

1            Constraints on neutron skin thickness and nuclear  
2 deformations using relativistic heavy-ion collisions from  
3            STAR\*

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7            In these proceedings, we present the measurements of neutron skin  
8 thickness and nuclear deformation using isobar  $^{96}_{44}\text{Ru}+^{96}_{44}\text{Ru}$  and  $^{96}_{40}\text{Zr}+^{96}_{40}\text{Zr}$   
9 collisions at  $\sqrt{s_{\text{NN}}} = 200$  GeV by the STAR detector. The significant devi-  
10 ations from unity of the isobar ratios of elliptic flow  $v_2$ , triangular flow  $v_3$ ,  
11 mean  $p_{\text{T}}$  fluctuations  $\langle\delta p_{\text{T}}^2\rangle/\langle p_{\text{T}}\rangle^2$ , and asymmetric cumulant  $\text{ac}_2\{3\}$  indi-  
12 cate large differences in their quadrupole and octuple deformations. The  
13 significant deviations of the isobar ratios of produced hadron multiplicity  
14  $N_{\text{ch}}$ , mean transverse momentum  $\langle p_{\text{T}}\rangle$ , and net charge number  $\Delta Q$  indicate  
15 a halo-type neutron skin for the Zr nucleus, much thicker than for the Ru  
16 nucleus, consistent with nuclear structure calculations. We discuss how we  
17 extract the neutron skin thickness, the symmetry energy slope parameter,  
18 and deformation parameters from data.

19            **1. Introduction**

20            The isobar collisions,  $^{96}_{44}\text{Ru}+^{96}_{44}\text{Ru}$  and  $^{96}_{40}\text{Zr}+^{96}_{40}\text{Zr}$ , were originally pro-  
21 posed to search for the chiral magnetic effect (CME) [1]. Based on the blind  
22 analysis with about 2 billion minimum-bias events for each species collected  
23 by the STAR experiment, the initial premise of isobar collisions for CME  
24 search was not realized [2]. The CME-related backgrounds were found to  
25 differ between isobar collisions [2], suggesting sizeable nuclear structure dif-  
26 ferences between the two isobar nuclei. Those nuclear structure differences  
27 and their consequences in experimental observables had been predicted by  
28 energy density functional theory (DFT) calculations [3, 4]. Owing to the  
29 large statistics collected by the STAR detector and robust cancellation of  
30 systematic uncertainties on the observable ratio  $R(X) \equiv \frac{X_{\text{RuRu}}}{X_{\text{ZrZr}}}$ , isobar col-  
31 lisions provide novel and accurate means to constrain the nuclear structures  
32 and strong force parameters.

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33 Nuclear deformation, a ubiquitous phenomenon for most atomic nuclei,  
 34 reflects collective motion induced by the interaction between valence nucle-  
 35 ons and shell structure. In most cases, the deformation has a quadrupole  
 36 shape that is characterized by overall strength  $\beta_2$  and triaxiality  $\gamma$ , and/or  
 37 an octupole shape  $\beta_3$ . In relativistic collisions of two nuclei such deforma-  
 38 tions enhance the fluctuations of bulk observables that are sensitive to initial  
 39 state geometry [5]. The deformation parameters can be constrained from  
 40 the precision measurements of the ratios between two isobar systems of  
 41 harmonic anisotropy coefficients  $v_2$ ,  $v_3$ , mean transverse momentum [ $p_T$ ]  
 42 fluctuations (mean, variance and skewness), and their Pearson correlation  
 43 coefficient  $\rho(v_2^2\{2\}, [p_T])$  [6]. Similar analysis has been performed in Au+Au  
 44 and U+U collisions [7, 8].

45 Neutron skin thickness  $\Delta r_{np}$  ( $\equiv \sqrt{\langle r_n^2 \rangle} - \sqrt{\langle r_p^2 \rangle}$ , the root mean square  
 46 difference between neutron and proton distributions) of nuclei can infer  
 47 nuclear symmetry energy. Such information is of critical importance to  
 48 the equation of state of dense nuclear matter in neutron stars.  $\Delta r_{np}$  has  
 49 traditionally been measured in low-energy hadronic and nuclear scatter-  
 50 ing experiments. Recent measurement using parity-violating electroweak  
 51 interactions by the PREX-II experiment has yielded a large neutron skin  
 52 thickness of Pb nucleus [9], at tension with the world-wide data established  
 53 in hadronic collisions. In isobar collisions at relativistic energies, neutron  
 54 skin was predicted [3, 4] to affect event multiplicity and elliptic anisotropy.  
 55 Measurements of those quantities can, in turn, offer an unconventional and  
 56 perhaps more precise means to probe the neutron skin [10]. Specifically,  
 57 the ratios between isobar collisions of the produced hadron multiplicities  
 58 ( $N_{ch}$ ) [10], the mean transverse momenta ( $\langle p_T \rangle$ ) [11], and the net charge  
 59 multiplicities ( $\Delta Q$ ) [12] can probe the neutron skin difference between the  
 60 isobar nuclei.

## 61 2. Nuclear deformation measurements in isobar collisions

62 Anisotropic flow in most central collisions is exquisitely sensitive to nu-  
 63 clear deformation. The deviations of  $R(v_2)$  and  $R(v_3)$  from unity observed  
 64 in most central isobar collisions [2] indicate a large quadrupole deforma-  
 65 tion in Ru nucleus and a large octupole deformation in Zr nucleus [5]. Fig-  
 66 ure 1 shows the  $R(v_2)$  and  $R(v_3)$  as a function of charged track multiplicity  
 67 ( $N_{trk}^{offline}$ ) with  $|\eta| < 0.5$ . We simulate events by a multi-phase transport  
 68 (AMPT) model with varying  $\beta_2$  for Ru (and fixed  $\beta_2 = 0.06$  for Zr) and  
 69 varying  $\beta_3$  for Zr (and fixed  $\beta_3 = 0$  for Ru) to match the data to extract the  
 70 best parameter values. The extracted quadrupole deformation parameter  
 71 for Ru is  $\beta_{2,Ru} = 0.16 \pm 0.02$  and the octupole deformation parameter for Zr is  
 72  $\beta_{3,Zr} = 0.20 \pm 0.02$ . The AMPT results with those deformation parameters

73 are also shown in Fig. 1. Those deformation parameters extracted from iso-  
 74 bar collisions are consistent with the measurements from traditional nuclear  
 75 structure experiments [13, 14]. We also show in Fig. 1 right panel the isobar  
 76 ratios of mean  $p_T$  fluctuations  $R(\langle\delta p_T^2\rangle/\langle p_T\rangle^2)$ . The trend can be qualita-  
 77 tatively described by the Glauber model with the deformation parameters for  
 Ru and Zr.

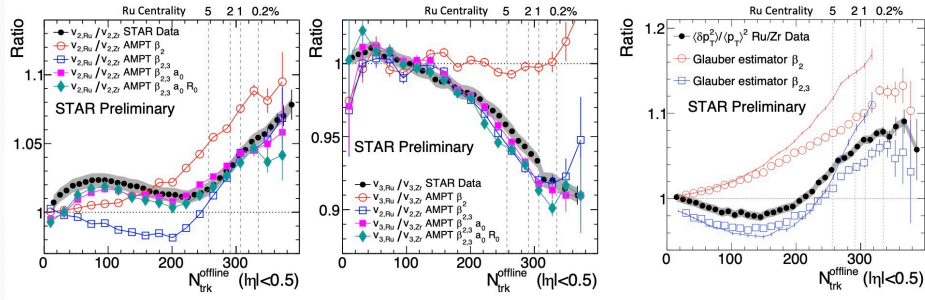


Fig. 1. The  $R(v_2)$  (left),  $R(v_3)$  (middle), and  $R(\langle\delta p_T^2\rangle/\langle p_T\rangle^2)$  (right) as a function of  $N_{\text{trk}}^{\text{offline}}$  in isobar collisions at  $\sqrt{s_{\text{NN}}} = 200$  GeV.

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The multi-particle correlations are also sensitive to nuclear deformation. In Fig. 2, we present the difference between isobar collisions of the three-particle asymmetric cumulant  $\text{ac}_2\{3\}$ . Here  $\text{ac}_2\{3\} \equiv \langle\langle e^{i(2\varphi_1+2\varphi_2-4\varphi_3)} \rangle\rangle$  is the average of three-particle azimuthal correlations over an ensemble of events. The  $R(\text{ac}_2\{3\})$  trend is similar to that of  $R(\langle v_2^4 \rangle)$  as shown in Fig. 2. The double ratio shown in the bottom panel of Fig. 2 indicates that the non-linear response coefficients  $\chi_{4,22} = \text{ac}_2\{3\}/\langle v_2^4 \rangle$  are almost identical for the two isobar systems. Both the trends of  $R(\text{ac}_2\{3\})$  and  $R(\chi_{4,22})$  can be reproduced by the AMPT model simulations. The  $R(\text{ac}_2\{3\})$  data, being extra sensitive to flow fluctuations, can help further constrain the deformation parameters of isobar nuclei [15].

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### 3. Neutron skin measurements

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Nuclear density distributions, and thus the neutron skin thicknesses, depend on the slope parameter  $L$  of symmetry energy as a function of nuclear density. With a given  $L$  of the nuclear interaction potential, the nuclear density can be calculated by the DFT framework. The event multiplicity produced in heavy-ion collisions is sensitive to the density distributions of the colliding nuclei, and thus the  $L$ . In Fig. 3, we present the ratio of the multiplicity distributions measured in isobar collisions within pseudo-rapidity  $|\eta| < 1$ , and those computed by the Monte Carlo (MC) Glauber model with the density distributions calculated by DFT with three

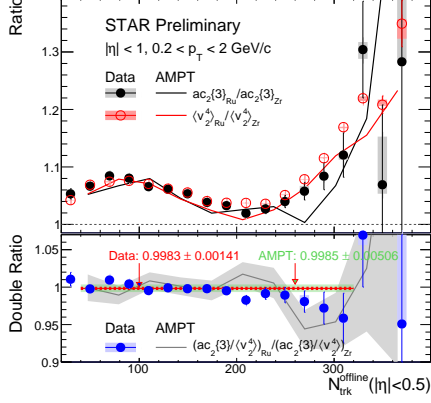


Fig. 2. The three-particle asymmetry cumulant ratio  $R(ac_2\{3\})$  (top) and non-linear response coefficient ratio  $R(\chi_{4,22})$  (bottom) as a function of  $N_{\text{trk}}^{\text{offline}}$  in isobar collisions at  $\sqrt{s_{\text{NN}}} = 200$  GeV. AMPT simulations are also shown for comparison.

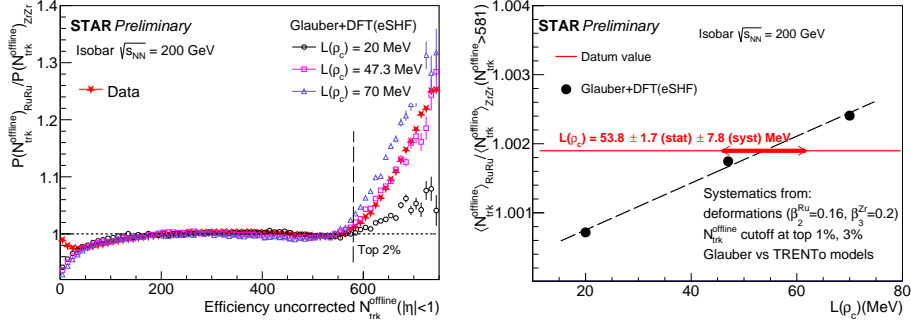


Fig. 3. Left: Ratio of the measured multiplicity ( $N_{\text{trk}}^{\text{offline}}$ ) distributions in isobar collisions at  $\sqrt{s_{\text{NN}}} = 200$  GeV. Also shown are Monte Carlo Glauber model results with nuclear densities calculated by DFT with three  $L(\rho_c)$  values. Right: Ratio of  $\langle N_{\text{trk}}^{\text{offline}} \rangle$  in top 2% central collisions as function of  $L(\rho_c)$  from Glauber model. The red line is the measured datum value 1.0019 with statistical uncertainties smaller than 0.0001.

100 values of  $L(\rho_c)$  [10], where  $\rho_c = 0.11\rho_0/0.16 \simeq 0.11 \text{ fm}^{-3}$  is nuclear sub-  
 101 saturation cross density. The model result with  $L(\rho_c) = 47.3$  MeV can  
 102 reasonably describe the data including the high multiplicity range. In the  
 103 right panel of Fig. 3 we compare the ratios of mean multiplicity at top  
 104 2% centrality between data and model calculations and obtain  $L(\rho_c) =$   
 105  $53.8 \pm 1.7(\text{stat.}) \pm 7.8(\text{syst.}) \text{ MeV}$ . The corresponding neutron skin thicknesses  
 106 for Ru and Zr are  $(\Delta r_{\text{np}})_{\text{Ru}} = 0.051 \pm 0.009 \text{ fm}$  and  $(\Delta r_{\text{np}})_{\text{Zr}} = 0.195 \pm 0.019$   
 107 fm, respectively. The systematic uncertainties are estimated with different  
 108 models (TRENTo vs. Glauber) and different  $N_{\text{trk}}^{\text{offline}}$  cutoffs, and considering  
 109 nuclear deformations which are the dominant contribution.

110 The mean transverse momentum  $\langle p_{\text{T}} \rangle$  also depends on the size of the  
 111 colliding nuclei. The sensitivity is the strongest in most central collisions.

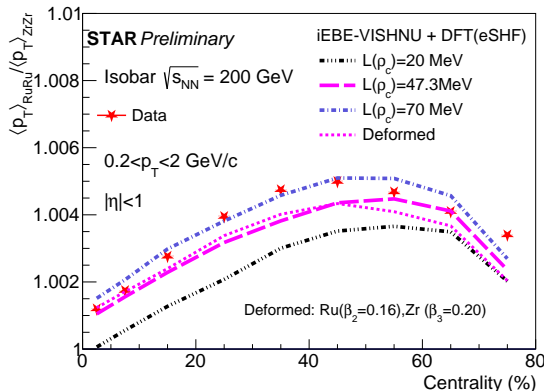


Fig. 4. The mean transverse momentum ratio  $R(\langle p_T \rangle)$  as a function of centrality in isobar collisions at  $\sqrt{s_{NN}} = 200$  GeV. The three thick curves are obtained from iEBE-VISHNU simulations with three different  $L(\rho_c)$  values, the thin-dashed curve denotes the model simulation with  $\beta_{2,\text{Ru}} = 0.16$  and  $\beta_{3,\text{Zr}} = 0.20$  under  $L(\rho_c) = 47.3$  MeV.

112 This provides another way to constrain the  $L(\rho_c)$  parameter. Although  
 113  $\langle p_T \rangle$  depends on bulk properties of the QGP medium, the  $R(\langle p_T \rangle)$  in isobar  
 114 collisions shows weak sensitivity to shear and bulk viscosities of the QGP  
 115 medium [11]. The  $R(\langle p_T \rangle)$  as a function of centrality is shown in Fig. 4. The  
 116 trend can be described by iEBE-VISHNU (Event-By-Event Viscous Israel  
 117 Stewart Hydrodynamics and UrQMD) model simulations, with the nuclear  
 118 densities obtained from the state-of-art energy density functional theory [10,  
 119 16]. Based on the  $R(\langle p_T \rangle)$  values at top 5% centrality, we extract the slope  
 120 parameter  $L(\rho_c) = 56.8 \pm 0.4(\text{stat.}) \pm 10.4(\text{sys.})$  MeV, and the corresponding  
 121 neutron skin thicknesses of the isobar nuclei of  $(\Delta r_{\text{np}})_{\text{Ru}} = 0.052 \pm 0.012$   
 122 fm and  $(\Delta r_{\text{np}})_{\text{Zr}} = 0.202 \pm 0.024$  fm. The systematic uncertainties are  
 123 also dominated by the nuclear deformation effect which can be improved in  
 124 the future. The results extracted from  $\langle p_T \rangle$  are consistent with those from  
 125 multiplicity distributions above.

126 We compare in Fig. 5 our  $L(\rho)$  (the slope parameter of symmetry energy  
 127 at saturation density  $\rho_0$ ) results with a compilation of world data from  
 128 traditional nuclear structure experiments [17]. Our results are consistent  
 129 with world data with comparable precision.

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#### 4. Summary

131 The isobar  ${}^{96}_{44}\text{Ru} + {}^{96}_{44}\text{Ru}$  and  ${}^{96}_{40}\text{Zr} + {}^{96}_{40}\text{Zr}$  collision data collected by STAR  
 132 at  $\sqrt{s_{NN}} = 200$  GeV provide novel means to probe the nuclear structure  
 133 and deformation of the isobar nuclei. From the isobar ratios of the mea-  
 134 sured anisotropic flow, multiplicity, and mean transverse momentum, with  
 135 the help of DFT calculations, we have extracted the nuclear deformation  
 136 parameters  $\beta_2, \beta_3$ , and neutron skin thicknesses  $\Delta r_{\text{np}}$  of the isobar nuclei,  
 137 and the density slope parameter of symmetry energy  $L(\rho_c)$ . The results are  
 138 consistent with world-wide data from traditional nuclear scattering experi-

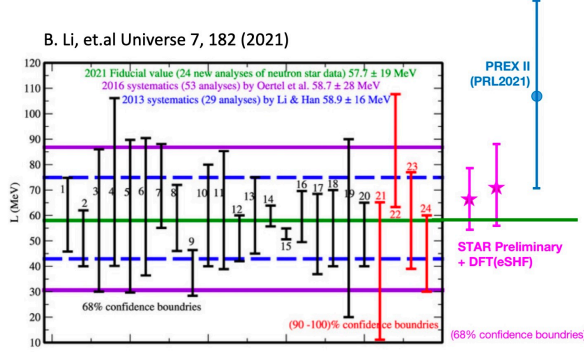


Fig. 5. The compilation of world data of  $L(\rho)$  from Ref. [17]. The  $L(\rho_c)$  values extracted from isobar data are converted into  $L(\rho)$  for comparison. The PREX-II data are taken from Ref. [9]

139 ments with comparable precision.

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## REFERENCES

- 143 [1] S. A. Voloshin, Phys. Rev. Lett. **105**, 172301 (2010).  
 144 [2] M. Abdallah *et al.* (STAR Collaboration), Phys. Rev. C **105**, 014901 (2022).  
 145 [3] H. j. Xu, X. Wang, H. Li *et al.*, Phys. Rev. Lett. **121**, 022301 (2018).  
 146 [4] H. Li, H. j. Xu, J. Zhao *et al.*, Phys. Rev. C **98**, no.5, 054907 (2018).  
 147 [5] C. Zhang and J. Jia, Phys. Rev. Lett. **128**, 022301 (2022).  
 148 [6] J. Jia and C. J. Zhang, [arXiv:2111.15559 [nucl-th]].  
 149 [7] J. Jia, S. Huang and C. Zhang, Phys. Rev. C **105**, 014906 (2022).  
 150 [8] C. Zhang [STAR], [arXiv:2203.13106 [nucl-ex]].  
 151 [9] D. Adhikari *et al.* (PREX Collaboration), Phys.Rev.Lett. **126**, 172502 (2021).  
 152 [10] H. Li, H. j. Xu, Y. Zhou *et al.*, Phys. Rev. Lett. **125**, 222301 (2020).  
 153 [11] H. j. Xu, W. Zhao, H. Li *et al.*, [arXiv:2111.14812 [nucl-th]].  
 154 [12] H. j. Xu, H. Li, Y. Zhou *et al.*, Phys. Rev. C **105**, L011901 (2022).  
 155 [13] T. Kibédi and R. h. Spear, Atom. Data Nucl. Data Tabl. **80**, 35-82 (2002).  
 156 [14] B. Pritychenko, M. Birch, B. Singh and M. Horoi, Atom. Data Nucl. Data  
 157 Tabl. **107**, 1-139 (2016) [erratum: *ibid.* **114**, 371-374 (2017)].  
 158 [15] S. Zhao, H. j. Xu, Y. X. Liu and H. Song, [arXiv:2204.02387 [nucl-th]].  
 159 [16] Z. Zhang and L. W. Chen, Phys. Rev. C **94**, no.6, 064326 (2016).  
 160 [17] B. A. Li, B. J. Cai, W. J. Xie and N. B. Zhang, Universe **7**, no.6, 182 (2021).