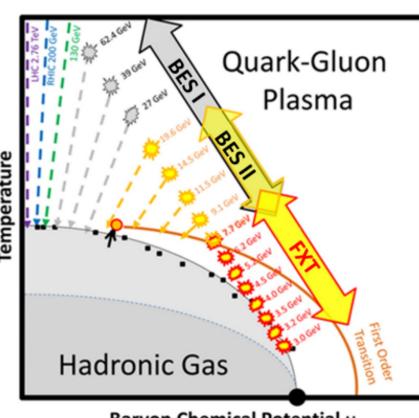
# **Charged Pion Production in** $\sqrt{s_{NN}}$ = 3.0 GeV Fixed-Target Collisions at STAR

# **Motivation**

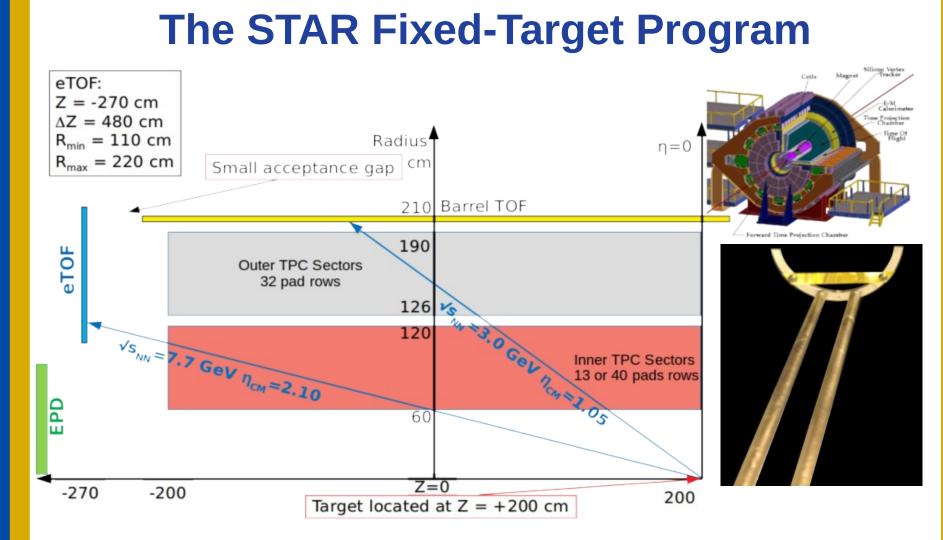
Charged particle spectra with be used can chemical equilibrium model to obtain thermal properties at chemical freeze-out<sup>[1]</sup>. Pions can provide information about the Coulomb potential from the positive interaction region<sup>[2]</sup>.

STAR



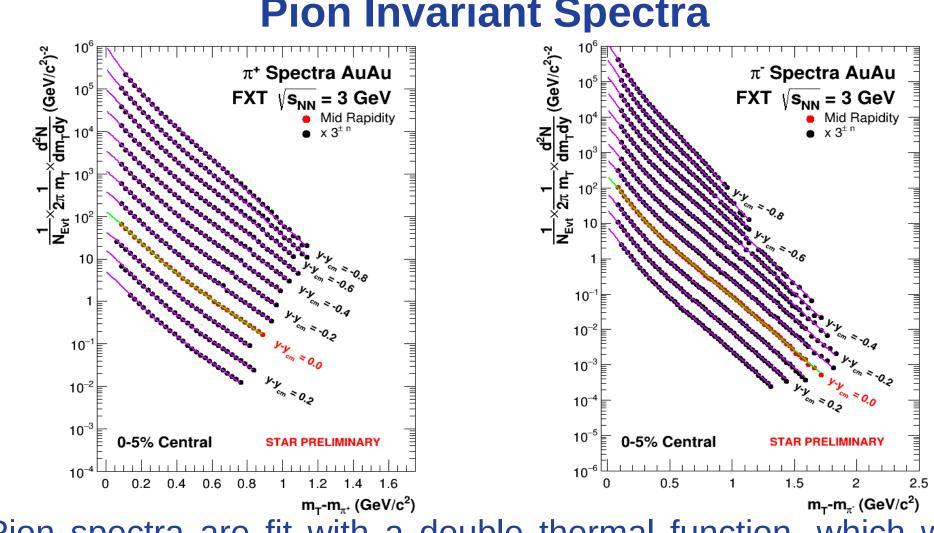
Baryon Chemical Potential µ<sub>p</sub>

<sup>[1]</sup>S. Wheaton et al., Comput. Phys. Commun. 180, 84 (2009) <sup>[2]</sup>D. Cebra et al., arXiv:1408.1369 [nucl-ex]



Target first installed in 2015 with successful test run. Target is 0.25 mm thick (1% interaction probability) and held 2 cm below the center of the beam axis. When running in Fixed Target (FXT) mode, 12 bunches are used in the collider.

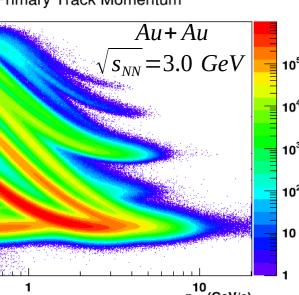
taken of dE/dx and  $1/\beta$ . Resulting distributions are fit with Gaussians in order to extract the raw yields. Detector efficiency was estimated using a full simulation of the STAR detector to determine the response. Efficiency corrected data are then fit and integrated to obtain yields.

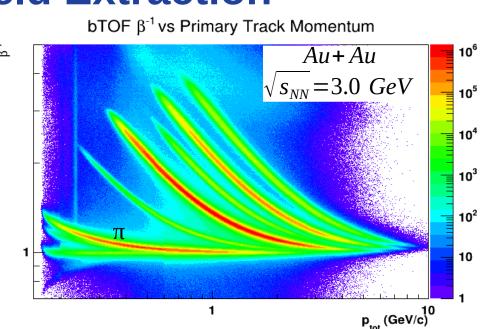


Pion spectra are fit with a double thermal function, which was shown by E895 to describe production from the  $\Delta$  resonance at low temperature and thermal production at high temperature<sup>[1]</sup>. <sup>[1]</sup>J. Klay et al. (E895 Collaboration), Phys. Rev. C 68, 054905 (2003)

Benjamin Kimelman University of California, Davis For the STAR Collaboration

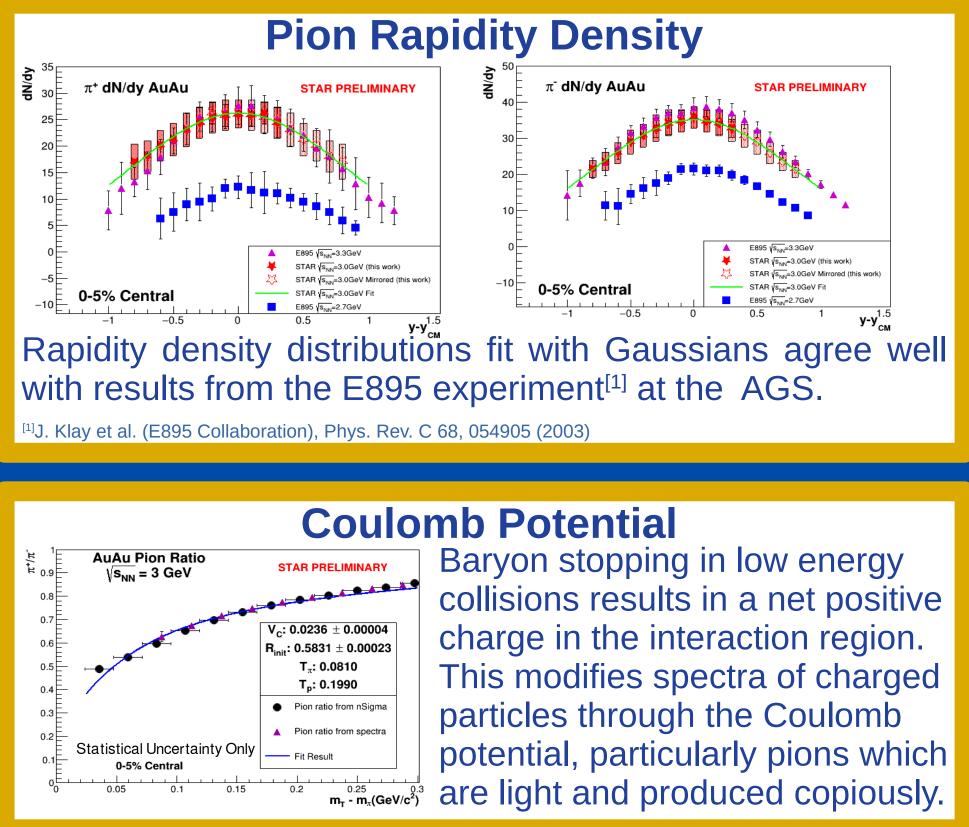
### **Spectra and Yield Extraction** dE/dx vs Primary Track Momentum





For each charge, slices in two dimensions  $(m_{\tau}-m_{0} \text{ and } y)$  are

**Pion Invariant Spectra** 



The first results from the STAR Fixed-Target program have been shown to be consistent with and extend upon those from the AGS experiments at similar energies. The Coulomb potential of other FXT energies will be analyzed and added to global systematics.



This material is based upon work supported by the National Science Foundation under Grant No. 1812398 (Cebra and Calderón de la Barca). Any opinions, findings, and conclusions or recommendations expressed in this material are those of the authors and do not necessarily represent the views of the National Science Foundation.



### **Conclusions and Outlook**

## Acknowledgments

