)$  (errors are estimated here by Poisson errors)
3. replace  $P_0(N)$  by  $P(N)$  and start again, until the value of  $\chi^2$  is small enough.

My experience is that this procedure converges in 5 to 10 cycles, using the  $\chi^2$  probability of 0.99 or 0.999999 to stop the iteration, without a significant difference ( $\pm 1$  iteration cycle). It's obvious from equation 4.5, that the correction is achieved by multiplying the measured distribution by an unfolding matrix:  $U_{ij} \equiv P(N_i|M_j)$ , defined by equation 4.4. Therefore, it's easy to estimate statistical errors on the resulting  $P(N)$  distribution, due to limited statistics of events in both simulation and data. Details can be found in [36].

In figure 4.6 the result of the correction are presented, together with measured (experimental) multiplicity distribution. Around  $N_{ch} = 17$  there is a strong fluctuation and within the errors, corrected distribution is not smooth as it should be. Therefore, something has to be added into the iterative algorithm to avoid this.

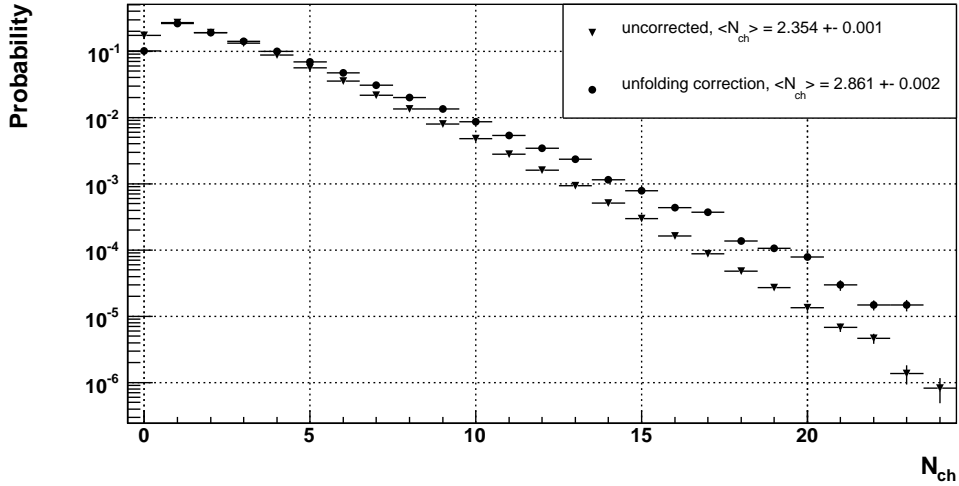


Figure 4.6: Uncorrected and corrected multiplicity distribution, without smoothing.  $\eta_c = 0.5$ ,  $p_{Tc} = 0.15$  GeV/c

One of the options is to fit the distribution in each iterative cycle. We tested Negative Binomial Distribution (NBD), which turned out to constrain the data too much (we don't know whether the resulting distribution must

follow the NBD). Next, a fit using a superposition of two independent NBDs was used, being more general. However, this sometimes prevents the iterative process from converging.

Finally, we decided to simply smooth the histogram (using `TH1::Smooth()` [34]), to avoid these problems. The resulting corrected distribution is in figure 4.7. Smoothing was applied in each cycle before calculating  $P(N_i|M_j)$  and  $P(N)$ , so it affects  $U_{ij}$  and not  $P(N)$  directly, so  $U_{ij}$  can be further used for example for error estimation.

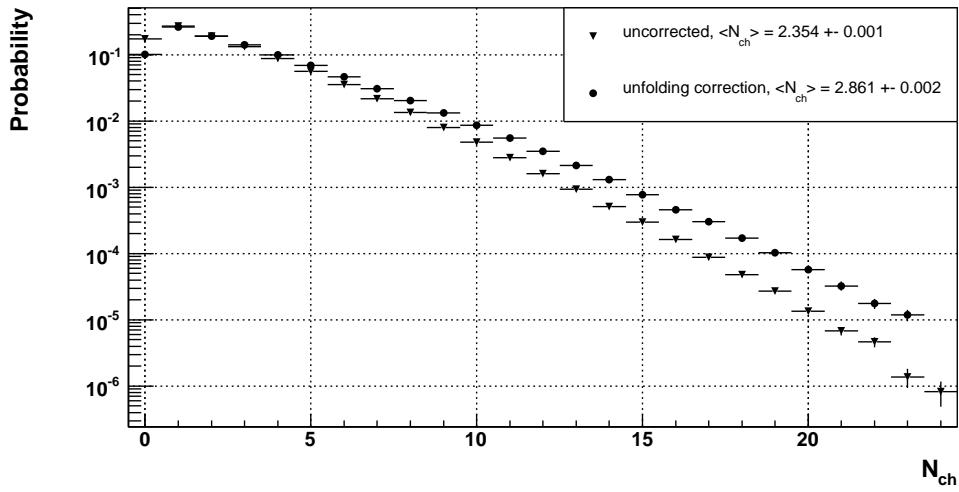


Figure 4.7: Uncorrected and corrected multiplicity distribution, with smoothing.  $\eta_c = 0.5$ ,  $p_{Tc} = 0.15$  GeV/ $c$

To test the whole unfolding procedure, it was applied to simulation. We corrected the reconstructed ( $M$ ) distribution and compared the result of correction to the original ( $N$ ) distribution. The two were identical (in terms of a  $\chi^2$  test), therefore we expect this procedure to work well on the data, too.

# Chapter 5

## Results

### 5.1 Data selection

I've used minimum bias (BBC coincidence trigger) p+p data from year 2001-2002. For efficiency studies and corrections, the already described Hi-jing+GEANT simulation (MC) was used. In table 5.1, corresponding numbers of events are shown. The *VertexZ* distribution in the data is approximately Gaussian with  $\sigma \approx 65$  cm.

Given the available statistics in simulation, we decided to use 3 vertex bins within 75 cm off the TPC center, as described in the previous chapter. We only use events with vertex found, and corrected for vertex finding efficiency, where necessary.

number of events	data	MC
all	10498 K	721 K
vertex found	9107 K	612 K
vertex found in $\pm 75$ cm	6350 K	556 K

Table 5.1: Number of events used for this analysis

We would like to divide the events into two classes, *soft* and *hard*. As was described in section 2.2, events with hard parton-parton scattering contain jets or mini-jets coming from parton fragmentations. To find these jets, a track clustering algorithm was used. A high- $p_T$  cluster ('mini-jet') is defined as one track with  $p_T > 0.7$  GeV/ $c$  with another track ( $p_T > 0.4$  GeV/ $c$ ) in a cone of radius  $R = 0.7$ . The cone radius is defined by means of pseudorapidity  $\eta$  and azimuth angle  $\phi$  as  $R = \sqrt{\Delta\eta^2 + \Delta\phi^2}$ . These two trigger tracks are good primary tracks, taken within full TPC acceptance (no cutting on  $|\eta|$ ).



*Hard* event is event containing at least one high- $p_T$  cluster, the remaining events are marked as *soft*.

A similar algorithm has been used by the CDF collaboration at Tevatron [37]. Here, the clusters have been searched for using a calorimeter with acceptance  $|\eta| < 2.4$ . In some limited phase space regions (calorimeter cracks) the above described track clustering algorithm has been used. I assume, that the parameters of calorimeter algorithm ( $E_T$  threshold,...) were chosen so, that the two algorithms are comparable.

To estimate the difference between the two approaches, I used the simulation. I made two different soft/hard selection criteria in it, both using track clustering algorithm. The difference is in what tracks are used. The first hard selection criterion, denoted as “TPC primaries”, uses reconstructed good primary tracks in the TPC acceptance (the same that is used for the data). The second uses Hijing generated charged tracks within  $|\eta| < 2.4$ .

In table 5.2 I present the fractions of soft and hard events in the minimum bias sample. The CDF results are presented for the two energies they’ve used, 630 corresponds to  $\sqrt{s} = 630$  GeV, 1800 to  $\sqrt{s} = 1800$  GeV. Based on the difference between the two triggers used for Hijing, STAR hard fraction would have been bigger, had the calorimeter algorithm been used. A clear rise of hard fraction with collision energy is observed, at  $\sqrt{s} = 200$  GeV (STAR), vast majority of events are soft events.

% of events	hard	soft
STAR	20	80
Hijing, TPC primaries hard sel.	13	87
Hijing, $ \eta  < 2.4$ hard sel.	28	72
CDF 630	43	57
CDF 1800	56	44

Table 5.2: Fraction of soft and hard events in STAR data, MC and CDF

## 5.2 Pseudorapidity spectra

A bin-to-bin correction was used, the effect of which can be seen in figure 4.4. Unlike corrections to  $p_T$  spectra and multiplicity, here we had to use 6 *VertexZ* bins, as the  $\eta$  spectra are obviously not symmetrical in  $\pm z$ .

We used  $p_{Tc} = 0.15$  GeV/ $c$  cut here, the same that was used in the examples in the previous chapter. As the left panel of figure 4.4 shows, in the outermost *VertexZ* bin, good primary tracks with  $p_T > 0.15$  GeV/ $c$  are

only detected for  $\eta < 1.0$ , and vice versa for the bin  $-75$  cm to  $-50$  cm. Therefore, the corrected  $\eta$  spectrum is only valid for  $|\eta| < 1.0$ .

In the left panel of figure 5.1 the corrected pseudorapidity spectrum (STAR) is compared to the simulation. Unlike STAR data, the simulation doesn't exhibit a central plateau. The ratio of the two spectra is shown in the right panel of figure 5.1, it's approximately flat in  $|\eta| < 0.8$ . As the two data samples (STAR and MC) have different mean multiplicities, the normalizations are different, so obviously the ratio can't be unity.

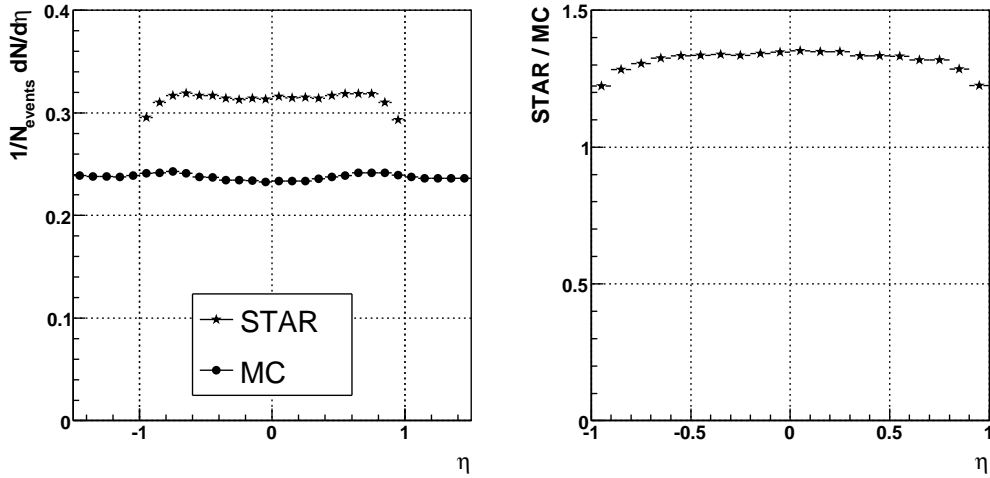


Figure 5.1: Left: Pseudorapidity spectra comparison,  $p_{Tc} = 0.15$  GeV/ $c$ . Right: Ratio of the two spectra.

No differences have been observed in the  $\eta$  spectra between the soft and hard data sample. We've observed no dependence on multiplicity, either. Obviously, the different collision dynamics between the soft and hard events, affects event  $p_T$  spectra and multiplicity, rather than pseudorapidity distribution.

### 5.3 Transverse momentum

As with the pseudorapidity spectra, we used a bin-to-bin correction, here with 3 *VertexZ* bins. Correction histograms were obtained separately for the two pseudorapidity cuts ( $\eta_c = 0.5, \eta_c = 1.0$ ), as the efficiencies are lower for higher  $|\eta|$ . All presented spectra are normalized to unity, so different mean multiplicities won't be seen here.

### 5.3.1 Comparison between different cuts

First, the  $p_T$  spectra for the three event samples are compared. We expect the  $p_T$  spectrum of hard events to contain a lot more high- $p_T$  tracks. In figure 5.2 the corrected  $p_T$  spectra from STAR are plotted for minimum bias, soft and hard event sample. The steep rise of hard spectra at  $p_T = 0.7$  GeV/ $c$  is caused by the trigger particle, defining the mini-jet. Starting at  $p_T \approx 0.7$  GeV/ $c$ , the ratio between hard and soft spectra is rising and saturating at about 3 GeV/ $c$ .

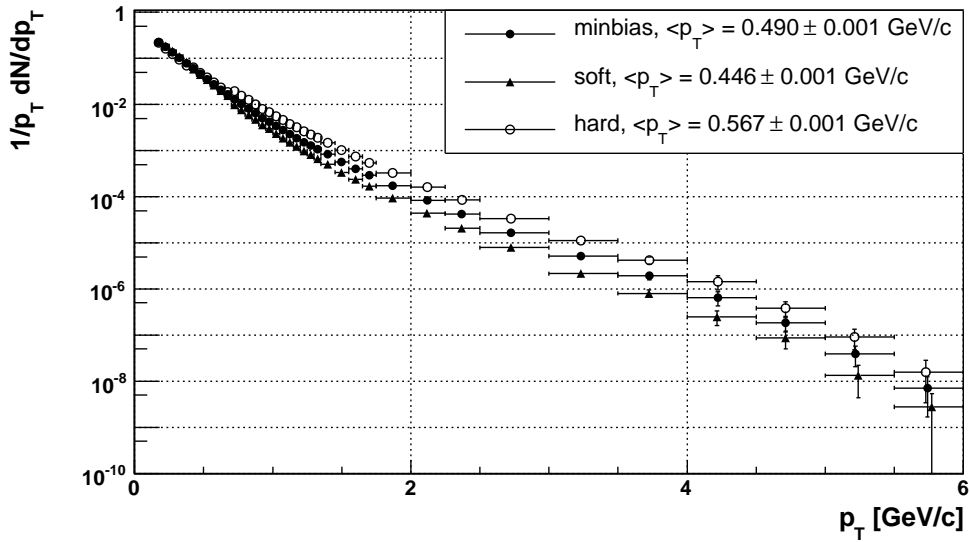


Figure 5.2: Corrected experimental  $p_T$  spectra for  $\eta_c = 0.5$

In figure 5.3 the ratios of  $p_T$  spectra (for  $\eta_c = 0.5$  / for  $\eta_c = 1.0$ ) are plotted. Except for a small difference at high  $p_T$ , the shapes of the spectra are independent of pseudorapidity cut used. As the error bars are mostly smaller than the markers, a  $\chi^2$  test was performed, showing good agreement between the two  $\eta_c$  used.

Transverse momentum spectra are known to get harder (higher  $\langle p_T \rangle$ ) at higher multiplicities, one of the reasons being obviously a rising fraction of hard events. Figure 5.4 shows  $p_T$  spectra for three multiplicity bins.

### 5.3.2 Fits of the $p_T$ spectra

We fitted the  $p_T$  spectra using the 'power-law' function (equation 2.8). Here,  $\eta_c = 0.5$  and inclusive spectrum for all multiplicities was used. A strong dependence of fit quality (in terms of  $\chi^2/ndf$ ) on fit range was observed.

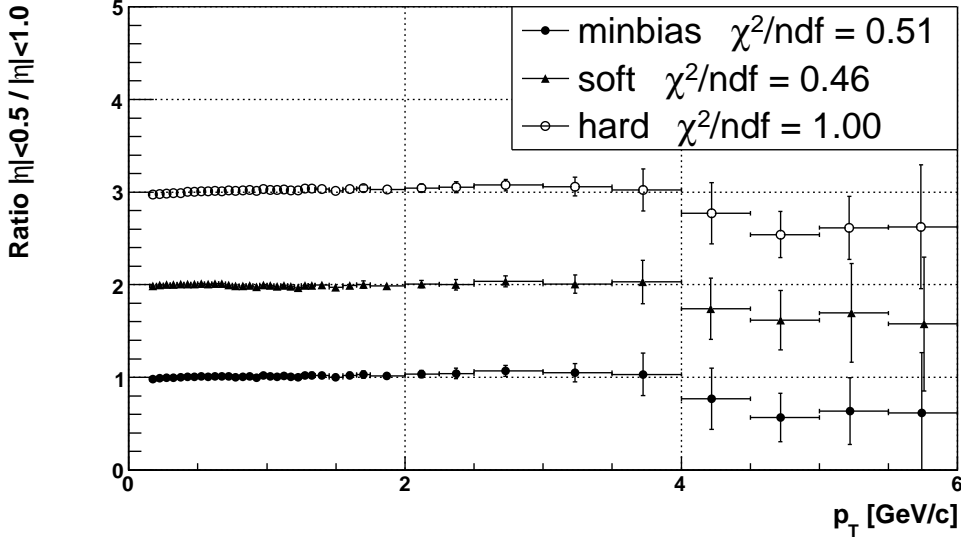


Figure 5.3: Ratios of  $p_T$  spectra, different pseudorapidity cuts.

First, soft and hard spectrum could only be fitted from  $p_T = 0.7$  GeV/ $c$  (to get reasonably low<sup>1</sup>  $\chi^2/ndf$ ), because of the structure at lower  $p_T$  caused by trigger particles (figure 5.2).

Fit results in range 0.7 – 6.0 GeV/ $c$  are shown in table 5.3. The  $\langle p_T \rangle$  is computed from extrapolating power-law function to zero and infinity - see equation 5.1. Note, that this value is a rough estimate - we don't know, how the spectra behave outside the fit range.

spectrum	$\chi^2/ndf$	$p_0$ [GeV/ $c$ ]	n	$\langle p_T \rangle$ [GeV/ $c$ ]
minbias	1.3	$1.85 \pm 0.07$	$11.8 \pm 0.3$	$0.42 \pm 0.02$
soft	1.3	$1.81 \pm 0.07$	$12.6 \pm 0.3$	$0.38 \pm 0.02$
hard	1.1	$2.16 \pm 0.09$	$12.2 \pm 0.3$	$0.47 \pm 0.02$

Table 5.3: Fits to power-law, fit range 0.7 – 6.0 GeV/ $c$ .

$$\langle p_T \rangle = \frac{\int_0^\infty p_T \frac{dN}{dp_T} dp_T}{\int_0^\infty \frac{dN}{dp_T} dp_T} = \frac{\int_0^\infty p_T^2 \left(1 + \frac{p_T}{p_0}\right)^{-n} dp_T}{\int_0^\infty p_T \left(1 + \frac{p_T}{p_0}\right)^{-n} dp_T} = \frac{2p_0}{n-3} \quad (5.1)$$

To see whether power-law describes the low  $p_T$  part of the spectrum, minimum bias spectrum was fitted, with varying fit range ( $p_{Tmin} - p_{Tmax}$ ).

<sup>1</sup> $\approx 1$ , fitting from 0.65 GeV/ $c$  gives  $\chi^2/ndf = 5$

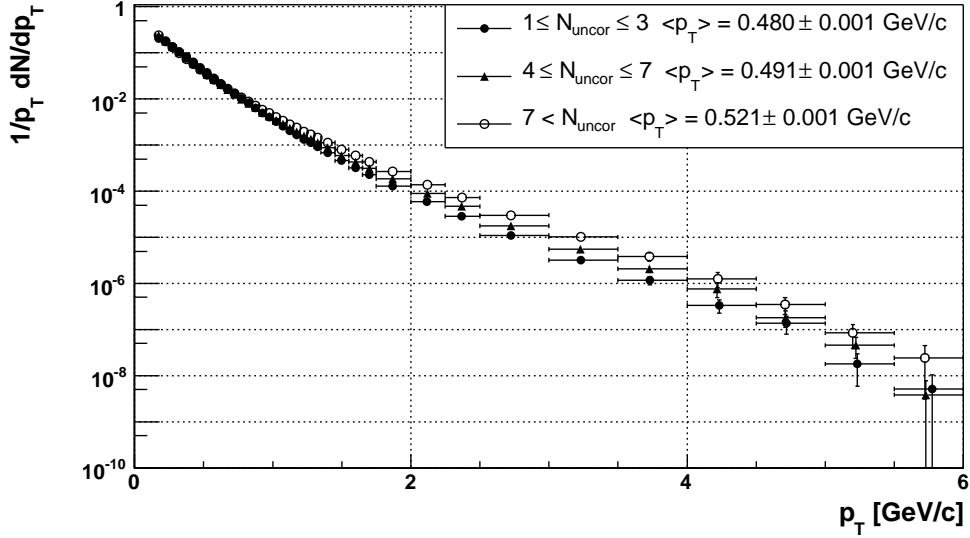


Figure 5.4: Transverse momentum spectra dependence on event multiplicity.  $N_{uncor}$  is uncorrected multiplicity,  $\eta_c = 0.5$ ,  $p_{Tc} = 0.15$  GeV/c.

The resulting fit parameters are shown in table 5.4. At lowest  $p_T$  power-law can't describe the data properly, and even for  $p_T > 0.5$  GeV/c fit parameters strongly depend on fit range used, therefore extrapolating the fit to lower  $p_T$  brings about a high systematic error.

For further analysis, we only use minimum bias  $p_T$  spectrum and the fit range 0.3 – 6.0 GeV/c was chosen, as an optimum to minimize both  $\chi^2/ndf$  and statistical errors of fit parameters. To see fit behaviour in more details, relative fit residuals  $(data - fit)/\sigma_{data}$  are plotted in figure 5.5. Between 1.5 and 4 GeV/c, the fit systematically underestimates the data and at highest  $p_T$ , power-law very likely breaks down. However, on the average power-law can describe STAR data well.

Next, we used fit ranges 0.3–6.0, 0.3–1.0 and 1.0–6.0 GeV/c and extrapolated to the measured  $p_T$  range (0.15 – 6.0 GeV/c). The ratios (data/fit) are plotted in figure 5.6. The fit parameters are shown at the bottom of table 5.4. It's remarkable, that the fit in range 0.3 – 1.0 GeV/c doesn't deviate substantially from the data even at highest  $p_T$ . In conclusion, power-law fit describes STAR data rather well.

$p_{Tmin}$ [GeV/c]	$p_{Tmax}$ [GeV/c]	$\chi^2/ndf$	$p_0$ [GeV/c]	n
0.15	6.0	5.8	$1.91 \pm 0.03$	$12.3 \pm 0.1$
0.2	6.0	2.4	$1.72 \pm 0.02$	$11.5 \pm 0.1$
0.3	6.0	1.6	$1.58 \pm 0.02$	$10.9 \pm 0.1$
0.4	6.0	1.5	$1.55 \pm 0.03$	$10.8 \pm 0.1$
0.5	6.0	1.5	$1.58 \pm 0.04$	$10.9 \pm 0.1$
0.7	6.0	1.3	$1.85 \pm 0.07$	$11.8 \pm 0.3$
0.9	6.0	1.0	$2.3 \pm 0.1$	$13.2 \pm 0.5$
0.3	1.0	1.2	$1.9 \pm 0.1$	$12.4 \pm 0.7$
1.0	6.0	0.9	$2.7 \pm 0.2$	$14.3 \pm 0.8$

Table 5.4: Fits to power-law, minimum bias.

### 5.3.3 Comparison to other experiments and to Hijing model

I will compare my results for minimum bias event sample to measurements in  $\bar{p}+p$  collisions at  $S\bar{p}pS$  and Tevatron colliders. Transverse momentum spectra were studied at  $\sqrt{s} = 200$  GeV by UA1 collaboration [38]. They measured  $p_T > 0.2$  GeV/c and used the same power-law functional form to fit it. The comparison can be found in table 5.5. Results by CDF collaboration at  $\sqrt{s} = 630$  and 1800 GeV, taken from [39] are also included in table 5.5. The acceptance of CDF detector's tracking system is  $p_T > 0.4$  GeV/c.

The pseudorapidity acceptance of UA1 (CDF) detector is  $|\eta| < 2.5$  ( $|\eta| < 1.0$ ), but as we could see from figure 5.3, different pseudorapidity cuts don't influence the shape of  $p_T$  spectra very much.

exp.	$\sqrt{s}$ [GeV]	$p_{Tmin}$ [GeV/c]	$\chi^2/ndf$	$p_0$ [GeV/c]	n	$\langle p_T \rangle$ [GeV/c]
STAR	200	0.2	2.4	$1.72 \pm 0.02$	$11.5 \pm 0.1$	$0.405 \pm 0.006$
UA1	200	0.2		$1.8 \pm 0.1$	$12.14 \pm 0.39$	$0.39 \pm 0.03$
STAR	200	0.4	1.5	$1.55 \pm 0.03$	$10.8 \pm 0.1$	$0.397 \pm 0.009$
CDF	630	0.4	0.71	$1.38 \pm 0.02$	$9.03 \pm 0.09$	$0.458 \pm 0.009$
CDF	1800	0.4	0.63	$1.24 \pm 0.01$	$8.09 \pm 0.04$	$0.487 \pm 0.005$

Table 5.5: Comparison of power-law fits to other experiments.  $\langle p_T \rangle$  according to equation 5.1

STAR data is in good agreement with UA1. Comparing to CDF results,

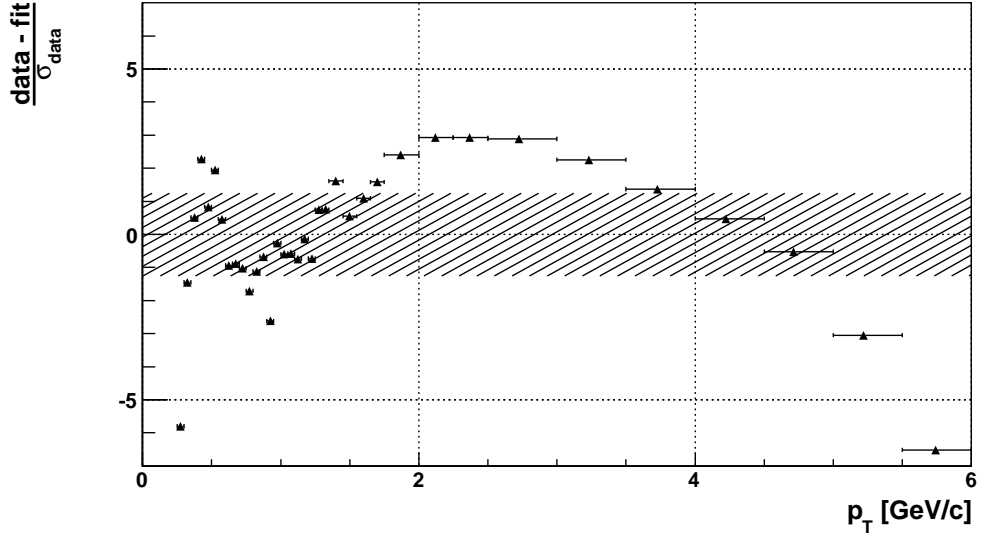


Figure 5.5: Relative fit residuals, fit range 0.3 – 6 GeV/ $c$ . The hatched band corresponds to  $\chi^2/ndf = 1.6$ .

there's a clear rise of  $\langle p_T \rangle$  with center-of-mass collision energy. On the other hand, the slope ( $n$ ) of the spectra at high  $p_T$  is decreasing, as well as  $p_0$  parameter. Unfortunately, there are no high-statistics p+p data at different collision energy from STAR available yet. Therefore a more detailed comparison is not possible.

Comparison to Hijing is shown in figure 5.7. The selection criterion for hard events in Hijing is “TPC primaries”. For the minimum bias and soft event sample, Hijing is clearly unable to describe STAR results properly. The agreement (except the lowest  $p_T$  bins) is relatively good for the hard sample, suggesting that mini-jet fragmentation in Hijing model is well tuned to the data.

We also tried fitting power-law to Hijing  $p_T$  spectrum, but for any reasonable fit range,  $\chi^2/ndf$  was always bigger than 2, therefore we conclude that Hijing cannot be described by power-law functional form at all.

## 5.4 Event charged multiplicity

In the following analysis good primary tracks in the following kinematic cuts:  $p_{Tc} = 0.15, 0.4$  GeV/ $c$  and  $\eta_c = 0.5, 1.0$  were used. To correct for detector efficiency the unfolding method was applied, which was described in the previous chapter. Again, the events were divided into 3 *VertexZ* bins. The

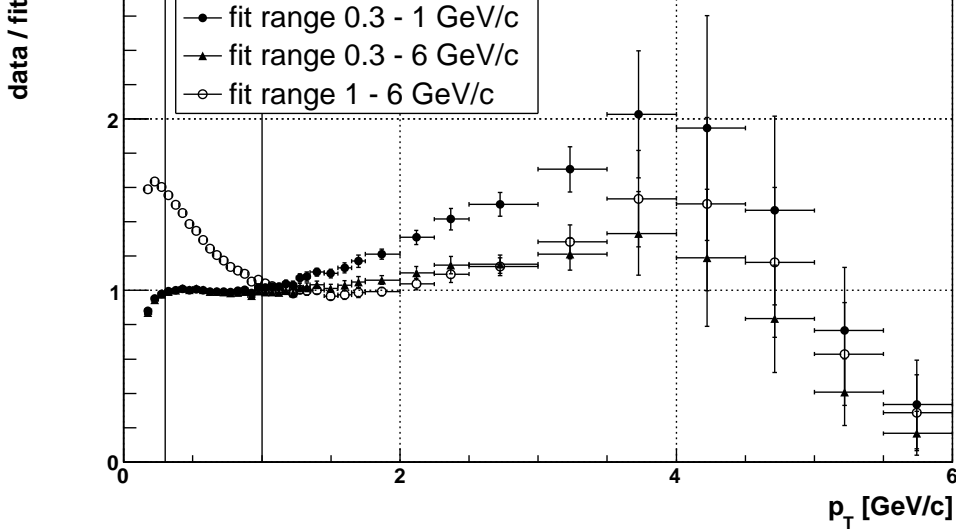


Figure 5.6: Data / fit for different fit ranges.

corrections differ between different cuts (single-particle tracking efficiencies depend on  $p_T$  and  $\eta$ ) and between event classes (Soft, Hard, MinBias) – the  $p_T$  spectra are different. In table 5.6 the averaged (over *VertexZ* bins and over all multiplicities) tracking efficiencies are shown. Therefore, for each kinematic cut and event class, correction factors were taken separately from simulation.

average tracking efficiency	$p_{Tc} = 0.15 \text{ GeV}/c$ $\eta_c = 0.5$	$p_{Tc} = 0.4 \text{ GeV}/c$ $\eta_c = 1.0$
Soft	0.81	0.81
Hard	0.86	0.88
MinBias	0.82	0.84

Table 5.6: Average tracking efficiencies for different kinematic cuts and event classes.

To compare to results by UA5 collaboration (to be shown later), we had to get corrected multiplicity without a  $p_T$  cut. As an input to unfolding procedure we used matrix  $P(M|N)$ , with  $N$  being number of particles from Hijing without a  $p_T$  cut and  $M$  number of reconstructed good primary tracks with  $p_T > 0.15 \text{ GeV}/c$ . We measured good primary tracks with  $p_T > 0.15 \text{ GeV}/c$ , and applying unfolding to  $P(M|N)$ , corrected multiplicity without a  $p_T$  cut was obtained.



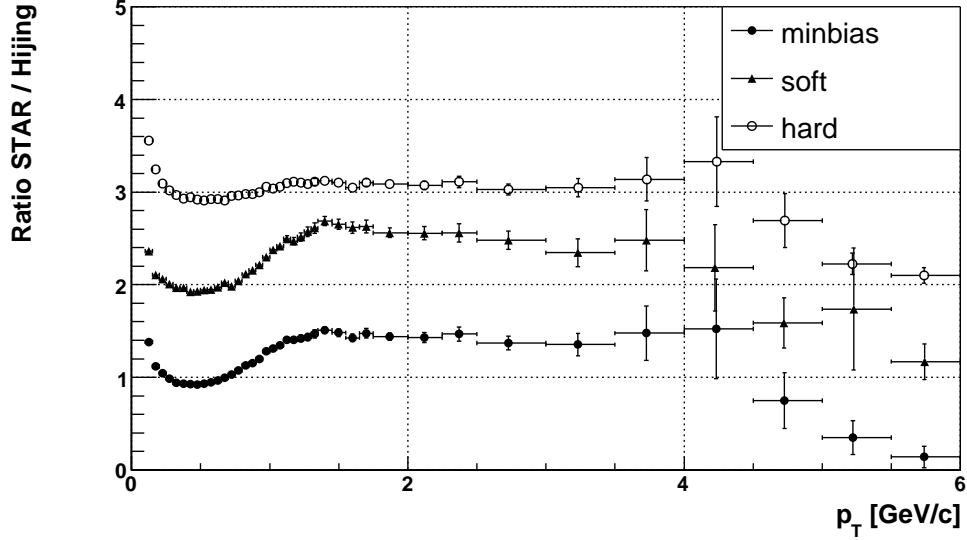


Figure 5.7: Comparison of  $p_T$  spectra to Monte Carlo model (Hijing),  $\eta_c = 0.5$ .

The output of correction is multiplicity for  $p_{Tc} = 0.0$  GeV/ $c$ , however it's only a rough estimate: there's no way to get the shape of the  $p_T$  spectrum from STAR down to  $p_T = 0.0$  GeV/ $c$ . It would be more precise, if we could use low- $p_T$  tracks measured by SVT detector (which is between the interaction point and the TPC), but SVT was not fully commissioned during the 2001-2002 running. Neither, is there any high-statistics data sample available at so called HalfField (magnetic field only 0.25 T, compared to standard 0.5 T), which would allow for low momentum tracking.

#### 5.4.1 General properties of multiplicity distributions

In the left panel of figure 5.8 the comparison between three event classes (minbias, soft, hard) is shown for  $p_{Tc} = 0.15$  GeV/ $c$ ,  $\eta_c = 0.5$ . The tail of minimum bias multiplicity distribution is dominated by hard events, whereas low multiplicity events are mostly soft. To see the actual shape of the distribution, regardless of its mean value, KNO variables are used in the right panel of figure 5.8. The soft events can achieve much higher values of  $z = \frac{N_{ch}}{\langle N_{ch} \rangle}$  than the hard ones.

In table 5.7 mean multiplicities ( $\langle N_{ch} \rangle$ ) are shown for different cuts and event classes. Their errors are of order of  $10^{-3}$ . There's about 10% rise of  $\langle N_{ch} \rangle$  from  $p_{Tc} = 0.0$  GeV/ $c$  to  $p_{Tc} = 0.15$  GeV/ $c$ , so the effect of using the above described "extrapolation" to  $p_{Tc} = 0.0$  GeV/ $c$  is not negligible.

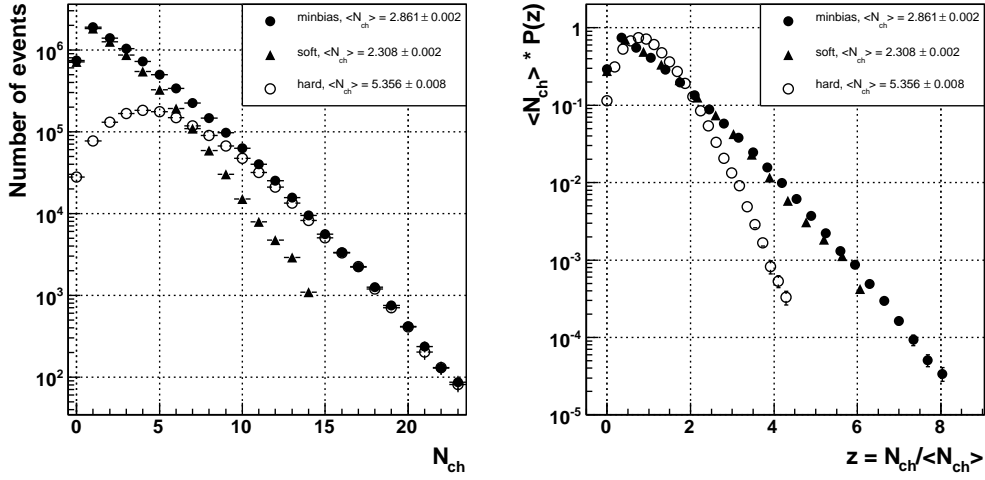


Figure 5.8: Comparison MinBias/Soft/Hard for  $p_{Tc} = 0.15$  GeV/ $c$ ,  $\eta_c = 0.5$

$p_{Tc}$ [GeV/ $c$ ]	$\eta_c$	Soft	Hard	MinBias
0.0	0.5	-	-	3.197
0.15	0.5	2.309	5.356	2.861
0.15	1.0	4.584	10.54	5.662
0.4	0.5	1.011	3.055	1.384
0.4	1.0	2.047	5.929	2.754

Table 5.7: Mean multiplicity for different kinematic cuts and event classes.

Comparison between kinematic cuts for minimum bias is displayed in figure 5.9. The abrupt fall at  $N_{ch} = 20$  for  $p_{Tc} = 0.0$  GeV/ $c$  is due to correction instability, showing that for this particular cut the correction is not working very well.

In figure 5.10 the same distributions are plotted in terms of the KNO variables. For low  $z$  there's an approximate scaling, which is broken at  $z > 3$ . Again we observe, that the higher  $\langle N_{ch} \rangle$ , the steeper is fall of  $\langle N_{ch} \rangle P(z)$  with  $z$ . This regularity has been reported also by the UA5 collaboration [40]. Quantitatively, this can be expressed with “relative dispersion”  $\frac{\sigma}{\langle N \rangle}$ . This gives (for cuts as in figure 5.10 from top to bottom): 1.10, 0.90, 0.87, 0.85, 0.76. Relative dispersion is related to  $k$  parameter of Negative Binomial Distribution (equation 2.5) (NBD):  $k^{-1} = \frac{\sigma^2 - \langle N \rangle}{\langle N \rangle^2}$ .

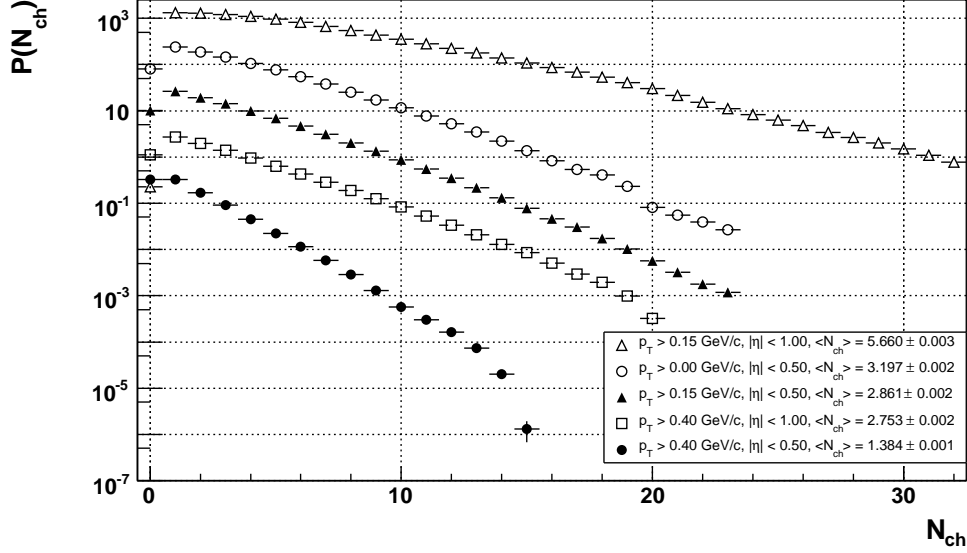


Figure 5.9: Comparison between kinematic cuts, minimum bias event sample. The distributions are scaled by  $1, 10, 10^2, 10^3, 10^4$ , from bottom to top.

### 5.4.2 Fits of the multiplicity distributions

For this part, we've chosen  $p_{Tc} = 0.15$  GeV/c,  $\eta_c = 0.5$  as it gives the best results, suggesting that the correction method works better than for other cuts. Negative binomial distribution was used for fitting. We've tried fitting this distribution for minimum bias event sample in the full range ( $N_{ch} = 0 - 23$ ). However, this fit gave  $\chi^2/ndf \approx 4500$ , so we decided to change the fit range ( $N_{min} - N_{max}$ ) - see the top of table 5.8.

$N_{min}$	$N_{max}$	$\chi^2/ndf$	$\langle N \rangle$	k
0	23	4529	$2.860 \pm 0.002$	$2.940 \pm 0.005$
1	23	79	$2.436 \pm 0.002$	$1.494 \pm 0.004$
2	23	13	$2.582 \pm 0.004$	$1.72 \pm 0.01$
3	23	11	$2.627 \pm 0.008$	$1.79 \pm 0.01$
4	23	7.0	$2.74 \pm 0.01$	$1.95 \pm 0.02$
2	11	1.7	$2.556 \pm 0.005$	$1.64 \pm 0.01$

Table 5.8: Fits to NBD for  $p_{Tc} = 0.15$  GeV/c,  $\eta_c = 0.5$ , minimum bias.

We've decided to use  $N_{min} = 2$  from now on. It seems that at  $N_{ch} = 0, 1$  NBD can't describe this data. Optimizing  $N_{max}$  to get better  $\chi^2/ndf$ , we

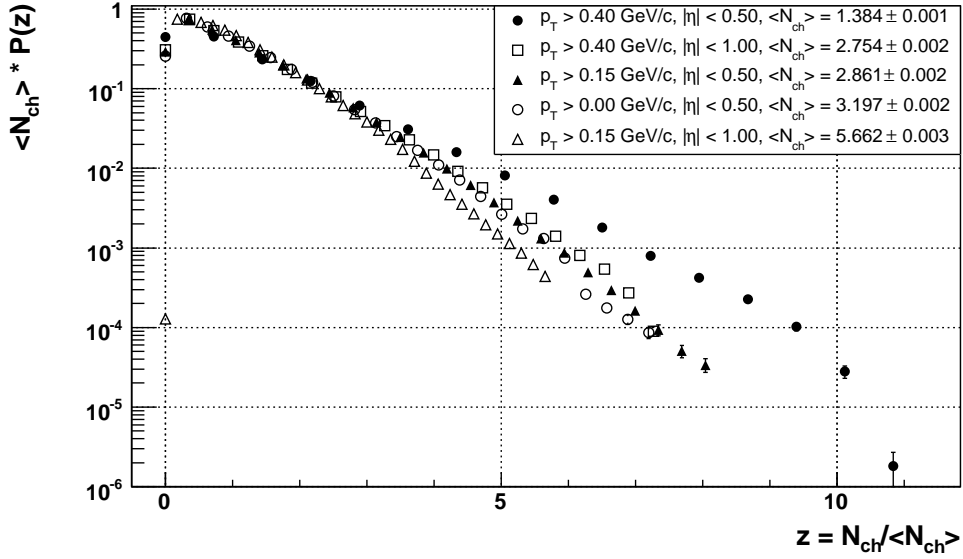


Figure 5.10: Comparison between kinematic cuts, minimum bias event sample, KNO variables.

found that the best fit range to use is  $N_{ch} = 2 - 11$ . The fit result are shown at the bottom of table 5.8. The resulting fit function is plotted in figure 5.11, extrapolated to the whole  $N_{ch}$  range (0–23). NBD overestimates the number of events with  $N_{ch} = 0$  by a factor of  $\approx 2.5$ . On the other hand, even this extrapolated function can approximately describe the tail of multiplicity distribution.

Let us note, that  $N_{ch} = 11$  in this cut corresponds on average to  $N_{ch} = 12$  for  $p_{Tc} = 0.0 \text{ GeV}/c$  (viz table 5.7), and this is the maximum multiplicity measured (and fitted) by the UA5 collaboration [41]. We’ll compare our results to UA5 later. It’s possible, that multiplicity distribution can have another component, which is significant at high multiplicities.

A straightforward generalization of NBD fit would be fit with two independent NBD functions (“double-NBD”):  $\alpha NBD(\langle N \rangle_1, k_1) + (1 - \alpha) NBD(\langle N \rangle_2, k_2)$ . One could expect the two components to coincide with soft and hard events, with  $NBD(\langle N \rangle_1 = \langle N \rangle_{soft}, k_1 = k_{soft})$  corresponding to the soft part.

Without constrains, this fit (still in the range 2 – 23) gives an excellent  $\chi^2/ndf = 1.13$ . While  $\alpha = 0.97$ , one would expect it to be about 0.8 (fraction of soft events - see table 5.2). Keeping  $\alpha = 0.8$  fixed and fitting only the 4 remaining parameters ( $\langle N \rangle_1, k_1, \langle N \rangle_2, k_2$ ) gives  $\chi^2/ndf = 1.18$ , which is only slightly worse.

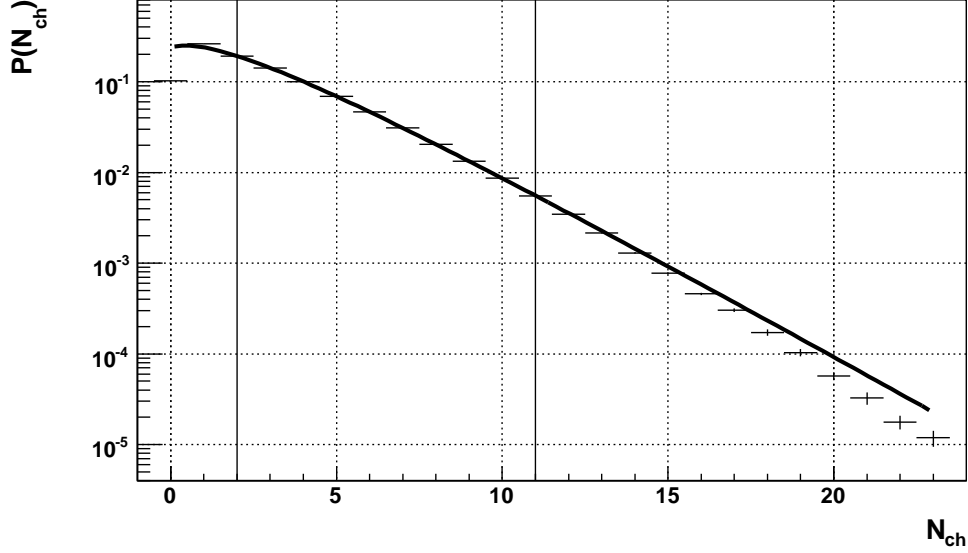


Figure 5.11: NBD fit to multiplicity distribution, minbias,  $p_{Tc} = 0.15 \text{ GeV}/c$ ,  $\eta_c = 0.5$ . Fit range 2–11.

Another option is to fit NBD to multiplicity distributions separately for soft and hard events. Again, we fitted starting from  $N_{ch} = 2$  and going to the maximum  $N_{ch}$  (23 for hard, 14 for soft). The results are shown in table 5.9.

data sample	$N_{min}$	$N_{max}$	$\chi^2/ndf$	$\langle N \rangle$	k
soft	2	14	25	$2.117 \pm 0.004$	$2.45 \pm 0.02$
hard	2	23	5.6	$5.355 \pm 0.005$	$6.48 \pm 0.04$

Table 5.9: Fits to NBD for  $p_{Tc} = 0.15 \text{ GeV}/c$ ,  $\eta_c = 0.5$ .

The fits are not good (in terms of  $\chi^2/ndf$ ), therefore our hypothesis is not completely correct, one can't interpret results of this fit as anything more than a rough approximation. Instead, we took the two  $\langle N \rangle$  parameters from these fits and put them into a double-NBD function as constraints. The resulting  $\chi^2/ndf = 1.7$  is remarkable, although it's worse than for previously performed double-NBD fits. Moreover,  $\alpha = 0.84$  is not far from the fraction of soft events (80 %). The result of this fit is plotted in figure 5.12. See table 5.10 for the summary of double-NBD fits.

Parameters  $k_1$  and  $k_2$  of the third fit are in a good agreement with those of NBD fits to soft and hard (shown in table 5.9). Therefore we conclude, that minimum bias  $N_{ch}$  distribution *can* be divided into two parts, described by

fit	$\chi^2/ndf$	$\langle N \rangle_1$	$k_1$	$\langle N \rangle_2$	$k_2$	$\alpha$
1	1.13	$2.45 \pm 0.07$	$1.83 \pm 0.06$	$6.5 \pm 0.6$	$12 \pm 7$	$0.97 \pm 0.02$
2	1.18	$2.043 \pm 0.008$	$2.25 \pm 0.09$	$4.98 \pm 0.06$	$5.4 \pm 0.2$	0.80 fixed
3	1.76	2.117 fixed	$2.35 \pm 0.06$	5.355 fixed	$6.50 \pm 0.06$	$0.836 \pm 0.003$

Table 5.10: Fits to double-NBD for  $p_{Tc} = 0.15$  GeV/ $c$ ,  $\eta_c = 0.5$ . For minimum bias event sample.

NBD. These two parts are very similar to our selection of soft and hard events, the difference may be caused by hard event selection that we use. When we tried to fit soft and hard by NBD in  $p_{Tc} = 0.4$  GeV/ $c$ ,  $\eta_c = 1.0$ , the resulting  $\chi^2/ndf$  were even worse than those shown in table 5.9. Here,  $N_{ch}$  distribution almost always contains the “trigger” high- $p_T$  particles. Therefore, we suggest using a narrow window in pseudorapidity to avoid this bias.

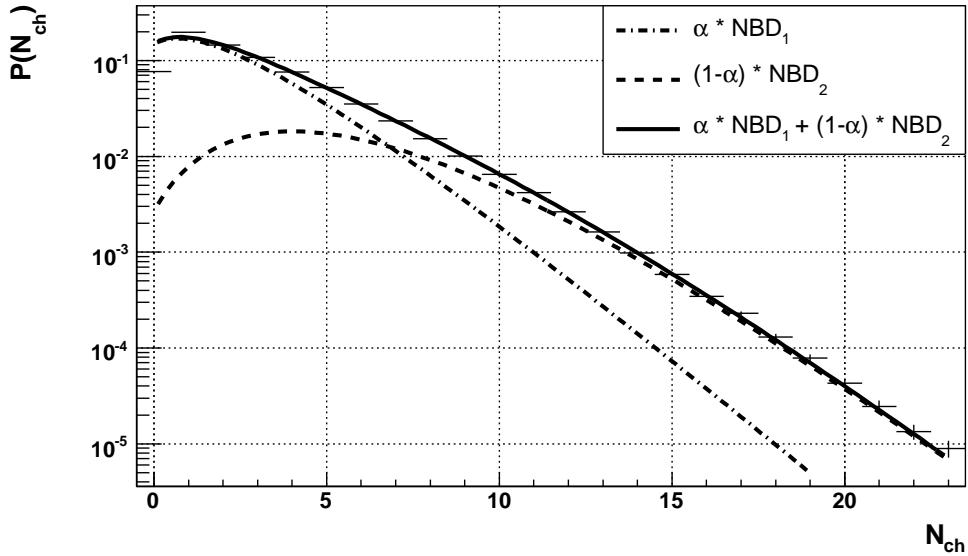


Figure 5.12: Double-NBD fit to multiplicity distribution, minbias,  $p_{Tc} = 0.15$  GeV/ $c$ ,  $\eta_c = 0.5$ .

To conclude, at low multiplicities ( $N_{ch} = 0, 1$ ) we found a strong disagreement between our results and theoretical expectations, based on NBD fit. The reason is unknown, although there is a possible explanation. This will be discussed in section 5.4.4. We found, that minimum bias multiplicity distribution for  $p_{Tc} = 0.15$  GeV/ $c$ ,  $\eta_c = 0.5$  is well described by double-NBD function, fitting in the range  $N_{ch} \geq 2$ .

### 5.4.3 Comparison to other experiments and to Hijing model

In the following, our multiplicity distribution at cut  $p_{Tc} = 0.0$  GeV/ $c$  and  $\eta_c = 0.5$  is compared to results by UA5 collaboration. These were taken at  $\sqrt{s} = 546$  GeV [40] and at  $\sqrt{s} = 200, 900$  GeV [41]. The main sub-detectors of UA5 experiment were two large streamer chambers, placed above and under the interaction point. As there was no magnetic field, they were able to measure even the lowest- $p_T$  particles with good accuracy.

In figure 5.13 the comparison is presented. Due to the already mentioned problem with  $N_{ch} = 0$  in the data (STAR), the average multiplicity for STAR is overestimated. Otherwise, we observe an obvious rise of  $\langle N_{ch} \rangle$  with center-of-mass energy. The slope of multiplicity distribution for UA5/200 GeV and Hijing is steeper than for STAR.

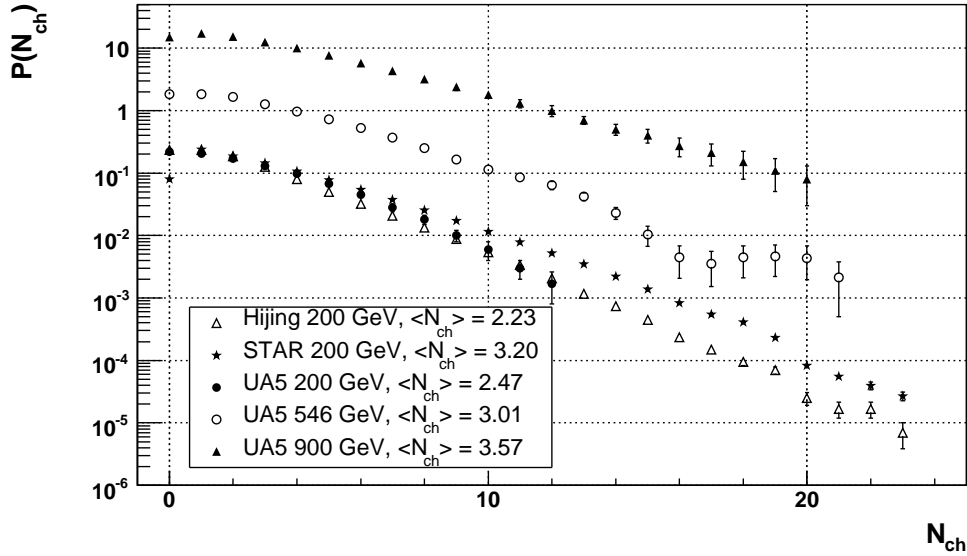


Figure 5.13: Comparison to Hijing and UA5. UA5 546 (900) GeV scaled by 10 (100).

To compare shapes of  $N_{ch}$  distributions, they're plotted in figure 5.14 in KNO form. To avoid problems with  $N_{ch} = 0$ , a truncated mean  $\langle N'_{ch} \rangle$  is used here: mean value is evaluated not in the range 0–maximum, but 1–maximum:

$$\langle N'_{ch} \rangle \equiv \frac{\sum_{N_{ch}=1}^{\infty} N_{ch} P(N_{ch})}{\sum_{N_{ch}=1}^{\infty} P(N_{ch})} \quad (5.2)$$

Values for  $N_{ch} = 0$  are not plotted, either. Also here  $\langle N'_{ch} \rangle$  of STAR is higher than that of UA5/200 GeV, caused by steeper slope of UA5/200 GeV described above. An approximate scaling holds at intermediate values of  $z$  (at lowest  $z$  the error bars are much smaller than the markers in the graph).

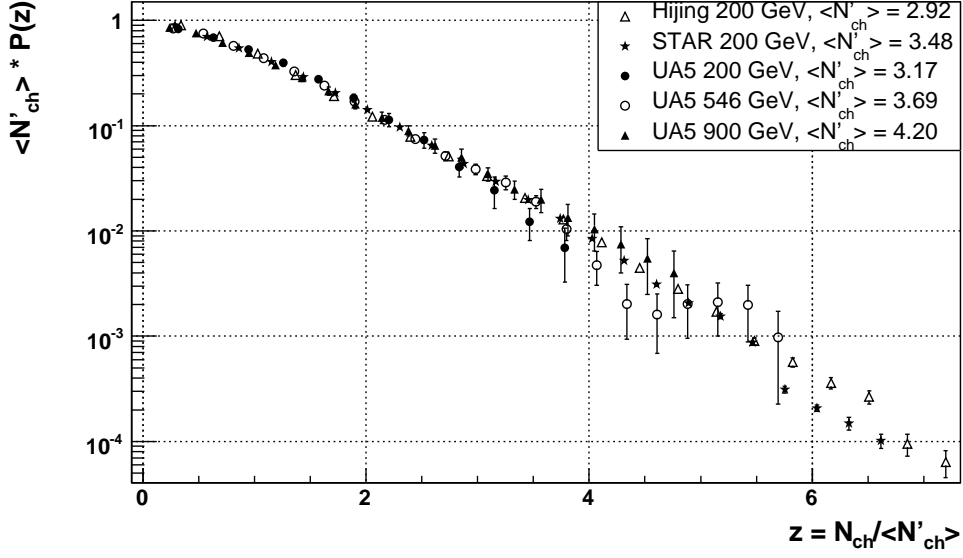


Figure 5.14: Comparison to Hijing and UA5 in KNO variables. Truncated means are used.

In the already mentioned analysis by CDF collaboration [37], the multiplicity distributions were studied in acceptance  $p_{Tc} = 0.4$  GeV/ $c$ ,  $\eta_c = 1.0$ . We compare our  $N_{ch}$  distribution in this acceptance to CDF results and to Hijing (using two hard events selection criteria, “TPC primaries” and “ $|\eta| < 2.4$ ” described in section 5.1). KNO form with truncated means is used.

In figure 5.15 one can see an approximate KNO scaling for minimum bias sample. The comparison for hard sample is displayed in figure 5.16. The discrepancy between STAR and CDF is caused by different hard events selection criteria - for STAR at  $\eta_c = 1.0$  the two high- $p_T$  tracks are almost always included in hard event multiplicity, whereas in the CDF algorithm ( $|\eta| < 2.4$ ) there are a lot more hard events with  $N_{ch} = 1$  (and  $N_{ch} = 0$ , not shown).

STAR results for hard event sample are in a good agreement with Hijing, “TPC primaries”. Similarly to the case of  $p_T$  spectra we conclude, that Hijing is able to describe STAR results well for the hard event sample. The difference in Hijing between the two hard selection criteria and the small



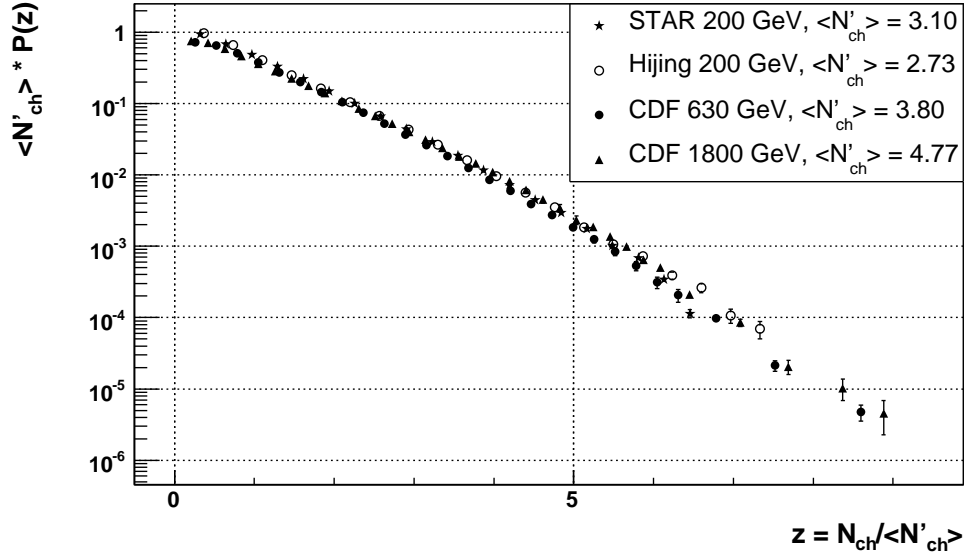


Figure 5.15: Minbias comparison to Hijing and CDF. Truncated means are used.

difference between Hijing, “ $|\eta| < 2.4$ ” and CDF 630 GeV suggests, that if there had been a calorimeter with acceptance  $|\eta| < 2.4$  in STAR experiment, the results for hard event sample at 200 GeV would be compatible with CDF measurements (i.e. weakly broken KNO scaling).

In figure 5.17 the same comparison is shown, but for the soft event sample. Different shape of STAR  $N_{ch}$  distribution is again due to different event selection criteria. KNO scaling holds between the two CDF energies, which was already mentioned in their publication [37]. Together with only a very weak energy dependence of  $\langle N_{ch} \rangle$  (or  $\langle N'_{ch} \rangle$ , respectively), this suggests that dynamics of soft interactions is almost energy-independent from RHIC to Tevatron energies.

#### 5.4.4 Low multiplicity events

For  $p_{Tc} = 0.0$  GeV/ $c$ ,  $\eta_c = 0.5$  we’ve observed too few events with  $N_{ch} = 0$  - figure 5.13. A similar behaviour ( $P(N_{ch} = 0) \ll P(N_{ch} = 1)$ ) can be seen for other kinematic cuts (figure 5.9). Also, for  $p_{Tc} = 0.15$  GeV/ $c$ ,  $\eta_c = 0.5$  Negative binomial distribution (figure 5.11) would predict much more events with  $N_{ch} = 0$  than our results show.

As both the corrections used (vertex finding efficiency correction and unfolding of the measured  $N_{ch}$  distribution) were successfully tested at Hijing

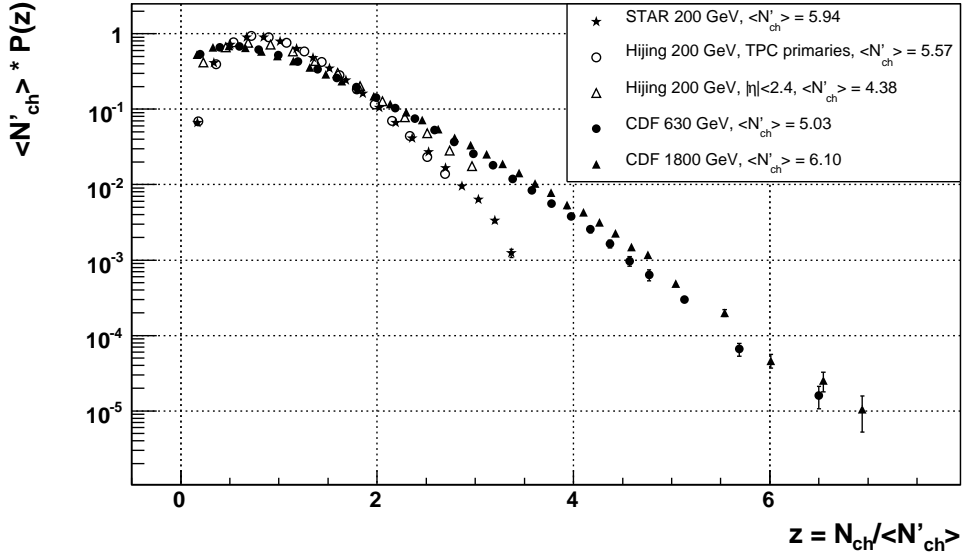


Figure 5.16: Hard comparison to Hijing and CDF. Truncated means are used.

simulation, we don't expect this effect to be an artifact of these corrections.

A possible explanation would be a lower sensitivity of STAR minimum bias trigger (used in 2001-2002 p+p running) to events with very low multiplicity in the TPC. This trigger employs 2 BBC detectors (described in Chapter 3) in coincidence, with acceptance  $3.3 < |\eta| < 5.0$ . This explanation would assume a long-range correlation between "multiplicity" (in a given kinematic cut, here we'll use  $p_{Tc} = 0.15$  GeV/c,  $\eta_c = 0.5$ ) and number of tracks at forward rapidities (in BBC acceptance).

Unfortunately, FTPC detectors, which could help estimate this forward multiplicity (their acceptance is  $2.7 < |\eta| < 3.9$ ) were not fully commissioned during the 2001-2002 p+p running. Therefore, Hijing simulation was used. Here, however, only tracks with  $|\eta| < 3.0$  are recorded, but it can serve as an approximation<sup>2</sup>. In figure 5.18 mean number  $\langle N_{forward} \rangle$  of charged primary tracks from Hijing with  $2.4 < |\eta| < 3.0$  is plotted versus Hijing multiplicity ( $p_{Tc} = 0.15$  GeV/c,  $\eta_c = 0.5$ ).

There is a strong correlation, so there could be substantial fraction of low multiplicity events have with no charged particles in BBC acceptance (i.e. in both detectors - for coincidence, to fire a minimum bias trigger), causing these events not to be recorded. The overall BBC coincidence trigger efficiency to detect a Non-Singly-Diffractive event is estimated 86% [26].

<sup>2</sup>if there is a correlation in multiplicity between TPC acceptance and BBC acceptance, there is likely to be a correlation between TPC acceptance and  $2.4 < |\eta| < 3.0$

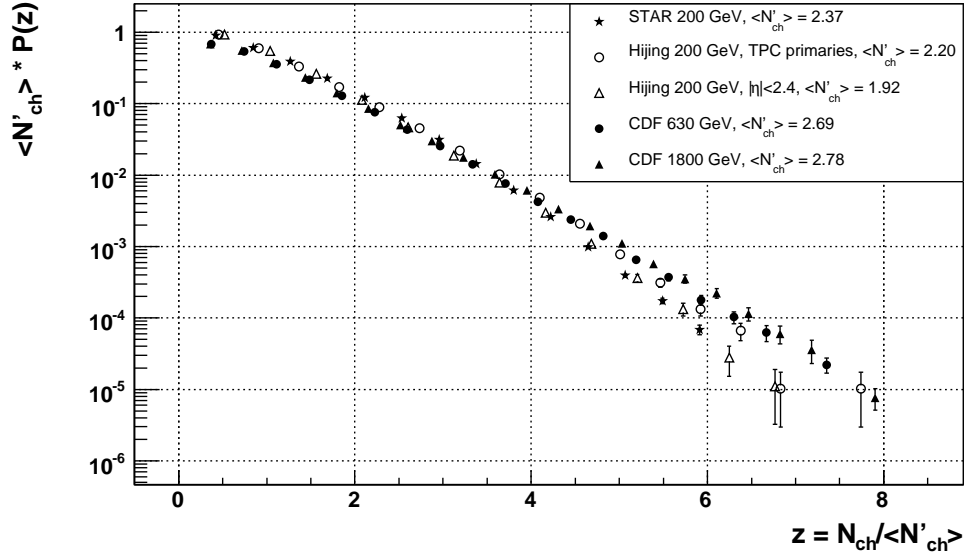


Figure 5.17: Soft comparison to Hijing and CDF. Truncated means are used.

In order to reproduce the 86% efficiency, the following approximation was made using a threshold: events with  $N_{forward} < N_{forward,thr.}$  will not fire BBC coincidence trigger, and those with  $N_{forward} \geq N_{forward,thr.}$  will always fire it. To achieve ( $N_{forward}$ -averaged) efficiency of 86%,  $N_{forward,thr.}$  has to be 1 with a good precision (the fraction of events with  $N_{forward} = 0$  is 14% = 100% – 86%). The ratio of  $N_{ch}$  distribution for  $N_{forward} \geq 1$  to  $N_{ch}$  distribution for all  $N_{forward}$  gives the efficiency of BBC coincidence trigger, estimated in this approximation. This is plotted in figure 5.19.

The efficiency for  $N_{ch} = 0$  is much lower than for  $N_{ch} > 0$ , so this could at least qualitatively explain, why we observe so few events with  $N_{ch} = 0$ . However, more elaborate analysis of forward multiplicities in STAR with respect to BBC coincidence trigger efficiency would be needed to make this more than an approximation. In future a similar method could be used to correct multiplicity distributions.

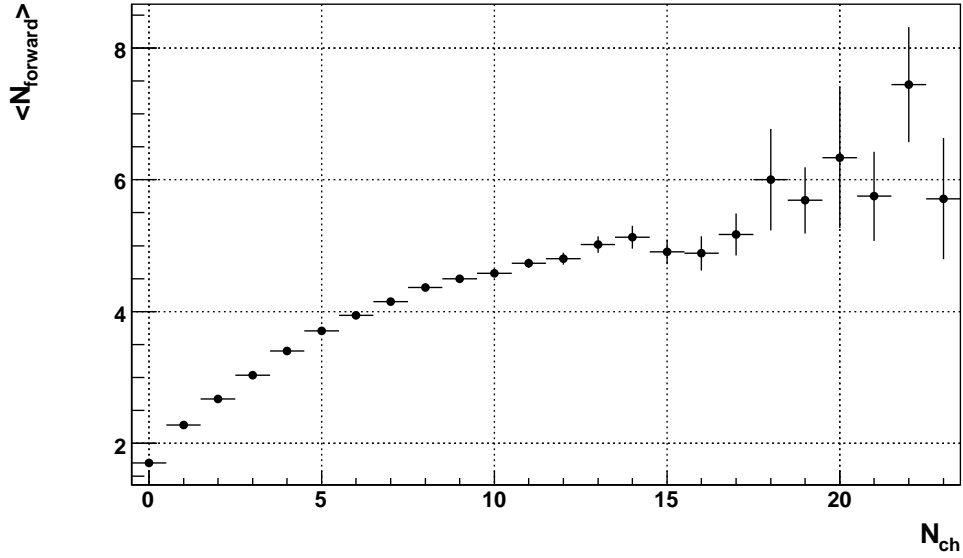


Figure 5.18: Hijing: mean number of tracks in  $2.4 < |\eta| < 3.0$  versus  $N_{ch}$  ( $p_{Tc} = 0.15$  GeV/ $c$ ,  $\eta_c = 0.5$ ).

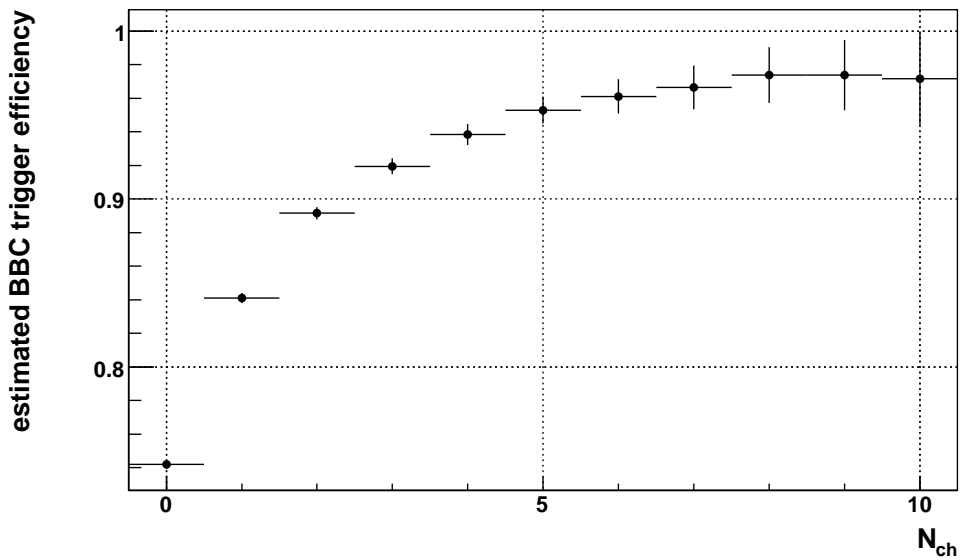


Figure 5.19: Estimated BBC coincidence trigger efficiency as a function of  $N_{ch}$  ( $p_{Tc} = 0.15$  GeV/ $c$ ,  $\eta_c = 0.5$ ). For  $N_{ch} > 10$  it's constant within the error.

# Chapter 6

## Conclusion

In this thesis, I have analysed 10.5 million minimum bias p+p events taken by the STAR experiment at RHIC collider in 2001-2002. The center-of-mass energy used (200 GeV) is currently the highest energy achieved in p+p collisions. A much higher energy will be delivered by the LHC p+p collider at CERN, currently under construction. Previous experiments at  $\text{S}\bar{\text{p}}\text{pS}$  collider (among them UA1, UA5) and at Tevatron (CDF, D0) studied proton-antiproton collisions.

My analysis was performed at RHIC Computing Facility, using standard STAR analysis framework, namely the FileCatalog, Scheduler and ROOT4STAR.

A GEANT model of STAR detector was used to estimate TPC tracking efficiency. Particles from Hijing simulated events were traced through this model and the output was embedded into background signal. Standard reconstruction algorithms were used to find tracks from hits produced in GEANT model. These tracks are compared to Hijing input, giving the efficiency.

The inclusive spectra and charged multiplicity distributions measured by STAR Time Projection Chamber were corrected for vertex finding efficiency and for tracking efficiency. The Central Trigger Barrel detector was used to reject tracks from pile-up events.

Bin-to-bin correction method was applied to pseudorapidity and transverse momentum spectra, as there's no significant inter-bin migration. To correct charged multiplicity distributions the unfolding method [36] based on Bayes' theorem was used.

For further analysis, events were divided into soft and hard classes, based on high- $p_{\text{T}}$  track clustering algorithm. Hard events contain at least one high- $p_{\text{T}}$  cluster, which may be due to underlying hard parton-parton scattering. Unlike pseudorapidity, transverse momentum spectra of charged particles

differ between soft and hard events.

For all considered event classes, transverse momentum spectra don't depend on pseudorapidity cut used. They are well described by "power-law" functional form, except for the lowest  $p_T$  (0.1–0.3 GeV/c), consistently with results by UA1 collaboration [38].  $p_T$  spectrum of hard events is in a good agreement with Hijing model, which fails to describe the  $p_T$  spectra from soft events.

Charged multiplicity distributions  $P(N_{ch})$  exhibit too few events with zero multiplicity, compared to Negative Binomial Distribution (NBD) fit prediction and to results by UA5 collaboration [40, 41]. This could be due to a lower sensitivity of STAR minimum bias trigger to events with zero multiplicity. This effect doesn't influence inclusive  $\eta$  and  $p_T$  distributions, as low multiplicity events don't contribute to these substantially.

For  $N_{ch} \geq 2$  the multiplicity distribution from minimum bias events can be fitted by NBD, except for the high multiplicity tail ( $N_{ch} \gg \langle N_{ch} \rangle$ ). A composite fit using two NBDs can describe even this tail and the two components coincide approximately with multiplicity distributions of soft and hard event samples.

STAR multiplicity distribution of hard events is well described by Hijing model. Except for the discrepancy due to different hard event selection criteria, it's compatible with results by CDF collaboration [37]. The multiplicity distribution of soft events shows KNO scaling with energy, when compared to CDF results.

# Bibliography

- [1] M. Harrison, T. Ludlam and S. Ozaki, *RHIC project overview*, Nucl. Instrum. Meth. A **499** (2003) 235.
- [2] P. Jacobs and X. N. Wang, *Matter in extremis: Ultrarelativistic nuclear collisions at RHIC*, Prog. Part. Nucl. Phys. **54** (2005) 443 [arXiv:hep-ph/0405125].
- [3] J. Letessier and J. Rafelski, *Hadrons and quark - gluon plasma*, Cambridge Monogr. Part. Phys. Nucl. Phys. Cosmol. **18** (2002) 1.
- [4] P. B. D. Collins, *Hadron Interactions*, Adam Hilger Ltd., 1984
- [5] J. Chýla, *Quarks, partons and Quantum Chromodynamics*, Lecture notes, available at <http://www-hep2.fzu.cz/~chyla/talks/mytalks/lecture.pdf>.
- [6] M. Gell-Mann, *Symmetries of baryons and mesons*, Phys. Rev. **125** (1962) 1067.
- [7] Y. Ne'eman, *Derivation of strong interactions from a gauge invariance*, Nucl. Phys. **26** (1961) 222.
- [8] M. Gell-Mann, *A Schematic Model Of Baryons And Mesons* Phys. Lett. **8** (1964) 214.
- [9] G. Zweig, *An SU3 model for strong interaction symmetry and its breaking*, CERN preprint CERN-TH-401.
- [10] Y. Nambu, in *Preludes in theoretical physics*, eds. A. de-Shalit, H. Feshbach and L. van Hove, North Holland, Amsterdam 1966, p. 133
- [11] C. Y. Wong, *Introduction to High-Energy Heavy-Ion Collisions*, World Scientific Publishing Co. Pte. Ltd. (1994).

- 
- [12] D. J. Gross and F. Wilczek, *Ultraviolet behavior of non-abelian gauge theories*, Phys. Rev. Lett. **30** (1973) 1343.
- [13] H. D. Politzer, *Reliable perturbative results for strong interactions*, Phys. Rev. Lett. **30** (1973) 1346.
- [14] M. Banner *et al.* [UA2 Collaboration], *Observation Of Very Large Transverse Momentum Jets At The Cern Anti-P P Collider*, Phys. Lett. B **118** (1982) 203.
- [15] W. Thome *et al.* [Aachen-CERN-Heidelberg-Munich Collaboration], *Charged Particle Multiplicity Distributions In P P Collisions At Isr Energies*, Nucl. Phys. B **129** (1977) 365.
- [16] G. Arnison *et al.* [UA1 Collaboration], *Some Observations On The First Events Seen At The Cern Proton - Anti-Proton Collider*, Phys. Lett. B **107** (1981) 320 [Erratum-ibid. **109B** (1982) 510].
- [17] L. V. Gribov, E. M. Levin and M. G. Ryskin, *Semihard Processes In QCD*, Phys. Rept. **100** (1983) 1.
- [18] X. Artru, *String Model With Baryons: Topology; Classical Motion*, Nucl. Phys. B **85** (1975) 442.
- [19] B. Andersson, G. Gustafson and B. Soderberg, *A General Model For Jet Fragmentation*, Z. Phys. C **20** (1983) 317.
- [20] K. Werner, *Strings, Pomerons, and the VENUS model of hadronic interactions at ultrarelativistic energies*, Phys. Rept. **232** (1993) 87.
- [21] H. U. Bengtsson and T. Sjostrand, *The Lund model for hadronic processes: Pythia version 4.8*, Comput. Phys. Commun. **46** (1987) 43.
- [22] X. N. Wang and M. Gyulassy, *HIJING: A Monte Carlo model for multiple jet production in p p, p A and A A collisions*, Phys. Rev. D **44** (1991) 3501.
- [23] V. N. Gribov, *A reggeon diagram technique*, Sov. Phys. JETP **26** (1968) 414 [Zh. Eksp. Teor. Fiz. **53** (1967) 654].
- [24] H. J. Drescher, M. Hladik, S. Ostapchenko, T. Pierog and K. Werner, *Parton-based Gribov-Regge theory*, Phys. Rept. **350** (2001) 93 [arXiv:hep-ph/0007198].



- 
- [25] X. Artru and G. Mennessier, *String model and multiproduction*, Nucl. Phys. B **70** (1974) 93.
- [26] J. Gans, *Inclusive Charged Hadron Transverse Momentum Spectra at Center of Mass Energy of 200 GeV for  $p+p$  and  $d+Au$  Collisions at the Relativistic Heavy Ion Collider*, PhD. Thesis, Yale University, 2004; available at [http://www.star.bnl.gov/STAR/central/theses/2004/gans\\_jon.pdf](http://www.star.bnl.gov/STAR/central/theses/2004/gans_jon.pdf).
- [27] Z. Koba, H. B. Nielsen and P. Olesen, *Scaling Of Multiplicity Distributions In High-Energy Hadron Collisions*, Nucl. Phys. B **40** (1972) 317.
- [28] G. J. Alner *et al.* [UA5 Collaboration], *A New Empirical Regularity For Multiplicity Distributions In Place Of Kno Scaling*, Phys. Lett. B **160**, 199 (1985).
- [29] G. Arnison *et al.* [UA1 Collaboration], *Transverse Momentum Spectra For Charged Particles At The Cern Proton Anti-Proton Collider*, Phys. Lett. B **118** (1982) 167.
- [30] G. Pancheri and Y. Srivastava, *Jets In Minimum Bias Physics*, Phys. Lett. B **159** (1985) 69.
- [31] L. Van Hove, *Multiplicity dependence of  $p(t)$  spectrum as a possible signal for a phase transition in hadronic collisions*, Phys. Lett. B **118** (1982) 138.
- [32] K. H. Ackermann *et al.* [STAR Collaboration], *STAR detector overview*, Nucl. Instrum. Meth. A **499** (2003) 624.
- [33] M. Anderson *et al.*, *The STAR time projection chamber: A unique tool for studying high multiplicity events at RHIC*, Nucl. Instrum. Meth. A **499** (2003) 659, arXiv:nucl-ex/0301015.
- [34] R. Brun and F. Rademakers, *ROOT: An object oriented data analysis framework*, Nucl. Instrum. Meth. A **389** (1997) 81, available at <http://root.cern.ch>.
- [35] The ROOT4STAR project, available at <http://www.star.bnl.gov/STAR/comp/root/index2.html>.
- [36] G. D'Agostini, *A Multidimensional unfolding method based on Bayes' theorem*, Nucl. Instrum. Meth. A **362** (1995) 487.

- 
- [37] D. Acosta *et al.* [CDF Collaboration], *Soft and hard interactions in  $p\bar{p}$  collisions at  $\sqrt{s} = 1800\text{-GeV}$  and  $630\text{-GeV}$* , Phys. Rev. D **65** (2002) 072005.
- [38] C. Albajar *et al.* [UA1 Collaboration], *A study of the general characteristics of proton - anti-proton collisions at  $\sqrt{s} = 0.2\text{ TeV}$  to  $0.9\text{ TeV}$* , Nucl. Phys. B **335** (1990) 261.
- [39] N. Moggi, *Soft multiparticle production in  $p\bar{p}$  interactions at 1800 and 630 GeV*, PhD. Thesis, Universita Degli Studi di Pavia, available at [http://www-cdf.fnal.gov/thesis/cdf5329\\_soft\\_multiparticle.ps.gz](http://www-cdf.fnal.gov/thesis/cdf5329_soft_multiparticle.ps.gz).
- [40] G. J. Alner *et al.* [UA5 Collaboration], *An Investigation Of Multiplicity Distributions In Different Pseudorapidity Intervals In Anti-P P Reactions At A Cms Energy Of 540-GeV*, Phys. Lett. B **160** (1985) 193.
- [41] R. E. Ansorge *et al.* [UA5 Collaboration], *Charged particle multiplicity distributions at 200 GeV and 900 GeV center-of-mass energy*, Z. Phys. C **43** (1989) 357.
- [42] S. Aronson, *The future of RHIC*, Nucl. Phys. A **774** (2006) 369.
- [43] STAR FileCatalog, available at <http://www.star.bnl.gov/STAR/comp/sofi/FileCatalog/user.html>.

# Appendix A

## STAR and RHIC rates

Table A.1 shows the summary of RHIC operation from its start-up until year 2005.

Year	Run plan	$\sqrt{s_{NN}}$	Sample	Running period
2000	Au+Au	130 GeV	$20 \mu\text{b}^{-1}$	6 weeks
2001/ 2002	Au+Au	200 GeV	$260 \mu\text{b}^{-1}$	16 weeks
	p+p	200 GeV	$1.4 \text{pb}^{-1}$	5 weeks
	Au+Au	19 GeV	$0.4 \mu\text{b}^{-1}$	1 day
2003	d+Au	200 GeV	$74 \text{nb}^{-1}$	10 weeks
	p+p	200 GeV	$5 \text{pb}^{-1}$	6 weeks
2004	Au+Au	200 GeV	$3740 \mu\text{b}^{-1}$	12 weeks
	Au+Au	62 GeV	$67 \mu\text{b}^{-1}$	3 weeks
	p+p	200 GeV	$100 \text{pb}^{-1}$	7 weeks
2005	Cu+Cu	200 GeV	$42 \text{nb}^{-1}$	8 weeks
	Cu+Cu	62 GeV	$1.5 \text{nb}^{-1}$	12 days
	Cu+Cu	22 GeV	$18 \mu\text{b}^{-1}$	39 hours
	p+p	200 GeV	$30 \text{pb}^{-1}$	10 weeks
	p+p	410 GeV	$0.1 \text{pb}^{-1}$	1 day

Table A.1: Summary of RHIC runs 1 through 5 (2000-2005). Taken from [42].  $\sqrt{s_{NN}}$  is the center-of-mass energy per nucleon–nucleon pair.

For p+p almost all RHIC runs have been at center-of-mass energy of 200 GeV, the same energy at which most of heavy ion runs have been performed. Recently, there's a tendency to reach higher energies in p+p, in connection with RHIC spin program. For heavy ions and d+Au there are two main energies used, namely 200 GeV and 62 GeV.

Both raw data and reconstructed ROOT files for physics analysis (MuDst) are stored in RCF storage systems. STAR uses FileCatalog [43] software to keep information about all files and to enable queries. The criteria for searching the files for given physics analysis are: collision system (p+p, Au+Au etc), energy, year, reconstruction software version, triggers, detectors used in data taking etc.

Table A.2 shows total number of events from years 2000–2005, stored in MuDst files, as found by FileCatalog in August 2006. More than 95 % of these events were taken when the TPC was ON, except for p+p 200 GeV, where about a half of all events was taken only using the fast detectors (mainly calorimeters). In August 2006, the data from 2006 running wasn't fully reconstructed, it's not included in table A.2.

Collision system	$\sqrt{s_{NN}}$ [GeV]	total # of events (millions)
p+p	200	324
p+p	410	2
d+Au	200	82
Cu+Cu	22	3
Cu+Cu	62	40
Cu+Cu	200	93
Au+Au	19	0.3
Au+Au	62	17
Au+Au	130	1
Au+Au	200	119

Table A.2: Total number of events in MuDst files, as in August 2006.

# Appendix B

## Kinematics

In experiments with colliding beams, Lorentz boosts are most important along the beam axis ( $z$ ). Therefore, it's convenient to introduce *rapidity*  $y$  defined as:

$$y = \frac{1}{2} \ln \left( \frac{E + p_z}{E - p_z} \right) = \frac{1}{2} \ln \left( \frac{E + p \cos \theta}{E - p \cos \theta} \right) \quad (\text{B.1})$$

Rapidity is Lorentz additive with respect to boosts along  $z$  axis, in other words, interval in rapidity  $\Delta y$  is Lorentz-invariant. *Pseudorapidity*  $\eta$  is an approximation of rapidity for mass-less particles:

$$\eta = \lim_{m \rightarrow 0} y = -\ln \tan \frac{\theta}{2} \quad (\text{B.2})$$

and  $\eta \approx y$  in mid-rapidity ( $y \approx 0$ ).

Using rapidity, one can simplify equation for invariant cross section 2.3, shown here again:

$$\sigma_{\text{pp}}^{\text{C}} = E_{\text{C}} \frac{d^3 \sigma}{dp_{\text{C}}^3} \quad (\text{B.3})$$

$$dp^3 = dp_z p_T dp_T d\phi \quad (\text{B.4})$$

$$dp_z = \left( \frac{dp_z}{dy} \right) dy \quad (\text{B.5})$$

$$\frac{dy}{dp_z} = \frac{1}{2} \frac{E - p_z}{E + p_z} \left( \frac{(E - p_z) \left(1 + \frac{dE}{dp_z}\right) - (E + p_z) \left(-1 + \frac{dE}{dp_z}\right)}{(E - p_z)^2} \right) \quad (\text{B.6})$$

$$\frac{dE}{dp_z} = \frac{d \left( \sqrt{m^2 + p_T^2 + p_z^2} \right)}{dp_z} = \frac{1}{2} \left( \frac{1}{\sqrt{m^2 + p_T^2 + p_z^2}} \right) 2p_z = \frac{p_z}{E} \quad (\text{B.7})$$

$$\frac{dy}{dp_z} = \frac{1}{E} \quad (\text{B.8})$$

$$E \frac{d^3\sigma}{dp^3} = \frac{1}{p_T} \frac{d^3\sigma}{d\phi dy dp_T} \quad (\text{B.9})$$

where  $p_T$  is transverse momentum and  $m$  is the rest mass. For unpolarized beams, one has azimuthal symmetry, therefore  $\phi$  can be integrated out, giving  $2\pi$ .

# Appendix C

## Luminosity and pile-up

In STAR, the primary luminosity measure in p+p running are the BBC detectors, used in coincidence to trigger Non-Singly-Diffractive (NSD) minimum bias events. The corresponding cross section is  $\sigma_{NSD} = 31$  mb and the efficiency of BBC coincidence trigger is 86%. Therefore,  $\sigma_{BBC} \approx 27$  mb. Then, the luminosity can be obtained as:

$$\mathcal{L} = \frac{f_{BBC}}{\sigma_{BBC}} \quad (\text{C.1})$$

where  $f_{BBC}$  is the BBC coincidence rate. In 2001–2002 p+p running, the average value of BBC coincidence rate was 6 kHz, which corresponds to luminosity  $\mathcal{L} \approx 0.2 \cdot 10^{30} \text{ cm}^{-2}\text{s}^{-1}$ .

The beam crossing rate for p+p running at RHIC is  $f_{cross} = 10$  MHz. So the probability of given beam crossing producing an interaction triggered by BBC coincidence is:

$$p = \frac{f_{BBC}}{f_{cross}} \quad (\text{C.2})$$

During 2001–2002 p+p running this was on average  $p = 6 \text{ kHz}/10 \text{ MHz} \approx 0.06\%$ . Therefore, even at peak luminosity (which was about four times higher than the average luminosity), the probability to have more than one interaction in single beam crossing is negligible.

However, as the TPC readout time is  $40 \mu\text{s}$ , another interaction can happen before the TPC is read out. Actually, the data read out from the TPC can contain tracks from pile-up events happening in time window  $\pm 40 \mu\text{s}$  around the triggered event. Therefore, the number of these pile-up events in the TPC is approximately:

$$N_{pile-up} = 80 \mu\text{s} \cdot f_{BBC} \quad (\text{C.3})$$

which gives on average  $N_{pile-up} = 80 \mu s \cdot 6 \text{ kHz} \doteq 0.48$ . This cannot be neglected. Note, that this is for average luminosity – at peak luminosity there are tracks from two more events present in the TPC, in addition to the triggered event. Therefore, the Central Trigger Barrel (CTB) detector has to be used in primary vertex finding, which is described in section 4.1.2.



# Appendix D

## Hijing settings

Hijing version 1.328 has been used, to simulate proton collisions with center-of-mass energy of  $\sqrt{s} = 200$  GeV.

Other important parameters were set this way:

- ' Impact parameter min/max (fm) ' 0. 5.0
- ' Jet quenching (1=yes/0=no) ' 0
- ' Hard scattering/pt jet ' 0 -2.25
- ' Max # jets per nucleon (D=10) ' 10
- ' Set ihpr2(11) and ihpr2(12) ' 1 1
- ' Set ihpr2(21) and ihpr2(18) ' 1 0
- ' set B production ' 1.5