

1 Production yield and azimuthal anisotropy measurements  
2 of strange hadrons from BES at STAR \*

3 ASWINI KUMAR SAHOO (FOR THE STAR COLLABORATION)

4 Indian Institute of Science Education and Research (IISER), Berhampur, India  
5 Email: aswinikumar.aks96@gmail.com/aswini96@rcf.rhic.bnl.gov

6 *Received July 28, 2022*

7 **Abstract:** We report the production and azimuthal anisotropy mea-  
8 surements of strange and multi-strange hadrons at STAR BES energies.  
9 The  $\Lambda/K_s^0$  ratio is reported at 3 GeV and observed to increase faster  
10 with transverse momentum than that at higher energies. The number-  
11 of-constituent quark (NCQ) scaling of  $v_2$  has been studied at 19.6 GeV  
12 (BES-II). The NCQ scaling holds for particles and anti-particles, which  
13 can be considered as an evidence of partonic collectivity. The production  
14 of  $K^{*0}$  resonance is also reported for 7.7-39 GeV (BES-I) and the  $K^{*0}/K$   
15 ratio suggests that hadronic re-scattering dominates over regeneration in  
16 central A+A collisions. Using the  $K^{*0}/K$  ratio, we also report the lower  
17 limit of hadronic phase lifetime ( $t_{kin} - t_{chem}$ ).

18 **1. Introduction**

19 Searching for the onset of the deconfinement is one of the main motiva-  
20 tions of the Beam Energy Scan (BES) program at RHIC. The production  
21 yield and azimuthal anisotropy of (multi-)strange hadrons is considered a  
22 good probe to study the properties of the matter produced in heavy-ion  
23 collisions. The (multi-)strange hadrons are expected to freeze out earlier  
24 than other light hadrons, such as  $\pi, K, p$  [1]. They are also expected to  
25 have smaller hadronic interaction cross sections compared to non-strange  
26 hadrons [2]. Hence the production of (multi-)strange hadrons should not be  
27 strongly affected by the later stage of heavy-ion collisions.

28 Resonances, like  $K^{*0}(892)$ , are short-lived particles produced in high  
29 energy collisions.  $K^{*0}$ , having a smaller lifetime ( $\sim 4.16$  fm/c) than the  
30 medium ( $\sim 10$  fm/c), is expected to be sensitive to the dynamics in the  
31 hadronic phase. In between chemical (CFO) and kinetic (KFO) freeze out,  
32 the daughter particles of  $K^{*0}(892)$  could undergo re-scattering and regener-  
33 ation. The final yield of the  $K^{*0}(892)$  depends on the interplay of these

---

\* Quark Matter-2022

34 effects, and can be used to study the hadronic phase of heavy-ion collisions  
 35 [3].

## 36 2. Data Sets and Analysis Details

37 In these proceedings, we report strange hadrons yield and elliptic flow in  
 38 Au+Au collisions at  $\sqrt{s_{NN}} = 3$  and 19.6 GeV, accumulated by the STAR  
 39 experiment in 2018 and 2019 as part of the RHIC BES-II program. The  
 40  $K^{*0}$  spectrum analysis is performed using Au+Au collisions at  $\sqrt{s_{NN}} =$   
 41 7.7-39 GeV collected in 2010, 2011 and 2014. For particle identification  
 42 both the Time Projection Chamber (TPC) and the Time Of Flight (TOF)  
 43 detector are used. In BES-II, the TPC detector has been upgraded for better  
 44 momentum resolution and wider pseudo-rapidity coverage ( $|\eta| < 1.5$ ).

## 45 3. Results

### 46 3.1. Probing partonic phase with strange hadrons

#### 47 3.1.1. $\Lambda/K_s^0$ ratio

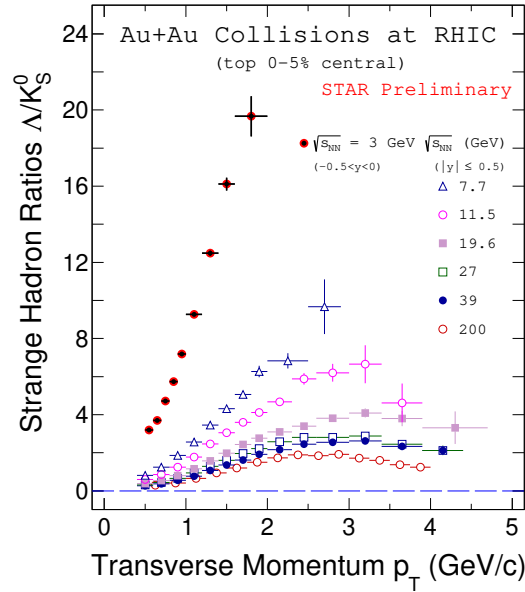


Fig.1.  $\Lambda/K_s^0$  as a function of  $p_T$  in Au+Au collisions at various beam energies [4] [5]. The bars indicate statistical uncertainties only .

48 The baryon-to-meson ratio can be used to investigate the particle pro-  
 49 duction mechanism in heavy-ion collisions. Figure 1 represents the variation  
 50 of  $\Lambda/K_s^0$  as a function of transverse momentum ( $p_T$ ) in central Au+Au col-  
 51 lision at various beam energies. We observe that  $\Lambda/K_s^0$  increases faster with  
 52  $p_T$  at 3 GeV compared to higher energies. According to the thermal model:

$$\frac{N(\Lambda)}{N(K_s^0)} \propto \exp\left(\frac{(1 - \sigma_s)}{T}\right) \quad \text{where } \sigma_s = \frac{\mu_s}{\mu_B} \quad (1)$$

53 Here T is the temperature and  $\mu_s$  and  $\mu_B$  are the strangeness chemical  
 54 potential and baryon chemical potential respectively. Since there is a signif-  
 55 icant difference in chemical potential at lower energies compared to higher  
 56 energies, this trend could be chemical potential driven.

### 57 3.1.2. Elliptic flow ( $v_2$ )

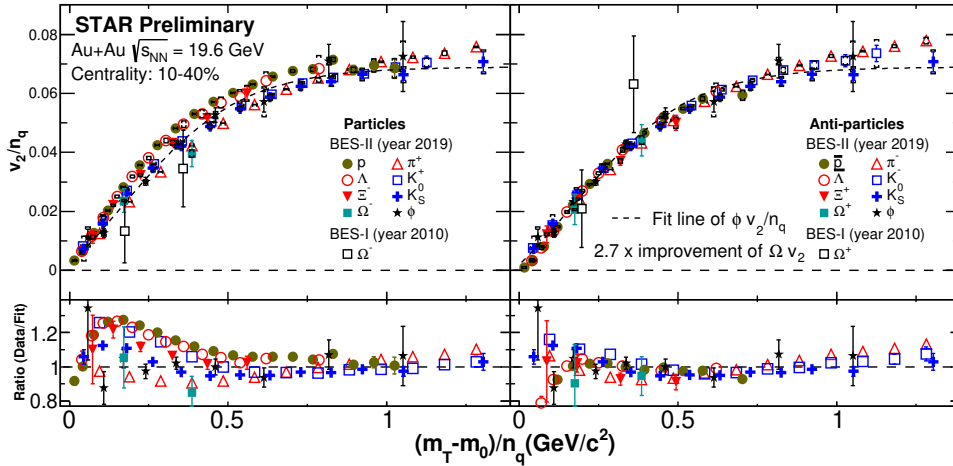


Fig. 2. The elliptic flow ( $v_2$ ) scaled by the number of constituent quarks ( $n_q$ ) as a function of  $(m_T - m_0)/n_q$  for particles and their corresponding anti-particles in Au+Au collisions at 19.6 GeV (BES-II) for 10-40% centrality. The bars and caps indicate statistical and systematic uncertainties respectively.

58 In the overlap region of two colliding nuclei the pressure gradient is  
 59 different in different direction that leads to momentum space anisotropy.  
 60 This anisotropy is the main cause for the development of elliptic flow ( $v_2$ ).  
 61 Hence  $v_2$  is sensitive to the initial dynamics of heavy-ion collisions. Figure 2  
 62 shows  $v_2$  divided by the number of constituent quarks ( $n_q$ ) as a function of

63  $(m_T - m_0)/n_q$ , where  $m_T = \sqrt{(p_T^2 + m_0^2)}$  is the transverse mass and  $m_0$  is  
 64 the rest mass of the hadron, at  $\sqrt{s_{NN}} = 19.6$  GeV (BES-II) for 10-40% cen-  
 65 trality. The NCQ scaling holds within 20% for particles and anti-particles,  
 66 which could be considered as a signature of partonic collectivity [6]. The  
 67 scaling holds better for anti-particles than for the particles, which might be  
 68 due to the transported quark effect.

### 69 3.2. Probing hadronic phase with $K^{*0}$ resonance

#### 70 3.2.1. $K^{*0}/K$ ratio and hadronic phase lifetime

71 The decay daughters of  $K^{*0}$  (i.e  $\pi$  and  $K$ ) may re-scatter with other par-  
 72 ticles during the hadronic phase of heavy-ion collisions. Meanwhile, pions  
 73 and kaons may regenerate  $K^{*0}$  via pseudo-elastic scattering. So the  $K^{*0}/K$   
 74 ratio can be used to probe the the relative contributions of these effects.

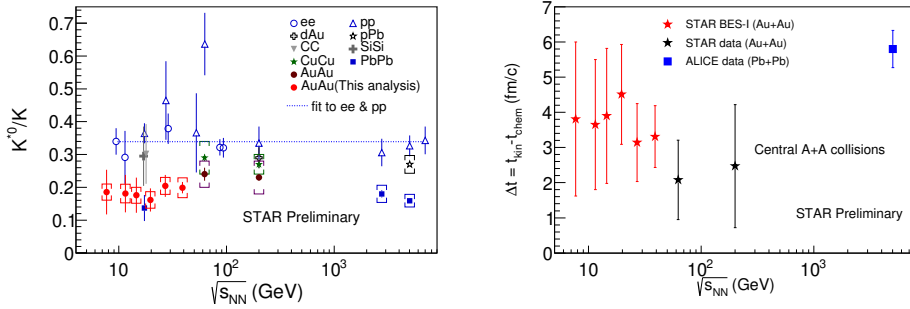


Fig. 3. Left panel:  $K^{*0}/K$  as a function of the collision energy [7, 8, 9, 10, 11, 12, 13, 14, 15, 16, 17, 18, 19, 20, 21, 22, 23]. Here the  $K^{*0}/K$  represents  $(K^{*0} + \bar{K}^{*0})/(K^+ + K^-)$ . The bars and caps indicate statistical and systematic uncertainties respectively. Right panel: Hadronic phase lifetime ( $\Delta t$ ) as a function of the collision energy. The result is compared with previous STAR [12, 16] and ALICE [21, 22, 23] results. The error bars are the quadratic sum of the statistical and systematic uncertainties.

75 In the left panel of Fig. 3 we have shown  $K^{*0}/K$  as a function of the col-  
 76 lision energy. Here we can see the ratio in central A+A collisions is smaller  
 77 than in elementary (e+e or p+p) collisions, indicating that re-scattering  
 78 might be dominant over re-generation in central A+A collisions.

79 We can also use the  $K^{*0}/K$  ratio to extract the lower limit of hadronic  
 80 phase lifetime [3] following [24], i.e

$$\left(\frac{K^{*0}}{K}\right)_{KFO} = \left(\frac{K^{*0}}{K}\right)_{CFO} \times e^{-\Delta t/\tau_{K^{*0}}}, \quad (2)$$

81 Here we have taken that the  $(K^{*0}/K)_{CFO}$  and  $(K^{*0}/K)_{KFO}$  are similar to  
 82 the  $K^{*0}/K$  ratios measured in elementary and heavy-ion collisions respec-  
 83 tively. We have assumed that (i) there is no  $K^{*0}$  regeneration taking place  
 84 between the chemical and kinetic freeze out, and (ii) all  $K^{*0}$  that decay  
 85 before the kinetic freeze out are lost due to the re-scattering effect

86 In the right panel of Fig. 3 we have shown the variation of  $\Delta t$  as a  
 87 function of  $\sqrt{s_{NN}}$ . Here we can see that measurements from RHIC seem  
 88 to be smaller than that at the LHC, However, more statistics is needed in  
 89 order to draw firm conclusions.

#### 90 4. Summary

91 The production yield and azimuthal anisotropy measurements of (multi-  
 92 )strange hadrons at STAR BES energies are reported. The rapid increase  
 93 in  $\Lambda/K_s^0$  ratio as a function of  $p_T$  at 3 GeV could be due to the change in  
 94 chemical potential at lower energies. The NCQ scaling of elliptic flow holds  
 95 for the particles and corresponding anti-particles, which could be due to  
 96 partonic collectivity. The suppression of  $K^{*0}/K$  ratio suggests that there is  
 97 a dominance of hadronic re-scattering in central heavy-ion collisions. Based  
 98 on the  $K^{*0}/K$  ratios, the extracted lifetime of the hadronic phase at RHIC  
 99 seems to be smaller than that at the LHC.

#### 100 5. Acknowledgements

101 Financial support from Department of Education, Government of India  
 102 is gratefully acknowledged.

#### REFERENCES

- 103 [1] H. van Hecke, H. Sorge, and N. Xu. *Phys. Rev. Lett.*, 81:5764–5767, 1998.  
 104 [2] Asher Shor. *Phys. Rev. Lett.*, 54:1122–1125, Mar 1985.  
 105 [3] C. Adler et al. *Phys. Rev. C*, 66:061901, 2002.  
 106 [4] Agakishiev et.al. *Phys. Rev. Lett.*, 108:072301, Feb 2012.  
 107 [5] J. et.al Adam. *Phys. Rev. C*, 102:034909, Sep 2020.  
 108 [6] Dénes Molnár and Sergei A. Voloshin. *Phys. Rev. Lett.*, 91:092301, Aug 2003.  
 109 [7] H. Albrecht et al. *Z. Phys. C*, 61:1–18, 1994.  
 110 [8] Yi-Jin Pei. *Z. Phys. C*, 72:39–46, 1996.  
 111 [9] Werner Hofmann. *Ann. Rev. Nucl. Part. Sci.*, 38:279–322, 1988.  
 112 [10] K. Abe et al. *Phys. Rev. D*, 59:052001, 1999.  
 113 [11] M. Aguilar-Benitez et al. *Z. Phys. C*, 50:405–426, 1991.

- 114 [12] J. Adams et al. *Phys. Rev. C*, 71:064902, 2005.  
115 [13] D. Drijard et al. *Z. Phys. C*, 9:293, 1981.  
116 [14] T. Akesson et al. *Nucl. Phys. B*, 203:27, 1982.  
117 [15] A. Adare et al. *Phys. Rev. C*, 90(5):054905, 2014.  
118 [16] M. M. Aggarwal et al. *Phys. Rev. C*, 84:034909, 2011.  
119 [17] B. I. Abelev et al. *Phys. Rev. C*, 78:044906, 2008.  
120 [18] T. Anticic et al. *Phys. Rev. C*, 84:064909, 2011.  
121 [19] Jaroslav Adam et al. *Eur. Phys. J. C*, 76(5):245, 2016.  
122 [20] Shreyasi Acharya et al. 10 2021.  
123 [21] Betty Bezverkhny Abelev et al. *Phys. Rev. C*, 91:024609, 2015.  
124 [22] Jaroslav Adam et al. *Phys. Rev. C*, 95(6):064606, 2017.  
125 [23] Shreyasi Acharya et al. *Phys. Lett. B*, 802:135225, 2020.  
126 [24] Subhash Singha, Bedangadas Mohanty, and Zi-Wei Lin. *Int. J. Mod. Phys.*  
127 *E*, 24(05):1550041, 2015.