



The XXIXth International Conference On Ultrarelativstic Nucleus-Nucleus Collisions

Femtoscopy of Proton, Light nuclei, and Strange hadrons in Au+Au Collisions at STAR

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- 1. Introduction
- 2. STAR Experiment & Datasets
- 3. Results
 - 1) Meson-Meson Correlation Function $(K_s^0 K_s^0)$
 - **2)** Baryon-Baryon Correlation Function (p-p, p-Ξ⁻)
 - 3) Light Nuclei Correlation Function (p-d, d-d)
- 4. Summary & Outlook



Introduction

Two-particle correlations at small relative momenta contain information about the space-time characteristics of the emitting source and final-state interaction effects

- 1. Meson-Meson Interaction -> neutral kaons correlation
 - 1) Kaon can provide complementary information to pions
 - 2) Kaon correlation measurement offers space-time evolution involving strangeness



- 2. Baryon-Baryon Interaction -> p-p and p-E⁻ correlations
 - 1) p-p interaction can be used as baseline for other systems
 - 2) Hyperon-Nucleon(Y-N) & Hyperon-Hyperon(Y-Y) interaction: Important for understanding the inner structure of compact stars and the formation of bound states



Lonardoni D, Lovato A, Gandolfi S, Pederiva F. Phys. Rev. Lett. 114:092301 (2015)



Two-particle correlations at small relative momenta contain information about the space-time characteristics of the emitting source and final-state interactions effects

- 3. Light Nuclei Correlation -> p-d, d-d correlations
 - A systematic measurement of p-p, p-d, and d-d correlation functions may tell us whether deuterons are directly emitted from the fireball or formed due to final-state interactions
 - A large amount of light nuclei produced at 3 GeV -> Allowing precision measurements



St. Mrówczyński and P. Słoń, Acta Physica Polonica B 51, 1739 (2020) St. Mrówczyński and P. Słoń, Physical Review C 104, 024909 (2021)





- ✓ Femtoscopy (HIC) is inspired by Hanbury Brown and Twiss interferometry method (Astronomy)¹
 - **Study the spatial and temporal extent of emission source**
 - Quantum Statistics (Fermi-Dirac, Bose-Einstein)
 - Final-state Interactions (Coulomb, Strong)
 - Collision Dynamics

✓ Two-particle correlation function:





STAR Detector and Datasets



Excellent Particle Identification > Large, Uniform Acceptance at Mid-rapidity

STAR Fixed-target Experiment Setup



Collider

39 GeV

200 GeV

2010

2010

~86

~230



Particle Identification and Acceptance at $\sqrt{s_{NN}}$ = 3 GeV



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Results



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- 1. Centrality dependence: $R_{0-10\%} > R_{10-70\%}$
- 2. Energy dependence: $R_{200GeV} > R_{39GeV}$
- 3. Significant difference in radii between QS and Lednicky & Lyuboshitz

models **→** Final state interactions

 Antonelli: eConf C020620, THAT06 (2002)
 Achasov2001: Phys. Rev. D 63, 094007 (2001)

 Achasov2003: Phys. Rev. D 68, 014006 (2003)
 Martin: Nucl. Phys. B 121, 514–530 (1977)

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Proton-E⁻ Femtoscopy in Au+Au Collisions at 3 GeV



- **1.** Large uncertainties due to limited proton- Ξ^- pairs at low energy
- 2. Modelled by hadronic transport model UrQMD + an afterburner, model results show a similar trend as data For details, see poster by Zhi Qin

CARB: <u>https://web.pa.msu.edu/people/pratts/freecodes/crab/home.html</u>

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- 1. Clear centrality dependence seen -> Smaller source size in peripheral collisions
- 2. More siginificant rapidity dependence in peripheral collisions





- 1. Clear depletion at small k* range seen in data
- 2. Data compared with Lednicky & Lyuboshitz model^{1,2}
 - ➤ A spherical source size with r = 3-4 fm is consistent with data

1.LednickýR,LyuboshitzV.Sov.J.Nucl.Phys.35:770(1982) 2.J. Arvieux, Nucl. Phys. A 221 (1974) 253–268





- 1. Clear depletion at small k* range seen in data
- 2. Data compared with Lednicky & Lyuboshitz model^{1,2}
 - > A spherical source size with r = 4-5 fm is consistent with data -> Larger than p-d

1,I.N. Filikhin and S.L. Yakovlev, Phys. Atom. Nucl. 63, 55 (2000) 2,I.N. Filikhin and S.L. Yakovlev, Phys. Atom. Nucl. 63, 216 (2000)





- 1. Compared with SMASH + Correlation after burner (CRAB) model
- 2. CF calculated with coalescence of deuterons is in better agreement with data
 - Support the deuteron formation at 3 GeV is dominated by coalescence
- 3. SMASH source size: 4.3 5.9 (fm) from peripheral to central collisions

SMASH: J. Weil et al. Phys.Rev.C 94 (2016) 5, 054905 Coalescence: W.Zhao et al. Phys. Rev. C.98 (2018) 5,054905



Summary

- 1. K_s⁰K_s⁰ Correlation Function (39 GeV & 200 GeV)
 - 1) Significant difference in radii between QS and L&L models -> Final state interactions
- 2. Baryon-Baryon Correlation Function (3 GeV)
 - 1) Strong centrality dependence found in p-p CF -> Smaller source size in peripheral
 - 2) p-E⁻: UrQMD+Crab model result shows similar trend as data
- 3. Light Nuclei Correlation Function (3 GeV)
 - 1) First measurement of p-d and d-d correlation functions from STAR
 - 2) p-d and d-d CF qualitatively described by L&L model -> d-d has larger emission source size than p-d
 - 3) d-d CF described better by the model including coalescence
 - Light nuclei are likely to be formed via coalescence

Outlook

➢ In the 2nd phase of BES, STAR has collected 10-20 times more data in Au+Au collisions at the energy range $\sqrt{s_{NN}}$ = 3 - 19.6 GeV. These data allow us to perform precision femtoscopy analysis.
Stay tuned for the RHIC BES-II !



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