D⁰-meson elliptic flow measurement in Au+Au collisions



at $\sqrt{s_{NN}}$ = 200 GeV from STAR

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Abstract

A recent result from the STAR experiment shows that in 10-40% central Au+Au collisions at the top RHIC energy the elliptic flow v₂ of D⁰-meson follows the Number-of-Constituent-Quark scaling in the same way as it does for light flavor hadrons. This suggests that charm quarks have gained sufficiently large collectivity through their interactions with the Quark-Gluon Plasma (QGP). In this poster, we present the centrality and transverse momentum dependences of the D⁰-meson v₂ measured in Au+Au collisions at $\sqrt{s_{NN}}$ = 200 GeV by the STAR experiment. The measurement is based on the combined datasets recorded in 2014 and 2016, which improve the precision by 30% compared to the previously published results.

Motivation	v ₂ Results	
Elliptic flow (v_2) - second order Fourier coefficient of the azimuthal distribution.	$\begin{bmatrix} 9 \\ 9 \\ 2500 \end{bmatrix} = \begin{array}{c} \text{STAR Au+Au } \sqrt{s_{NN}} = 200 \text{ GeV} \\ 2016, 10-40\%, 2.5 < p_m < 3.0 \text{ GeV} \\ \end{array} = \begin{array}{c} 0.3 \\ >^{N} \end{array} = \begin{array}{c} \text{STAR Au+Au } \sqrt{s_{NN}} = 200 \text{ GeV}, 10-40\% \\ >^{N} \end{array}$	



- $D^0 v_2$ studies the heavy quark-medium interaction and the degree of thermalization of the charm quarks
- Results on $D^0 v_2$ using 2014 data indicate significant flow and are consistent with that of light hadrons in 10-40% central collisions.



D⁰ Reconstruction

- STAR Au+Au $\sqrt{s_{NN}}$ = 200 GeV 2014 + 2016 dataset: ~0.9B (2014)+1.2B (2016) events
- - (HFT)

 - TOF PID: $|1/\beta 1/\beta_K| < 0.03$, $|1/\beta 1/\beta_\pi| < 0.03$
- D^0 topological cuts: \succ
 - bin separately, and also for each centrality bin separately for year 2016 data





Fig. 1: D⁰ yields are measured in bins of the relative azimuthal angle with respect to the event plane angle (ϕ - Ψ_2). D⁰ decayed daughters are excluded from event plane reconstruction.

 v_2 is measured by fitting the azimuthal distribution with the function :

A * $(1 + 2 * v_2 \cos(2 * (\phi - \Psi_2)))$





Fig. 2: $D^0 v_2$ as a function of p_T from 2014 and 2016 STAR data. Data points are shifted along x-axis for clarity.

• v_2 results from 2014 and 2016 are combined as:



• Systematic uncertainties are assumed fully correlated between 2014 and 2016 dataset



Conclusions

- The D⁰ v₂ in Au+Au collisions at $\sqrt{s_{NN}}$ = 200 GeV is measured by STAR utilizing Heavy Flavor Tracker. Results using 2014 and 2016 dataset are consistent with each other and are combined.
- The D⁰ v₂/n_a is consistent with other hadrons at $(m_T m_0)/n_a < 2.5 \text{ GeV/c}^2$ in 0-10% and 10-40% centrality intervals. Results suggest that charm quarks exhibit the same strong collective behavior as light hadrons and may be close to thermal equilibrium in Au+Au collisions at $\sqrt{s_{NN}}$ = 200 GeV.
- **Outlook** Further study : Enlarge η gap to further reduce non-flow effect, and to study D⁰ v₂ in the peripheral (40-80%) interval.
- Study higher Fourier coefficient of the azimuthal distribution, such as triangular flow (v_3) , and also correlations with light hadron flow.

References

[1] M. He, R. J. Fries, and R. Rapp, Phys. Rev. C86, 014903 (2012); Phys. Rev. Lett. 110, 112301 (2013); and private communication. [2] V. Ozvenchuk, J. Torres-Rincon, P. Gossiaux, J. Aichelin, and L. Tolos, Phys. Rev. C90, 054909 (2014); M. Nahrgang, J. Aichelin, S. Bass, P. B. Gossiaux, and K. Werner, Phys. Rev. C91, 014904 (2015); and private communication. [3] S. S. Cao, G. Qin, and S. A. Bass, Phys. Rev. C92, 024907(2015); and private communication [4] S. Cao, T. Luo, G.-Y. Qin, and X.-N. Wang, Phys. Rev. C 94, 014909 (2016).

[5] H. Berrehrah, P. B. Gossiaux, J. Aichelin, W. Cassing, J. M. Torres-Rincon, and E. Bratkovskaya, 90, 051901 (2014); T. Song, H. Berrehrah, D. Cebrera, J. M. Torres-Rincon, L. Tolos, W. Cassing, and E. Bratkovskaya, 92, 014910 (2015).

[6] L. G. Pang, Y. Hatta, X. N. Wang, and B. W. Xiao, Phys. Rev. D 91, 074027 (2015); L. G. Pang, Q. Wang, and X. N. Wang, Phys. Rev. C 86, 024911 (2012); (private communication).

[7] Francesco Scardina, Santosh K. Das, Vincenzo Minissale, Salvatore Plumari, Vincenzo Greco, arXiv:1707.05452

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