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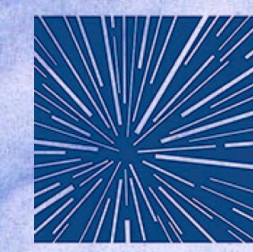


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Jets at STAR

Hot Jets:

Advancing the Understanding of High-Temperature QCD with Jets

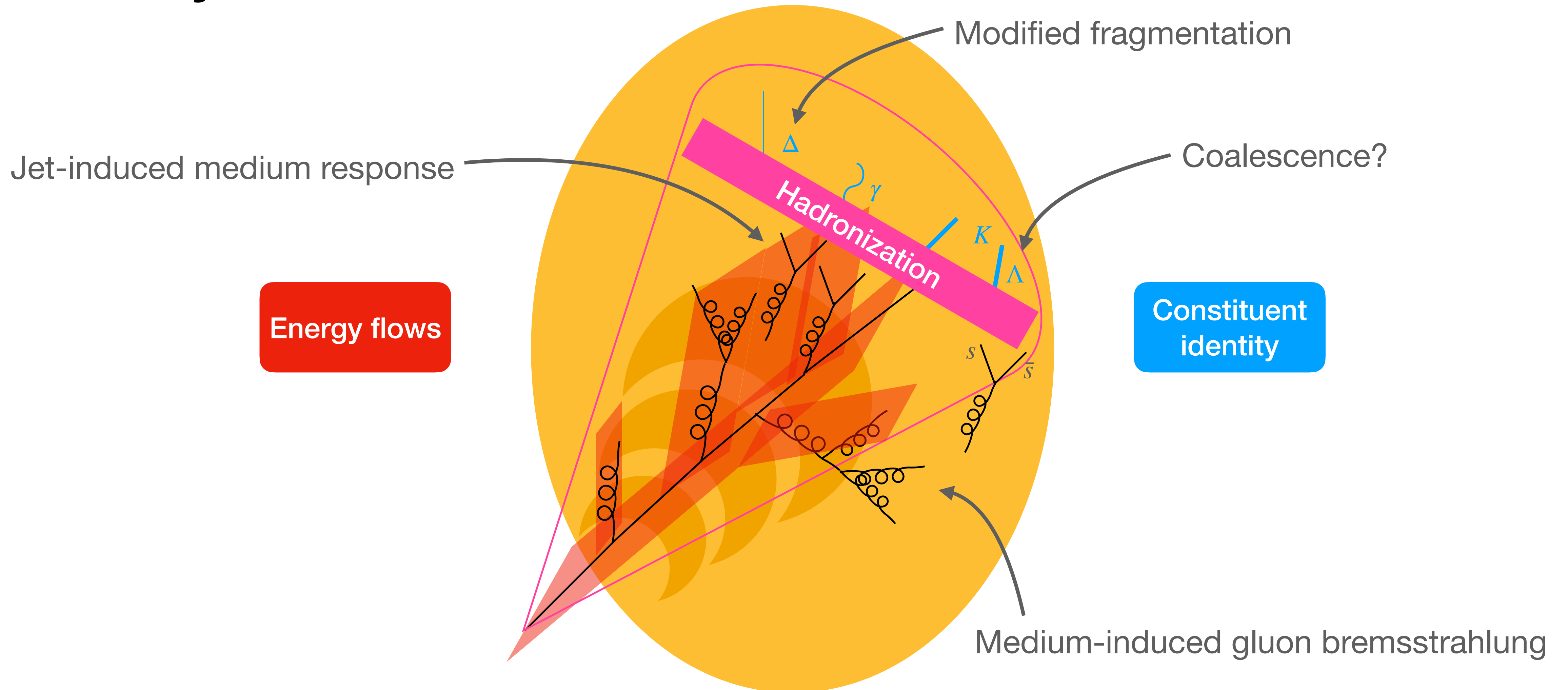
University of Illinois, Champaign-Urbana, IL
January 9, 2025

Isaac Mooney (Yale University, BNL) for the STAR Collaboration, isaac.mooney@yale.edu



How to understand jet evolution in media

Two ways: the How and the What



Solenoidal Tracker at RHIC (STAR)

Image: [NSWW](#)

Main subdetectors, as of mid-2010s

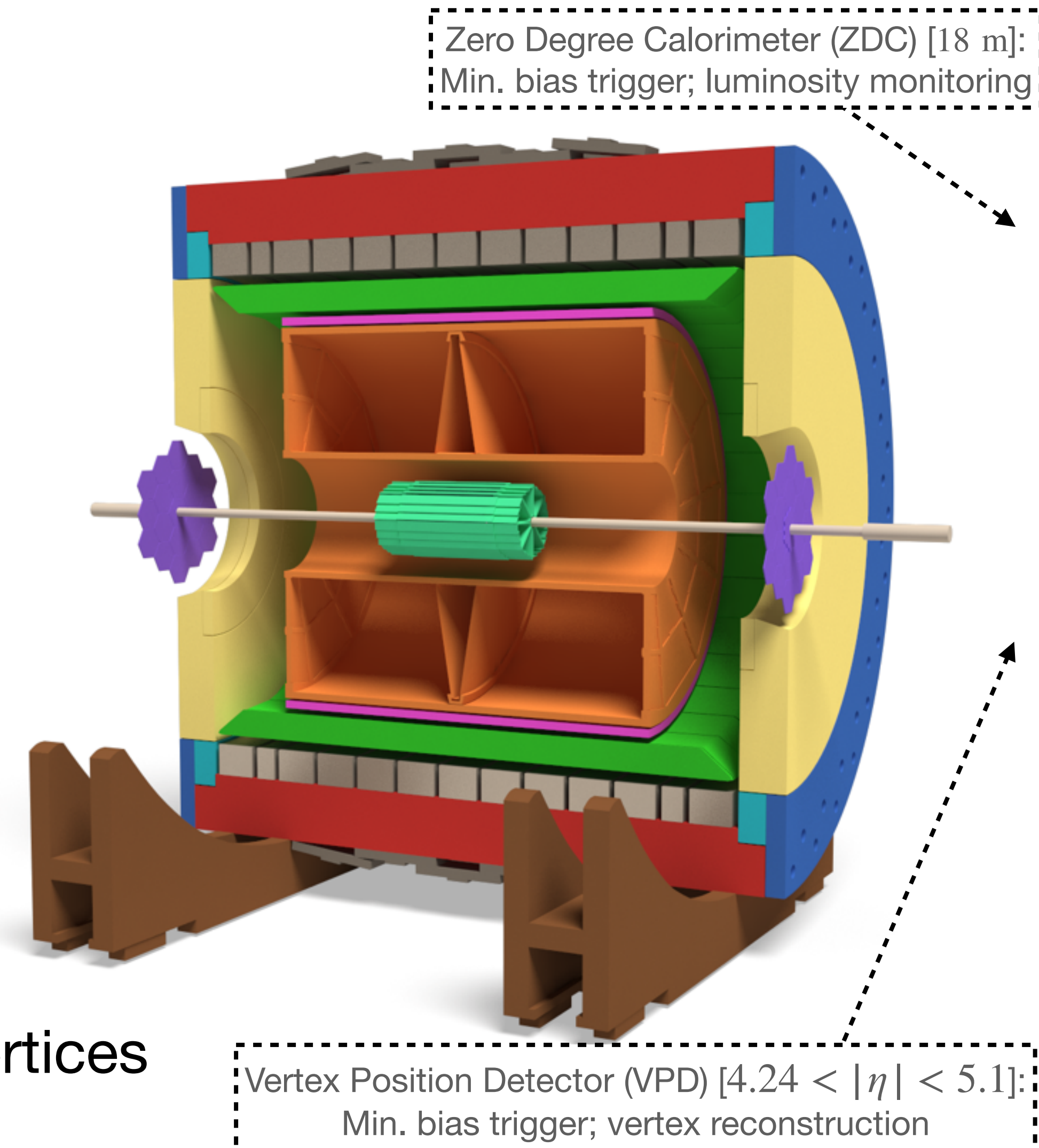
Relativistic Heavy Ion Collider (RHIC)
collides $p+p$, isobars ($Zr+Zr$, $Ru+Ru$), $Au+Au$, etc.
from $\sqrt{s_{NN}} = 3$ to 510 GeV

Time Projection Chamber (**TPC**) [$|\eta| < 1$]:
momenta of charged tracks + centrality + PID

Barrel Electromagnetic Calorimeter (**BEMC**) [$|\eta| < 1$]:
neutral energy deposits + online trigger

Time of Flight (**TOF**) [$|\eta| < 0.9$]: PID + pileup mitigation

Heavy flavor tracker (**HFT**) [$|\eta| < 1$]: displaced decay vertices



Precision QCD; exploring the Lund plane
with *multi-dimensional jet substructure*

Path-length dependence of jet energy
loss in medium with *jet anisotropies*
(with respect to event plane)

Separating p-QCD and np-
QCD with *energy correlators*

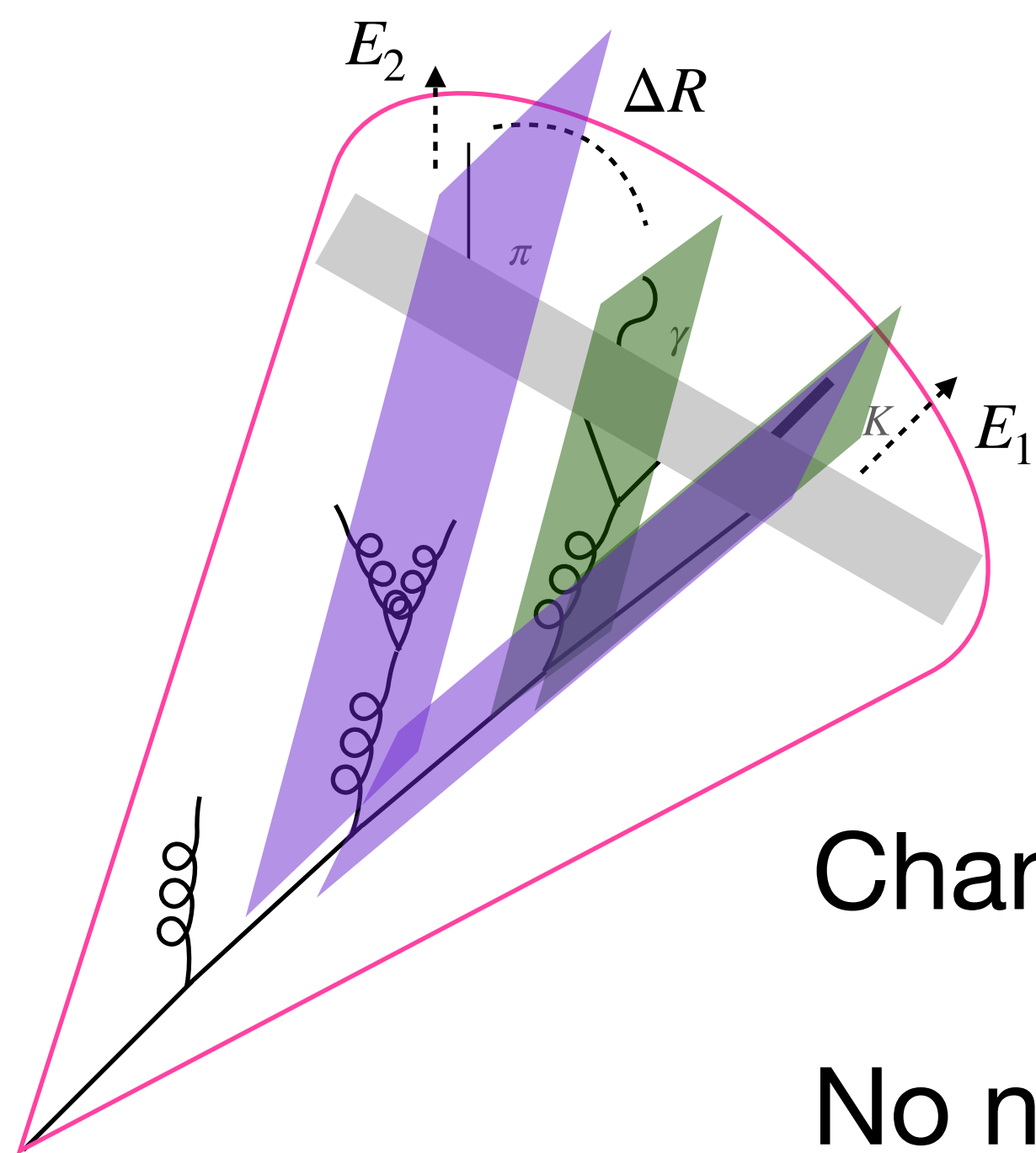
Energy flows

.....

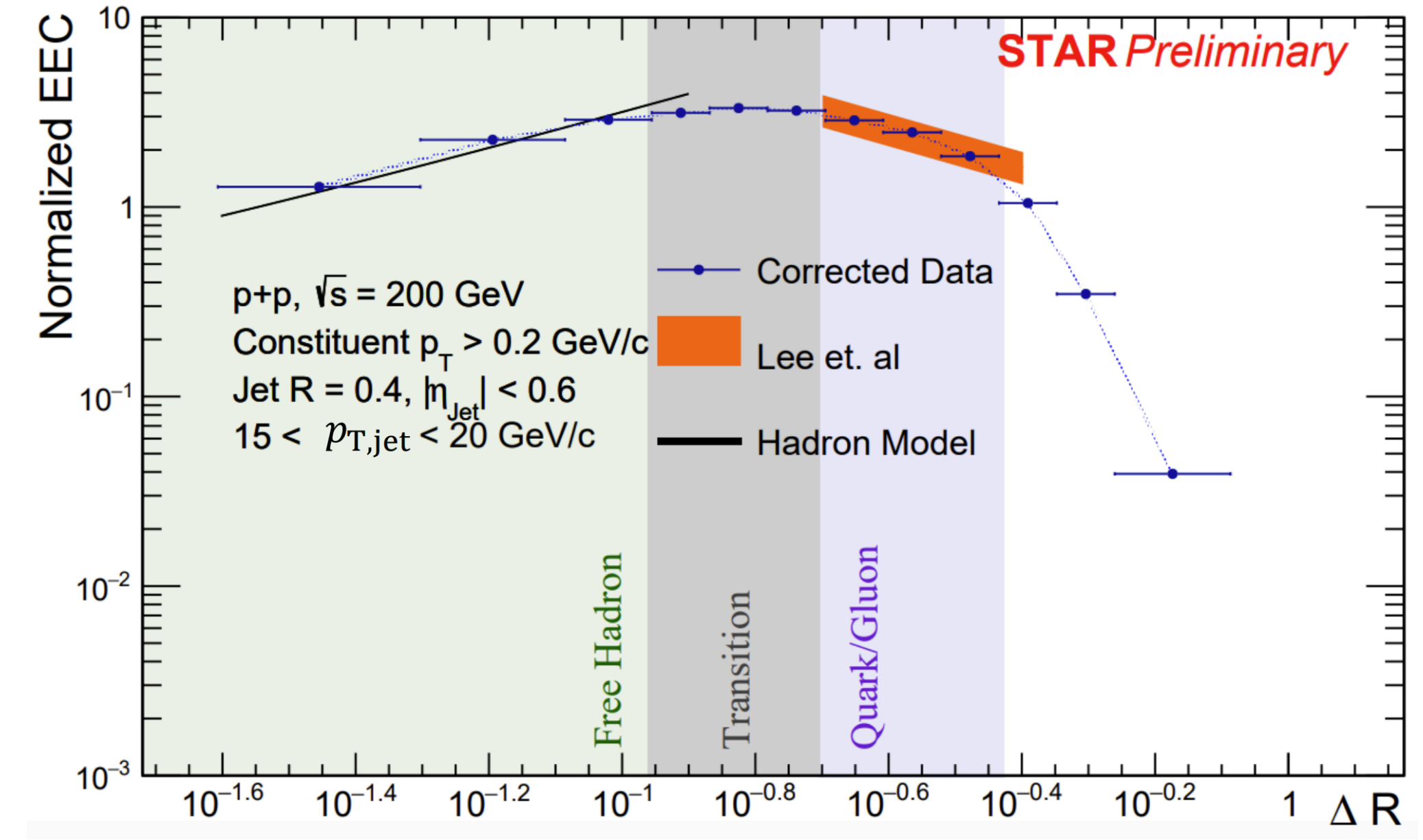
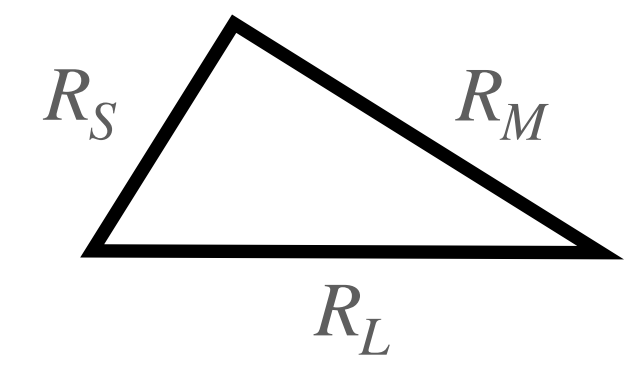
Energy-density dependence of jet energy loss in medium;
angular distribution of radiation in quenched jets
with *inclusive/semi-inclusive jet & high- p_T hadron yields*

Energy correlators

$$\text{ENC}(R_L) = \left(\prod_{k=1}^N \int d\Omega_{\vec{n}_k} \right) \delta(R_L - \Delta \hat{R}_L) \frac{1}{(E_{\text{jet}})^N} \langle \mathcal{E}(\vec{n}_1) \mathcal{E}(\vec{n}_2) \dots \mathcal{E}(\vec{n}_N) \rangle^{1,2,3}$$



$$\text{EEC}(\Delta R) = \frac{1}{\mathcal{O}} \frac{d\mathcal{O}}{d(\Delta R)}, \quad \mathcal{O} = \sum_{\text{jets}} \sum_{i \neq j} \frac{E_i E_j}{p_{T,\text{jet}}^2}$$



Change in scaling when virtuality $\sim p_T R_L \sim \Lambda_{\text{QCD}}$ so $R_L^{\text{transition}} \propto 1/p_T$

No need to recluster or remove npQCD contributions

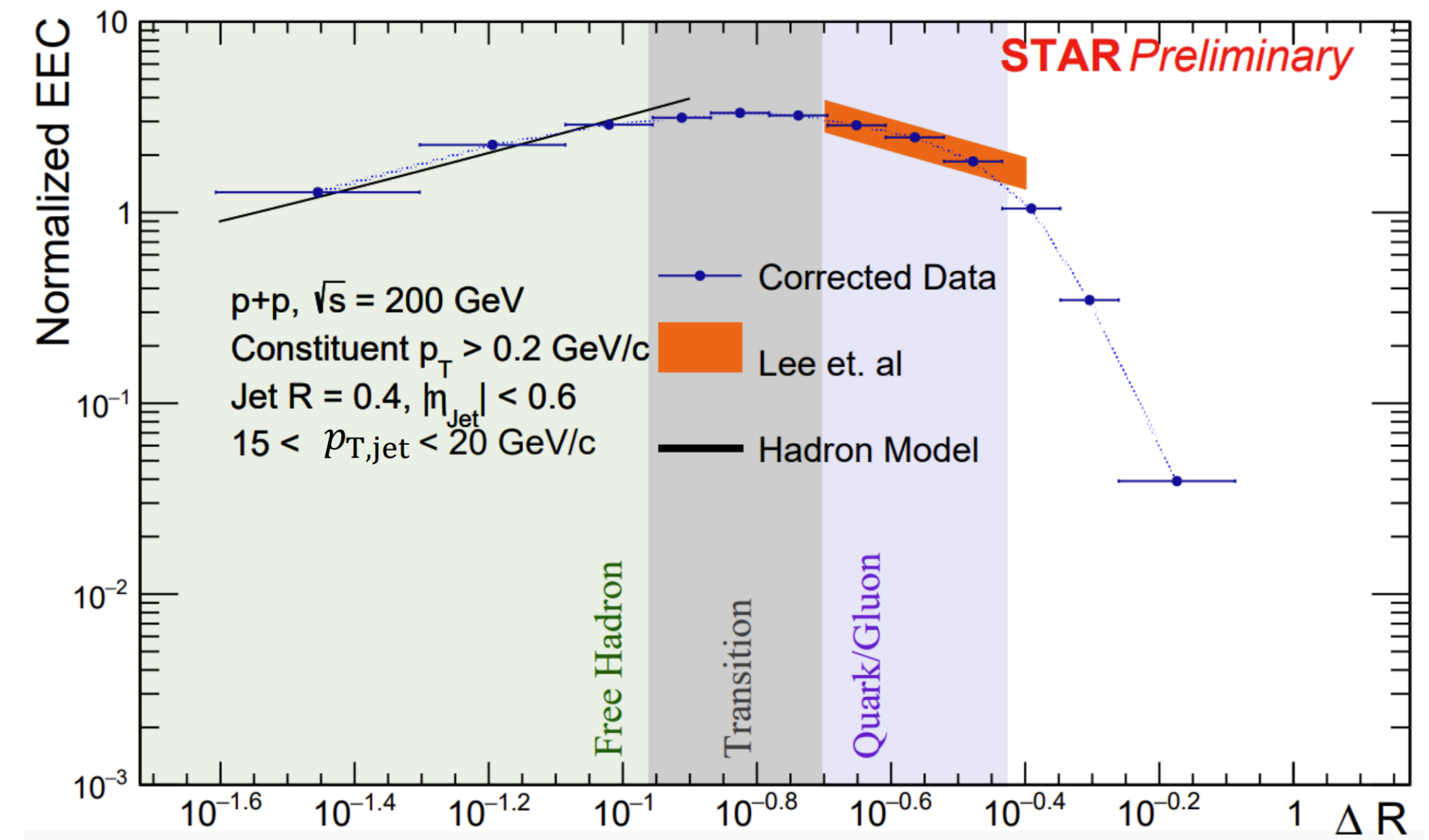
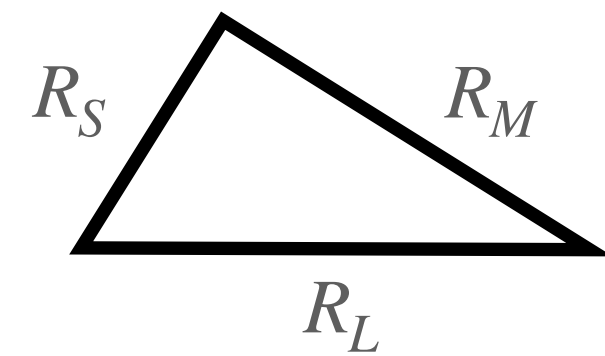
Simple scaling in the **hadronic** and **partonic** regimes

¹Basham, Brown, Ellis, Love, [PRL 41 \(1978\), 1585](#)
²Chen, Moulton, Zhang, Zhu, [PRD 102 \(2020\) 5, 054012](#)
³Komiske, Moulton, Thaler, Zhu, [PRL 130 \(2023\) 5, 051901](#)

Energy correlators

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Data agree well with NLL pQCD calculation* (& MC model, not shown)

Data agree well with model assuming non-interacting hadrons

¹Basham, Brown, Ellis, Love, [PRL 41 \(1978\), 1585](https://doi.org/10.1103/PhysRevLett.41.1585)

²Chen, Moulton, Zhang, Zhu, [PRD 102 \(2020\) 5, 054012](https://doi.org/10.1103/PhysRevD.102.054012)

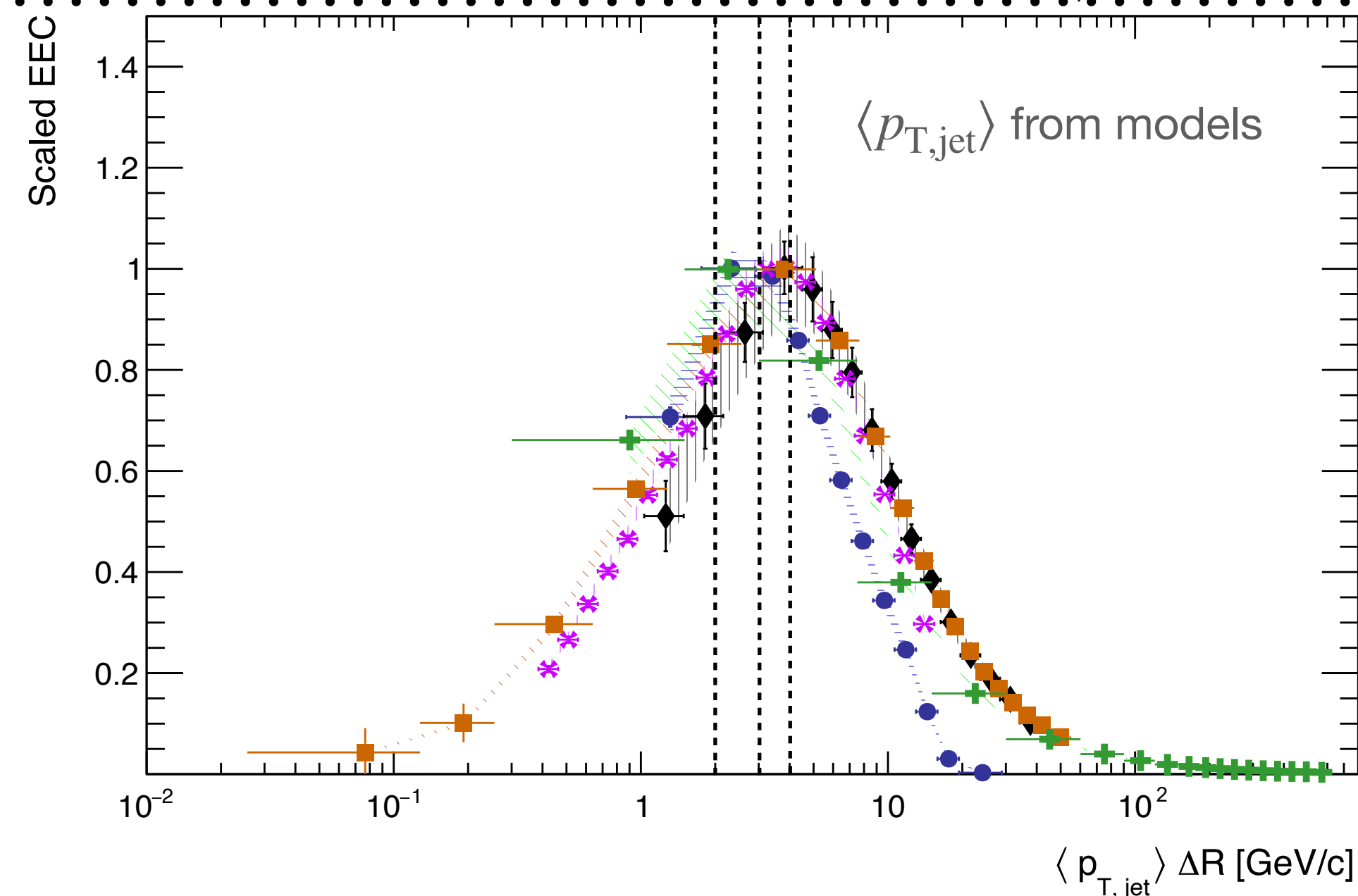
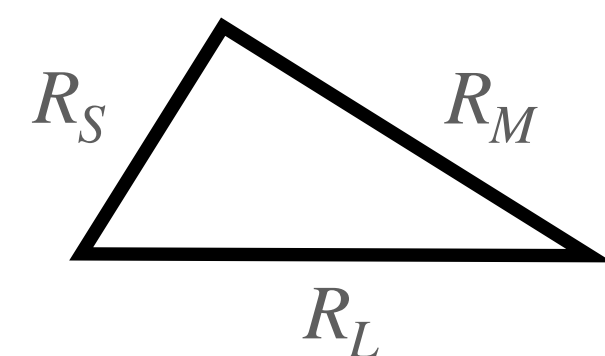
³Komiske, Moulton, Thaler, Zhu, [PRL 130 \(2023\) 5, 051901](https://doi.org/10.1103/PhysRevLett.130.051901)

Energy correlators

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- STAR Preliminary: $\sqrt{s} = 200 \text{ GeV}$ $30 < \text{Full Jet } p_T < 50 \text{ GeV}/c$
- * ALICE Preliminary: $\sqrt{s} = 5.02 \text{ TeV}$, $20 < \text{Charged Jet } p_T < 40 \text{ GeV}/c$
- ◆ ALICE Preliminary: $\sqrt{s} = 13 \text{ TeV}$, $60 < \text{Charged Jet } p_T < 80 \text{ GeV}/c$
- CMS Preliminary: $\sqrt{s} = 13 \text{ TeV}$ $97 < \text{Full Jet } p_T < 220 \text{ GeV}/c$
- + CMS Preliminary: $\sqrt{s} = 13 \text{ TeV}$, $1410 < \text{Full Jet } p_T < 1784 \text{ GeV}/c$

$$\text{EEC}(\Delta R) = \frac{1}{\mathcal{O}} \frac{d\mathcal{O}}{d(\Delta R)}, \quad \mathcal{O} = \sum_{\text{jets}} \sum_{i \neq j} \frac{E_i E_j}{p_{T,\text{jet}}^2}$$



Testing universality of transition region by comparing to LHC data:

~ 2 orders of magnitude in \sqrt{s} and $p_{T,\text{jet}}$ from STAR \rightarrow ALICE \rightarrow CMS,
transition region ~ 2 – 4 GeV/c

STAR more similar to CMS high- p_T (high- x) jets than ALICE or CMS low- p_T jets — **q vs. g** differences

¹Basham, Brown, Ellis, Love, [PRL 41 \(1978\), 1585](#)

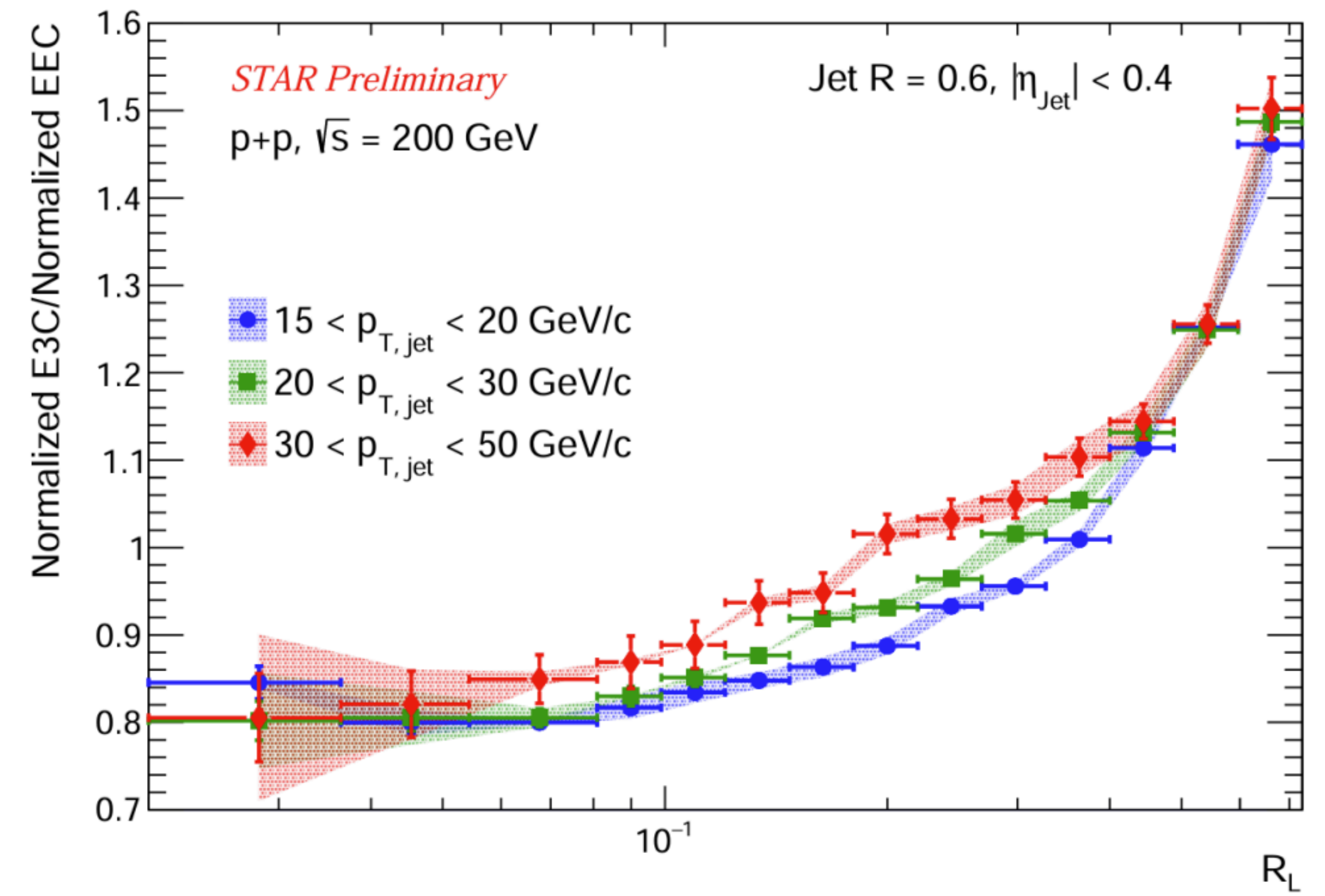
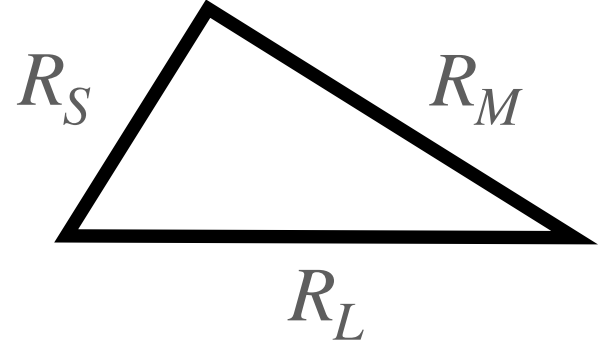
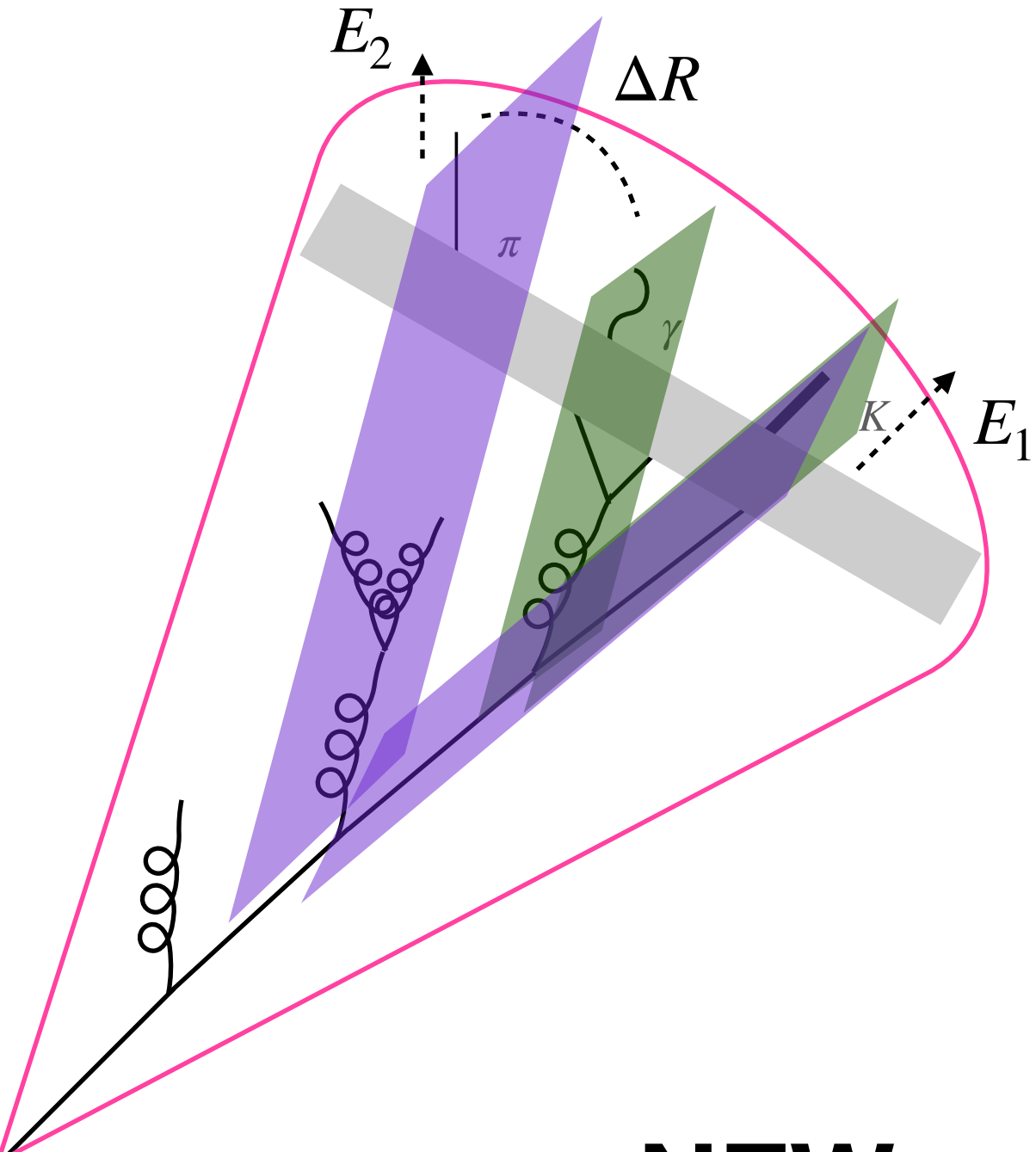
²Chen, Moul, Zhang, Zhu, [PRD 102 \(2020\) 5, 054012](#)

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NEW: extended to E3C — running of α_s causes change in slope with p_{T^*}

SoftDrop and the Lund Plane

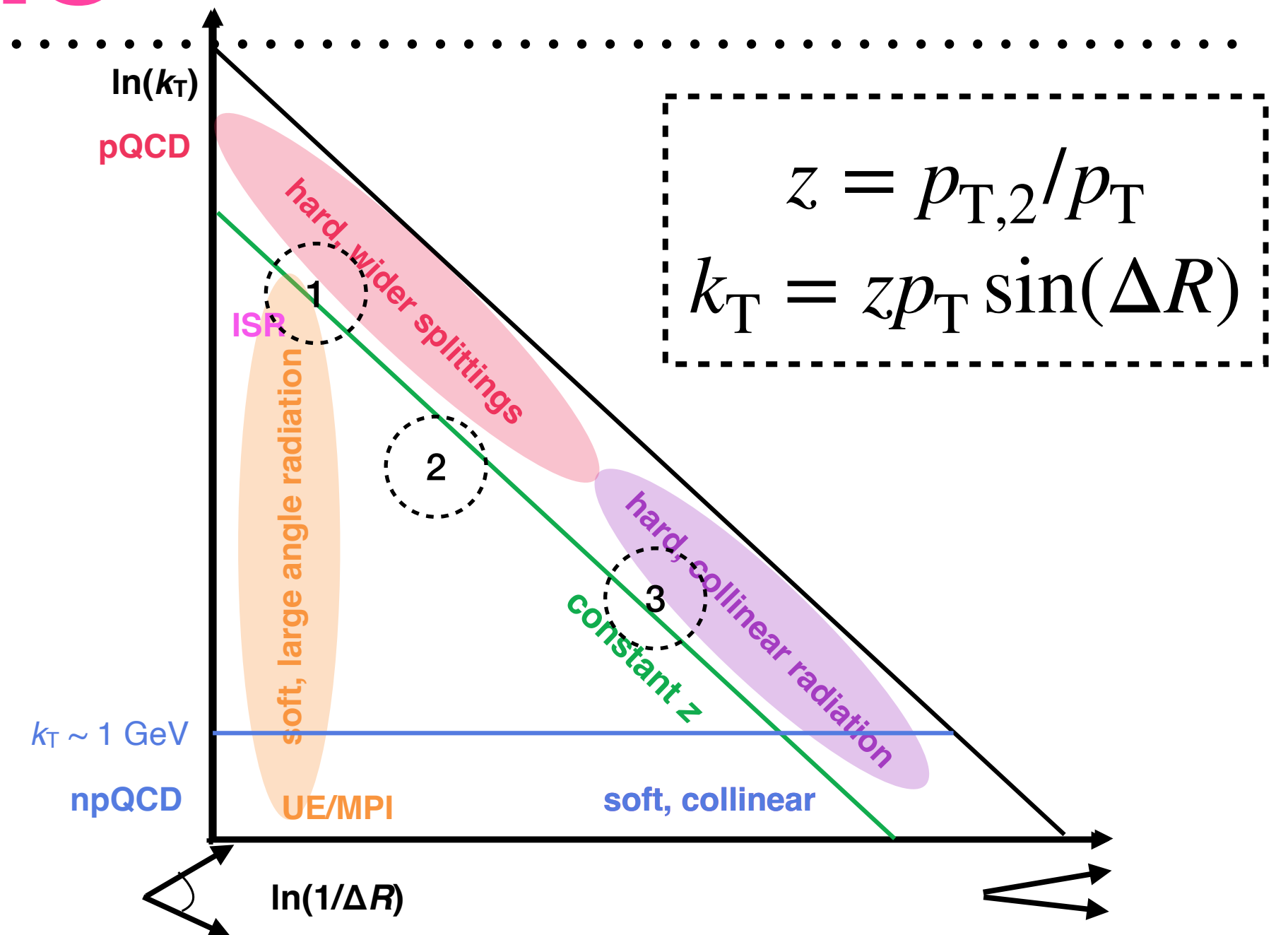
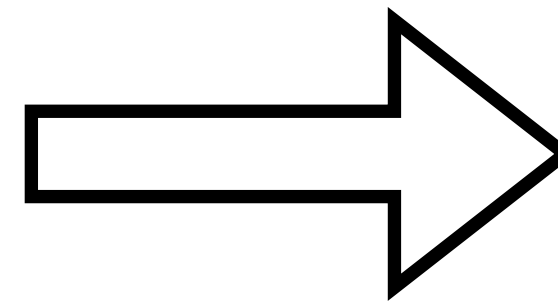
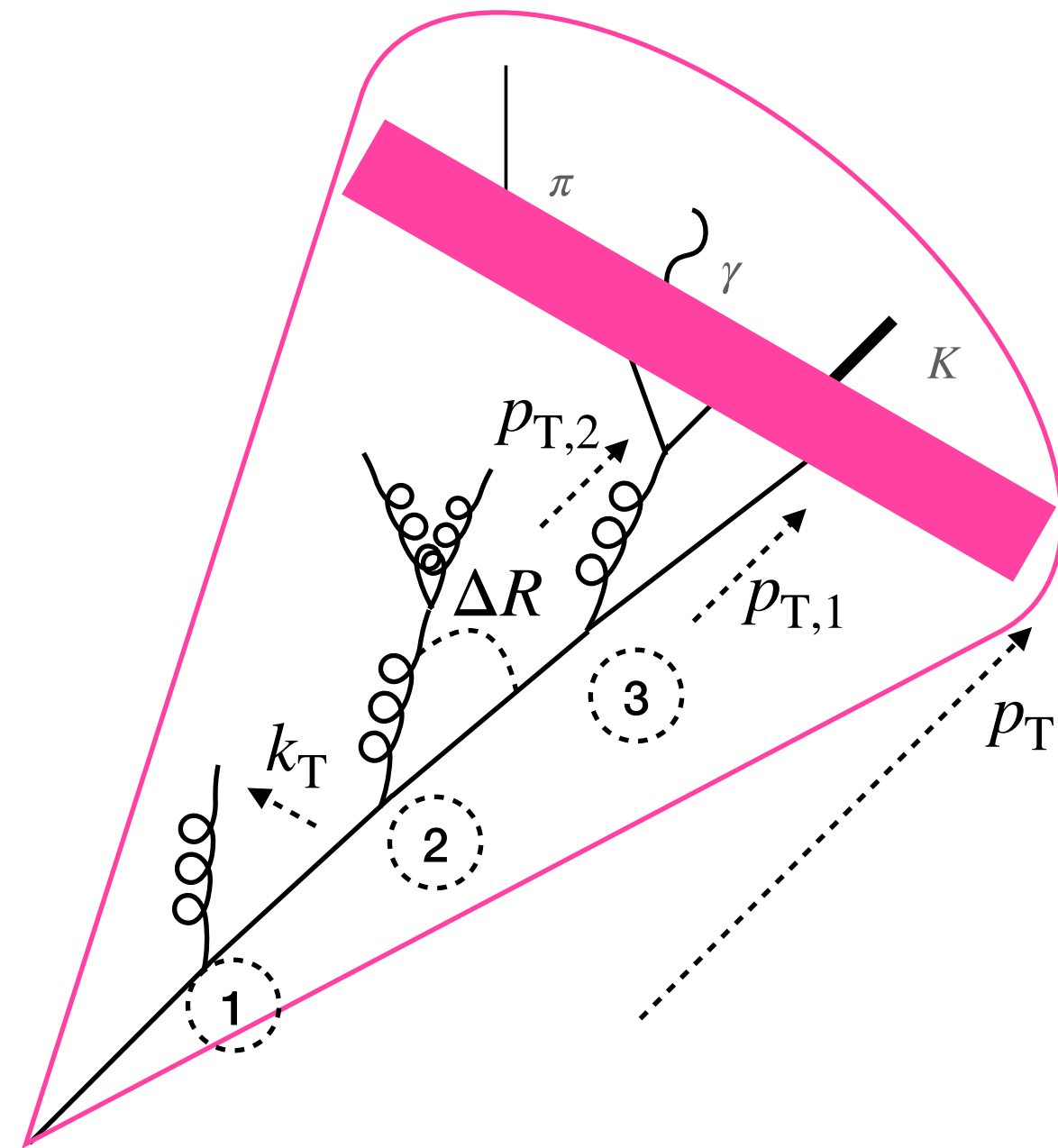
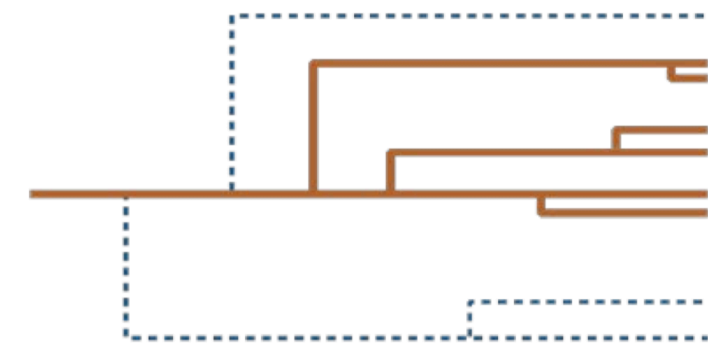


Image: Laura Havener, modified from Andrews et al., [J.Phys.G 47 \(2020\) 6, 065102](#)

$$\frac{\min(p_{T,i}, p_{T,j})}{p_{T,i} + p_{T,j}} > z_{\text{cut}} \left(\frac{\Delta R_{ij}}{R} \right)^\beta$$

$$z_g = \frac{\min(p_{T,1}, p_{T,2})}{p_{T,1} + p_{T,2}}$$



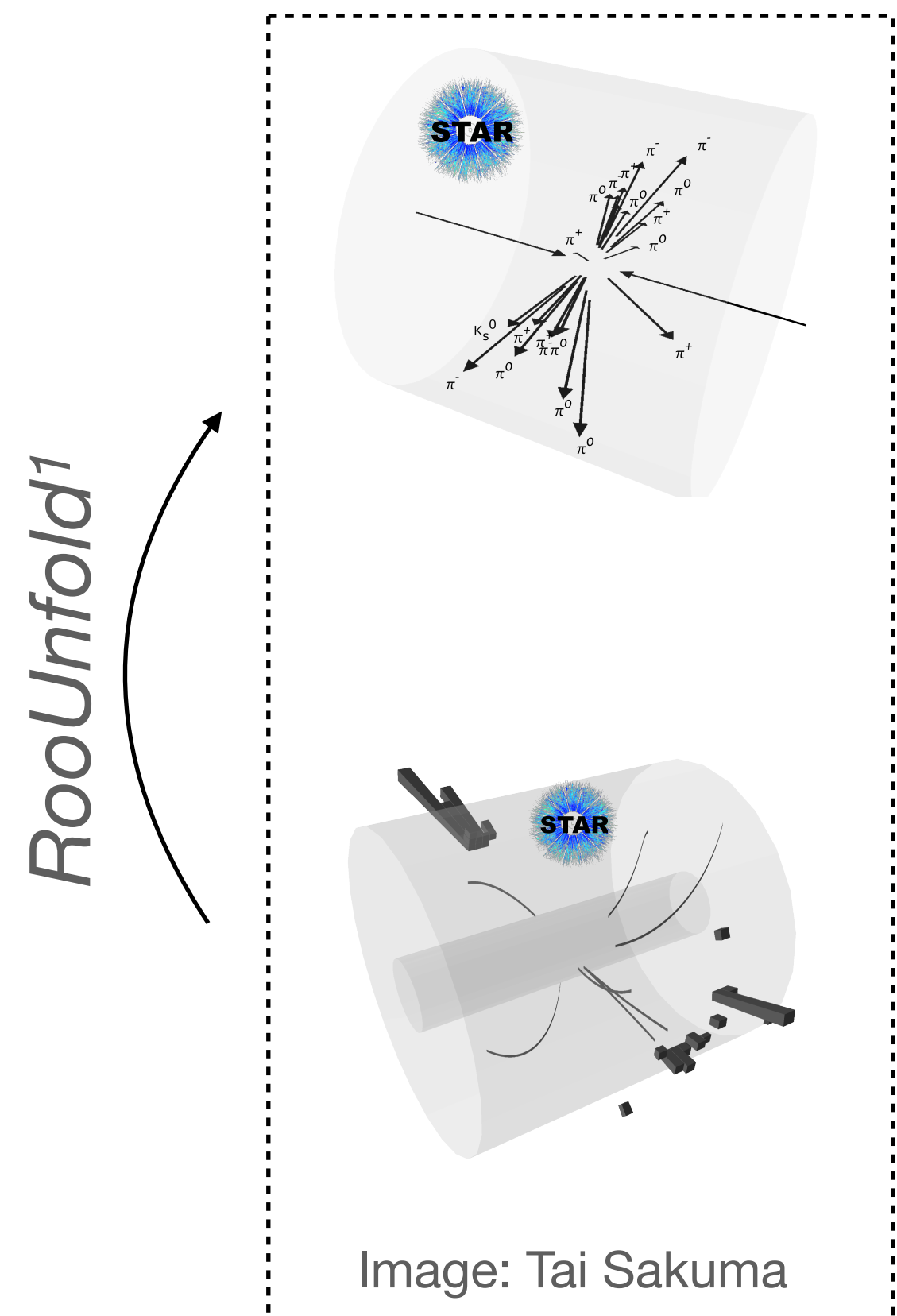
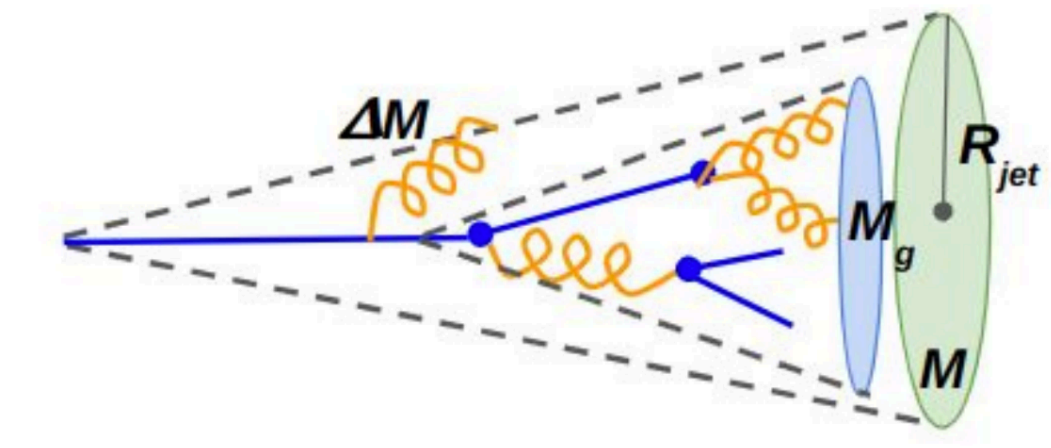
$$M_g = \left| \sum_{i \in J_g} p_i \right|$$

Image: Larkoski, Marzani, Thaler, Xue, [PRL 119 \(2017\) 13, 132003](#)

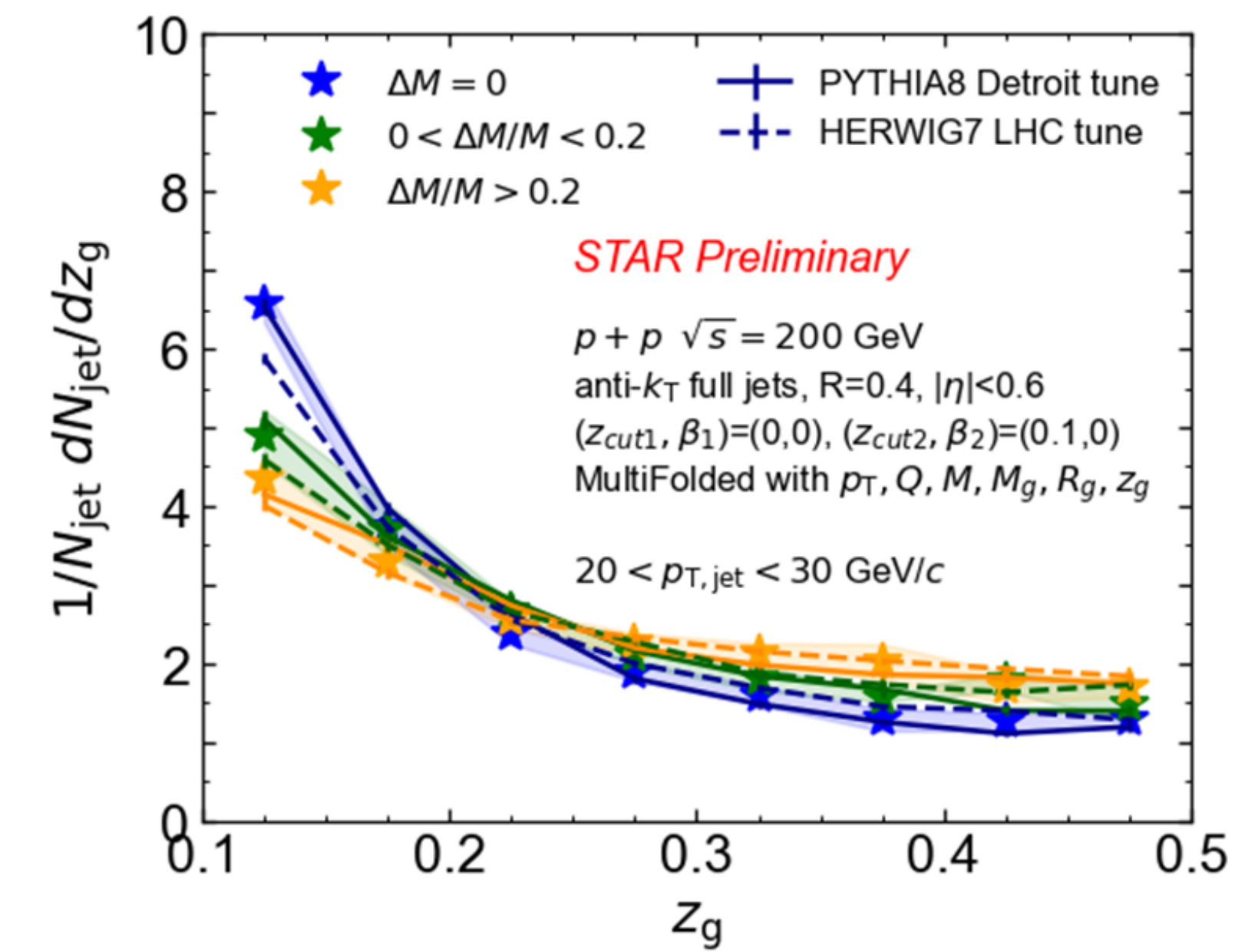
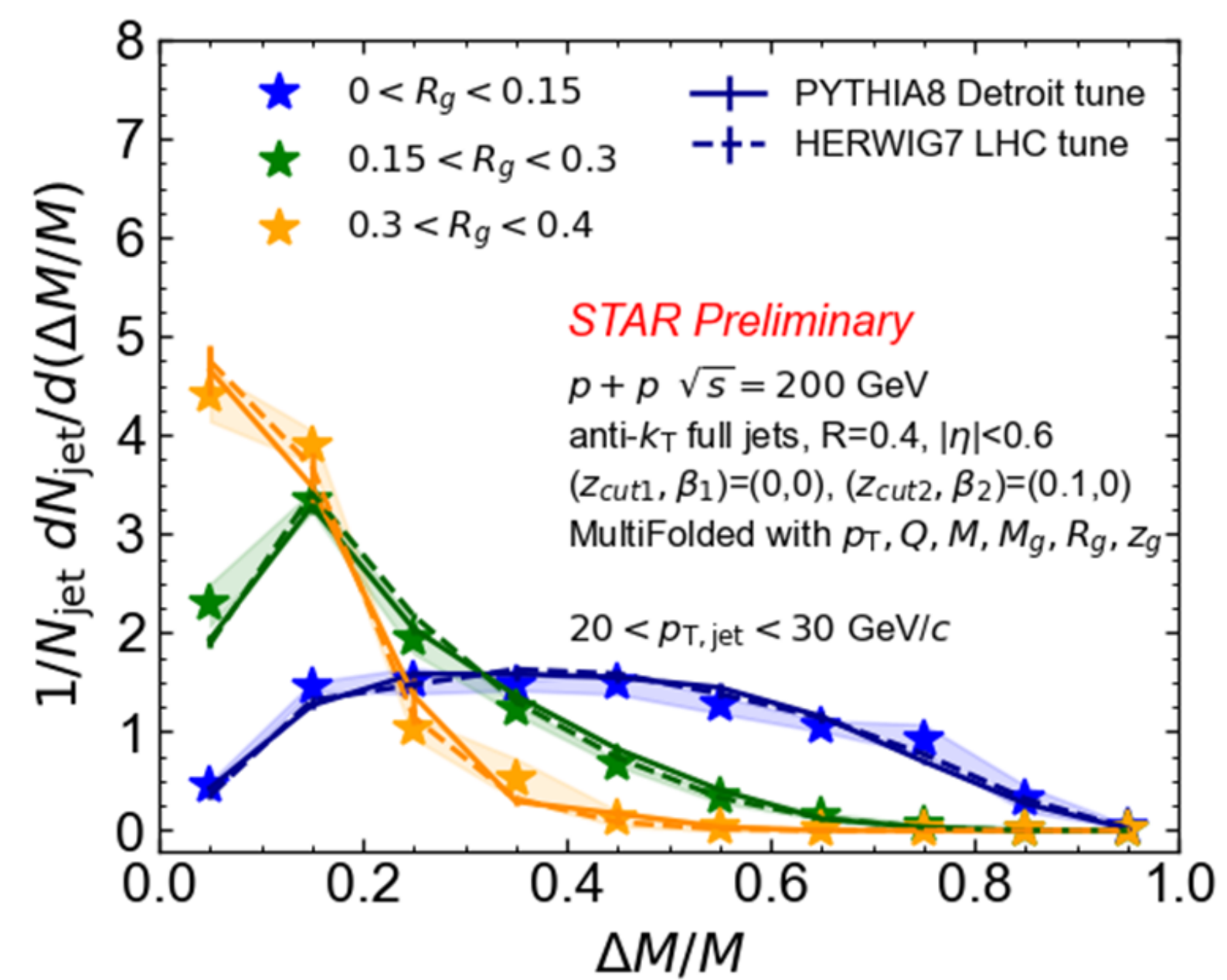
SoftDrop¹ grooming: reduce soft non-perturbative contribution
 → better theoretical control

N -dimensional substructure

With MultiFold



$$\Delta M = M - M_g^*$$



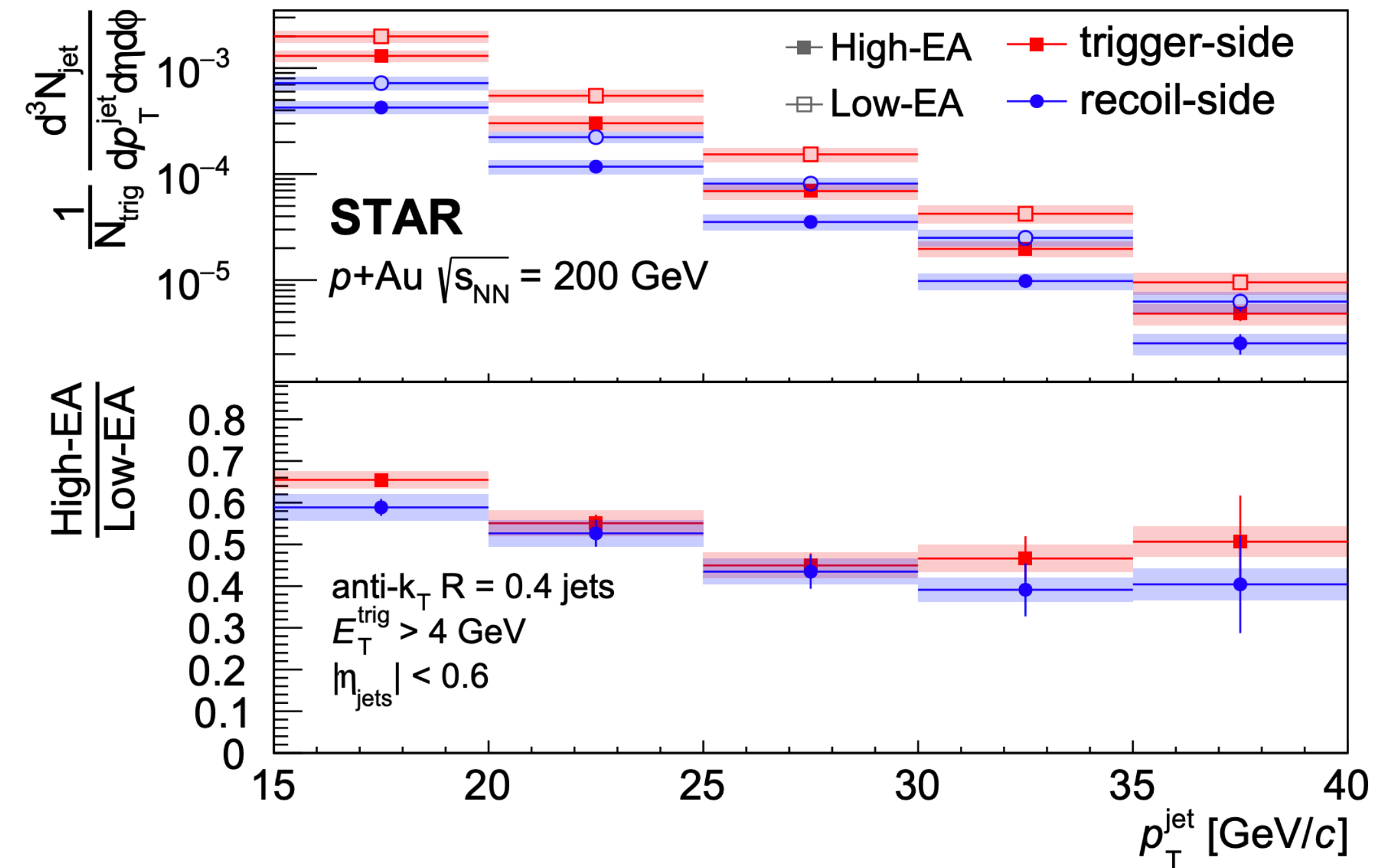
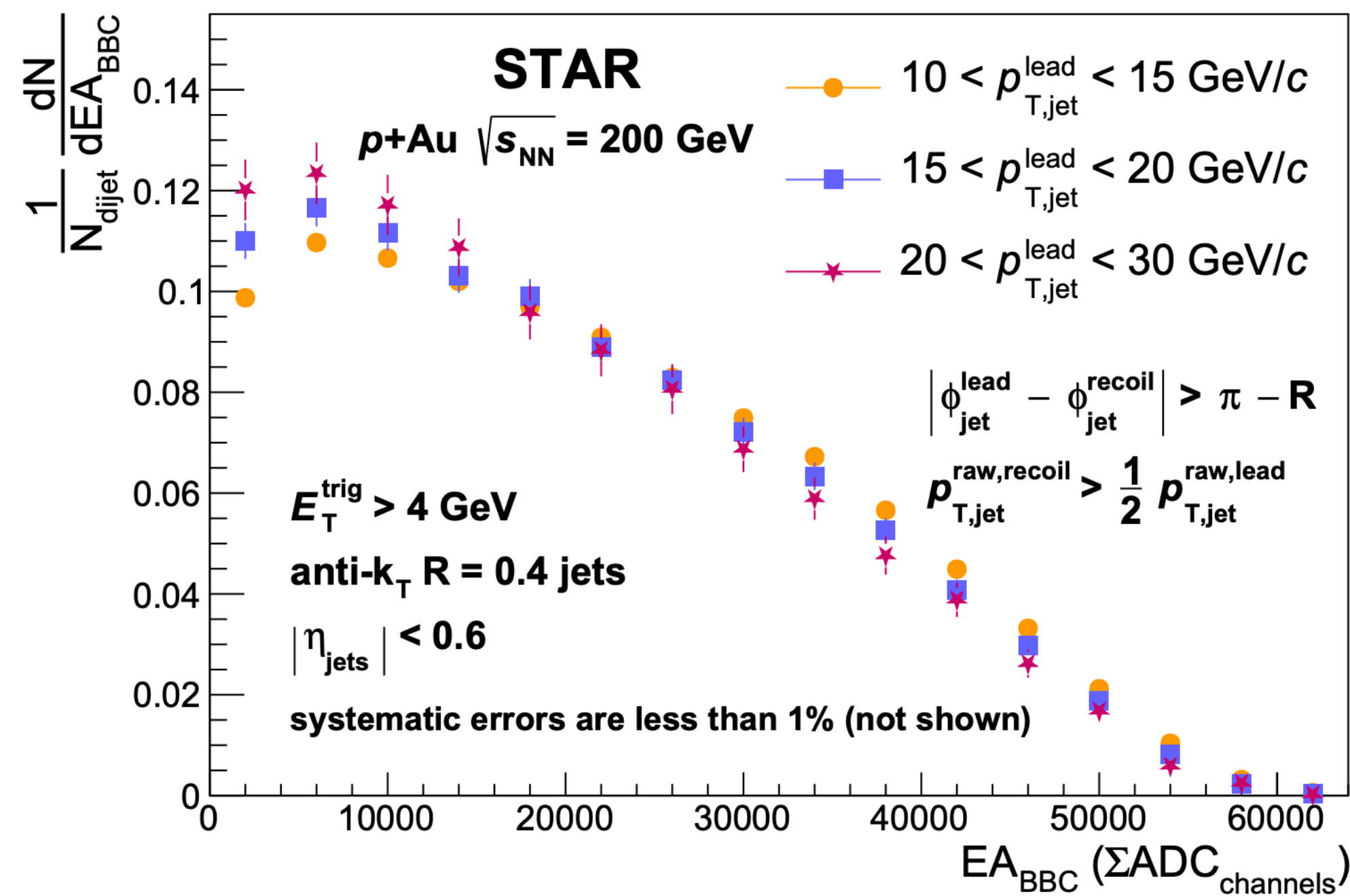
First application of MultiFold at RHIC

Correlation between R_g , z_g and ΔM consistent with angular ordering + kinematic constraint between early and late time splittings

Interlude: quenching in small systems?

Recently published in PRC! – [PRC 110 \(2024\) 4, 044908](#)

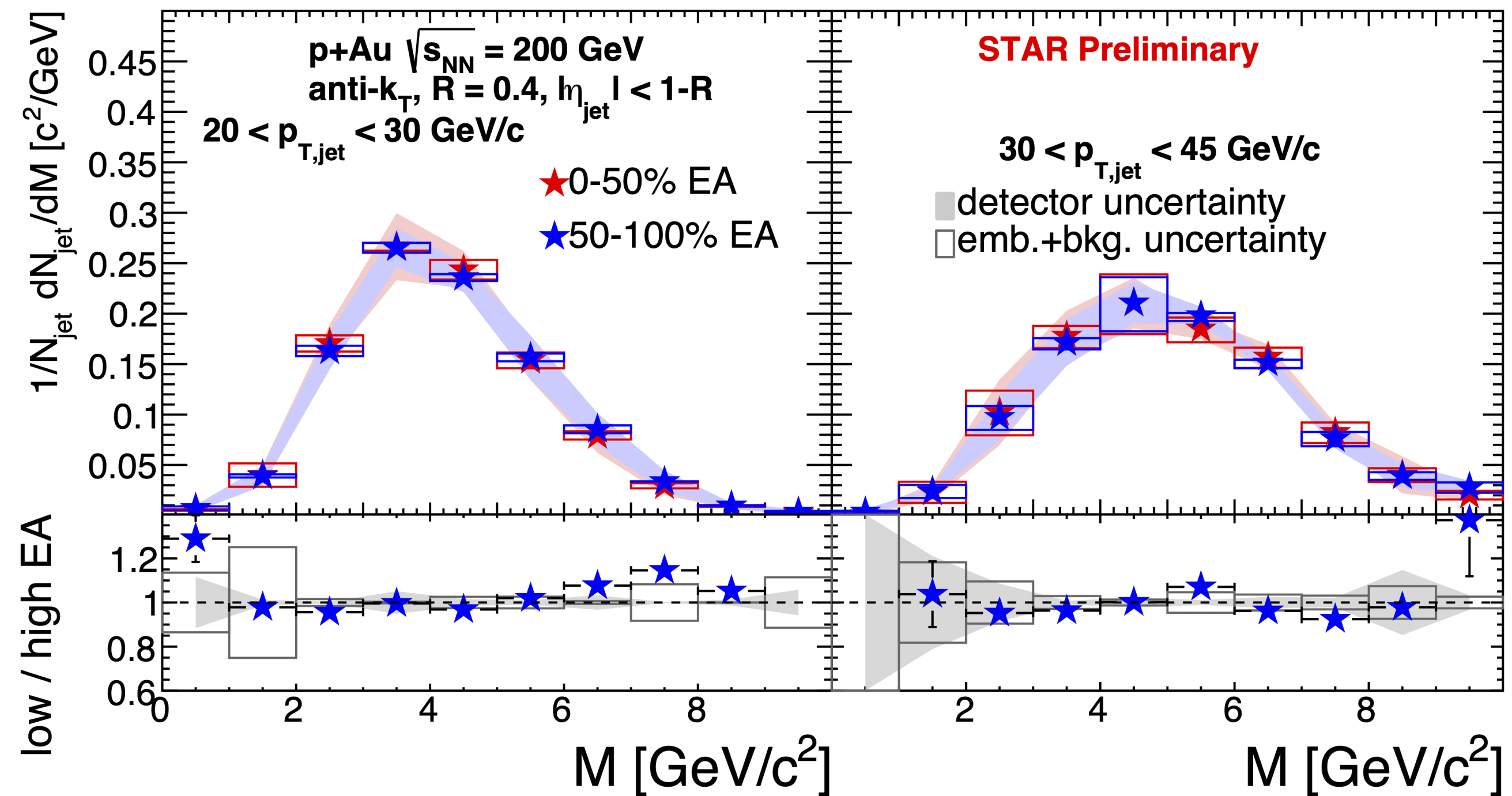
Short answer: disfavored at RHIC in p+Au collisions by this set of yield measurements from STAR



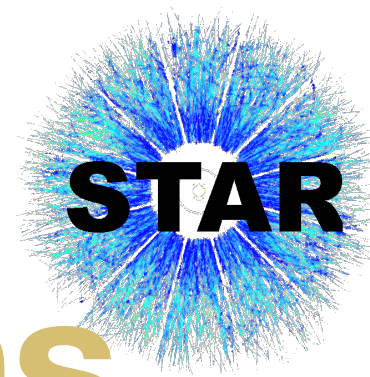
Suppression of yields but not corresponding to typical **surface bias** picture of medium-modification

Interlude: quenching in small systems?

...and by corresponding set of substructure measurements from STAR

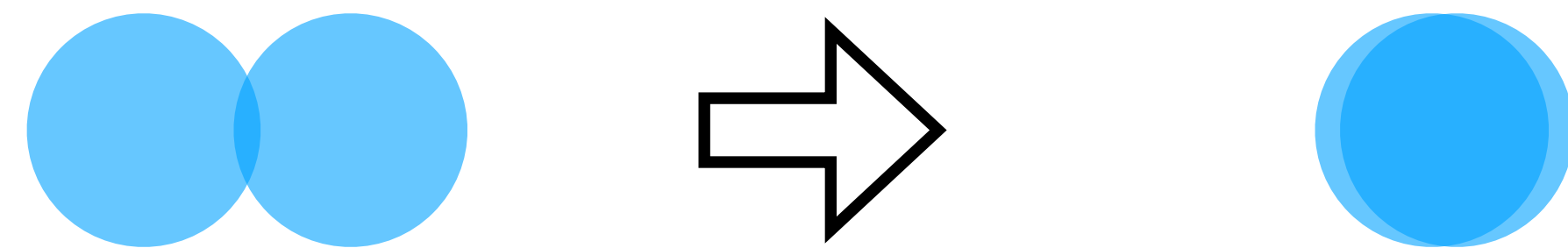
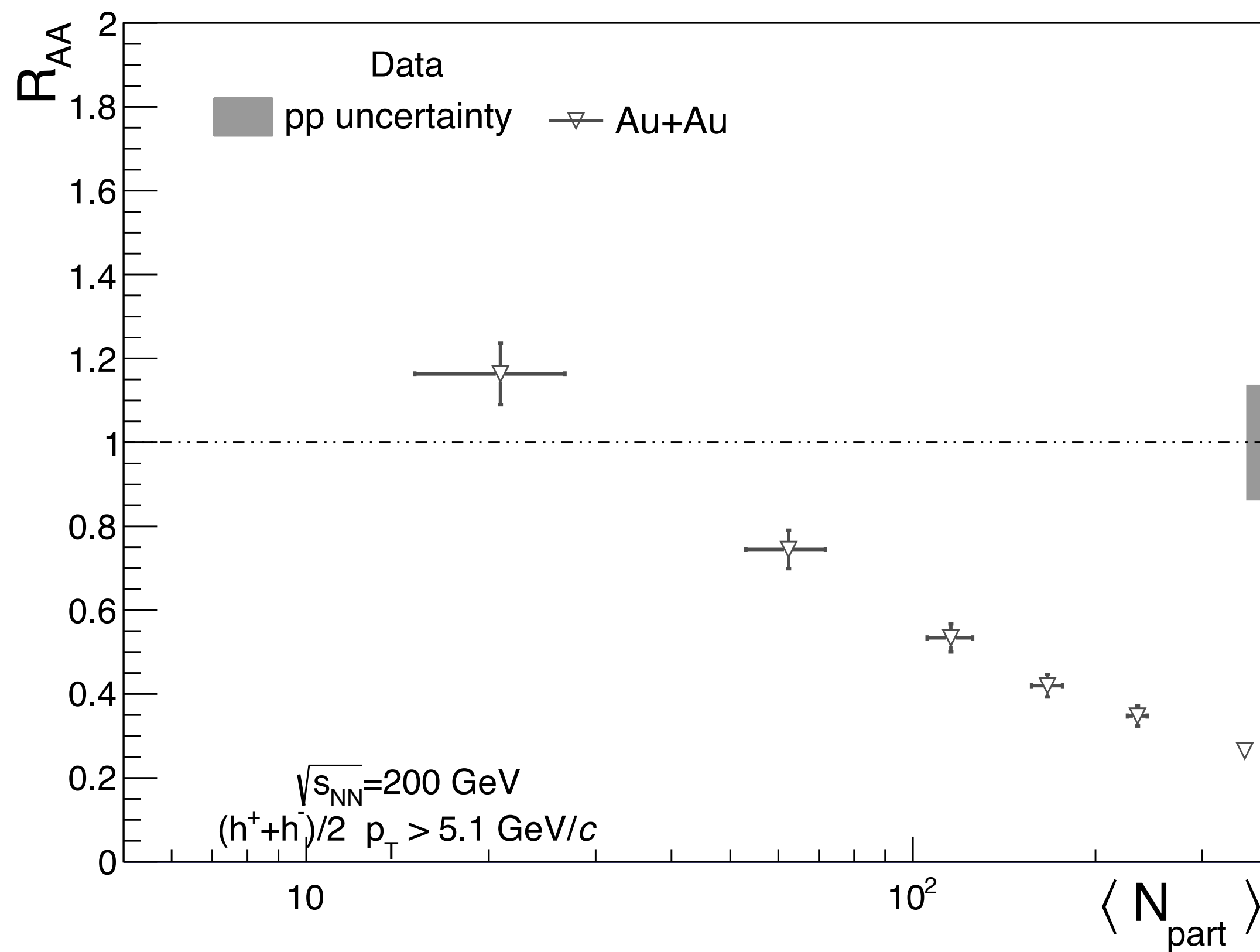


Rather, modifications likely due to **early-time dynamics** and/or **initial state configuration**



Inclusive yield modification in heavy-ion collisions

$$R_{AA} = \frac{1}{N_{ev}^{AA}} \frac{d^2 N^{AA} / d\eta dp_T}{\langle T_{AA} \rangle d^2 \sigma^{NN} / d\eta dp_T}, \quad \langle T_{AA} \rangle = \langle N_{coll} \rangle / \sigma_{inel}^{NN}$$

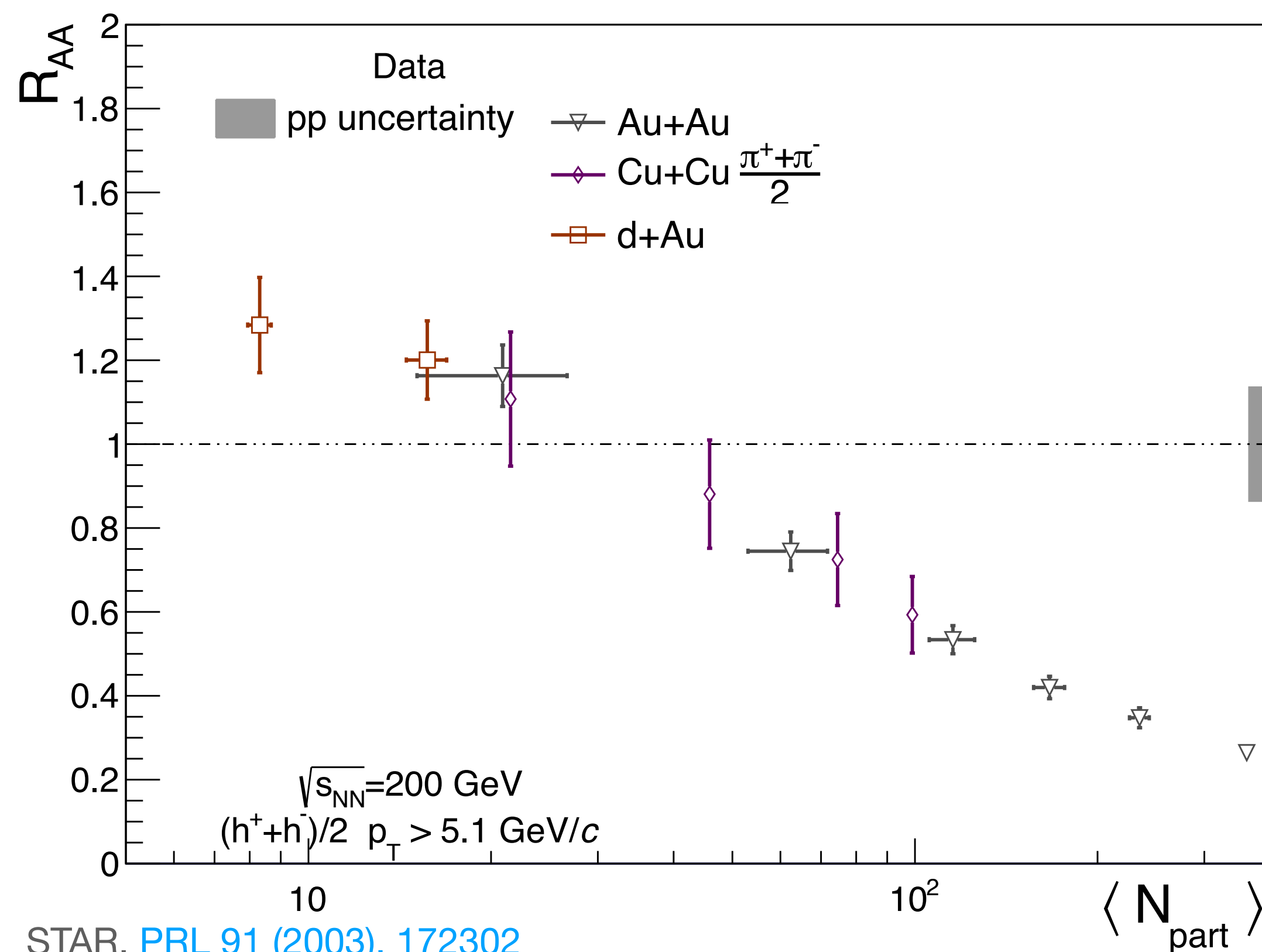
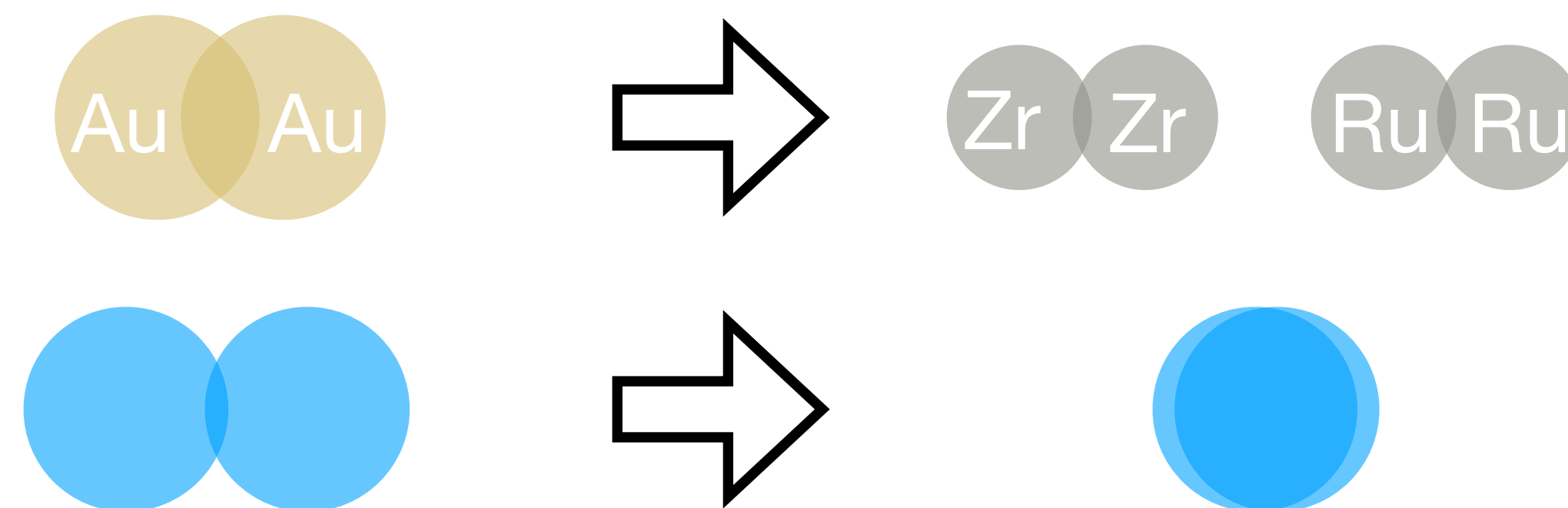


Suppression of charged hadrons strongly increases with $\langle N_{part} \rangle$

STAR, [PRL 91 \(2003\), 172302](https://arxiv.org/abs/0305195)

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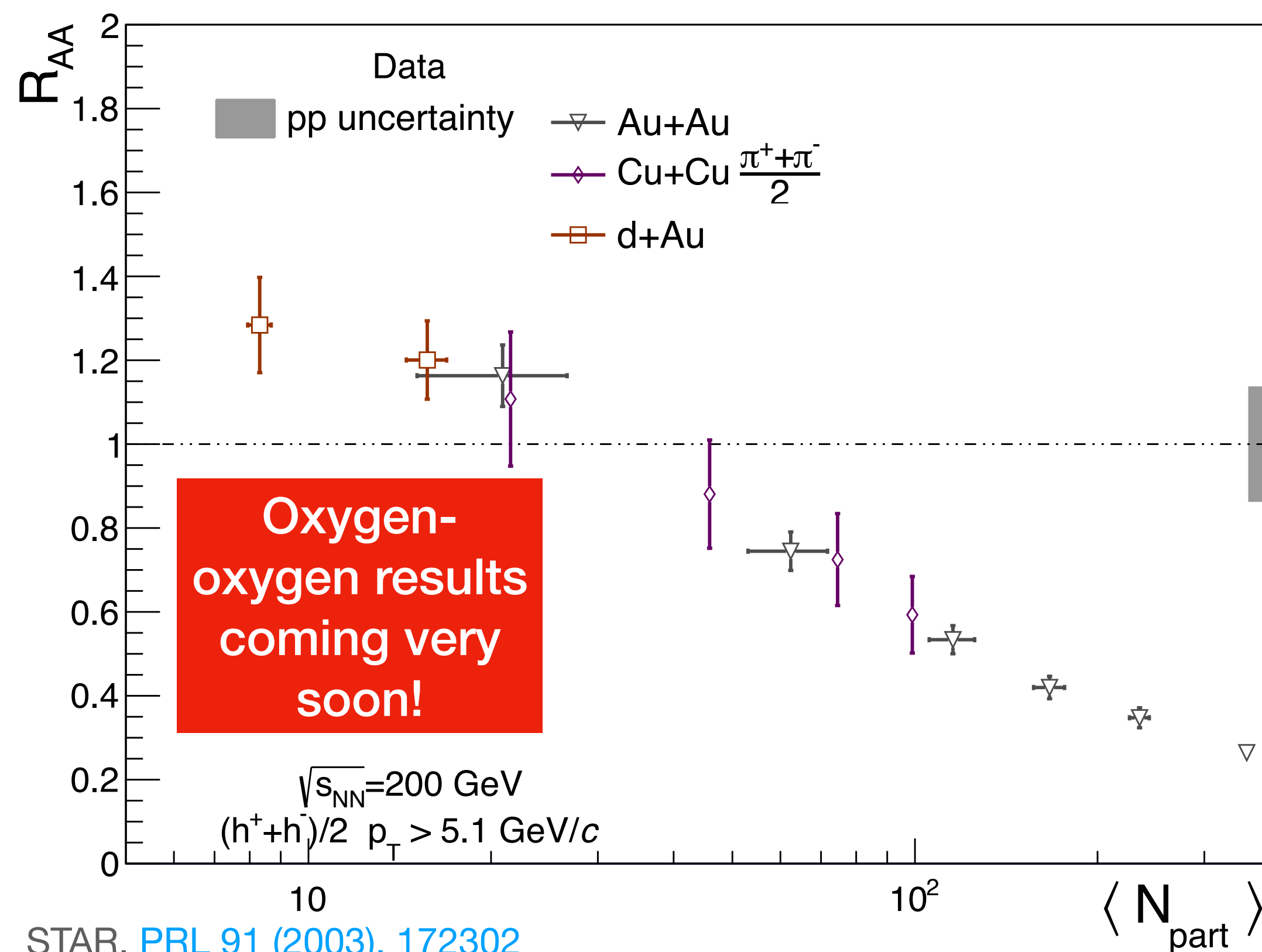
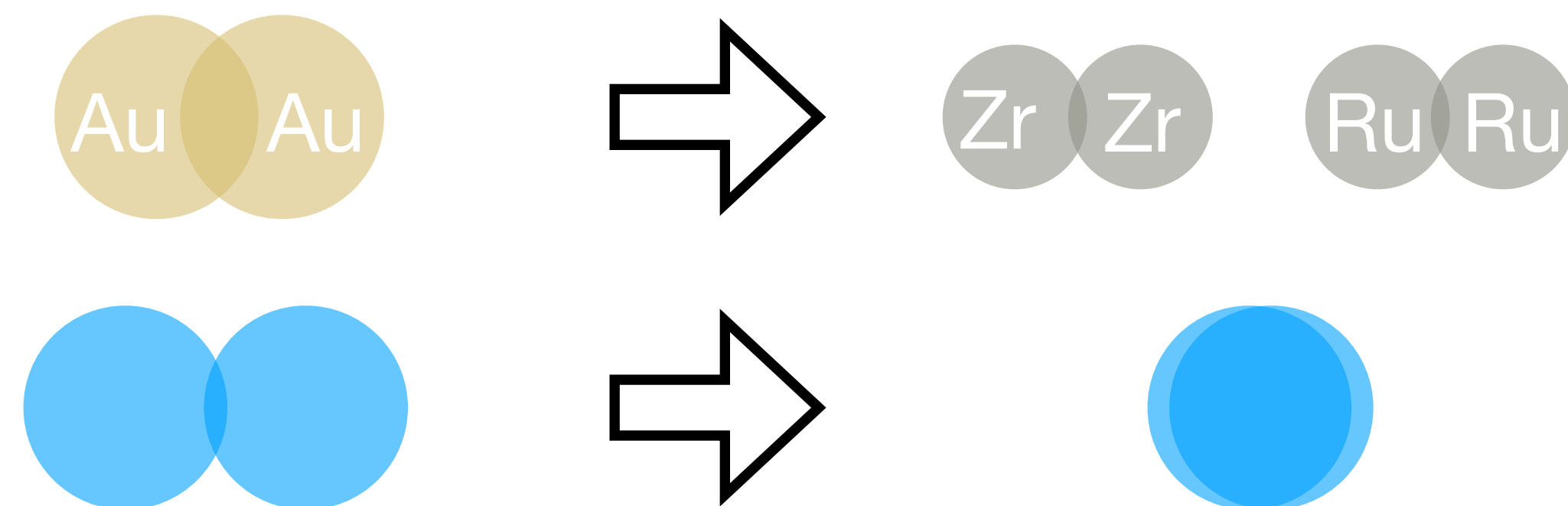


R_{AA} falls with $\langle N_{part} \rangle$ independent of collision species (system size)

STAR, [PRL 91 \(2003\), 172302](#)
 STAR, [PRL 91 \(2003\), 072304](#)
 STAR, [PRC 81 \(2010\), 054907](#)

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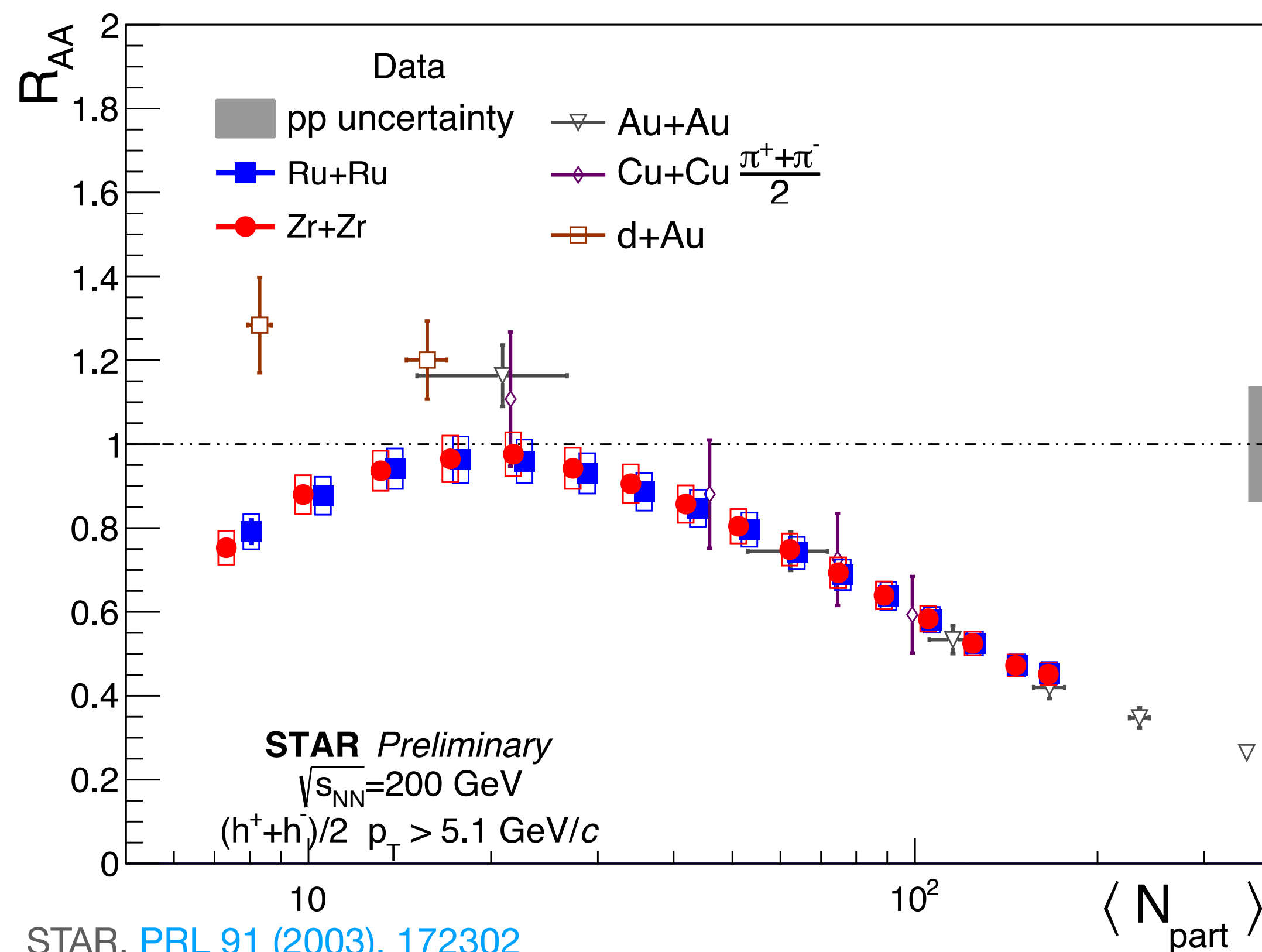


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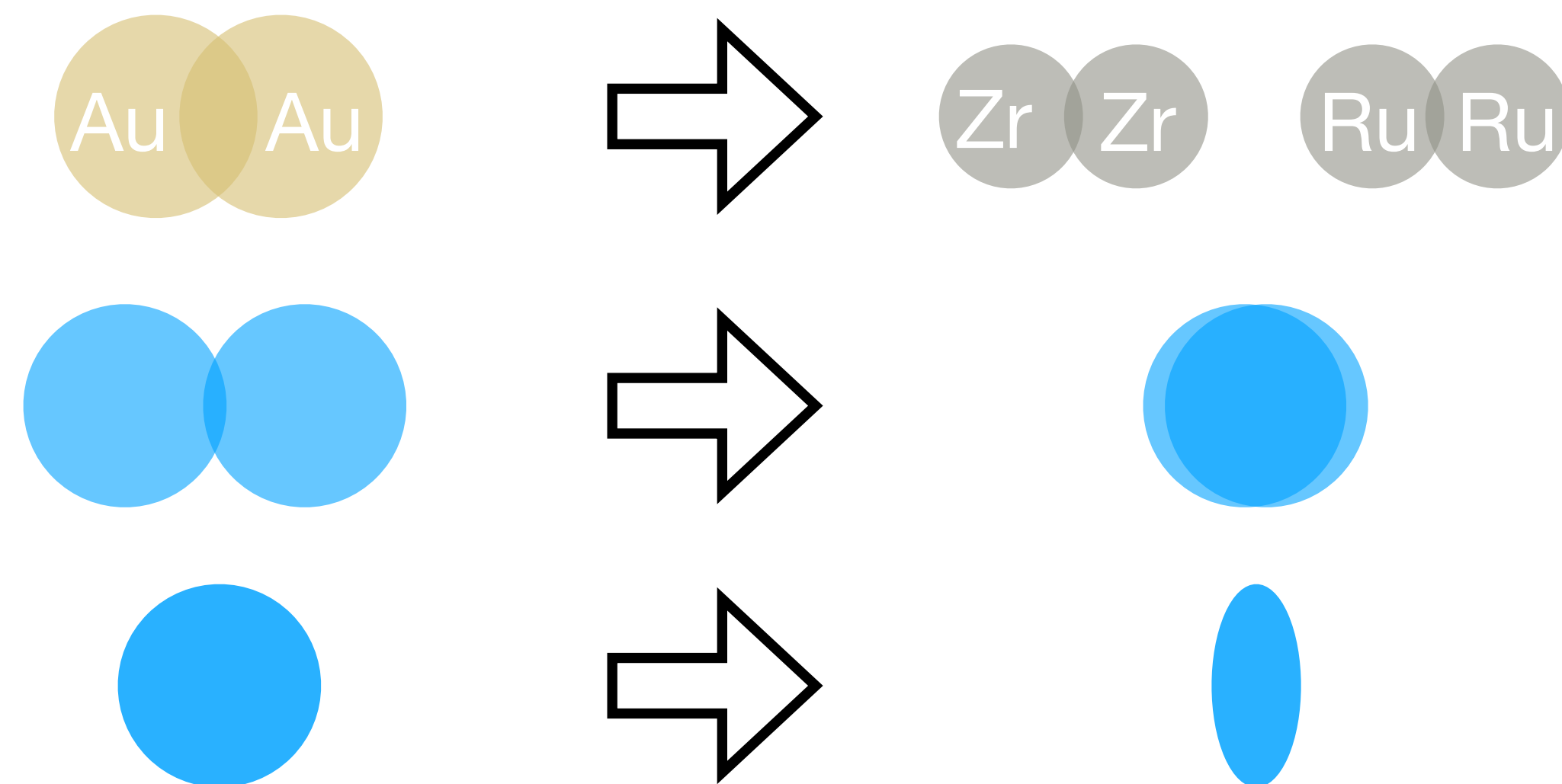
STAR, [PRL 91 \(2003\), 172302](#)
 STAR, [PRL 91 \(2003\), 072304](#)
 STAR, [PRC 81 \(2010\), 054907](#)

Inclusive yield modification in heavy-ion collisions

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STAR, [PRL 91 \(2003\), 172302](#)
 STAR, [PRL 91 \(2003\), 072304](#)
 STAR, [PRC 81 \(2010\), 054907](#)



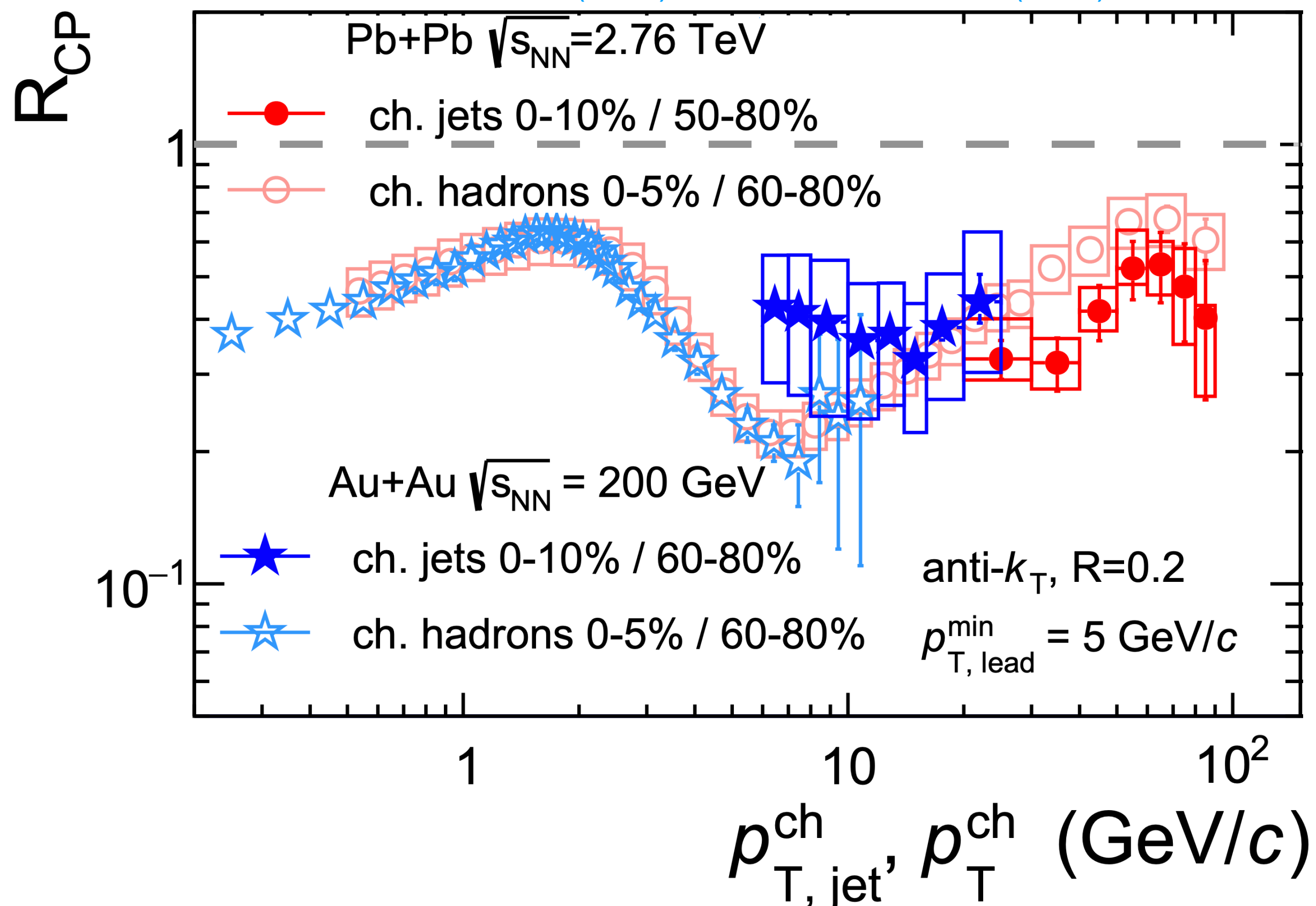
R_{AA} falls with $\langle N_{part} \rangle$ independent of collision species (system size), above ~ 20

Later: for given $\langle N_{part} \rangle$, how does geometry influence E -loss?

Inclusive yield modification of jets

$$R_{CP} = \frac{N_{ev}^P \langle T_{AA,P} \rangle \frac{d^2N^C}{d\eta dp_T}}{N_{ev}^C \langle T_{AA,C} \rangle \frac{d^2N^P}{d\eta dp_T}}, \quad \langle T_{AA} \rangle = \langle N_{coll} \rangle / \sigma_{inel}^{NN}$$

STAR, [PRC 102 \(2020\) 5, 054913](#), STAR, [PRL 91 \(2003\) 172302](#),
ALICE, [JHEP 03 \(2014\) 013](#), ALICE, [JHEP 09 \(2015\) 050](#)



Jet R_{AA} consistent with hadron R_{AA}

Strong suppression across p_T

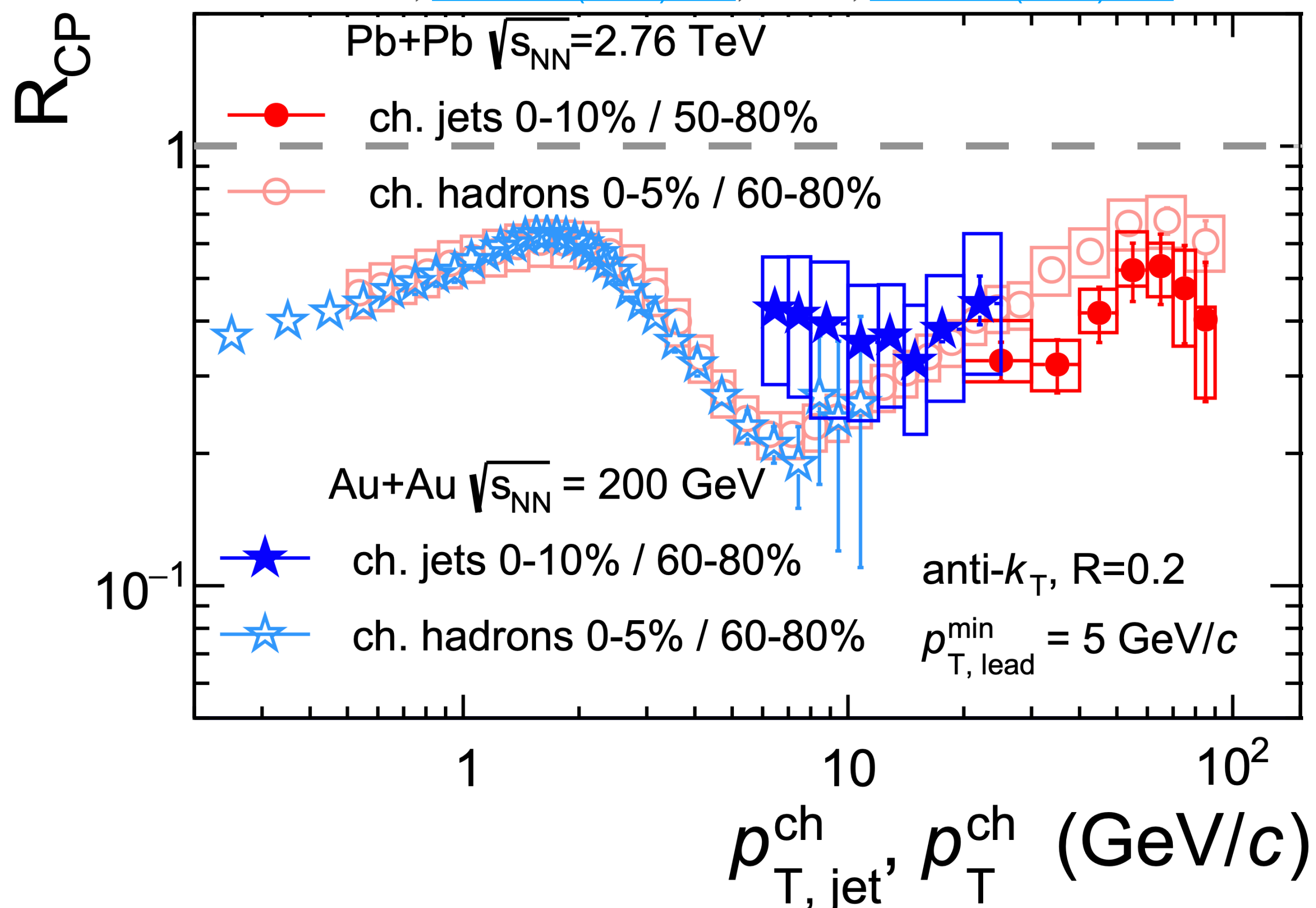
RHIC and LHC jets already have kinematic overlap

Similar quenching?

Inclusive yield modification of jets

$$R_{CP} = \frac{N_{ev}^P \langle T_{AA,P} \rangle \frac{d^2 N^C}{d\eta dp_T}}{N_{ev}^C \langle T_{AA,C} \rangle \frac{d^2 N^P}{d\eta dp_T}}, \quad \langle T_{AA} \rangle = \langle N_{coll} \rangle / \sigma_{inel}^{NN}$$

STAR, [PRC 102 \(2020\) 5, 054913](#), STAR, [PRL 91 \(2003\) 172302](#),
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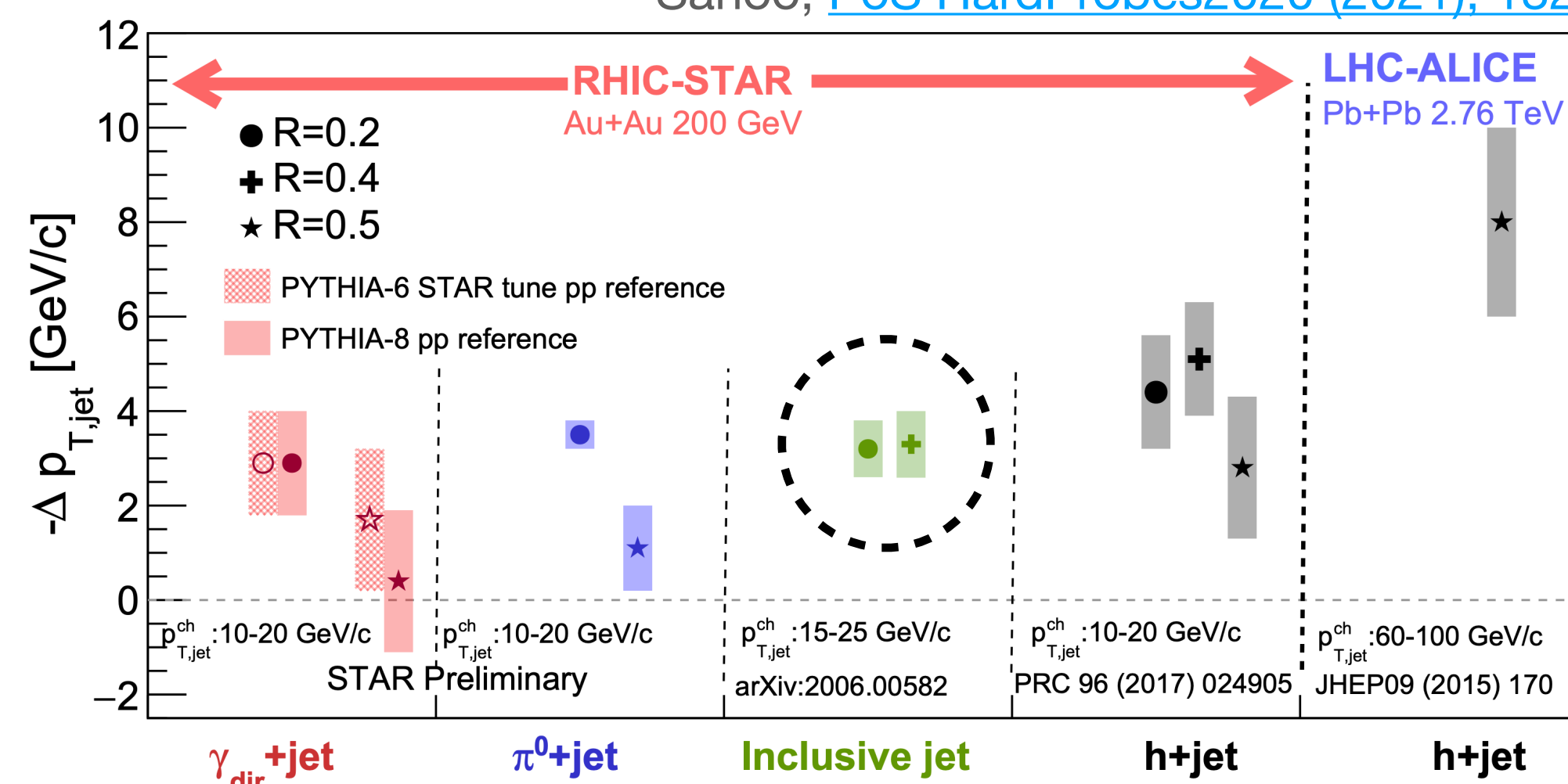
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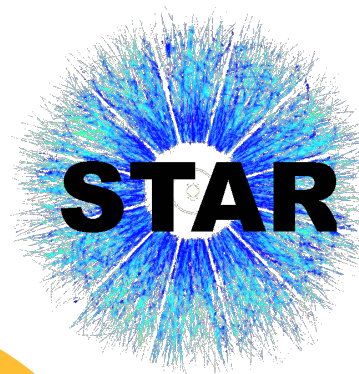
RHIC and LHC jets already have kinematic overlap

Similar quenching?

Absolute, smaller. Relative, *larger!*

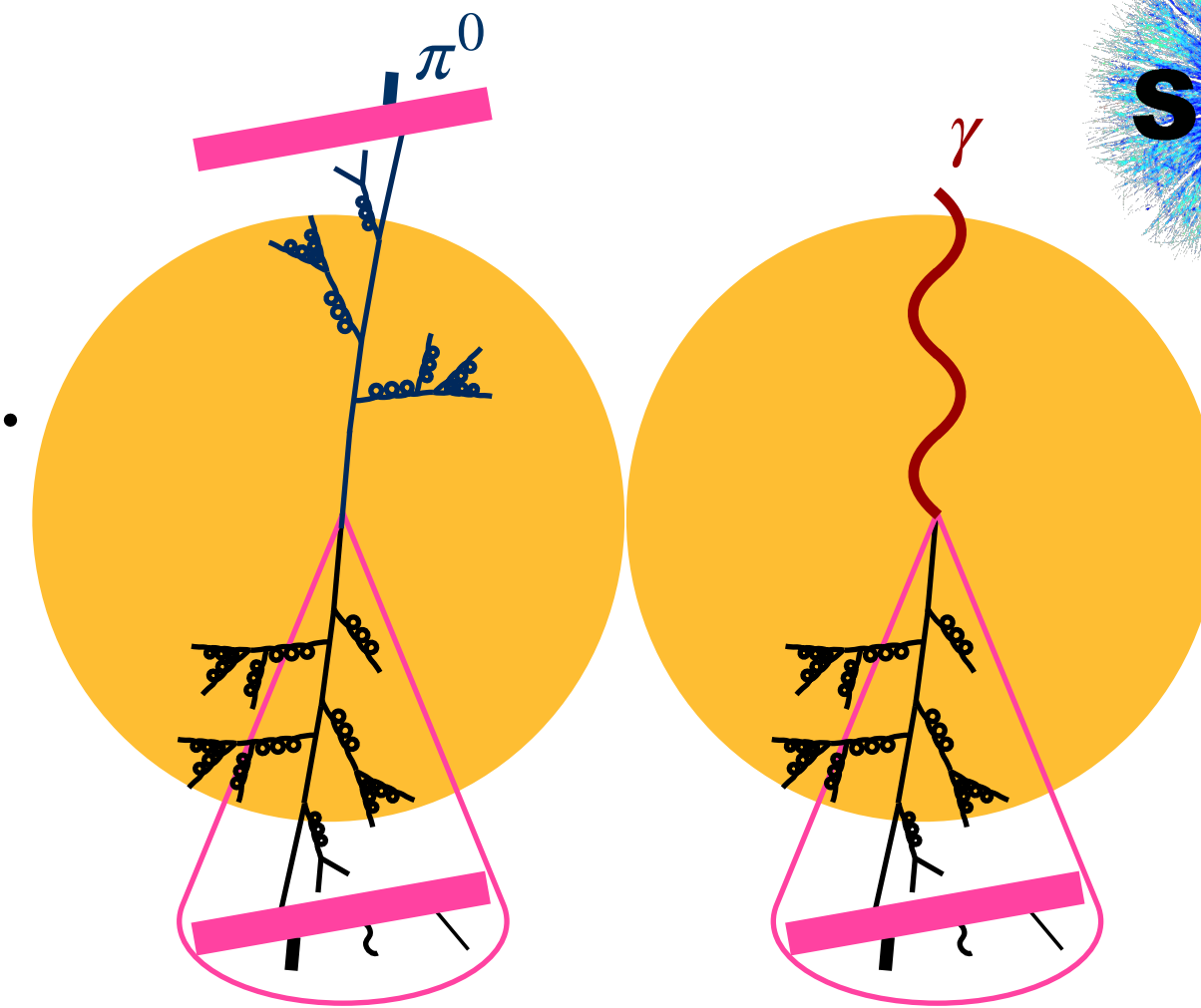
Sahoo, [PoS HardProbes2020 \(2021\), 132](#)



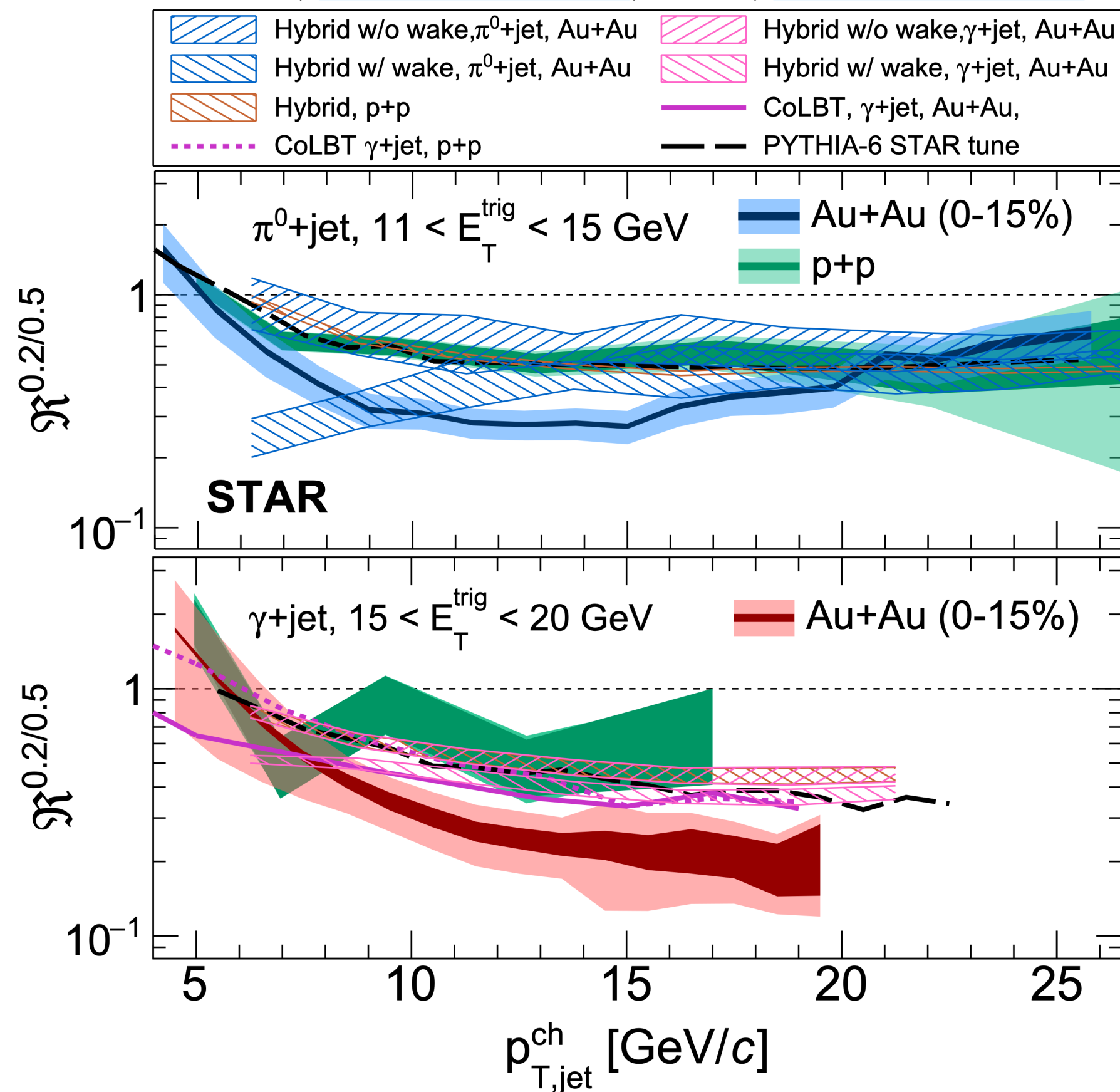


Semi-inclusive yield modification

$$I_{AA} = \frac{Y^{AA}(p_{T,jet}^{ch}, R)}{Y^{pp}(p_{T,jet}^{ch}, R)}, Y(p_{T,jet}^{ch}, R) = \frac{1}{N_{trig}} \int_{3\pi/4}^{5\pi/4} d\Delta\phi \left[\frac{d^2 N_{jet}(R)}{dp_{T,jet}^{ch} d\Delta\phi} \right]_{E_T^{trig} \in [E_T^{min}, E_T^{max}]}$$



STAR, [arXiv:2309.00145](https://arxiv.org/abs/2309.00145), STAR, [arXiv:2309.00156](https://arxiv.org/abs/2309.00156)

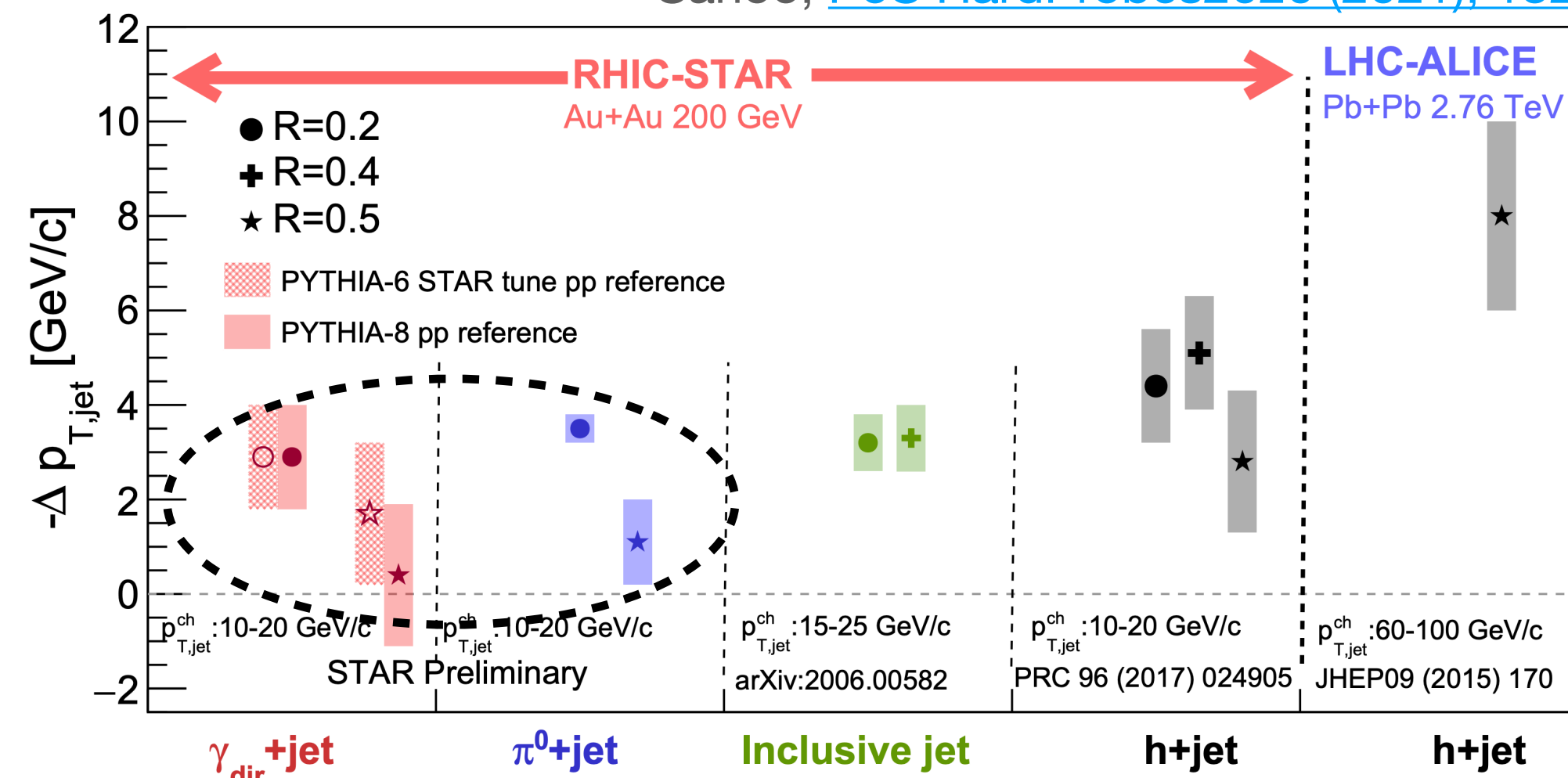


Recoil jet yield suppression in Au+Au, stronger in small R jets

Clear observation of *intra-jet broadening*

Models unable to quantitatively describe the effect

Sahoo, [PoS HardProbes2020 \(2021\), 132](https://arxiv.org/abs/2012.0132)



Path-length-dependent quenching

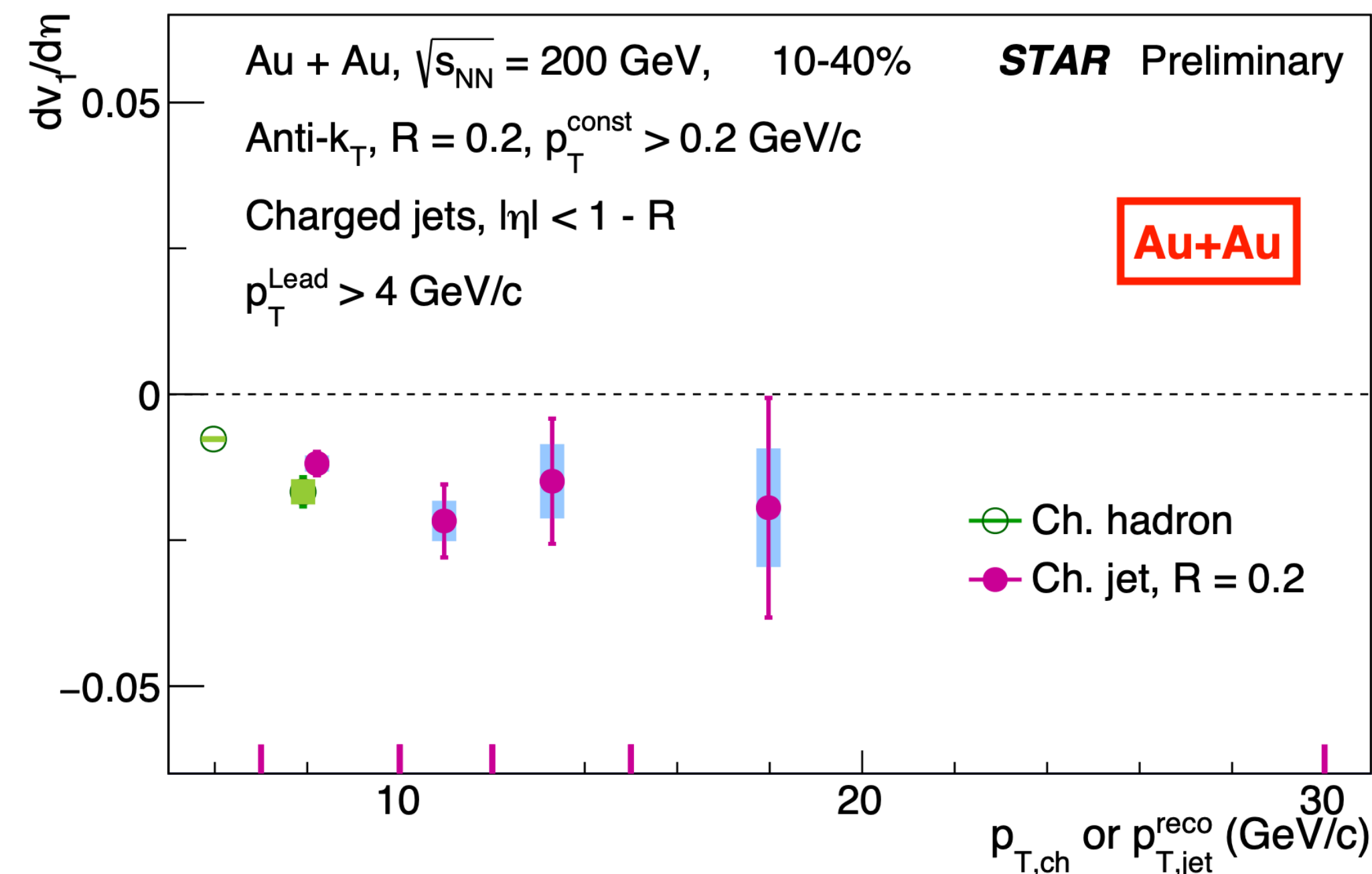
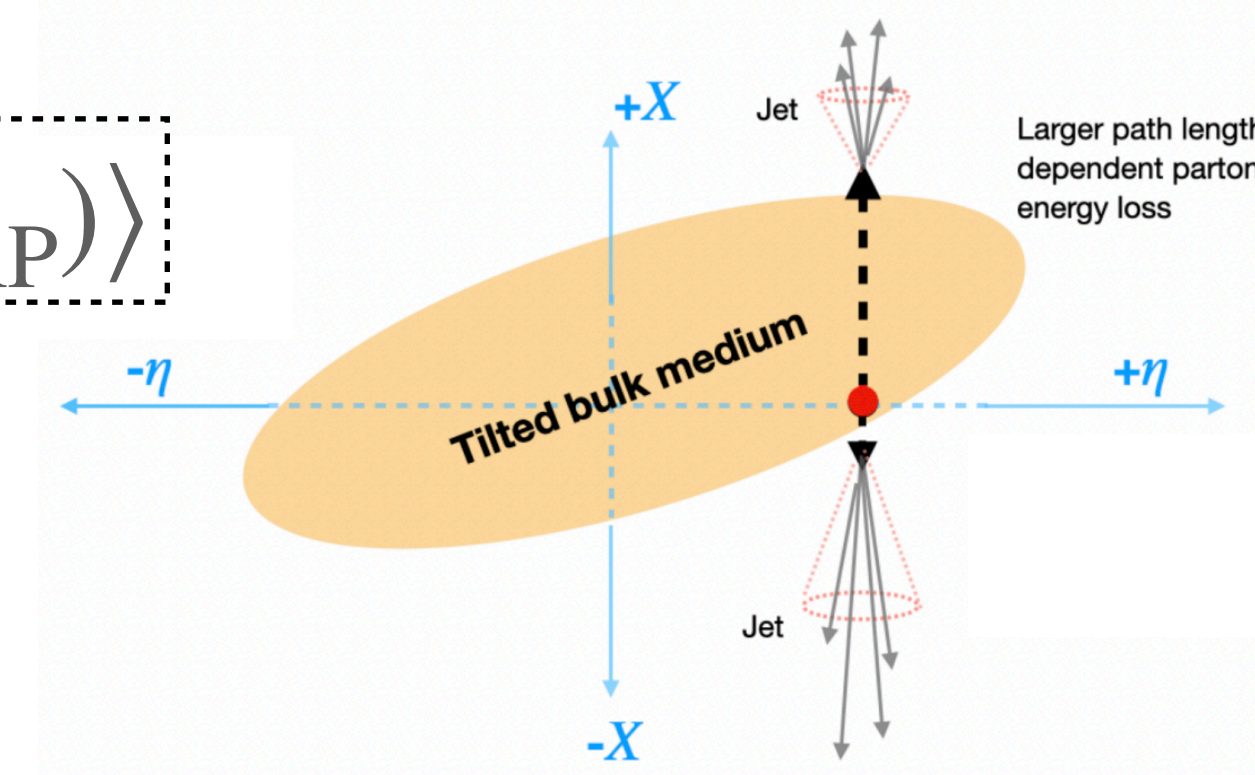
Bulk is tilted in heavy-ion collisions^{1,2} causing asymmetric paths for isotropically produced hard probes

Jet v_1 : a new observable to probe *pathlength-dependent energy loss* in QGP

Clear v_1 signal for all studied jet R , p_T : 7 – 20 GeV/c, in **Au+Au** data, similar for **isobar** systems as well

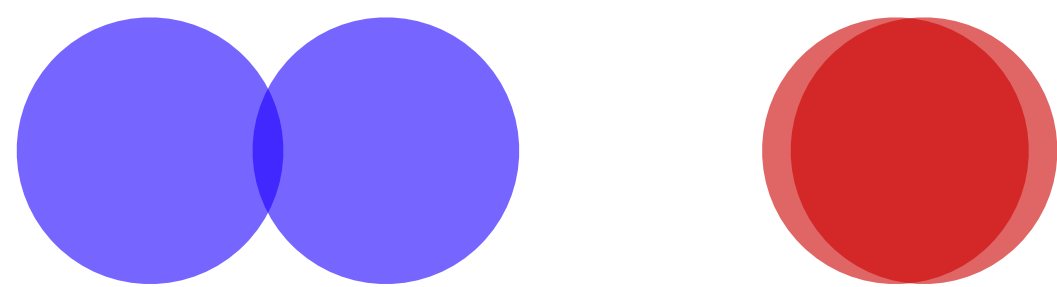
Outlook: event-shape engineering with multiplicity fluctuations

$$v_1(p_T, y) = \langle \cos(\phi - \Psi_{RP}) \rangle$$

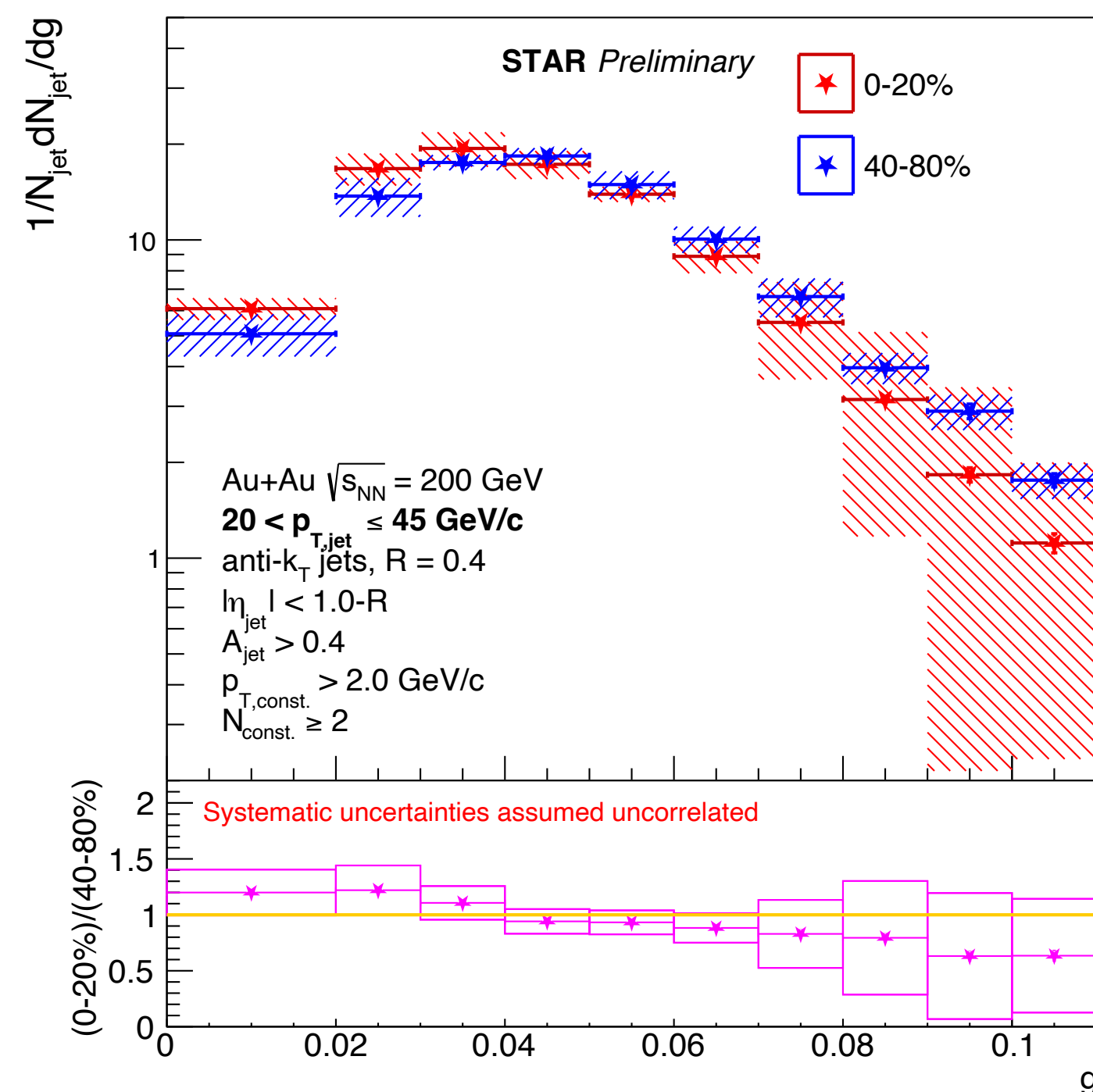


Generalized angularities

With MultiFold

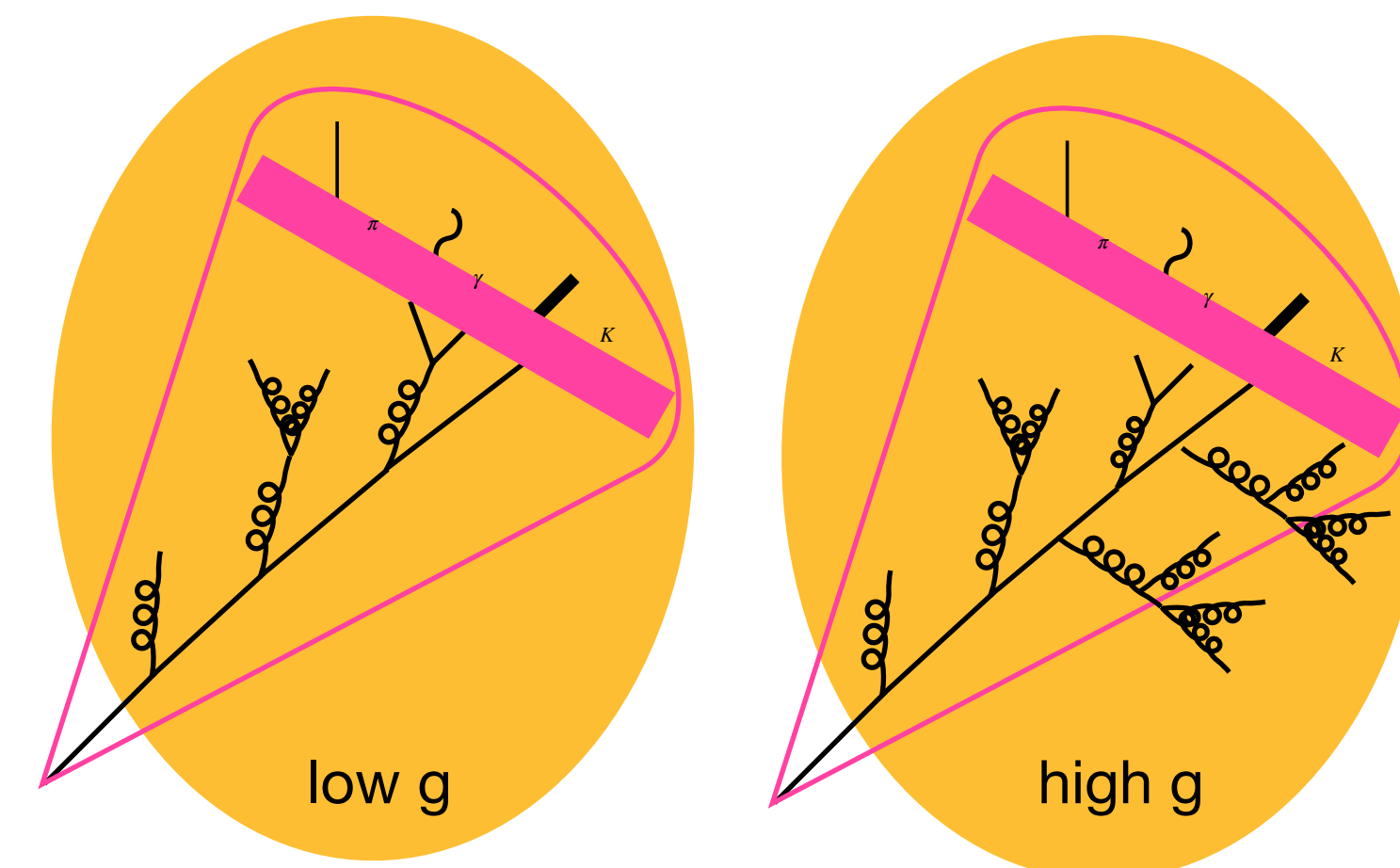


Data corrected using MultiFold in 7D



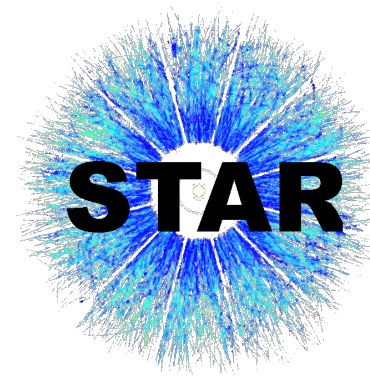
$$\lambda_{\beta}^{\kappa} = \sum_{\text{cons} \in \text{jet}} \left(\frac{p_{\text{T,cons}}}{p_{\text{T,jet}}} \right)^{\kappa} \Delta R (\text{cons, jet})^{\beta}$$

$$g = \lambda_1^1 = \frac{\sum_{\text{trk} \in \text{jet}} p_{\text{T,trk}} \Delta R}{p_{\text{T,jet}}}$$



Generalized angularities allow tunable contribution of momentum, angular scales in IRC safe way

With conservative systematic uncertainties *in biased pop.*, girth in peripheral and central collisions are consistent



Charm quark energy loss,
diffusion, fragmentation
modification in medium
with *charmed-jet yields*

Hadronization
mechanism
with *flavor correlators*

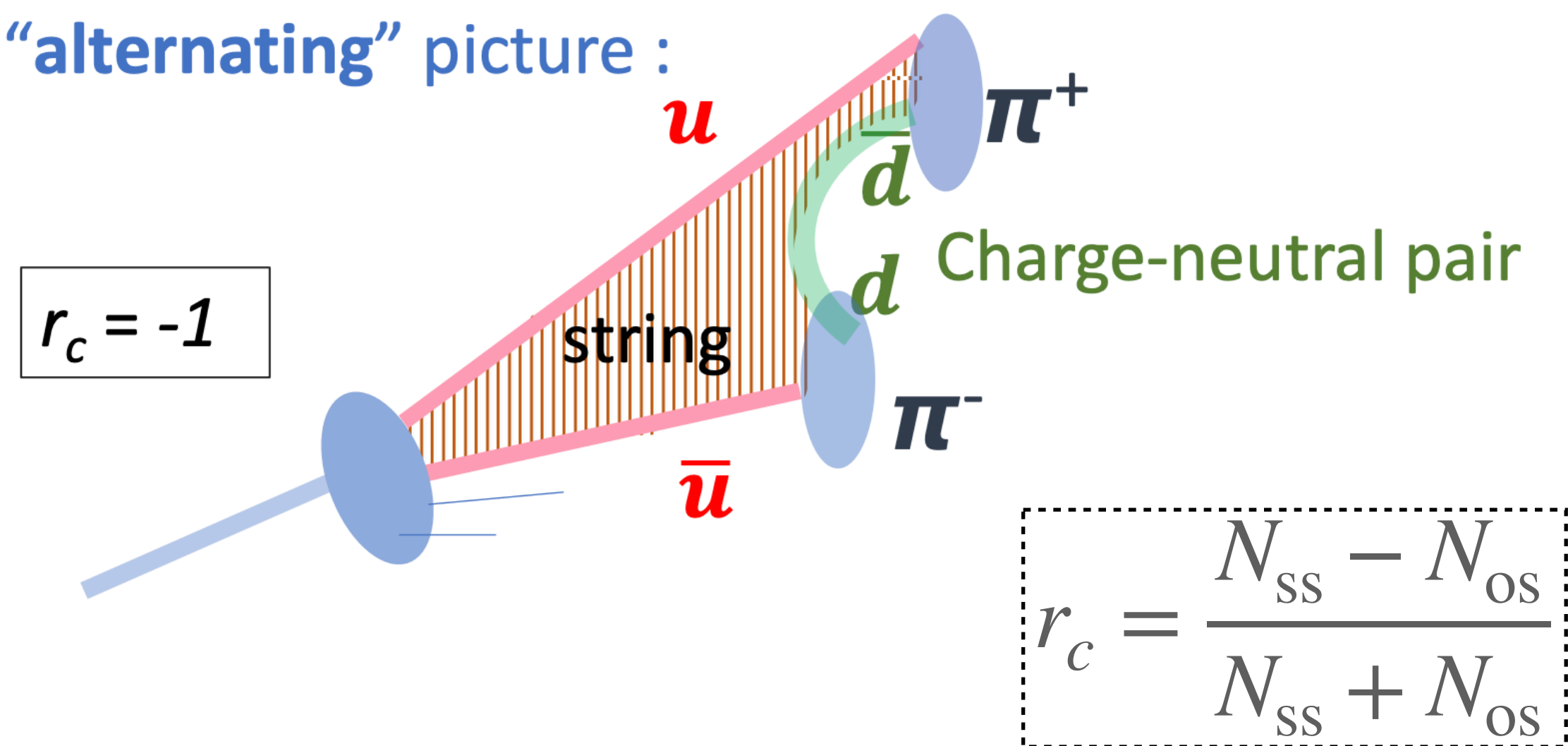
Constituent identity

.....

Hadrochemistry modification
via medium response
with *baryon-to-meson ratios*

Assessing fragmentation mechanism in jets

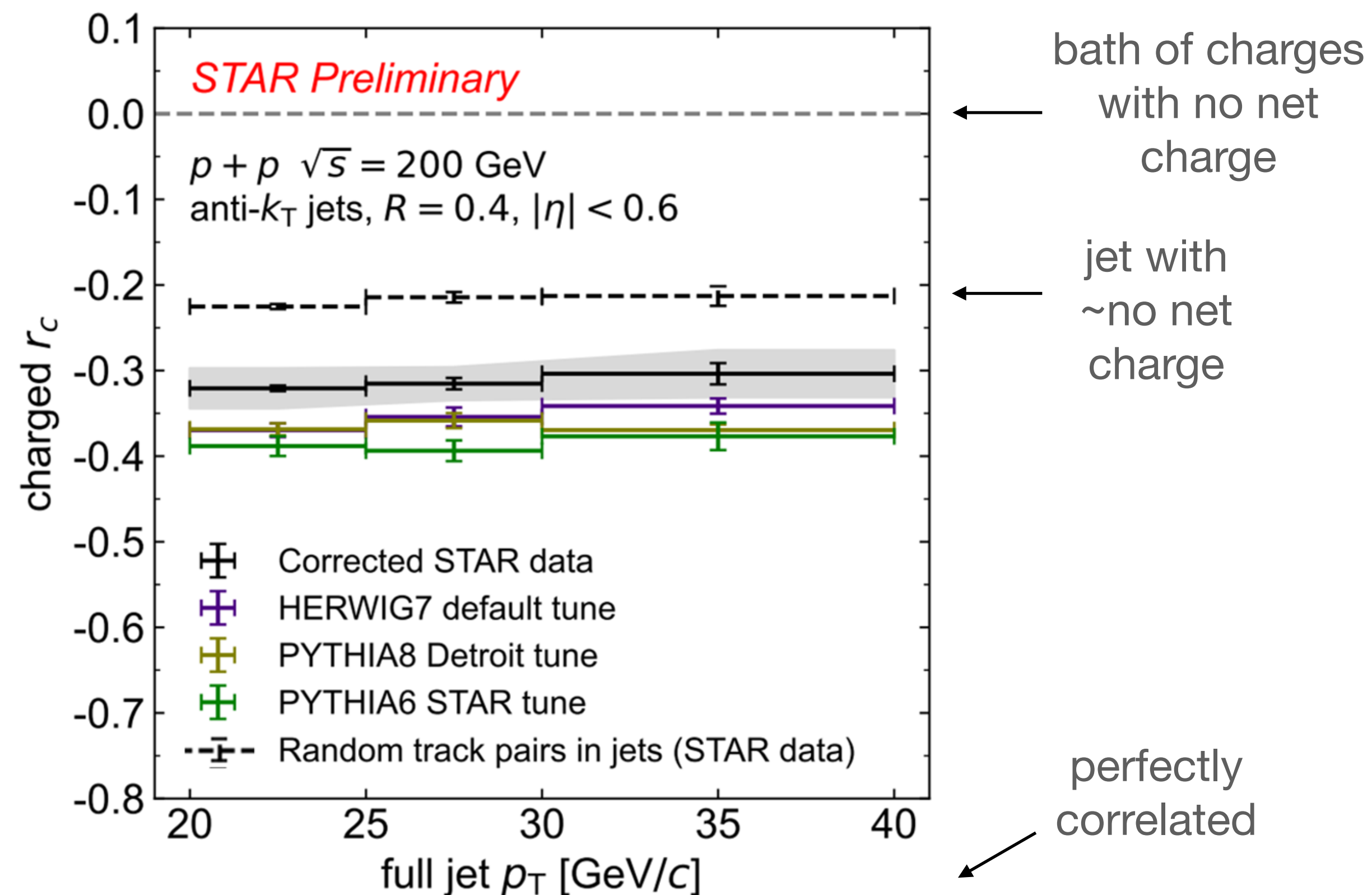
“alternating” picture :



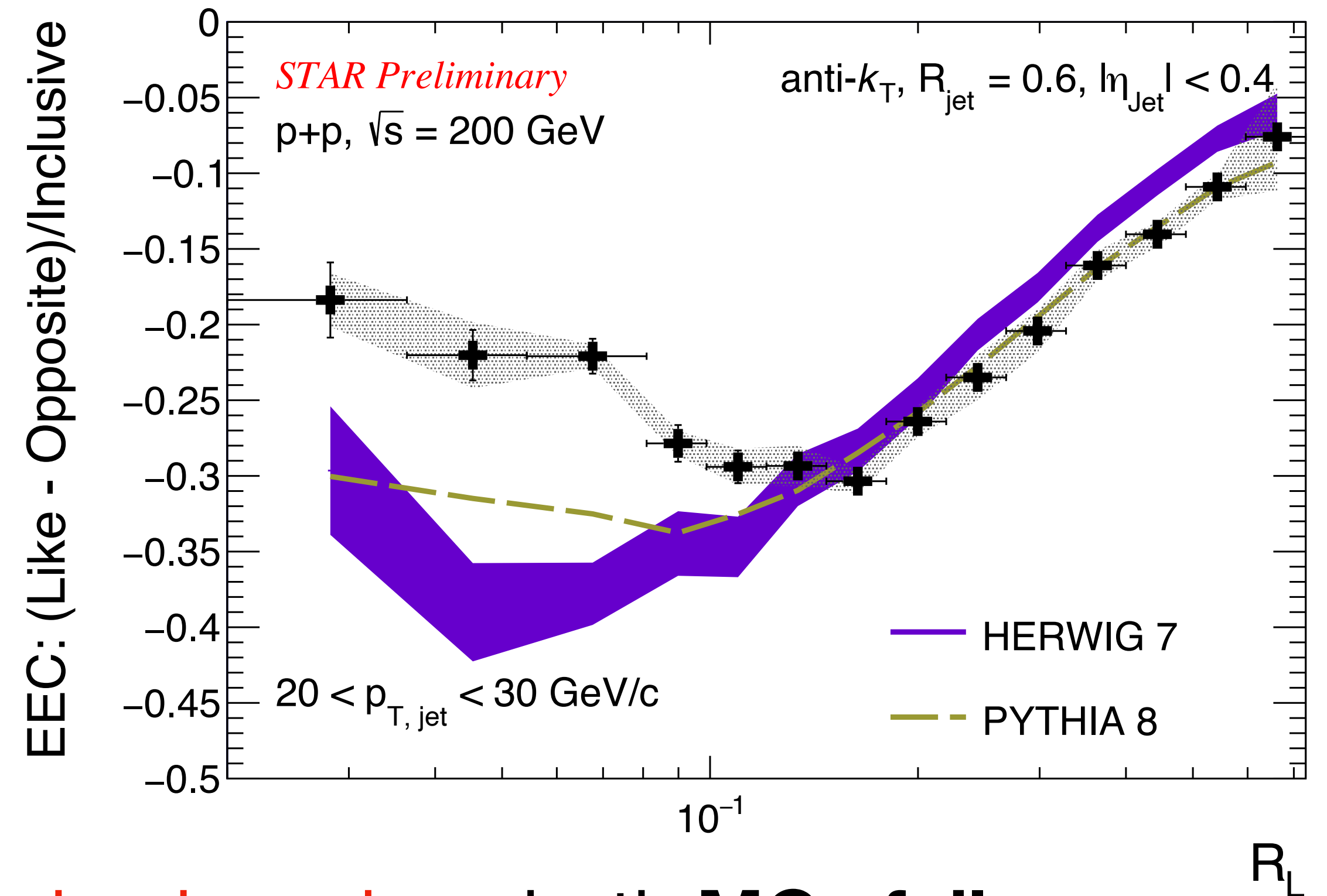
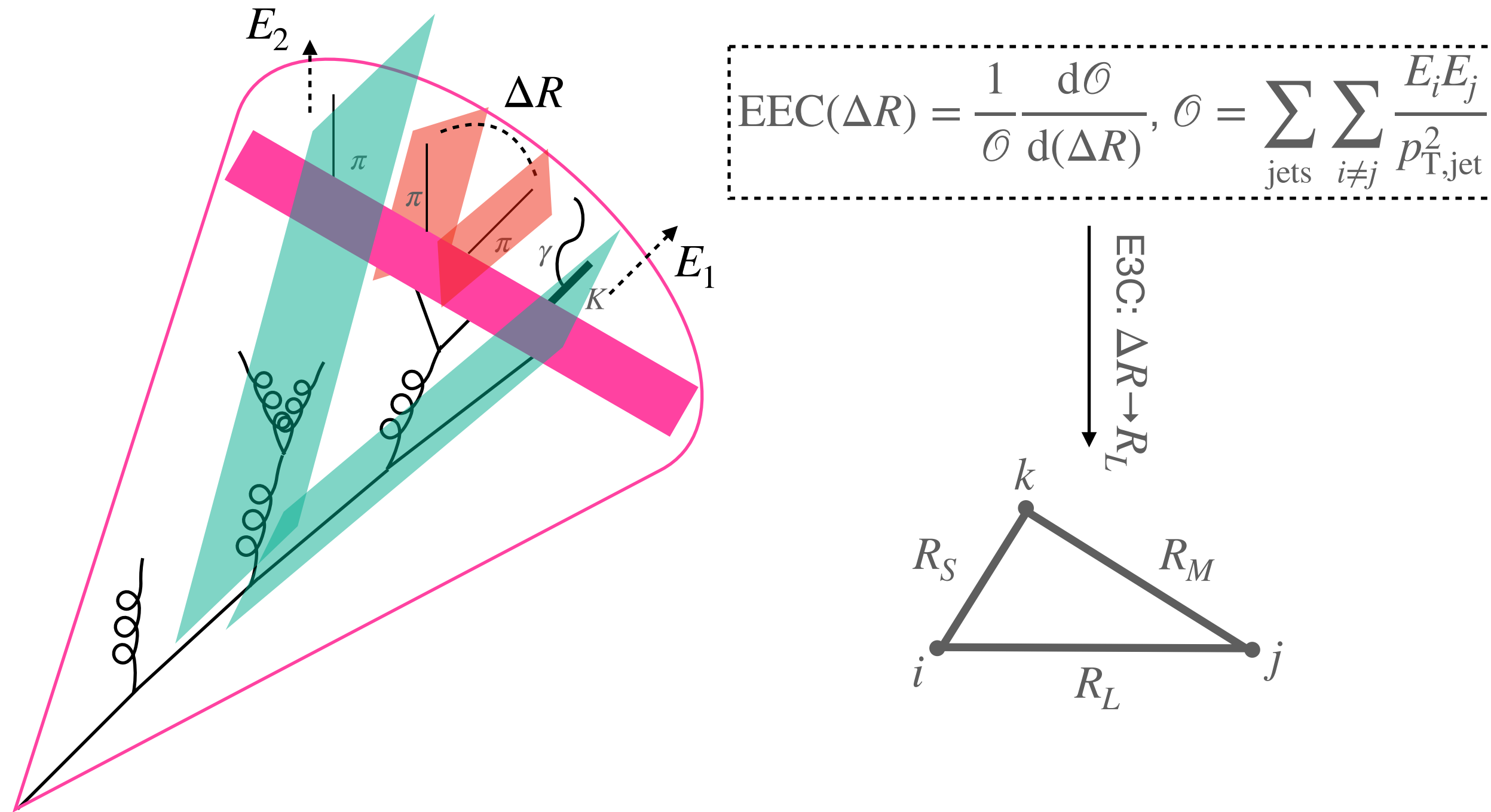
Leading charge correlator, r_c , can probe contribution of string-like fragmentation

First pp measurement: MCs predict more charge correlation than supported by data

Outlook: extension to heavy-ion collisions ongoing



Assessing fragmentation mechanism in jets

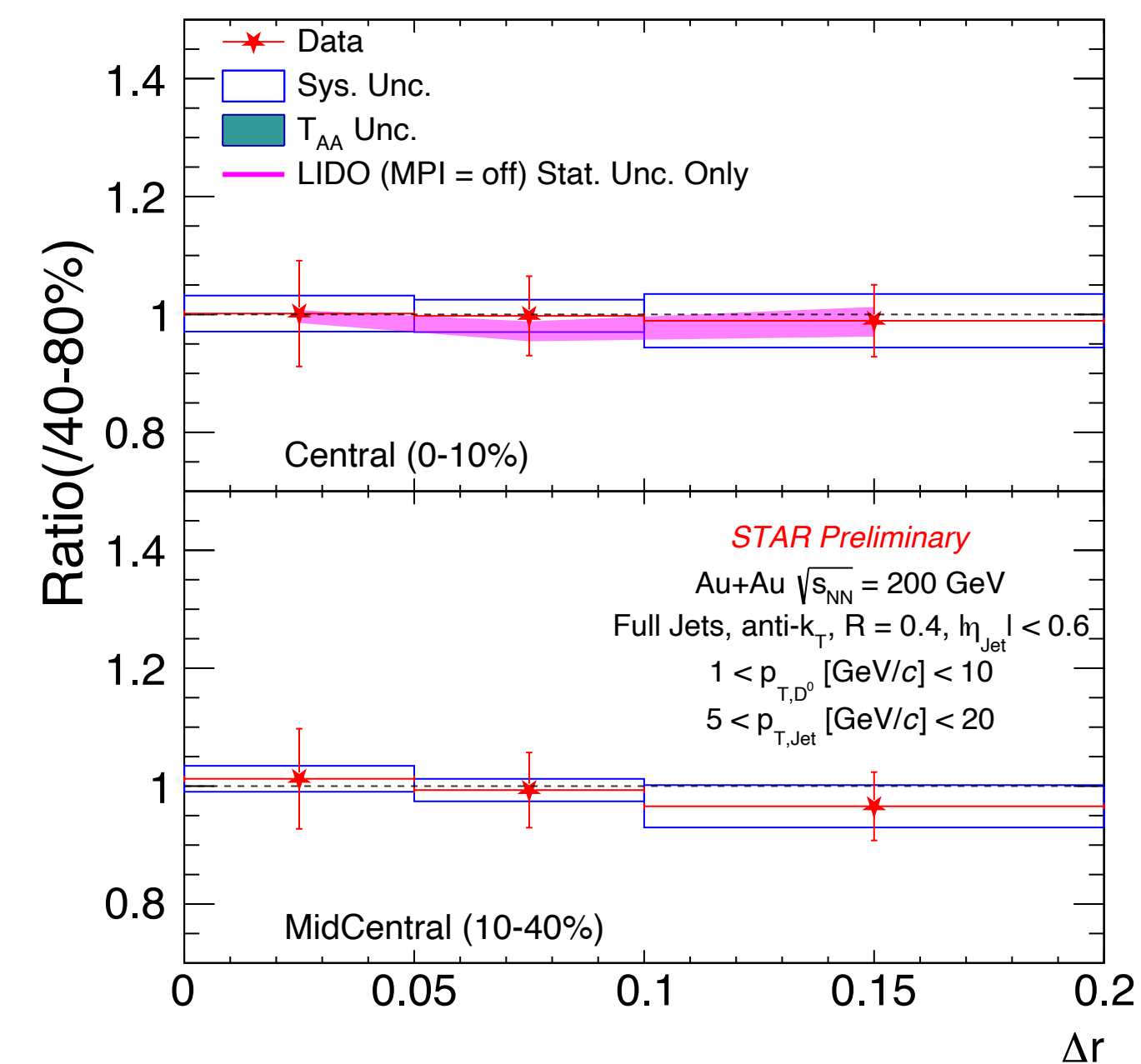
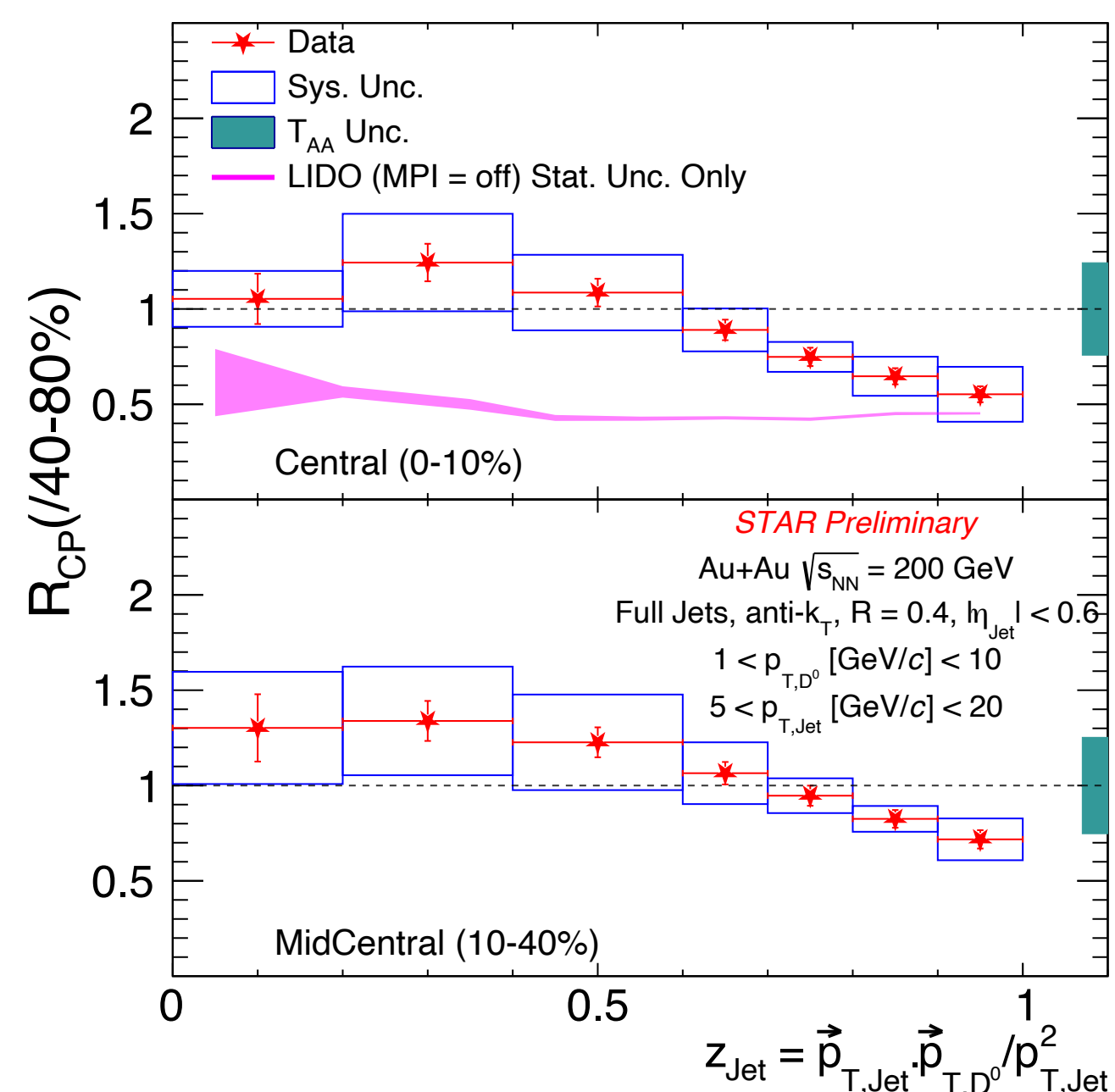
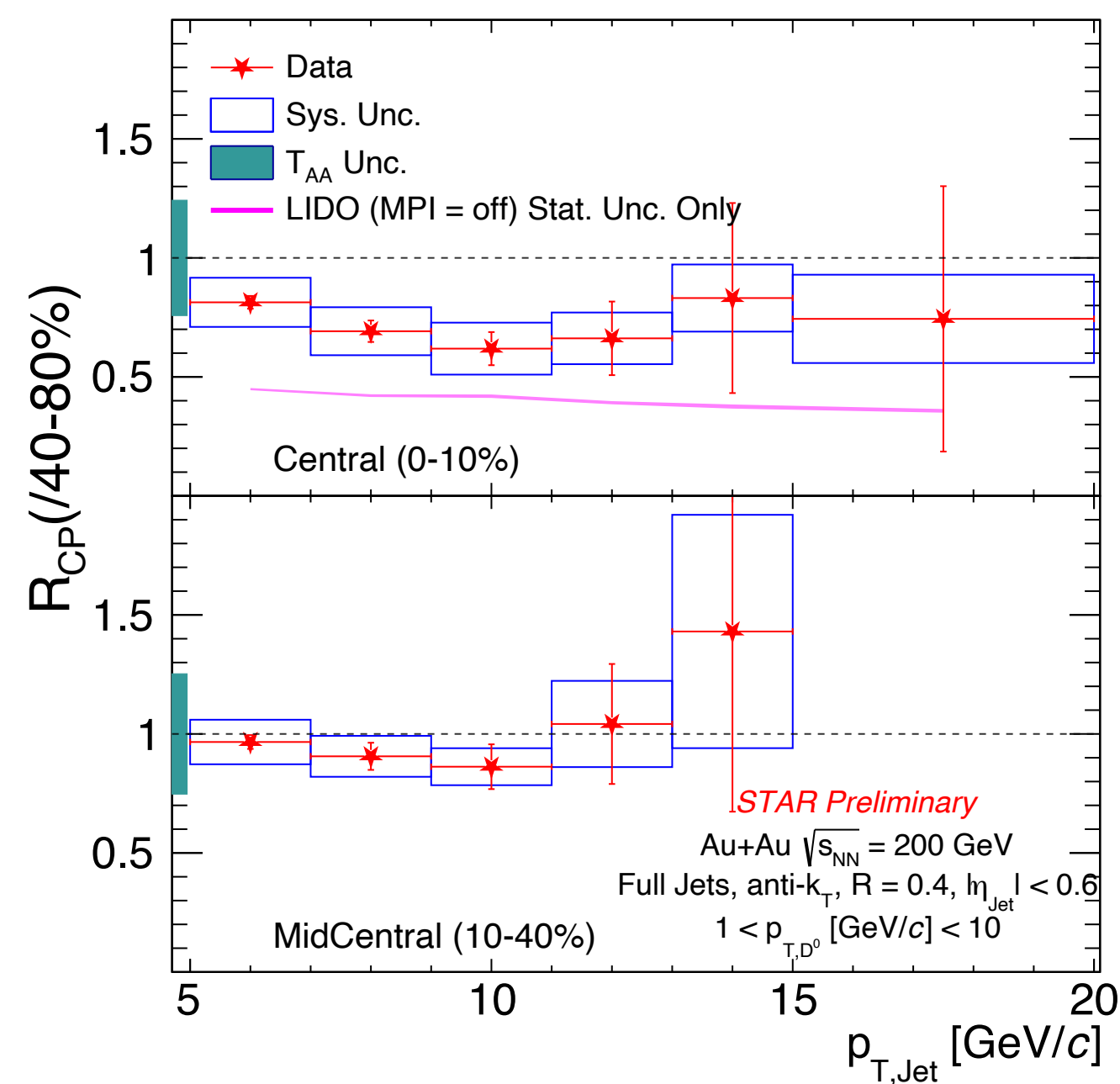


New charge-dependent EEC & E3C — in **hadronic regime**, both **MCs fail to capture data**; qualitatively consistent with behavior seen in r_c

Outlook: Extension to heavy-ion collisions ongoing

$$D^0 = c\bar{u}$$

D⁰-jet spectra, profile, fragmentation

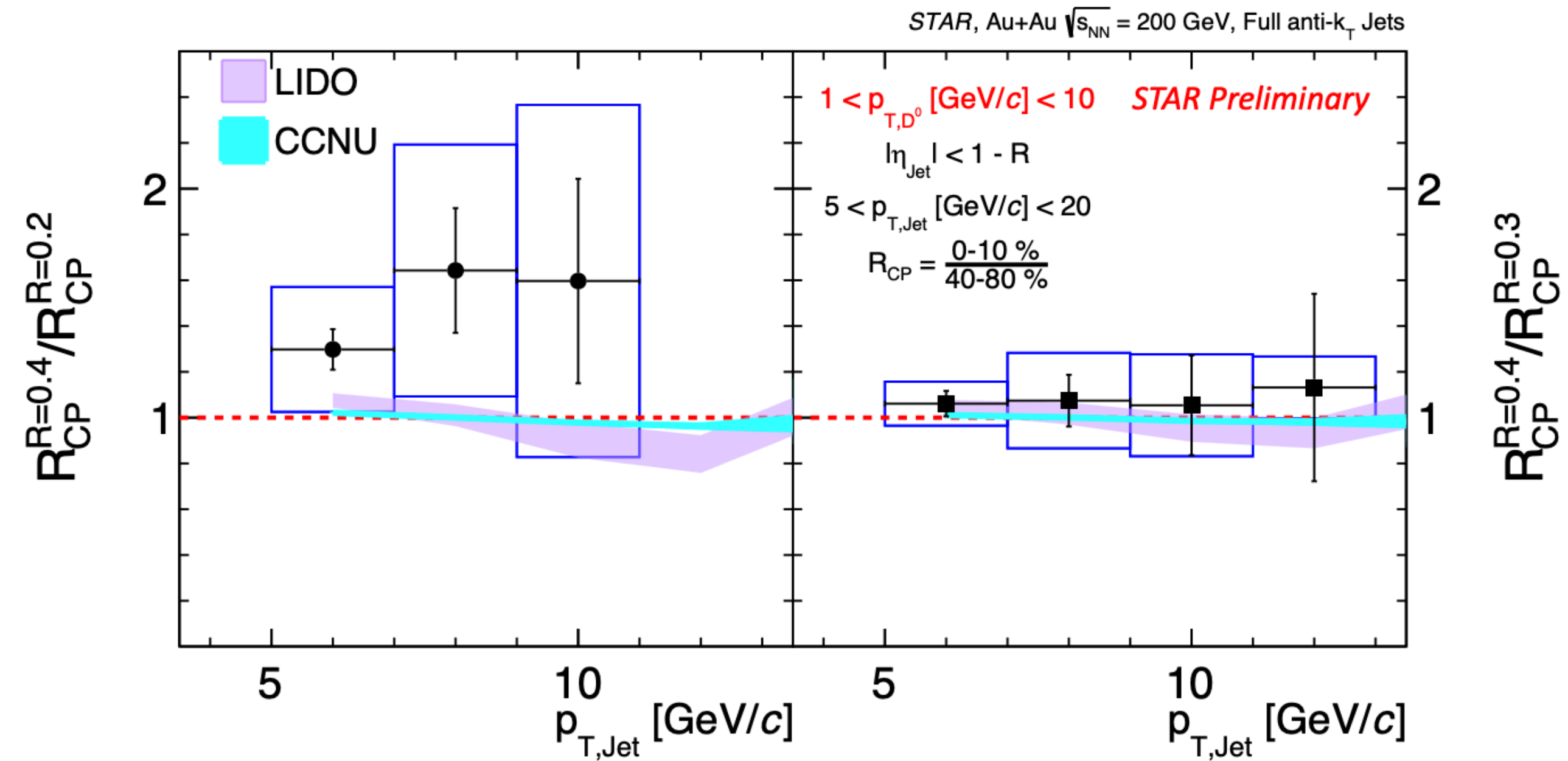
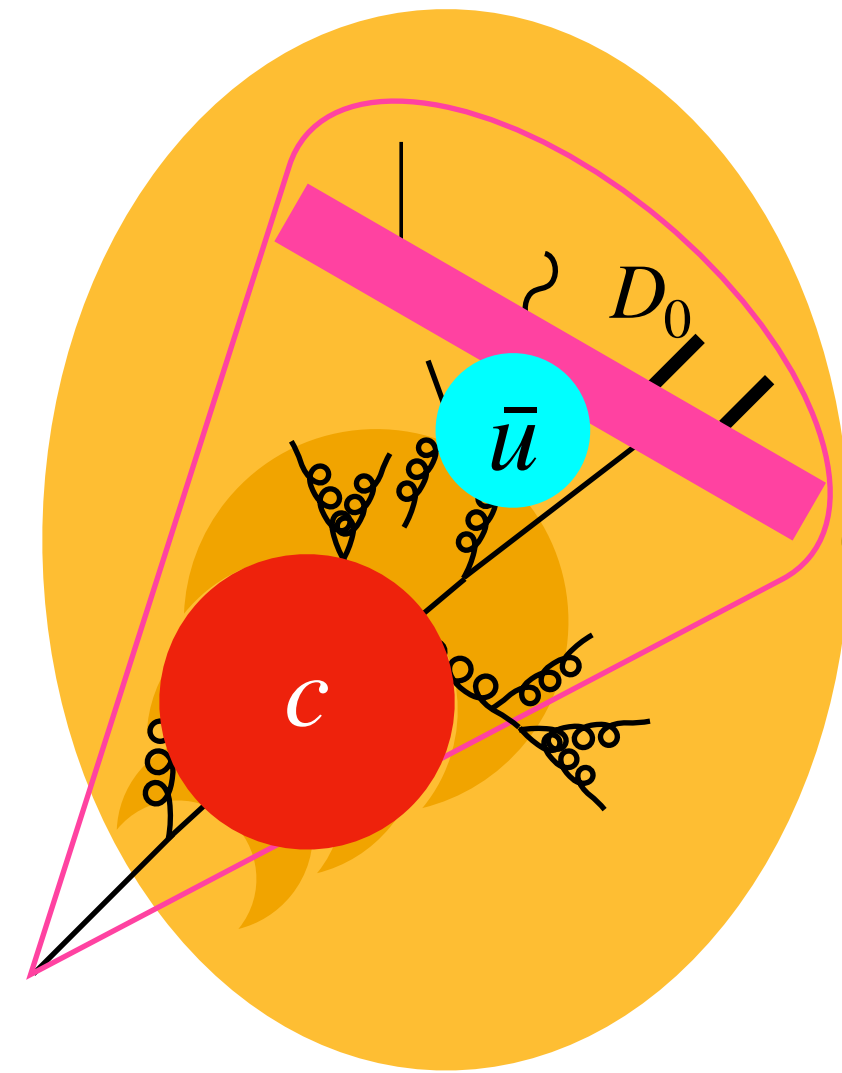
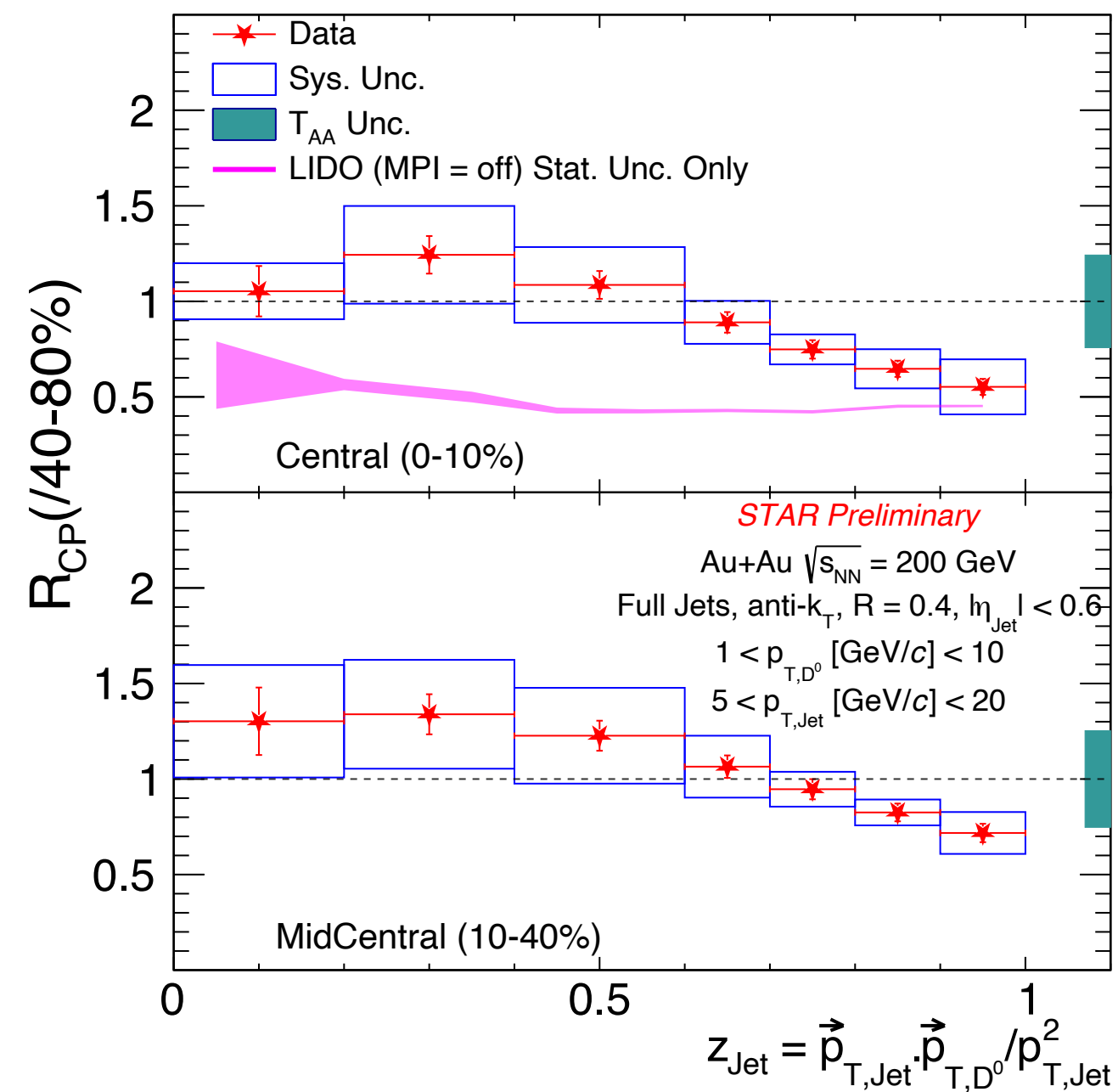


Testing charm quark energy loss, diffusion, and fragmentation modification

Hint of yield suppression in central. Hard-fragmented charm jets suppressed. No diffusion.

Model including radiative and collisional energy loss during heavy quark evolution underpredicts central yields — MPI might be important for D⁰ p_T this low

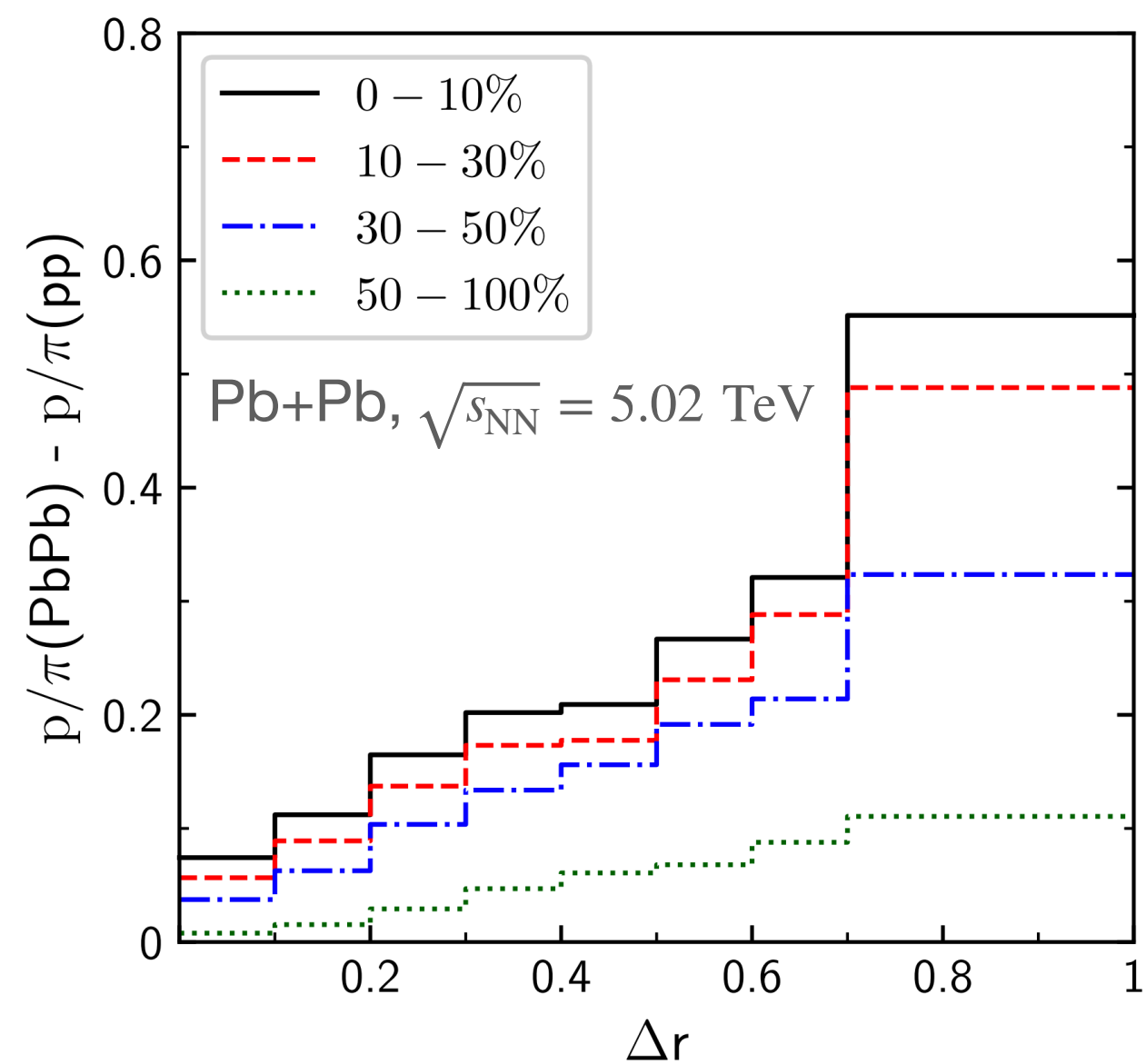
Recovering charm-associated radiation



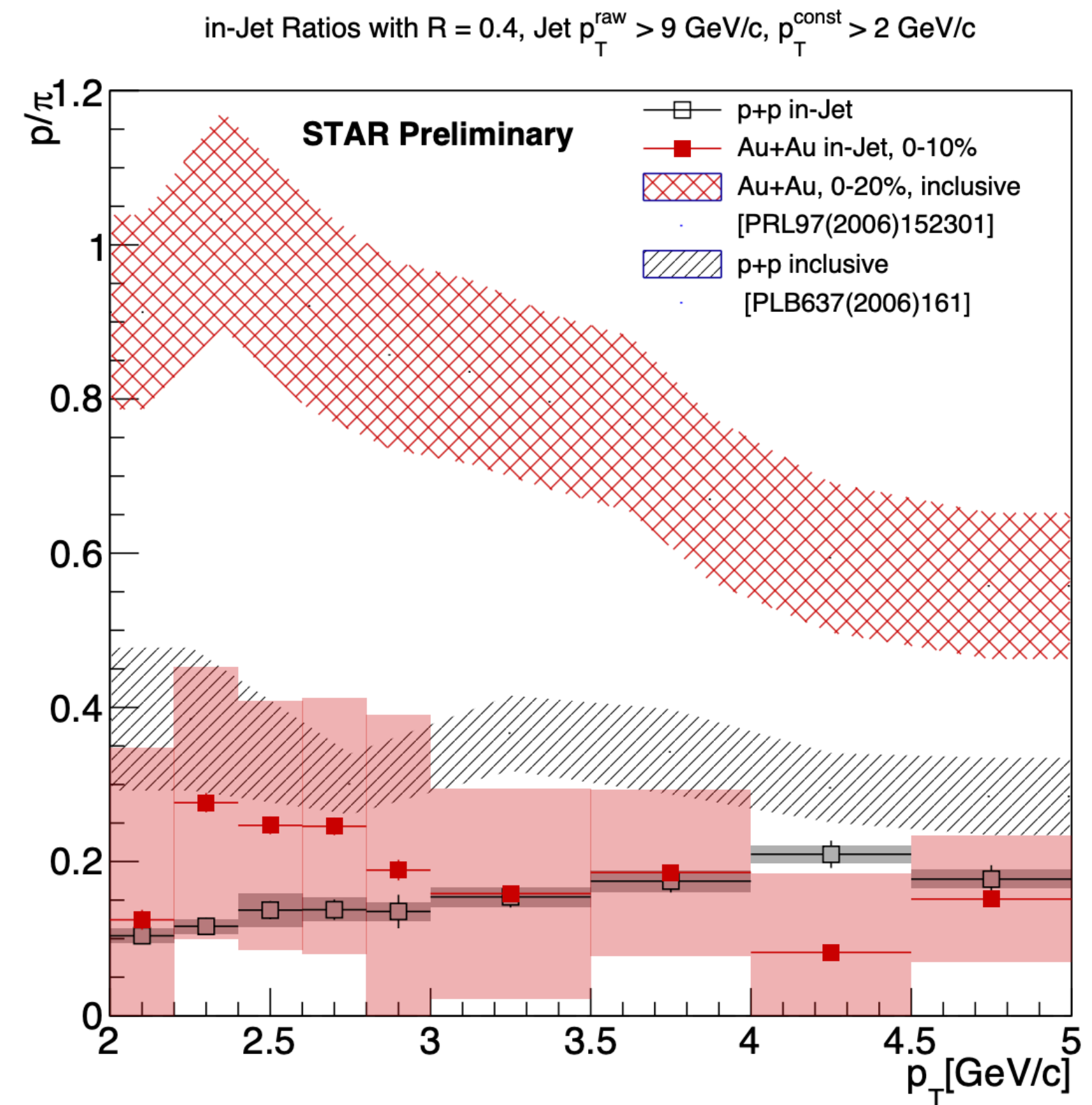
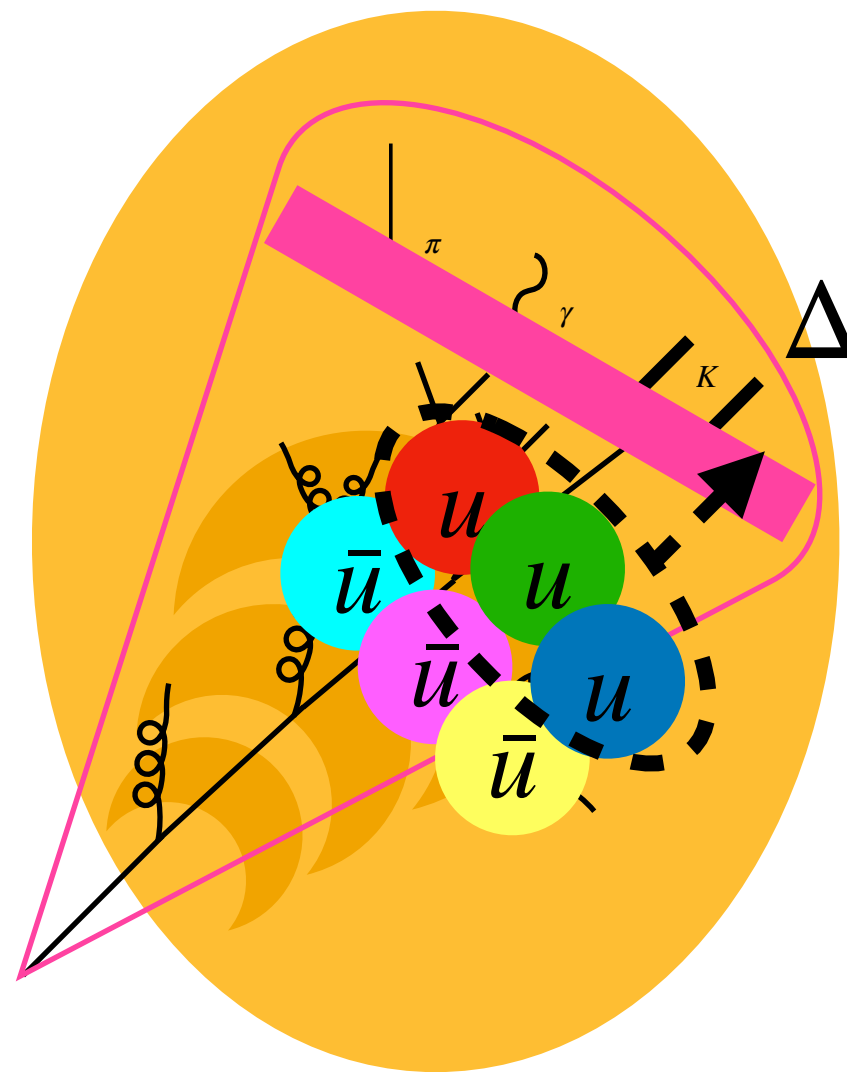
Wider jets \rightarrow more medium interaction/ E -loss $\implies ratio < 1$,
 but recover more energy + more potential for medium response $\implies ratio > 1$
 Observe: **No radius dependence of R_{CP}** within uncertainties.
 Agrees with models predicting minimal R -dependence of suppression.

Outlook: measuring generalized angularities

Searching for medium response



Luo, Mao, Qin, Wang, Zhang,
[PLB 837 \(2023\) 137638](#)



Possible expectation of *parton coalescence in jet*: enhanced baryon-to-meson ratio in A+A (left)

No observed modification of *in-jet* p/π ratio for $R = 0.2 - 0.4$ jets, after extension to lower constituent threshold (right)

Outlook: finalizing for publication in near future

Physics from the STAR jet program

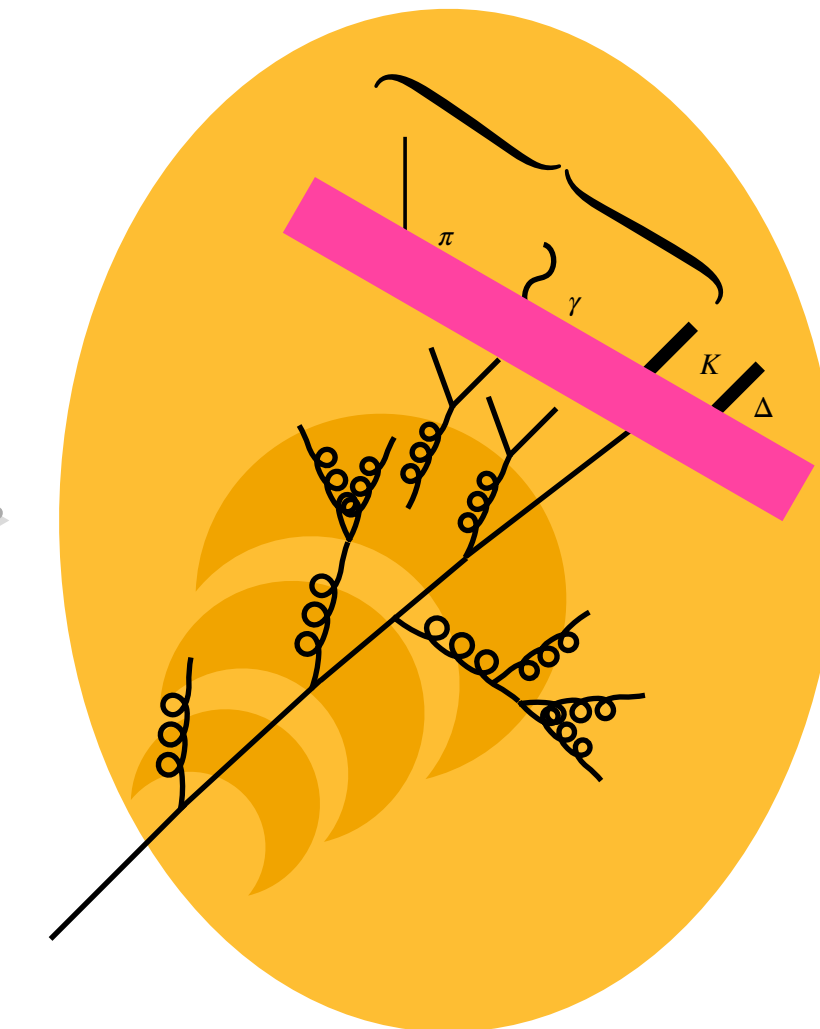
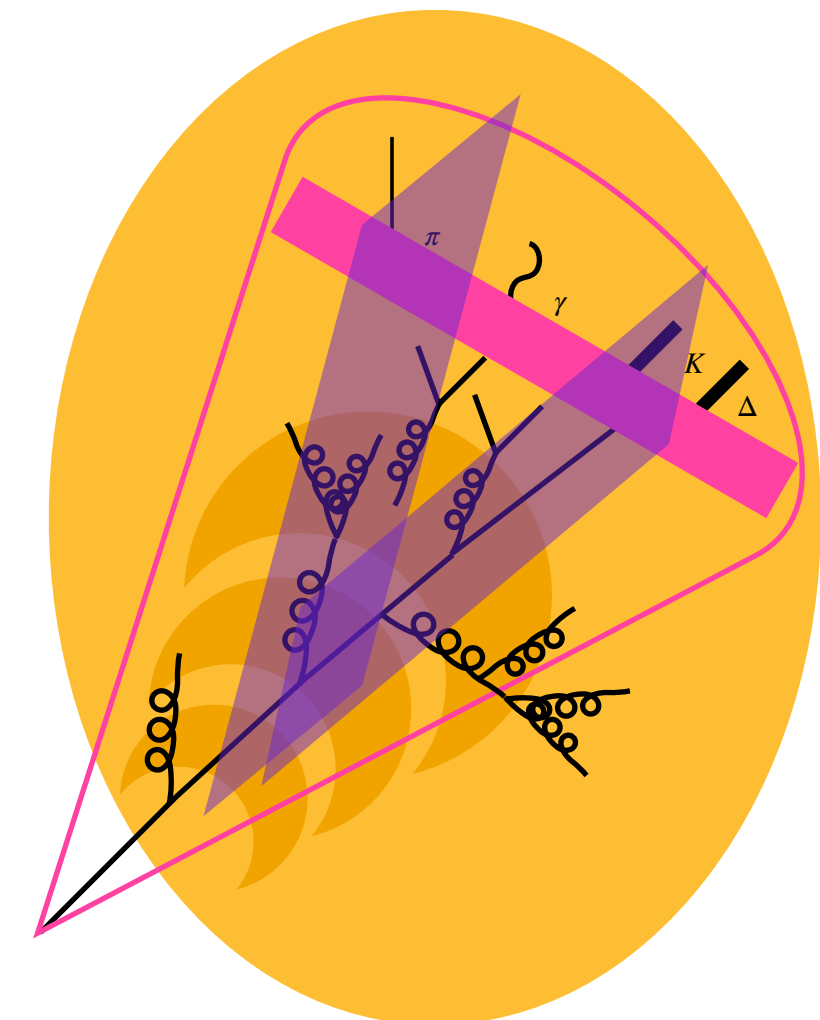
Pathlength dependence of energy loss

Testing limits of models' description of charge flow of hadronization

Suppression of jets with hard-fragmenting charm hadrons

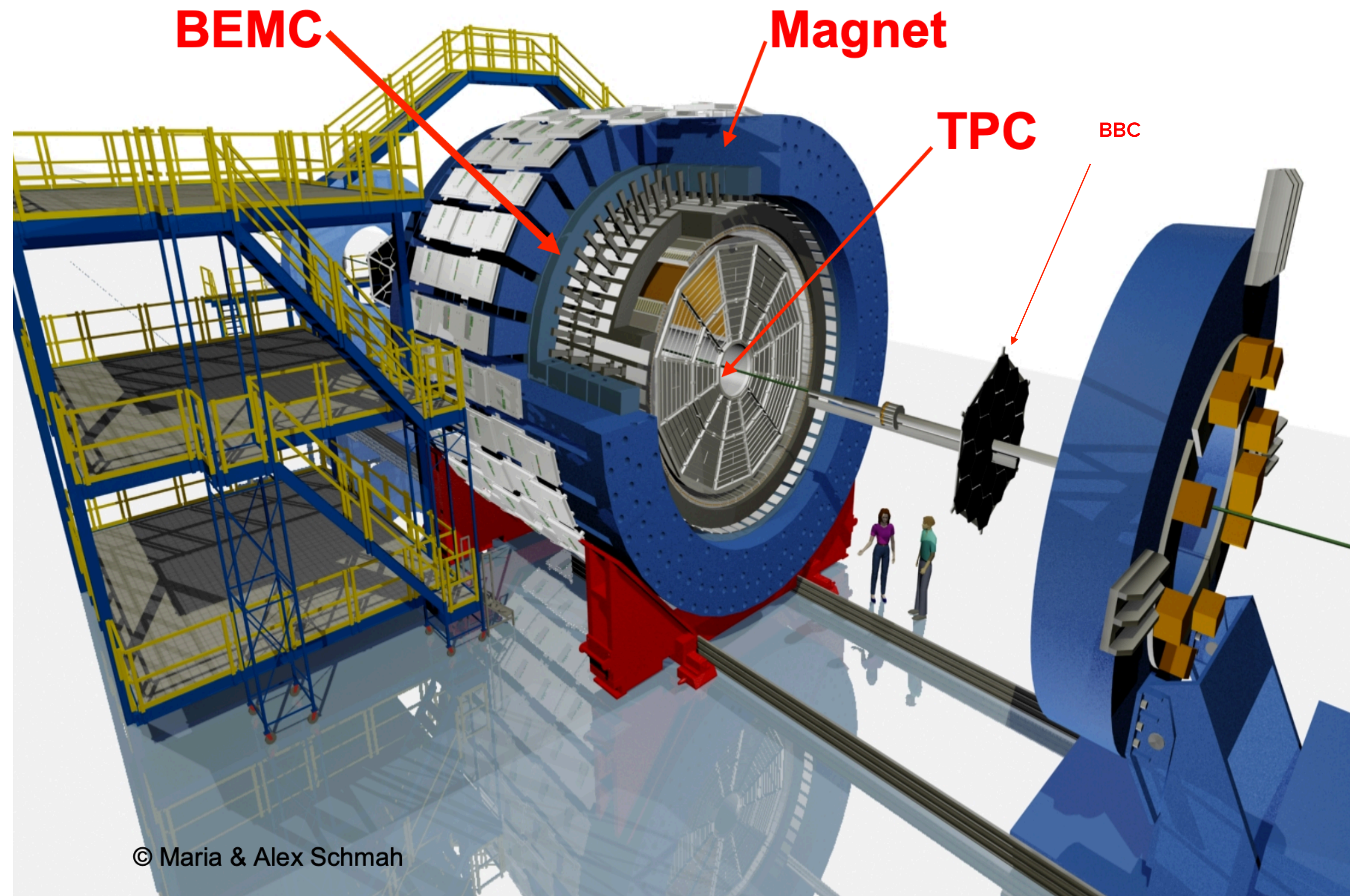
Medium-induced modification of substructure not observed

Medium-induced hadrochemistry effect not observed in jets



Hard probes at STAR

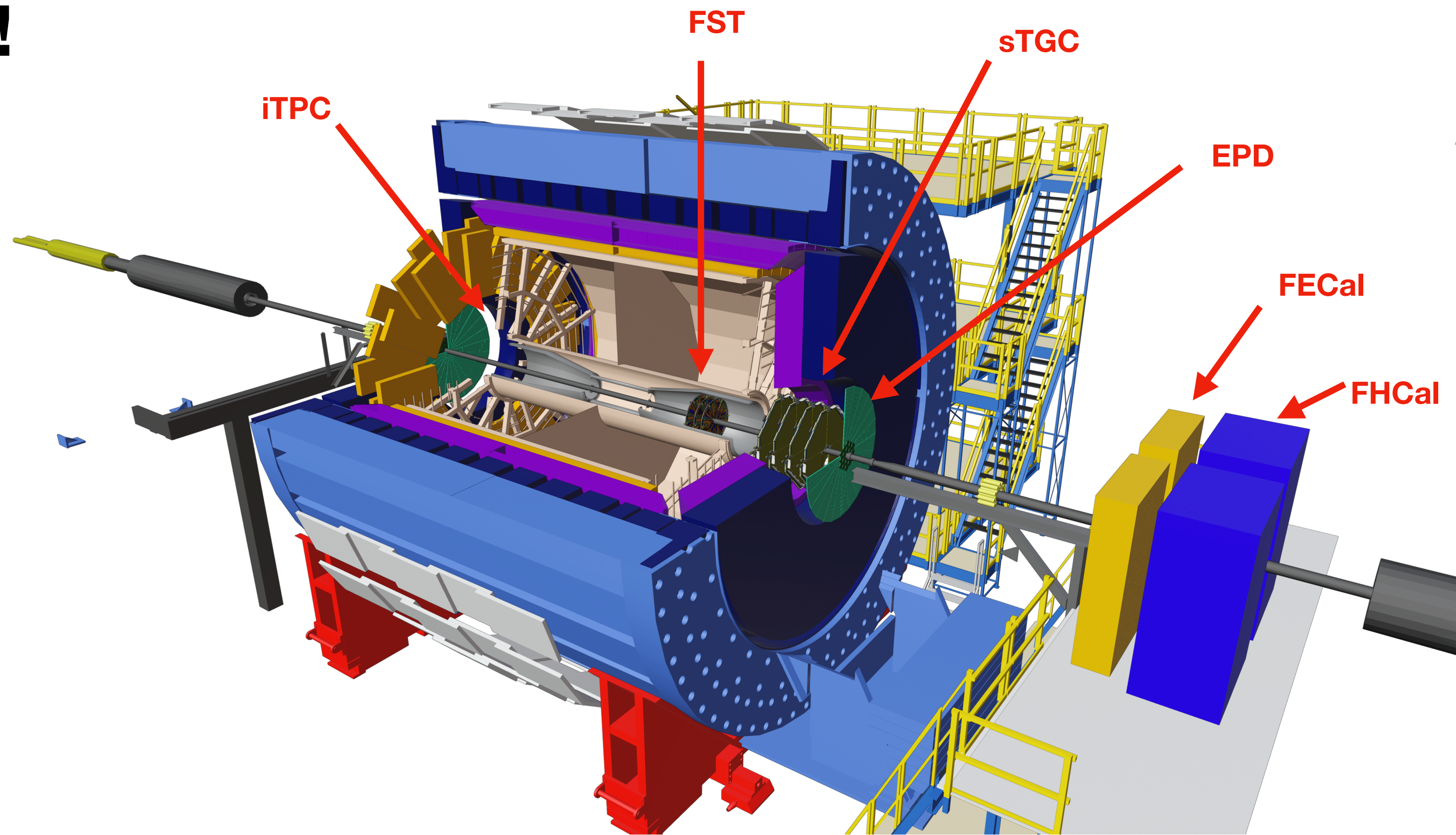
In the 2010s



© Maria & Alex Schmah

Hard probes at STAR

In the 2020s!



Precision tracking

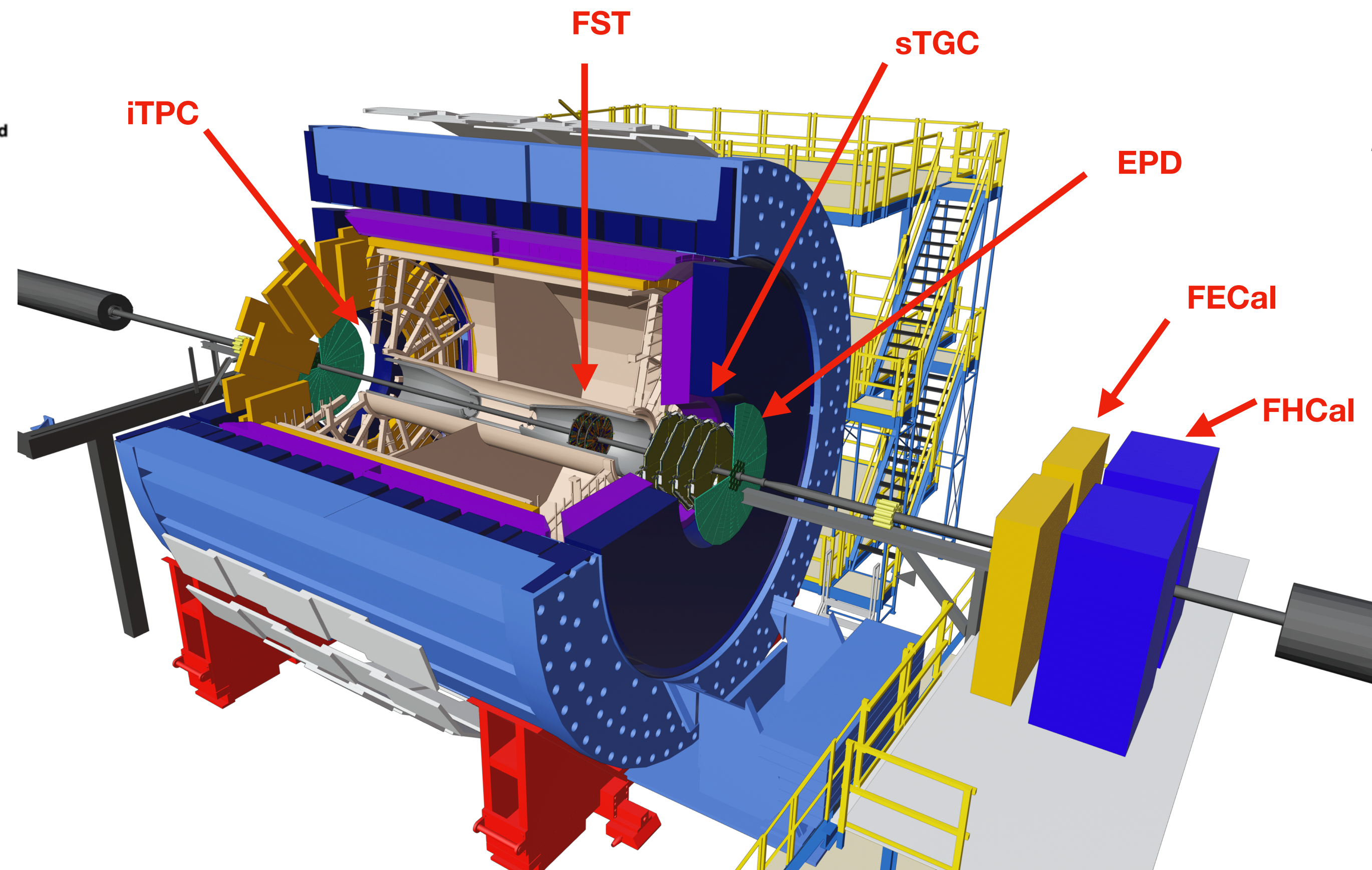
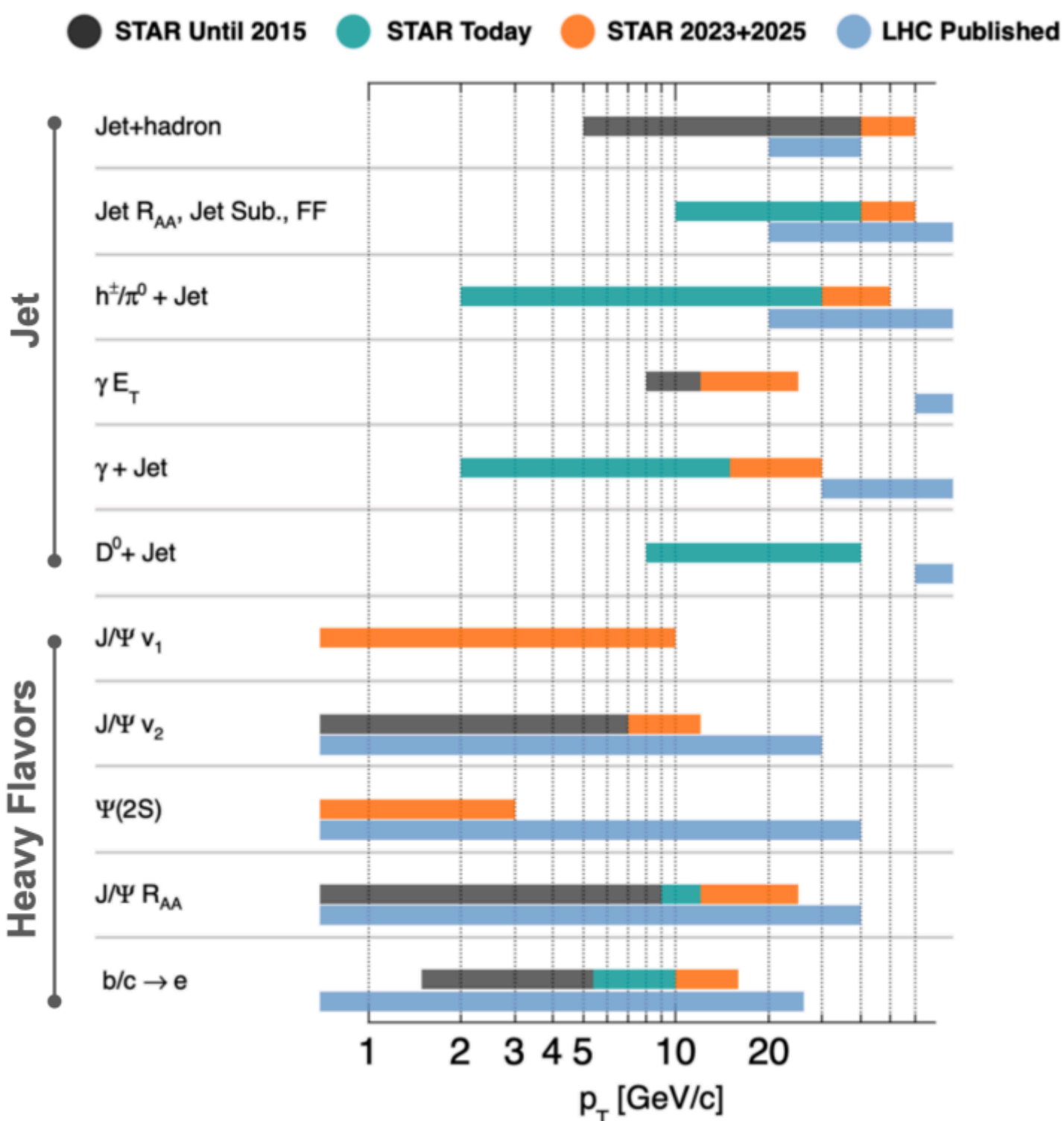
*Forward jets →
different x ; q v. g*

*Unbiased
centrality/EP
determination*

DAQ rate: 5 kHz

Hard probes at STAR

In the 2020s!



Precision tracking

*Forward jets \rightarrow
different x ; q v. g*

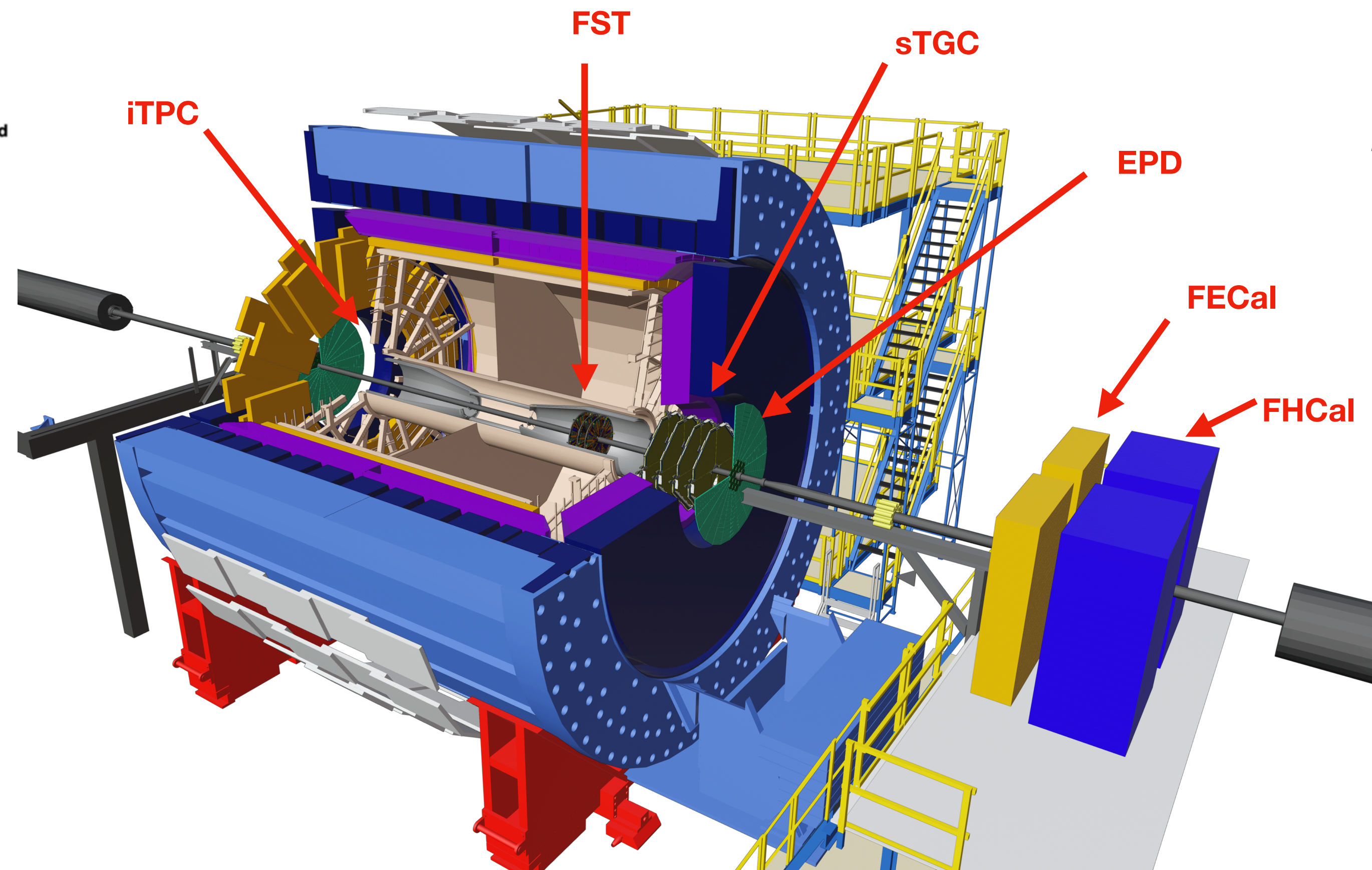
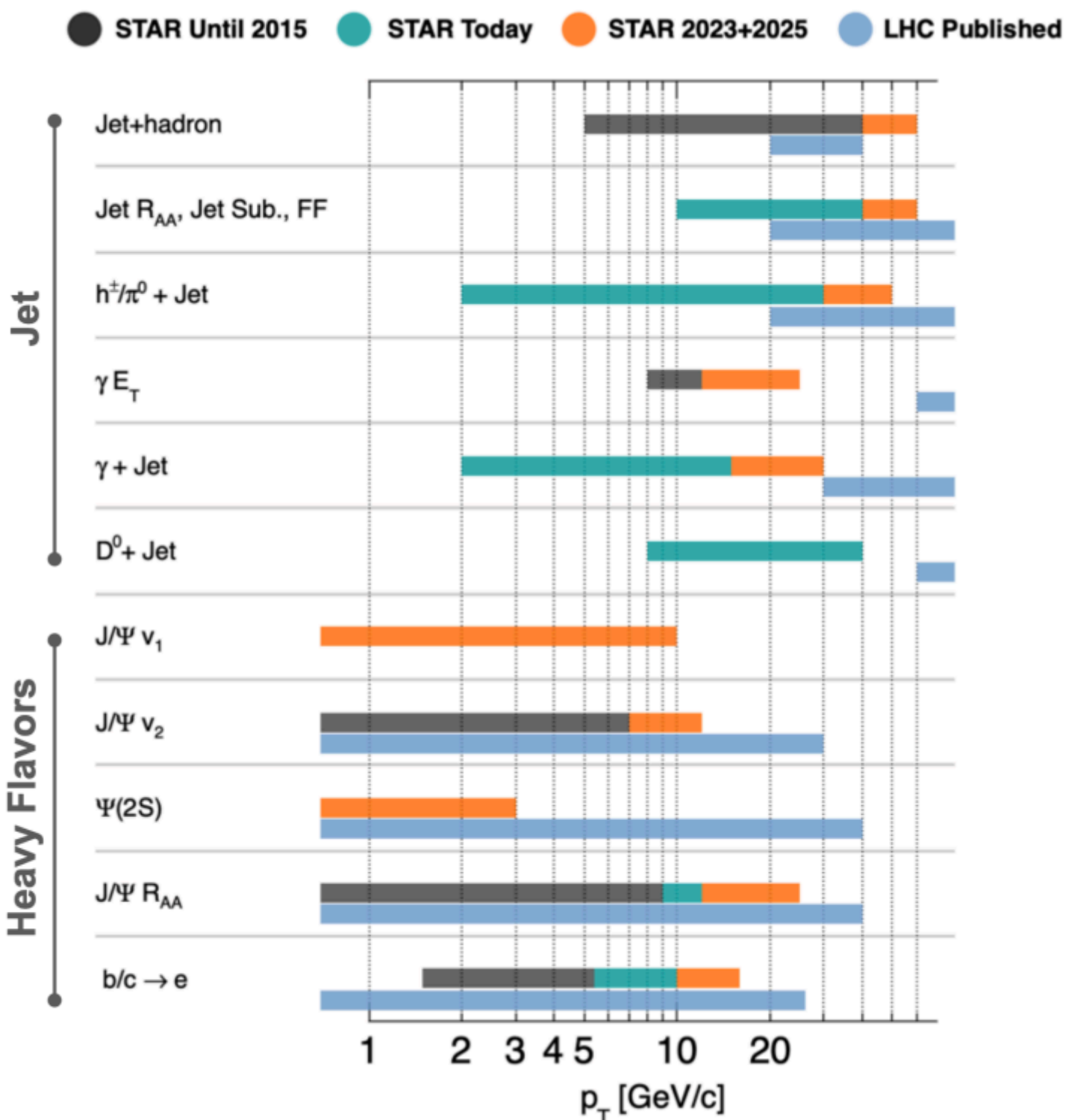
*Unbiased
centrality/EP
determination*

DAQ rate: 5 kHz

Runs 23 (Au+Au) & 24 (pp): Took $872 \mu\text{b}^{-1}$ and 168pb^{-1} of high- p_T triggers

Hard probes at STAR

In the 2020s!



Precision tracking

*Forward jets \rightarrow
different x ; q v. g*

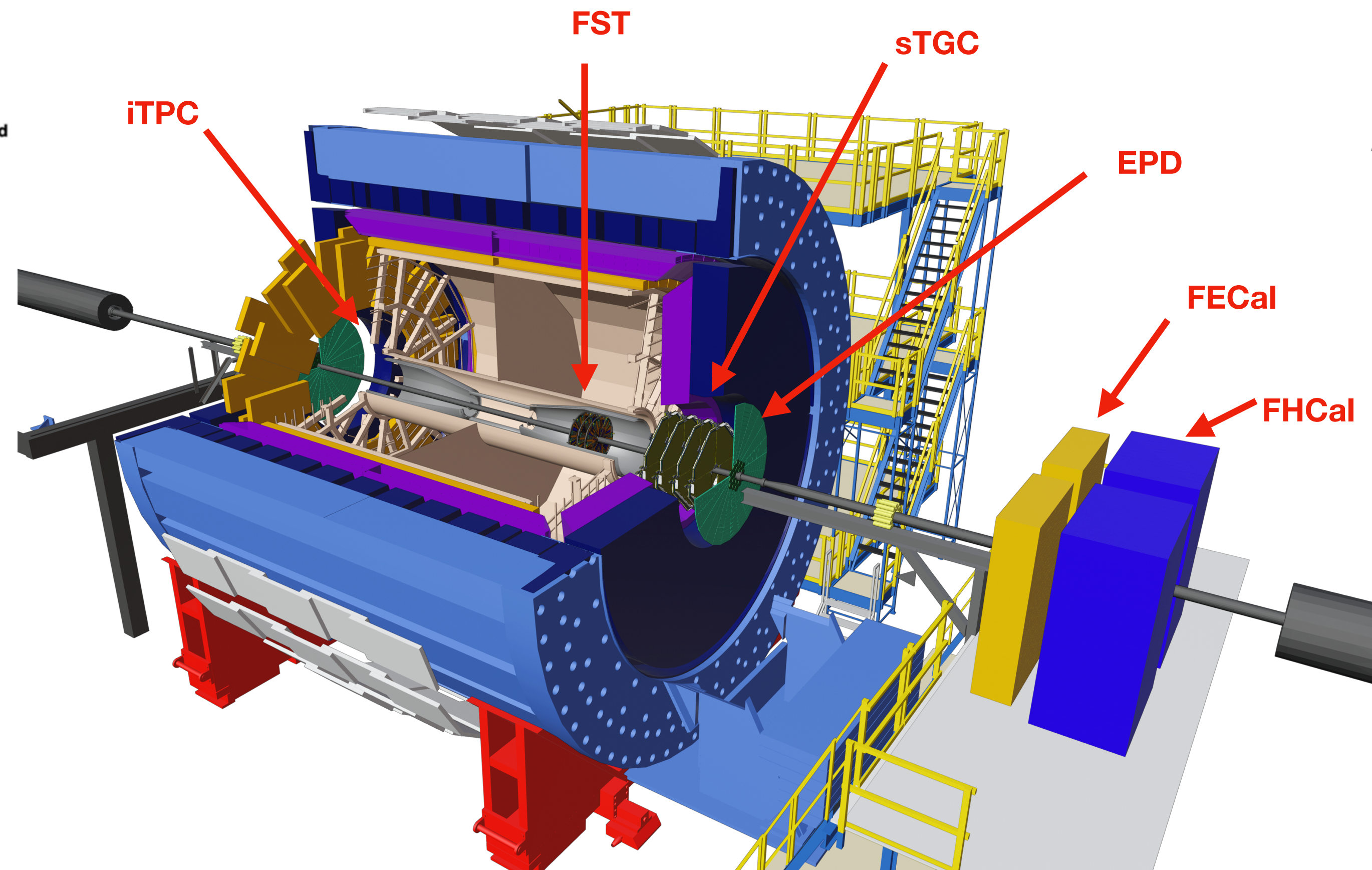
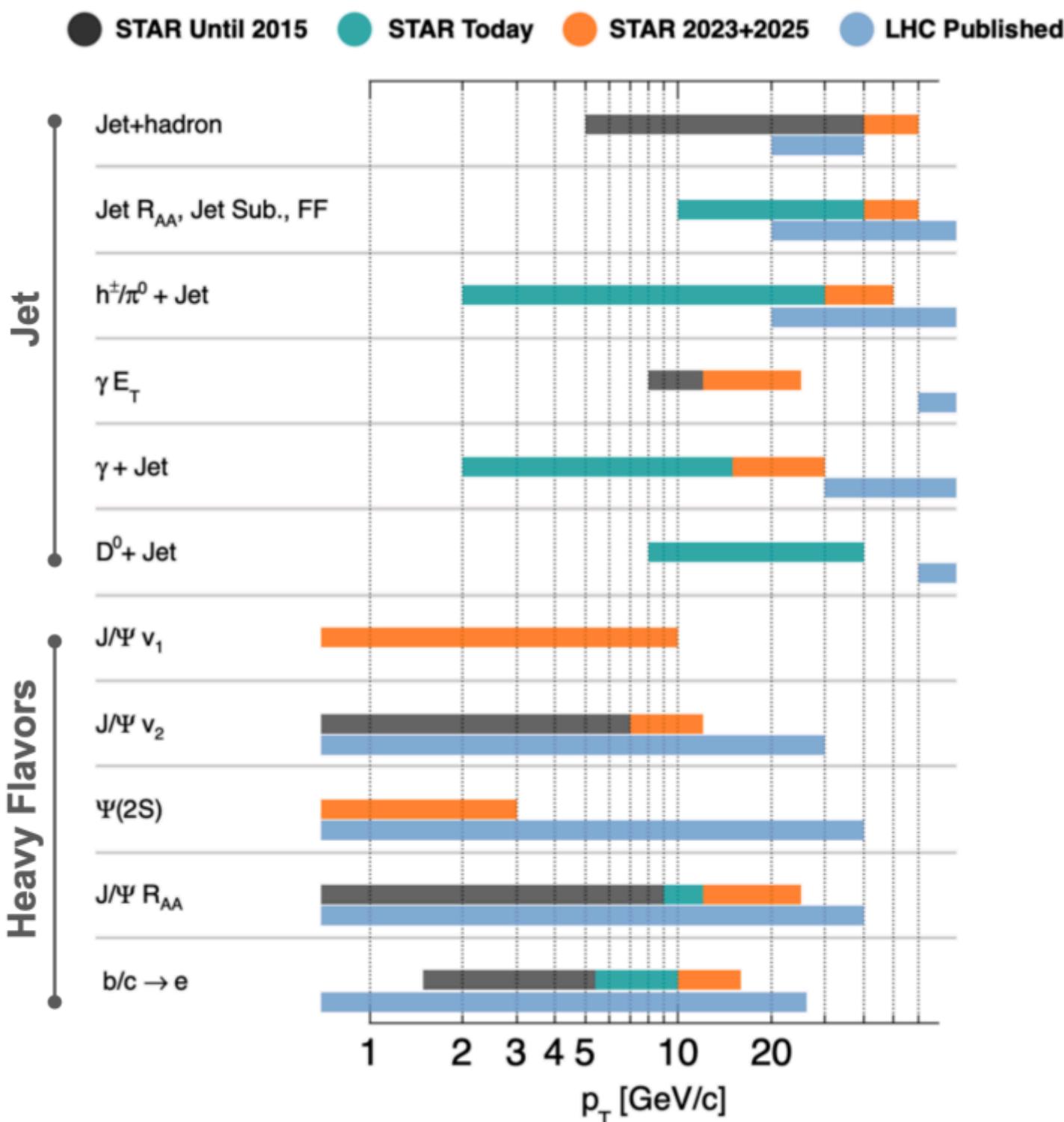
*Unbiased
centrality/EP
determination*

DAQ rate: 5 kHz

*Runs 23 + 25^{1,2}: expected $\sim 3 \times$ increase in statistics for hard probes measurements relative to current Au+Au analyses w/ Run 14 \rightarrow improved uncertainties & **kinematic reach / overlap w/ LHC***

Hard probes at STAR

In the 2020s!



Precision tracking

*Forward jets \rightarrow
different x ; q v. g*

*Unbiased
centrality/EP
determination*

DAQ rate: 5 kHz

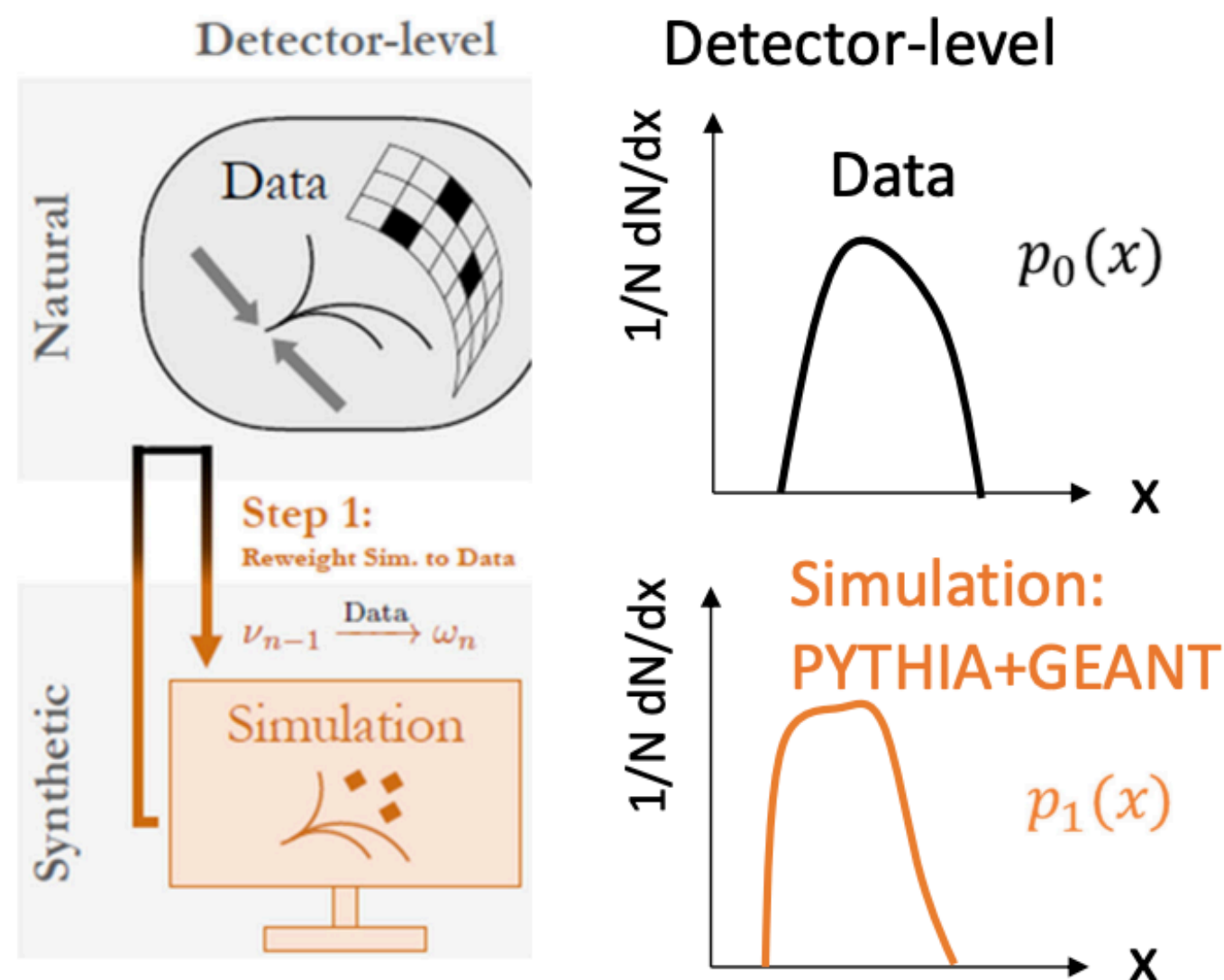
Run 25++? Opportunity if goals met to take pAu data with upgrades!

Thank you!

MultiFold



Method: machine learning



Where does the machine learning part come in?

E.g., Iteration 1, step 1:

Weights: $w(x) = p_0(x)/p_1(x)$ Ok for 1D
 $\approx f(x)/(1 - f(x))$ [\(Andreassen and Nachman PRD 101, 091901 \(2020\)\)](#)

where $f(x)$ is a neural network and trained with the binary cross-entropy loss function

to distinguish jets coming from data vs from simulation

Unfolding \rightarrow Reweighting histograms
 \rightarrow Classification \rightarrow Neural network

MultiFold

Method: machine learning

- Architecture: Dense neural network
- Activation function for dense layers: Rectified linear unit
- Activation function for output layer: Sigmoid
- Loss function: Binary cross entropy

$$\text{loss}(f(x)) = - \sum_{i \in 0} \log f(x_i) - \sum_{i \in 1} \log(1 - f(x_i))$$

- Optimization algorithm: Adam
<https://arxiv.org/pdf/1412.6980.pdf>
- Nodes per dense layer: [100,100,100]
- Output dimension: 2
- Input dimension: 6
- All hyperparameters are default:
<https://energyflow.network/docs/archs/#dnn>

