Jet-like azimuthal correlations in p+p and p+A collisions at forward rapidity with STAR

Mriganka M. Mondal (for the STAR Collaboration)

Institute of Physics, Bhubaneswar, India

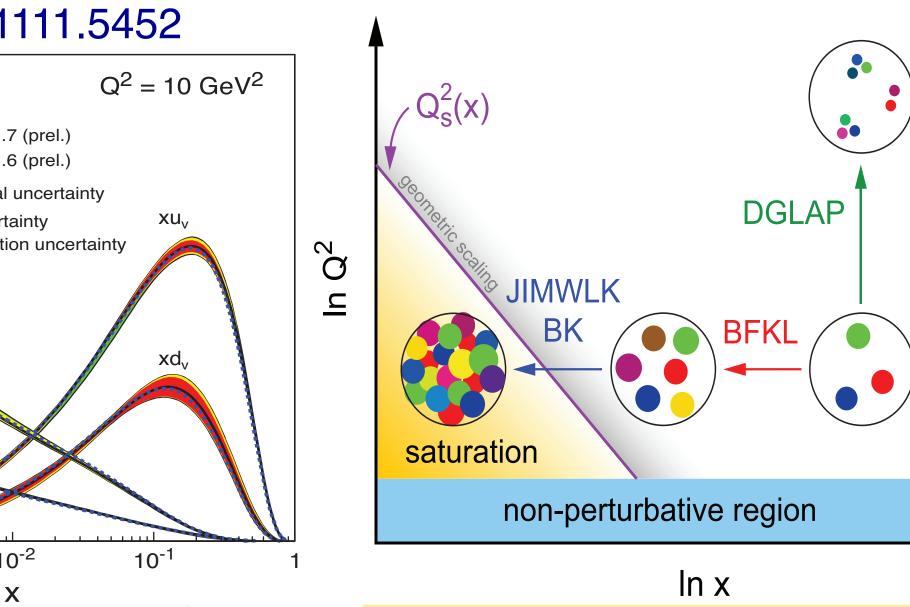
At very low Bjorken x, parton densities in a nucleon/nucleus reach a point where the gluon density is expected to transit to a saturation regime. The saturation scale (Q_s) is defined as the inverse of typical transverse inter-partonic distance and can be expressed as $Q_s \sim (A/x)^{1/3}$. So far no clear experimental observation of the onset of gluon saturation has been made. Observables at forward rapidity at STAR probe gluons at very small x (.001< x <.005) in heavy nuclei where parton saturation is expected. Scanning the azimuthal correlations of jets as a function of rapidity and transverse momentum has the potential to yield critical insight into gluon saturation. During the 2015 RHIC run, STAR recorded data in p+p, p+Al and p +Au collisions at $\sqrt{s_{NN}}$ = 200 GeV, with mid-rapidity calorimeters and the Forward Meson Spectrometer (FMS), which cover a wide acceptance of -1.0 to 2.0 and 2.5 to 4.0, respectively, in pseudorapidity over the full azimuth. The status of determining jet-like azimuthal correlations from these data sets is reported.

Expect to see at RHIC?

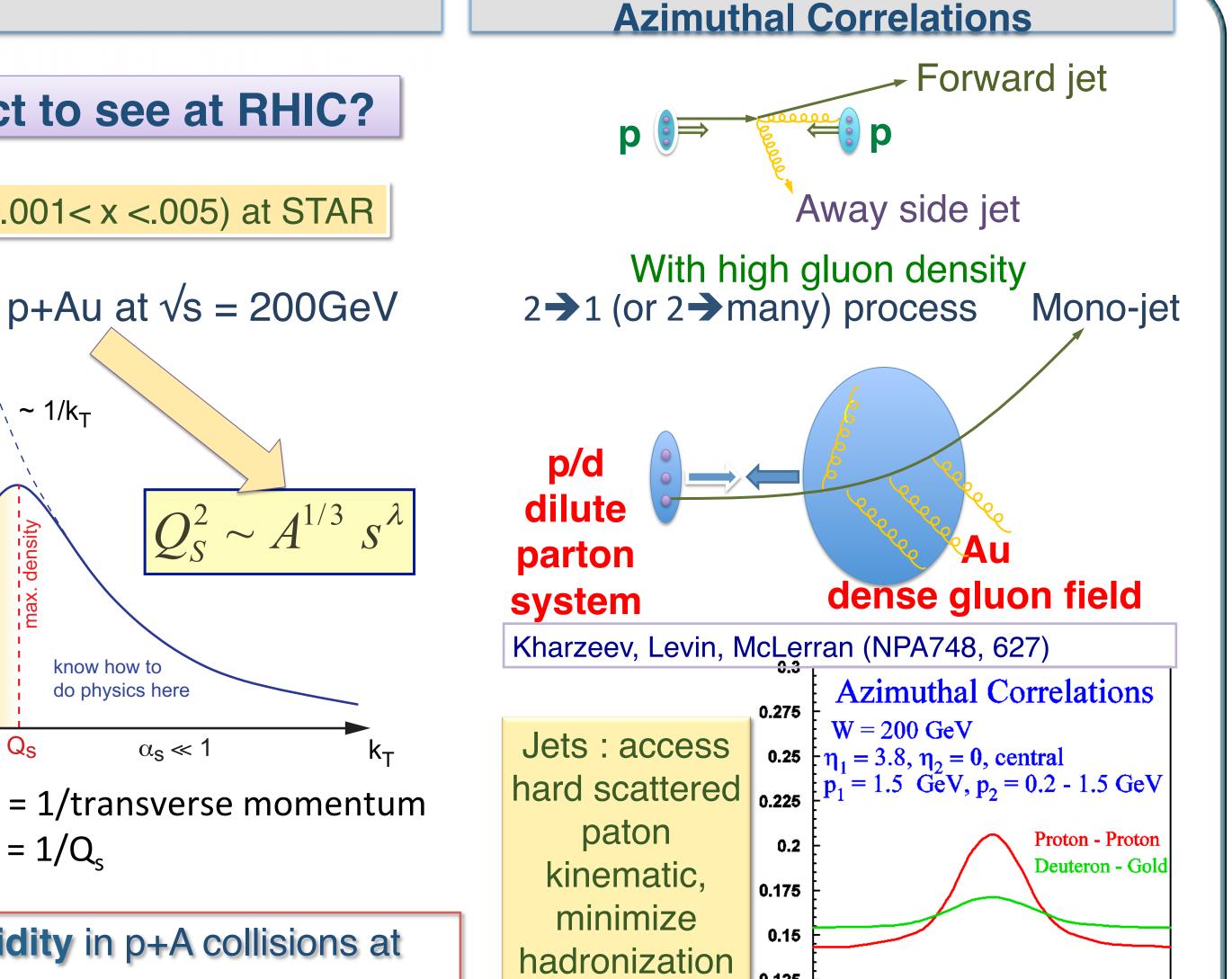
small x (.001< x <.005) at STAR

 $\alpha_{\rm S} \ll 1$

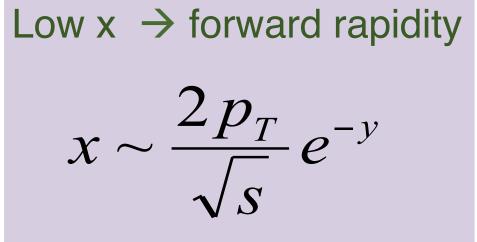
Gluon Saturation Densities of gluons and sea quarks are high at low x. Parton density evolution is linear (BFKL) at low parton density. As density grows, gluons start to recombine and the evolution become nonlinear, leading to the saturation of parton density, called Color Glass Condensate. arXiv:1111.5452 $Q^2 = 10 \text{ GeV}^2$ HERA



 $\sim 1/k_T$ know how to do physics here $\alpha_s \sim 1$ Λ_{QCD} Gluon "size" = 1/transverse momentum $= 1/Q_{s}$ forward rapidity <---- midrapidity



LC 2017



xG (× 0.05)

xS (x 0.05)

STAR

- The saturation region may be reachable at forward rapidity in p+A collisions at RHIC
- Scan in p_T , rapidity and nucleus size (A) may allow to cross saturation scale $Q_s^2(x)$

STAR at RHIC Transversely **Polarized** beam **Forward Meson Spectrometer (FMS)** Midrapidity calorimeters : BEMC : η (-1.0, 1.0) & **EEMC**: η (1.0, 2.0)

FMS: η (2.5,4.0), π^0/η --> $\gamma\gamma$, direct photons, Drell-Yan

Working on FMS gain

uniformity

Pre-shower detector installed in 2015 Post-shower detector installed in 2017

Jet Recontruction

- Underlying event contributions in jets for p+p in STAR is well studied
- For p+Au collisions the underlying event background is expected to be larger.
- Off-axis cone correction done for p+p and p+Au

Off-axis correction: Jet p_T correction from p_T density calculated for fixed cone at the same η and $\Delta \varphi$ $=\pi/2$ away from Φ_{iet} Jet-like $\Delta \phi$ correlations (parton vs. detector)

Particle vs. Detector Jets comparisons - PYTHIA 2000 Simulation Mean 0.0009133 Particle Jet Energy (GeV) $\Delta\Phi$ (particle jet – detector jet)

Jets reconstruction ("FastJet" package):

Forward jets: reconstructed photons (photon

candidates) from FMS; Midrapidity jets: towers

in BEMC and EEMC; Anti-kt: R = 0.6 for FMS

and R=0.5 for B/EEMC; Off-Axis cone radius =

Simulation

• Jets-parton

influence

PYTHIA events after **GEANT** simulation of STAR: Good performance for forward jet-jet correlations!

 Jets from FMS (mostly electromagnetic) retain the directional information significantly With FMS triggers: Detector Jets retain substantial fraction of jet energies

Scanning in p_T and x (π^0 mass peaks in different p_T bin) RHIC 2015 $\vec{p} + p$, $L_{\text{int}} = 40 + 50 \text{ pb}^{-1}$ $\vec{p} + Al$, $L_{\rm int} = 1.0 \text{ pb}^{-1}$ Mass[GeV] $\vec{p} + Au$, $L_{\rm int} = 0.45 \ \rm pb^{-1}$ STAR year 2015 data being SLAC CCFR analyzed for π^0 - π^0 and EM m^{trig}NoCut, scaled jet-EM jet correlation p+p, p+Al, p+Au (d+Au in 2016)

Akio Ogawa: QM2017: https://indico.cern.ch/event/433345/contributions/2358413/

Current Status for 2015 STAR data

2015 □FMS-Preshower □RP Phase-II* Run: □p+p 200 GeV longitudinal & transverse □p ↑ +Au/Al 200 GeV transverse 2017 □FMS-Postshower

Upgrades Completed



The STAR Collaboration: drupal.star.bnl.gov/STAR/presentations

Run:

With Year2015 data's statistics, $\sqrt[6]{-6.1}$ $\sqrt{6.2}$ $\sqrt{6.3}$ $\sqrt{6.4}$ >3GeV

STAR can study evolution of Q_s² (x) with A