

# Spin and Forward Physics with the STAR detector: Measurements and Future Plans

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## Abstract

The spin program at the STAR experiment at the Relativistic Heavy Ion Collider (RHIC) has explored many interesting topics and has helped our understanding of nuclear and nucleon structures. In particular, non-vanishing transverse single-spin asymmetry measurements at RHIC and other experiments have shown that there is a rich substructure of the nucleon that needs further exploration in both theory and experiment. The STAR forward upgrade will utilize RHIC's unique capability of colliding polarized proton and heavy ion beams to carry out measurements of Drell-Yan, jets, hadrons in jets, and dijets, among others with improved precision. The new forward system will be in operation for the  $pp$ ,  $pA$  and  $AA$  runs starting in Fall 2021 and utilize the latest developments in detector technologies so that they are ready for the Electron-Ion Collider (EIC). The forward upgrade will cover  $2.5 < \eta < 4.0$ , by installing two new forward tracking systems and a new calorimeter system. The tracking systems will consist of silicon disks and small-strip thin gap chambers. The calorimeter system will consist of a preshower hodoscope, an electromagnetic calorimeter and a hadronic calorimeter. These proceedings will show some of the recent results from STAR's spin program as well as the design and capabilities of the forward upgrade and how it will complement measurements from a future EIC.

## 1 Introduction

One of the main open questions in nuclear physics today is what is the origin of the proton spin. This question arose from Deep Inelastic Scattering (DIS) experiments that showed that the spin of the quarks is not sufficient to account for the total spin of the proton [1]. The results shown here will highlight some of the work that has been done by the Solenoidal Tracker at RHIC (STAR) to constrain the contribution of the quarks and gluons to the total spin of the proton. Also, I will discuss how the STAR forward upgrade can be used to even further constrain these quantities. Another important question is how we can describe the multi-dimensional landscape of nucleons and nuclei. Transverse momentum dependent PDFs (TMD) are a key aspect of this question. TMDs address how a parton's transverse momentum inside the proton can be related to physics

34 observables. STAR has measured many TMDs and these proceedings will present the results from  
35 only one such TMD, transversity via the Collins fragmentation function, and how the forward  
36 upgrade aims to improve upon that result.

## 37 2 The Relativistic Heavy Ion Collider

38 The relativistic heavy ion collider (RHIC) at Brookhaven National Laboratory (BNL) is the only  
39 polarized  $pp$  collider in the world. In addition, as its name suggests, it can also collide various  
40 heavy ion species at a range of energies in both collider and fixed target mode. In the last two  
41 decades RHIC has collided various mixtures of p, d,  $^3\text{He}$ , Al, Cu, Zr, Ru, Au, and U at various  
42 energies. Proton-proton collisions have taken place at center-of-mass energies ( $\sqrt{s}$ ) of 62, 200,  
43 500, and 510 GeV. Center-of-mass energies per nucleon-nucleon pair ( $\sqrt{s_{NN}}$ ) in  $pA$  and  $AA$  colli-  
44 sions have reached up to 200 GeV, while fixed target experiments at STAR have reached as low as  
45  $\sqrt{s_{NN}} = 3$  GeV.

46 The RHIC facility consists of the Brookhaven Linear Accelerator (LINAC), a booster, an Al-  
47 ternating Gradient Synchrotron (AGS), two main storage rings, and an Electron Beam Ion Source  
48 (EBIS). Polarized  $pp$  beams start by inserting transversely polarized protons ( $\sim 95\%$  polarization)  
49 into the LINAC followed by the booster that then feeds into the AGS. The AGS is used to further  
50 increase the beam energy and then routes the beam into one of the two main storage rings, where  
51 the beams may be further accelerated; the two opposite going beams are named blue and yellow.  
52 The polarization is maintained both during acceleration and at collision energy using Siberian  
53 Snakes located on the AGS and the main storage rings. To reduce systematic uncertainties, the  
54 polarization pattern is chosen from a set of well-defined fill patterns that alternate the polarization  
55 direction for consecutive bunches or pairs thereof. In addition, spin rotators are located around  
56 the interaction region to allow for either transversely or longitudinally polarized  $pp$  or  $pA$  col-  
57 lisions. Polarimeters within the ring allow continuous measurements of the beam polarization  
58 during a RHIC fill, which typically lasts 8 hours. Unpolarized heavy ion beams start in the EBIS  
59 that generates the initial ions that go into the LINAC and then follow the same process [2].

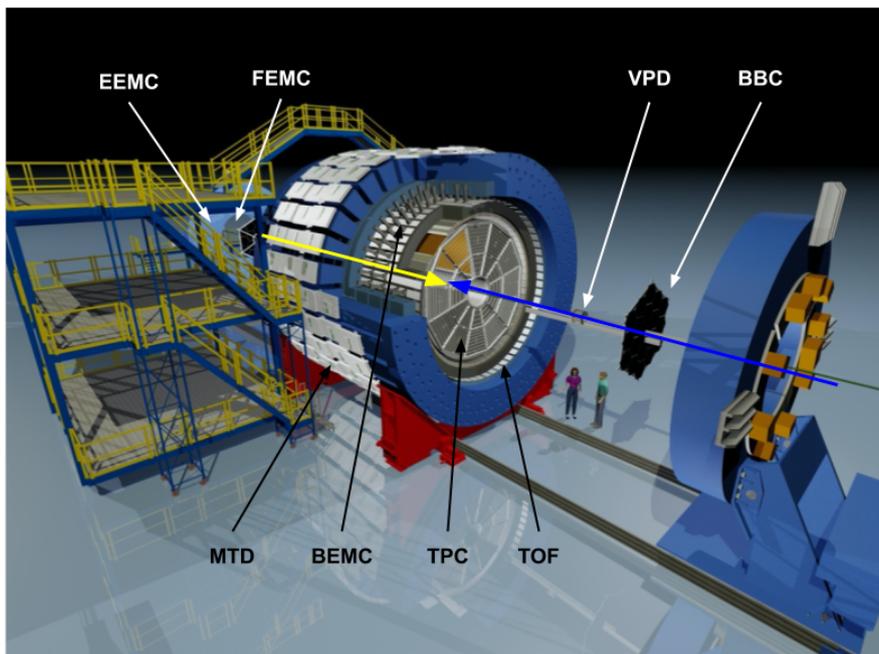
60 The latest transversely polarized  $pp$  run was RHIC Run 17 at  $\sqrt{s} = 510$  GeV. It has the highest  
61 delivered luminosity per week for all  $pp$  runs to date. RHIC has been able to provide highly  
62 polarized beams, achieving an average polarization of about 50-60%. Also, of interest, is Run 15  
63 that had a mix of longitudinal and transverse spin  $pp$  collisions at  $\sqrt{s} = 200$  GeV.

## 64 3 STAR Detector

65 The STAR detector is shown in Fig. 1 [3]. STAR has a Time Projection Chamber (TPC) at mid-  
66 rapidity  $|\eta| < 1.0$  that covers a full  $2\pi$  in azimuth [4]. It is used for charged particle reconstruction  
67 and identification. Just outside the TPC is the Barrel Electromagnetic Calorimeter (BEMC). The  
68 BEMC has the same coverage as the TPC and is used for measuring the energies of electrons and  
69 photons. The other two detectors at mid-rapidity are the Time Of Flight (TOF), used to improve  
70 particle identification from the TPC, and the Muon Telescope Detector (MTD), which is used to  
71 detect muons. In addition, there exist global detectors that have multiple functionalities. These are  
72 the Beam-Beam Counter (BBC) located at  $3.3 < |\eta| < 5.2$  and the Vertex Position Detector (VPD)  
73 located at  $4.24 < |\eta| < 5.1$ . Lastly, the detectors in forward pseudorapidity (blue beam direction)  
74 with full  $2\pi$  azimuth coverage are the Endcap Electromagnetic Calorimeter (EEMC)  $1 < \eta < 2$ ,  
75 and the Forward Electromagnetic Calorimeter (FEMC)  $2.5 < \eta < 4.0$ .

76 There were some upgrades completed in 2019 that are not visible in Fig. 1. The inner TPC  
77 was upgraded which improved  $\frac{dE}{dx}$  resolution and increased coverage to  $|\eta| < 1.5$ . An Endcap

78 TOF extends coverage to  $1.05 < \eta < 1.7$ . Lastly, installed in 2018, the Event Plane Detector (EPD)  
 79 covers  $2.1 < |\eta| < 5.1$ . It is used to provide event triggers and improve event plane resolution.



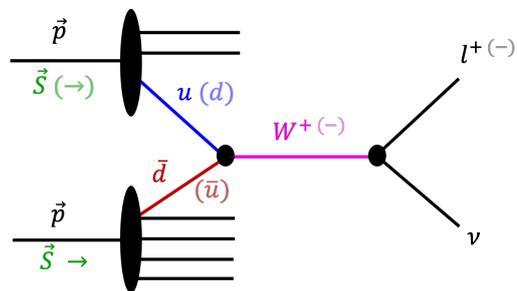
**Figure 1:** STAR detector and its various detector subsystems. Information on each subsystem can be found in the text.

#### 80 4 Helicity Structure of Proton

81 The longitudinal spin of the proton can be decomposed into the various angular momenta of the  
 82 quarks and gluons that make it up. One such decomposition, the Jaffe-Manohar spin sum rule,  
 83 decomposes the spin of the proton according to equation (4.1)

$$\frac{1}{2} = \frac{1}{2}\Delta\Sigma + \Delta G + L_z \quad (4.1)$$

84 where  $\Delta\Sigma$  is the quark polarization,  $\Delta G$  is the gluon po-  
 85 larization and  $L_z$  is the orbital angular momentum of the  
 86 quark-gluon system. The quark polarization can be fur-  
 87 ther broken down into the valence quark and sea quark  
 88 polarization. The quark polarization has been measured  
 89 using DIS experiments and accounts for only about 30%  
 90 of the total proton spin in a limited  $x$  range [1]. The  
 91 EIC will provide better constraints on the valence quark  
 92 polarization. The sea quark polarization can be probed  
 93 using the parity violating  $W^\pm$  production at RHIC. The  
 94 diagram in Fig. 2 shows how  $W$ 's are produced. This  
 95 production is maximally parity violating so the quarks  
 96 must have opposite helicities. An up quark from one

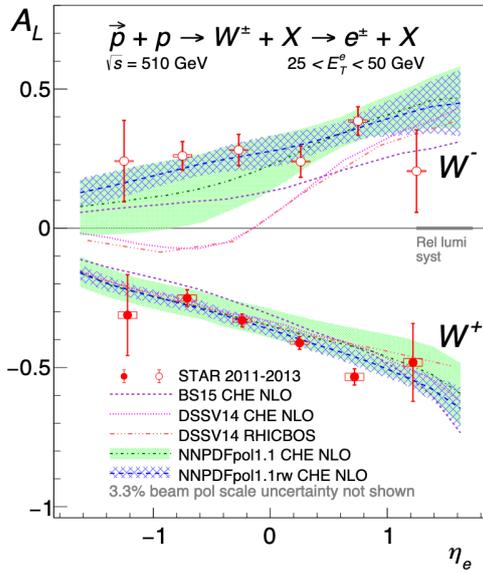


**Figure 2:** Feynman diagram showing parity violating  $W$  production. It requires the  $u(d)$  quark must interact with a  $\bar{d}(\bar{u})$  quark.

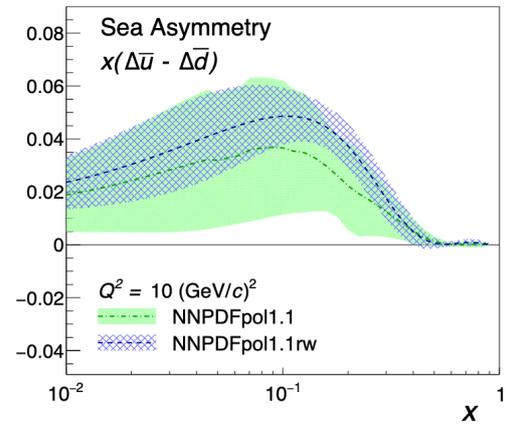
97 proton must interact with an anti-down quark from the other proton to produce a  $W$  or vice versa.  
 98 This measurement is complementary to Semi-Inclusive DIS (SIDIS) as there is no fragmentation  
 99 to tag the flavor. STAR measures the single-spin asymmetry  $A_L$  (equation (4.2)) of  $W$ 's via its  
 100 leptonic decay in longitudinal  $pp$  collisions,

$$A_L^{W^\pm} = \frac{\sigma^+ - \sigma^-}{\sigma^+ + \sigma^-} \sim \frac{\Delta\bar{d}(x_1)u(x_2) - \Delta u(x_1)\bar{d}(x_2)}{\bar{d}(x_1)u(x_2) + u(x_1)\bar{d}(x_2)} \quad (4.2)$$

101 where the “+” denotes positive helicity and “-” denotes negative helicity. The  $A_L$  measurement for  
 102  $W^\pm$  from STAR in the region  $|\eta| < 1.2$  is shown in Fig. 3. The lepton pseudorapidity is used to dial  
 103 into the sea quark polarization. These data show a positive  $\Delta\bar{u}$  in  $0.05 < x < 0.25$ . Furthermore,  
 104 the NNPDF reweighting of the new data, shown in Fig. 4 in the blue band, shows a clear sea quark  
 105 polarization asymmetry  $\Delta\bar{u} > \Delta\bar{d}$  [5].



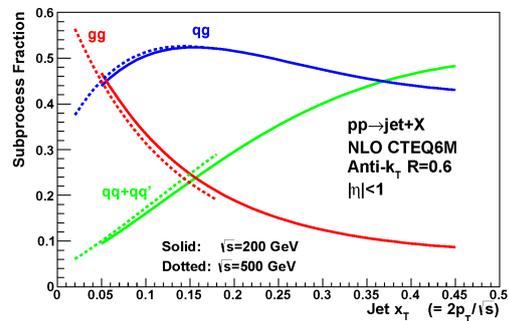
**Figure 3:** Measured  $A_L$  for  $W^\pm$  vs. lepton pseudorapidity ( $\eta_e$ ) together with theory expectations (curves and bands). Data shown are combined from 2011, 2012 and 2013 [5].



**Figure 4:** Difference in light sea-quark polarizations as a function of  $x$  at  $Q^2 = 10 \text{ GeV}^2$ . The green band shows the NNPDFpol1.1 results before the 2013 STAR data shown in Fig. 3. The blue hatched band is a reweighting of the PDF after the 2013 STAR data was included. A clear sea quark polarization asymmetry of  $\Delta\bar{u} > \Delta\bar{d}$  is seen [5].

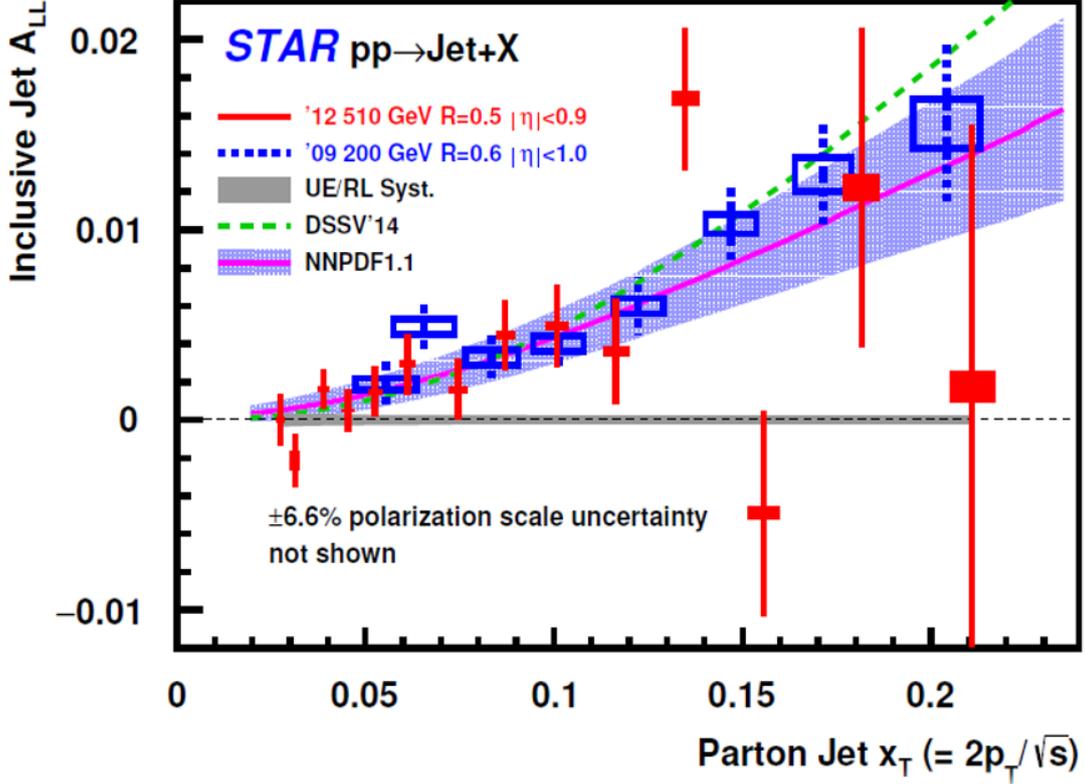
106 The gluon polarization can be measured using jet, dijet and  $\pi^0$  production at RHIC. Figure 5  
 107 shows the relative fractions of different processes that contribute to inclusive jet production as a  
 108 function of  $x_T = 2p_T/\sqrt{s}$ . It shows that at low  $x_T$  gluon-gluon subprocesses dominate over the  
 109 quark-quark subprocesses [6]. The longitudinal double-spin asymmetry ( $A_{LL}$ ) for inclusive jets,  
 110 defined in equation (4.3), in this regime is sensitive to the gluon polarization.  
 111  
 112  
 113  
 114  
 115

$$A_{LL} = \frac{\sigma^{++} - \sigma^{+-}}{\sigma^{++} + \sigma^{+-}} = \frac{1}{P_1 P_2} \frac{N^{++} - RN^{+-}}{N^{++} + RN^{+-}} \quad (4.3)$$



**Figure 5:** Subprocesses that dominate inclusive jet production as a function of  $x_T$ . At low  $x_T$  gluon-gluon scattering dominates [6].

116 In equation (4.3)  $\sigma$  is the inclusive jet cross section,  $P_1$  and  $P_2$  are the polarization of beam 1  
 117 and 2 respectively,  $R$  is the relative luminosity and  $N$  is the number of events with “++” denoting  
 118 same helicity and “+-” denoting opposite helicity. Measurements of  $A_{LL}$  of inclusive jets at STAR  
 119 mid-rapidity ( $|\eta| < 1.0$ ) are shown in Fig. 6 for both  $\sqrt{s} = 200$  GeV and 510 GeV energies. At  
 120 both energies there is a clear asymmetry at low  $x_T$ . These results will provide important new  
 121 constraints on the magnitude of the gluon polarization [6].



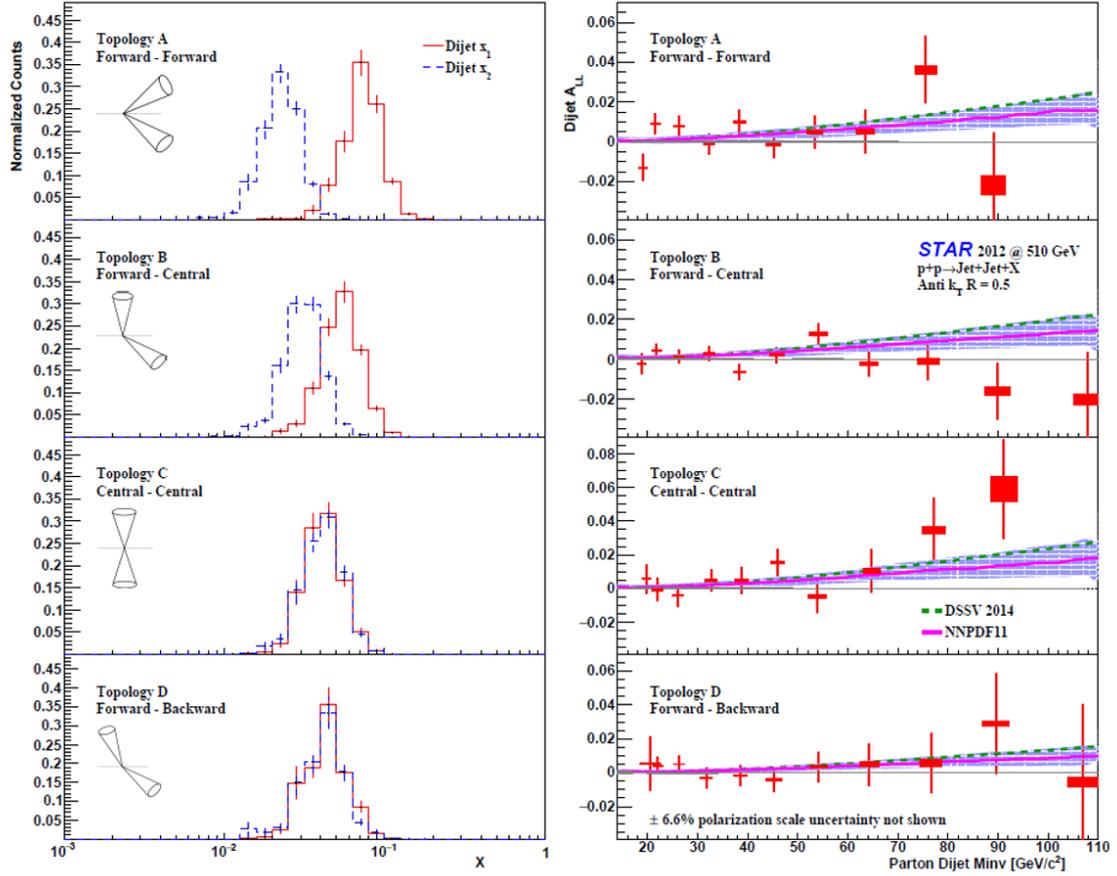
**Figure 6:**  $A_{LL}$  of inclusive jets as a function of  $x_T$ , where  $x \approx x_T e^{\pm\eta}$ . Data are for both 200 GeV and 510 GeV  $pp$  collisions. A clear asymmetry exists in the low  $x_T$  region [6].

Inclusive di-jet  $A_{LL}$  was also measured at STAR at various pseudorapidity bins or topologies. Figure 7 shows the measured  $A_{LL}$  in each topology where the designation central corresponds to the pseudorapidity bin  $|\eta| < 0.3$  and forward/backward corresponds to  $0.3 < |\eta| < 0.9$ . It also shows the various  $x$  regions probed by the different bins, where

$$x_1 = \frac{1}{\sqrt{s}}(p_{T,3}e^{\eta_3} + p_{T,4}e^{\eta_4}) \quad (4.4a)$$

$$x_2 = \frac{1}{\sqrt{s}}(p_{T,3}e^{-\eta_3} + p_{T,4}e^{-\eta_4}) \quad (4.4b)$$

122 The  $x$  values as shown in the left panel of Fig. 7 become more equal as you go from forward-  
 123 forward to central-central and forward-backward. These  $A_{LL}$  measurements, when combined  
 124 with global analyses, will help to constrain the shape of  $\Delta g(x)$  [6].



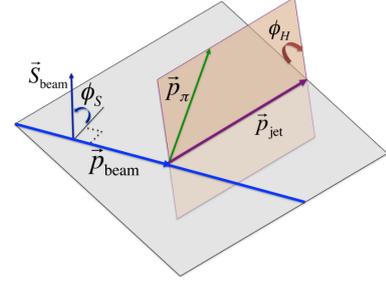
**Figure 7:**  $A_{LL}$  of inclusive dijets as a function of parton dijet invariant mass ( $Minv$ ) at various  $\eta$  topologies. Central corresponds to the pseudorapidity bin  $|\eta| < 0.3$  and forward/backward corresponds to  $0.3 < |\eta| < 0.9$ . Left panel shows the  $x$  of the two partons in the various pseudorapidity bins. As you go from forward-forward to central-central and forward-backward the  $x$  of the two partons becomes equal. These measurements will help to constrain the shape of  $\Delta g(x)$  [6].

## 125 5 Transverse Structure of Proton

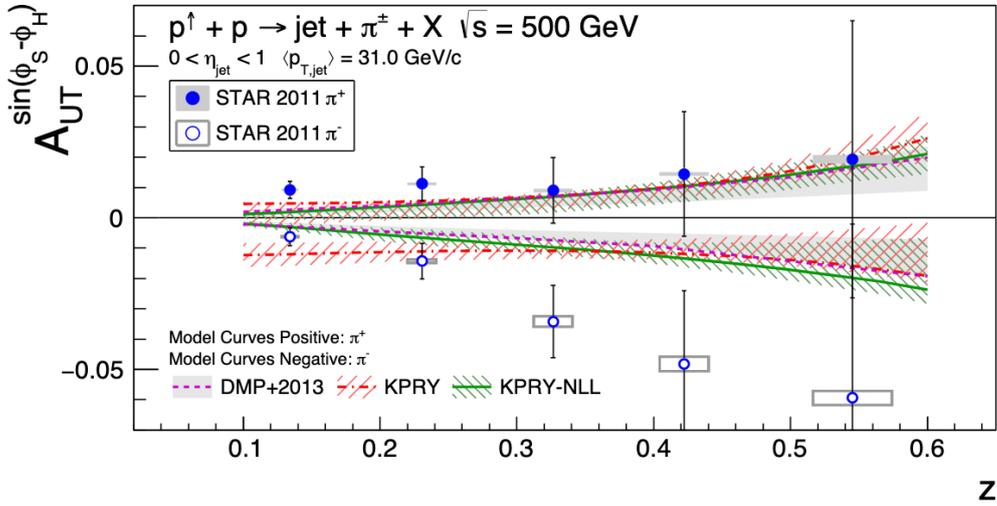
126 Transverse momentum dependent PDFs (TMD) are used to go beyond the one-dimensional pic-  
 127 ture of the nucleon by adding more degrees of freedom. This allows a three-dimensional picture  
 128 of the proton momentum to be constructed. STAR has made measurements that are sensitive to  
 129 several of the TMDs, one of which is transversity. The transversity TMD relates the transverse  
 130 quark spin to the transverse nucleon spin  $\delta q(x)$ . Transversity is chiral odd and therefore needs an  
 131 additional chiral-odd function (Collins fragmentation function  $H_1^\perp$ ) to be accessible in a physics  
 132 observable. The Collins asymmetry is the azimuthal distribution of hadrons inside a jet and is  
 133 defined in equation (5.1)

$$A_{UT}^{\sin(\phi_s - \phi_h)} = \frac{d\sigma^\uparrow - d\sigma^\downarrow}{d\sigma^\uparrow + d\sigma^\downarrow} \quad (5.1)$$

134 Here  $\phi_s$  is the angle of the proton spin with respect to  
 135 the proton-jet momentum plane and  $\phi_h$  is the angle of the jet-  
 136 pion momentum plane to the proton-jet momentum plane as  
 137 depicted in Fig. 8;  $d\sigma^{\uparrow(\downarrow)}$  is the cross section when spin is up  
 138 (down) with respect to the proton momentum. The Collins  
 139 asymmetry for  $p^\uparrow + p \rightarrow jet + \pi^\pm + X$  at  $\sqrt{s} = 500$  GeV  
 140 at  $0 < \eta < 1$  is shown in Fig. 9, where  $z$  is the fractional hadron  
 141 momentum to the jet momentum i.e.  $z = \frac{\text{hadron momentum}}{\text{jet momentum}}$ . The  
 142 asymmetry is with respect to the azimuthal distribution of pi-  
 143 ons inside jets. Figure 9 shows the asymmetry growing as  $z$   
 144 increases which is the first sign that TMDs survive at high  $Q^2$   
 145 [7].



**Figure 8:** The angles used in the Collins asymmetry definition in equation (5.1) [7].



**Figure 9:** Measured Collins asymmetry  $A_{UT}$  as a function of  $z$ . This shows the first sign that TMDs survive at high  $Q^2$  [7].

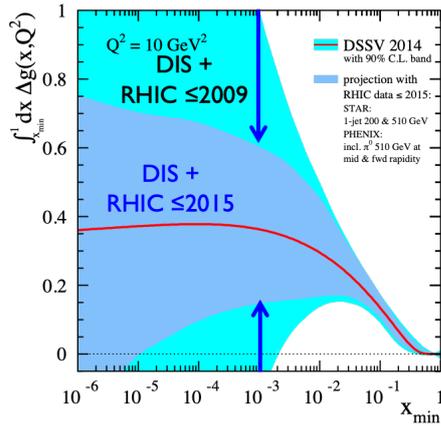
## 146 6 STAR Forward Upgrade Capabilities

147 To improve and expand on the measurements already discussed, STAR is in the process of upgrad-  
 148 ing several detector subsystems. RHIC data as of 2009 show  $\int_{0.05}^1 \Delta g dx \sim 0.2 \pm_{0.07}^{0.06}$  at  $Q^2 = 10$  GeV<sup>2</sup>  
 149 which is also depicted as the light blue band in Fig. 10. The dark blue band in Fig. 10 shows a  
 150 projection with RHIC data up to and including the 2015 run. In order to constrain  $\Delta g(x)$  at lower  
 151  $x$  either one has to go to higher  $\sqrt{s}$ , which is not feasible at RHIC, or larger pseudorapidity. The  
 152 STAR forward upgrade will do the latter and extend STAR's forward capabilities in the region  
 153  $2.5 < \eta < 4.0$  to go to lower  $x$ . Figure 11 shows the projected  $x_1$  and  $x_2$  for inclusive dijets with the  
 154 proposed upgrade. In addition, the Collins asymmetry measurement shown in Fig. 9 can also be  
 155 improved. Figure 12 shows the projected Collins asymmetry precision as well as the asymmetries

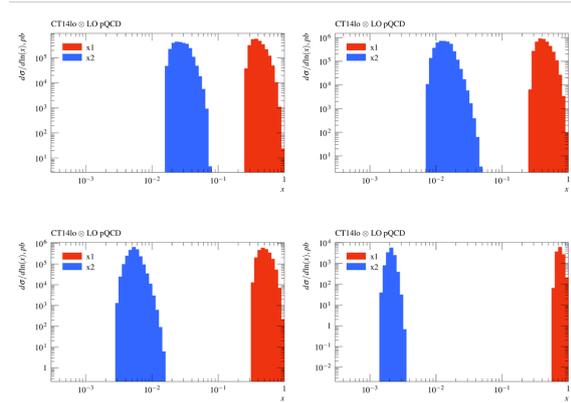
Detector	$pp$ and $pA$	AA
EM calorimeter	$\sim 10\%/\sqrt{E}$	$\sim 20\%/\sqrt{E}$
Hadron calorimeter	$\sim 60\%/\sqrt{E}$	
Tracking system	Charge separation; photon suppression	$0.2 < p_T < 2 \text{ GeV}/c$ with 20-30% $\ast 1/p_T$

**Table 1:** Table of hardware requirements for STAR forward upgrade to achieve physics goals

156 obtained from transversity extractions for one jet  $p_T$  and pseudorapidity bin [8]. The black tri-  
157 angle points represent the uncertainties while the red curve indicates the asymmetries for  $\pi^+$  and  
158 the blue curve indicates  $\pi^-$ . In fact, this projection shows only one of several  $x$  and  $Q^2$  bins that  
159 the STAR forward upgrade will be able to access. Figure 13 shows both the current data on TMDs,  
160 which come from DIS experiments, and the projected  $x$  and  $Q^2$  accessible with the STAR forward  
161 upgrade at RHIC as black filled squares. At  $\sqrt{s} = 500 \text{ GeV}$  the new kinematic coverage of STAR  
162 will range from 0.05 to 0.5 in  $x$  and 10 to 100  $\text{GeV}^2$  in  $Q^2$ . In order to accomplish these tasks, the  
163 forward upgrade requires a tracking system to deliver good electron-hadron separation, as well  
164 as electromagnetic and hadronic calorimeters to provide hadron,  $\pi^0$ , and photon identification.  
165 Table 1 shows the individual requirements broken down by species and hardware performance  
166 [8].



**Figure 10:**  $\Delta G$  from DIS and RHIC data in light blue until 2009. Projected  $\Delta G$  from RHIC data up to and including 2015 in darker blue.

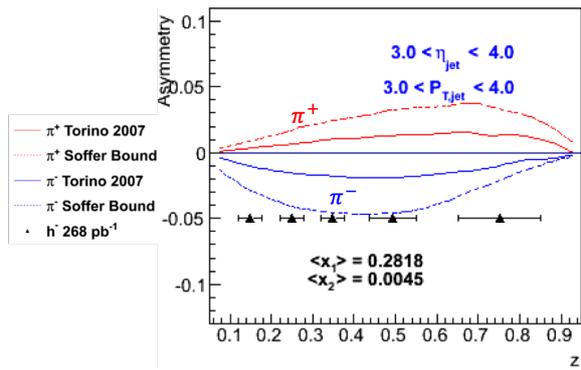


**Figure 11:** The projected dijet  $x_1$  and  $x_2$  range to be accessed by the STAR forward upgrade, where  $x_1$  and  $x_2$  are defined in equation (4.4).

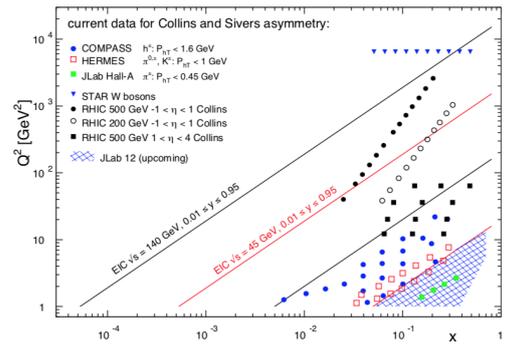
## 167 7 STAR Forward Upgrade Design

168 The design of the STAR forward upgrade side view can be seen in Fig. 14. It spans  $2.5 < \eta < 4.0$   
169 with nearly  $2\pi$  coverage. The trackers consist of three layers of silicon disks and 4 layers of  
170 small-strip Thin Gap Chamber (sTGC). The calorimeters consist of a preshower, an electromag-  
171 netic calorimeter (ECal), and a hadronic calorimeter (HCal).

172 The silicon disks will be located between 140 to 170 cm from the STAR interaction region (IR).  
173 Each disk contains 12 modules. They consist of an inner and outer portion that are connected via  
174 mechanical structures. A cooling system will also be installed. The sTGC modules consist of 4



**Figure 12:** Projected Collins asymmetry uncertainties with pion asymmetries based on transversity extractions as a function of fractional energy  $z$ . The projected asymmetry uncertainties are the black triangles with the asymmetries of  $\pi^+$  in red and  $\pi^-$  in blue. Solid and dashed designate the two different extractions. Only one jet  $p_T$  and pseudorapidity bin is shown.



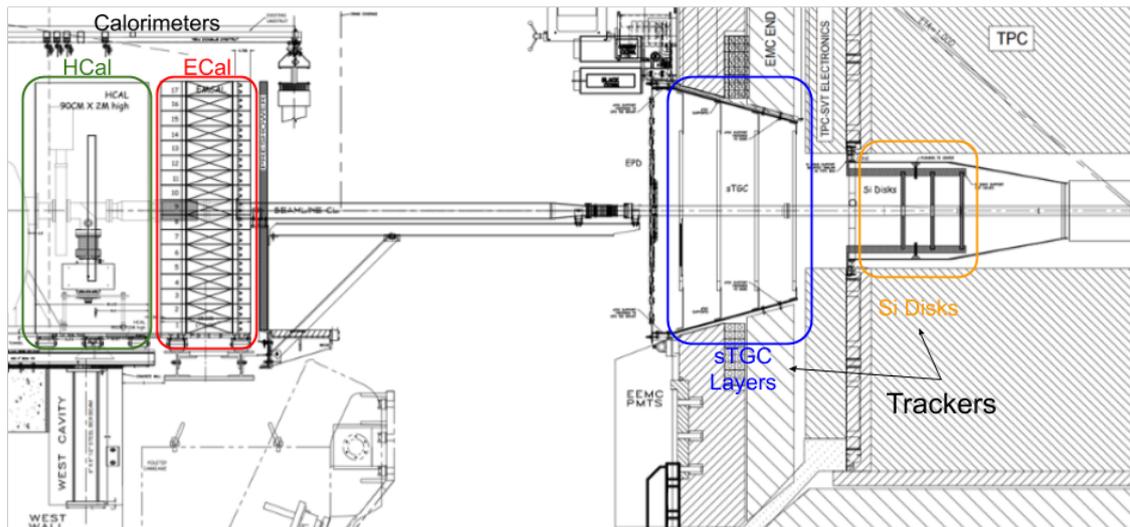
**Figure 13:** Current TMD data from SIDIS with the STAR forward upgrade projected region of coverage in black squares. One set of the black squares ( $x = 0.05$  and  $Q^2 = 10 \text{ GeV}^2$ ) represents the kinematic region for which the Collins asymmetry was projected as shown in Fig. 12.

175 layers and will be located between 270 to 370 cm. Each layer is double sided to provide  $x, y$  and  
 176 diagonal ( $45^\circ$  with respect to  $x, y$ ) coordinates. It has a position resolution of about  $100 \mu\text{m}$ . It has  
 177 almost  $2\pi$  coverage as there needs to be room for the beam pipe support. The sTGC uses the same  
 178 technology as the ATLAS design [9].

179 The ECal is a Pb/Sc sandwich that was repurposed from PHENIX. It has been modified to use  
 180 SiPM readout. It is split into two halves that are located on North side of STAR (right of blue going  
 181 beam) and South side of STAR (left of blue going beam) with no coverage above and below the  
 182 beam pipe. It is positioned 7 m from the STAR IR and at a slight angle so that the front face of the  
 183 ECal is oriented towards the IR to mitigate incident angle effects at this distance. It is  $18X_0$  lengths  
 184 long. The preshower will be a scintillator hodoscope. The hadronic calorimeter will be used for  
 185 the first time at STAR and needs to be built from scratch. It is a steel (Fe) scintillator (Sc) sandwich  
 186 with 20 mm Fe/3 mm Sc. It will also utilize SiPM readout and will be located directly behind the  
 187 EM calorimeter and can be seen in Fig. 14. It is  $\sim 4.5\lambda$  long and the lateral size of each HCal tower  
 188 is  $10 \times 10 \text{ cm}^2$ , i.e. one HCal tower covers an area roughly equal to a  $2 \times 2$  set of towers in ECal [8].

## 189 8 Conclusions

190 STAR has made key measurements that have helped in our understanding of the proton struc-  
 191 ture. Results from longitudinally polarized  $pp$  collisions have shown a clear sea quark polarization  
 192 asymmetry. Also, they have provided constraints for the magnitude of the gluon polarization and  
 193 the shape of  $\Delta g(x)$  using inclusive jets, and dijets at mid-rapidity. Measurements of the Collins  
 194 asymmetry have shown the first sign that TMDs survive at high  $Q^2$ . The STAR forward upgrade  
 195 plans to both improve on these measurements as well as explore a region of  $x$  and  $Q^2$  that has  
 196 yet to be probed with any facility. The STAR forward upgrade will accomplish this by installing  
 197 a new tracking system and a new calorimeter system. This upgrade will utilize the newest avail-  
 198 able technology and build on RHIC and STAR's unique capabilities to carry out measurements in  
 199 polarized  $pp$  collisions.



**Figure 14:** Side view of proposed STAR forward upgrade showing the various components to be installed and some of their details.

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