



Charge Asymmetry Dependence of Pion Azimuthal Anisotropy in Au + Au Collisions at STAR

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CME & CSE

• Chiral Magnetic Effect (CME): nonzero axial charge density induces a vector (electric) current along external magnetic field.

$$\boldsymbol{j}_V = \frac{N_c e}{2\pi^2} \mu_A \boldsymbol{B} \quad \Rightarrow \quad \text{electric charge separation along } \boldsymbol{B} \text{ field}$$

• Chiral Separation Effect (CSE): nonzero vector charge density induces an axial current along external magnetic field.

$$\boldsymbol{j}_A = \frac{N_c e}{2\pi^2} \mu_V \boldsymbol{B} \Rightarrow$$
 chiral charge separation along *B* field

D. E. Kharzeev, L. D. McLerran, and H. J. Warringa, Nuclear Physics A **803**, 227 (2008) T. Son, D. and A. R. Zhitnitsky, Phys. Rev. D **70**, 074018 (2004)

CME in Heavy Ion Collisions



- QCD vacuum transition in heavy ion collision $\rightarrow Q_w \neq 0$
- Extremely strong magnetic field created in heavy ion collision
- CME causes out-of-plane electric charge separation
- Charge asymmetry w.r.t. reaction plane as a signature of LPV.

CME in Heavy Ion Collisions



- Charge particle azimuthal angle distribution
 - $\Rightarrow \quad \frac{\mathrm{d}N_{\alpha}}{\mathrm{d}\phi} \propto 1 + 2v_{1,\alpha}\cos(\Delta\phi) + 2v_{2,\alpha}\cos(2\Delta\phi) + \cdots$

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+ 2a_{1,\alpha}\sin(\Delta\phi) + 2a_{2,\alpha}\sin(2\Delta\phi) + \cdots
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 $\diamond \quad \text{Correlations} \langle a_{\alpha} a_{\beta} \rangle \text{ can be measured}$

$$\diamond \quad \alpha, \beta = +, -$$

• The three-particle correlations are directly sensitive to predicted local *P*-violation in heavy-ion collisions.

$$\left<\cos\left(\phi_{\alpha}+\phi_{\beta}-2\Psi_{\rm RP}\right)\right>$$

- Out-of-plane charge separation same charge correlation < 0 opposite charge correlation > 0
- The observed signal cannot be described by models (HIJING, HIJING + v_2 , URQMD, MEVSIM)

S. A. Voloshin, Phys. Rev. C **70**, 057901 (2004) STAR Phys. Rev. Lett. **103**, 251601 (2009), Phys. Rev. C **81**, 054908 (2010)



Chiral Magnetic Effect

$$\boldsymbol{j}_V = \frac{N_c \ e}{2\pi^2} \mu_A \boldsymbol{B}$$

Chiral Separation Effect

$$\boldsymbol{j}_A = rac{N_c \ e}{2\pi^2} \mu_V \boldsymbol{B}$$

Chiral Magnetic Wave: coupling two effects

important at high beam energy:

enough energy for "jumping" over barriers to create initial axial charge density

important at low beam energy:

there is sizable initial vector charge density to set off chiral magnetic waves

$$\left(\partial_0 \mp \frac{N_c e B \alpha}{2\pi^2} \partial_1 - D_L \partial_1^2\right) j_{L,R}^0 = 0$$



D. Kharzeev and H.-U. Yee, Physical Review D 83, 085007 (2011)

Chiral Magnetic Wave



Y. Burnier, D. E. Kharzeev, J. Liao and H-U Yee, Phys. Rev. Lett. 107, 052303 (2011)

STAR Observable for Chiral Magnetic Wave

$$\frac{\mathrm{d}N_{\pm}}{\mathrm{d}\phi} = N_{\pm} \left[1 + 2v_2 \cos(2\phi)\right] \qquad \qquad A_{ch} = \frac{N_{\pm} - N_{\pm}}{N_{\pm} + N_{\pm}}$$
$$\approx \bar{N_{\pm}} \left[1 + 2v_2 \cos(2\phi) \mp A_{ch}r \cos(2\phi)\right] \qquad \qquad r = 2\left(\frac{q_e}{\bar{\rho}_e}\right)$$
$$v_2^{\pm} = v_2 \mp \left(\frac{q_e}{\bar{\rho}_e}\right) A_{ch}$$

$$v_2^{\pm} = v_2 \mp \left(\frac{q_e}{\bar{\rho}_e}\right) A_{ch}$$

- $v_2(\pi^-) > v_2(\pi^+)$
- $v_2(\pi^-)$ and $v_2(\pi^+)$ have opposite • trend as a function of A_{ch}
- difference between $v_2(\pi^-)$ and $v_2(\pi^+)$ • has a linear relationship with A_{ch}





$$v_2^- - v_2^+ = 2\left(\frac{q_e}{\bar{\rho}_e}\right)A_{ch}$$
 $r = 2q_e/\bar{\rho}_e$



Y. Burnier, D. E. Kharzeev, J. Liao and H-U Yee, Phys. Rev. Lett. 107, 052303 (2011) and arXiv: 1208.2537



STAR Experiment





Ionization energy loss in the Time Projection Chamber





Analysis Details

- Data Set
 - Au + Au
 - Minimum Bias trigger
 - 0 80% centrality

√s _{NN} (GeV)	Events (M)
200	238
62.4	63
39	104
27	46
19.6	23
11.5	10
7.7	4

- Pion Selection
 - PID: $|n\sigma_{\pi}| < 2$
 - $0.15 < p_{\rm T} < 0.5 \, {\rm GeV/c}$
 - |η| < 1.0

Particles for charge asymmetry

- charged particle
- $0.15 < p_{\rm T} < 12 \, {\rm GeV/c}$
- |η| < 1.0
- exclude (anti-)protons with $p_{\rm T} < 0.4 \text{ GeV/c}$

STAR Observed Charge Asymmetry



• N_+ and N_- : number of positive and negative particles

• observed
$$A_{ch}$$

$$A_{ch} = \frac{N_{+} - N_{-}}{N_{+} + N_{-}}$$

• Each bin has roughly the same number of events

Integrated Elliptic Flow



- $v_2(\pi^+)/v_2(\pi^-)$ integrated over 0.15 < $p_T < 0.5 \text{ GeV}/c$
- $v_2(\pi) > v_2(\pi)$
- v₂(π⁻) and v₂(π⁺) have opposite trend as a function of A_{ch}
- difference between v₂(π⁻) and v₂
 (π⁺) has a linear relationship with A_{ch}

All of the above are consistent with calculation based on Chiral Magnetic Wave

Integrated Elliptic Flow



- $v_2(\pi^+)/v_2(\pi^-)$ integrated over 0.15 < $p_T < 0.5 \text{ GeV}/c$
- $v_2(\pi) > v_2(\pi^+)$
- v₂(π⁻) and v₂(π⁺) have opposite trend as a function of A_{ch}
- difference between $v_2(\pi^-)$ and $v_2(\pi^+)$ has a linear relationship with $A_{\rm ch}$

All of the above are consistent with calculation based on Chiral Magnetic Wave

- $v_2 \approx 0.1 \times p_{\rm T}$ with $p_{\rm T} < 1 {\rm GeV}/c$
- Mean p_T difference between π⁻ and π⁺ gives about 1% of the observed v₂ difference

STAR Detector inefficiency Correction on A_{ch}



Fit the Slope Parameter





Slope vs. Centrality



- Similar trends between data and theoretical calculations with CMW, which use different magnetic field calculations, i.e. DH and BS(not shown here), in arXiv: 1208.2537.
- There is no specific beam energy input for the theoretical calculation.
- UrQMD cannot reproduce the slopes at $\sqrt{s_{NN}}$ = 200 GeV.

Slope vs. Centrality in U+U Collisions



• A similar centrality dependence of the slope parameter has been seen in U + U collisions at $\sqrt{s_{NN}}$ = 193 GeV

Slope vs. Centrality for v_2 of K^{\pm}



- Theoretical calculations suggest the slope of $v_2(A_{\rm ch})$ for *K* can be different from that of π
- The slopes for $v_2(A_{ch})$ for *K* have the same magnitude as that from π in the middle centralities with large error bars. More experimental data are needed to make the measurements of *K* conclusive.

STAR Slope vs. Centrality from v_3 of π^{\pm}



- The trend of $\Delta v_3(\pi)$ slope is similar to that of $\Delta v_2(\pi)$, which needs more study to understand
- The $\Delta v_3(\pi)$ slope is consistent with zero in mid-central collisions in contrast to finite values

A. Bzdak and P. Bozek, Physics Letters B 726 (2013) 239

Energy Dependence



- The slope parameter *r* shows a rise and fall feature from central to peripheral collisions
- Slope as a function of centrality shows weak energy dependence down to 27GeV
- Similar trend to the theoretical calculations with CMW

Y. Burnier et al., arXiv:1208.2537 [hep-ph]



Summary

- ► The difference between $v_2(\pi)$ and $v_2(\pi^+)$ shows a linear dependence on charge asymmetry in Au + Au collisions at $\sqrt{s_{NN}}$ = 200, 62.4, 39 and 27 GeV and in U + U collisions at $\sqrt{s_{NN}}$ = 193 GeV. The UrQMD model calculations cannot reproduce this feature in Au + Au collisions at $\sqrt{s_{NN}}$ = 200 GeV.
- As a function of collision centrality, the slope parameter r for π^{\pm} shows a rise and fall feature from central to peripheral collisions and the energy dependence seems weak from 200 to 27 GeV.
- > The centrality dependence of the $\Delta v_2(\pi)$ slope shows a similar trend of the calculations based on Chiral Magnetic Wave at 200GeV. However, model calculations do not reproduce the centrality dependence exactly.
- > The $v_3(\pi^{\pm})$ as a function of A_{ch} has also been studied in Au + Au and U + U collisions. The centrality dependence of $\Delta v_3(\pi)$ slope is similar to that of $\Delta v_2(\pi)$ slope. However, the $\Delta v_3(\pi)$ slope is consistent with zero in mid-centrality collisions.
- > At low energies, i.e. $\sqrt{s_{NN}}$ = 19.6, 11.5 and 7.7 GeV, the slopes are consistent with zero with large bar errors.