

Femtoscopy of Proton, Light nuclei, and Strange hadrons in Au+Au Collisions at STAR *

KE MI (FOR THE STAR COLLABORATION)

Key Laboratory of Quark & Lepton Physics (MOE) and Institute of Particle Physics, Central China Normal University, Wuhan, 430079, China

Physics Institute, Heidelberg University, Heidelberg, 69117, Germany

Received July 29, 2022

1 In these proceedings, we present the measurements of proton, light nu-
2 clei, and strange particle with neutral kaons correlation functions in Au+Au
3 collisions at the BES program and top RHIC energy. The experimental
4 results will be compared with model calculations to extract the size of
5 emitting source and the properties of final state interactions. The collision
6 energy and centrality dependence of the source size will be studied. Fur-
7 ther, the implications for the production mechanism of light nuclei will be
8 discussed.

9 1. Introduction

10 Measurements of the correlation function for a pair of particles with
11 small relative momenta can provide insight into the geometry and lifetime
12 of the particle-emitting source in relativistic heavy-ion collisions [1]. By
13 measuring correlation at low relative velocities, one could access the smallest
14 sizes in nature, which corresponds to the size of a nucleon (1 fm), therefore
15 such a two-particle correlation method is called 'Femtoscopy' [2].

16 The fundamental observable in femtoscopy is the correlation function
17 $C(k^*)$. The theoretical definition is expressed as a function of the relative
18 distance between two particles \mathbf{r}^* and their reduced relative momentum, k^*
19 $= \frac{1}{2} * |\mathbf{p}_1^* - \mathbf{p}_2^*|$ in the pair rest frame, with $\mathbf{p}_1^* = -\mathbf{p}_2^*$, by the Koonin-Pratt
20 formula [3]:

$$C(k^*) = \int S(\mathbf{r}^*) |\psi(\mathbf{r}^*, \mathbf{k}^*)|^2 d^3r, \quad (1)$$

* Presented at Quark Matter 2022, Krakow, Poland, April 4-10, 2022

21 where the first term $S(\mathbf{r}^*)$ describes the source that emits the particles; the
 22 second term contains the interaction part via the two-particle wave function
 23 $\psi(\mathbf{r}^*, \mathbf{k}^*)$.

24 The experimental correlation function is defined as a ratio of the proba-
 25 bility of registering two particles simultaneously (in the same event) to the
 26 product of registering probabilities of such particles independently (in the
 27 mixed event):

$$C(k^*) = \frac{N_{\text{same}}(k^*)}{N_{\text{mixed}}(k^*)}. \quad (2)$$

28 The shape and amplitude of the correlation function are determined
 29 by the quantum statistical effect (identical pair) and final state interaction
 30 (strong and Coulomb interactions).

31 2. Particle Identification and Reconstruction

32 Particle identification (PID) of protons and deuterons are done by using
 33 the energy loss (dE/dx) information measured by Time Projection Cham-
 34 ber (TPC) (Fig. 1 (a)) and particles' m^2 information from Time Of Flight
 35 (TOF) (Fig. 1 (b)). A combination of dE/dx and m^2 criteria is used to
 36 identify the particles with purity higher than 96%.

37 The K_s^0 particles are reconstructed via a weak decay channel: $K_s^0 \rightarrow$
 38 $\pi^+ + \pi^-$ (Fig. 1(c)). A set of topological cuts were used to ensure the
 39 reconstructed K_s^0 purity higher than 95%.

40 3. Results

41 3.1. $K_s^0-K_s^0$ Correlation Functions at $\sqrt{s_{\text{NN}}} = 39 \text{ GeV}$ and 200 GeV

42 Figure 2 shows the $K_s^0-K_s^0$ correlation functions as a function of q_{inv}
 43 ($q_{\text{inv}} = 2k^*$) at $\sqrt{s_{\text{NN}}} = 39 \text{ GeV}$ and 200 GeV with three centrality classes
 44 (0-10%, 10-70%, and 0-70%) in Au+Au collisions. The correlation functions
 45 show dip structures around $q_{\text{inv}} = 0.1 \text{ GeV}/c$, which is caused by the near-
 46 threshold $f_0(980)$ and $a_0(980)$ resonances. The parameterization is done by
 47 using Gaussian function which only includes quantum statistics (QS):

$$C(q_{\text{inv}}) = 1 + \lambda e^{[-R^2 q_{\text{inv}}^2]}, \quad (3)$$

48 and Lednický-Lyuboshitz model [2] (L-L model) which includes both QS
 49 and strong interaction (SI) :

$$C(q_{\text{inv}}) = 1 + \lambda \left(e^{[-R^2 q_{\text{inv}}^2]} + \frac{1}{2} \left[\left| \frac{f(k^*)}{R} \right|^2 + \frac{4\Re f(k^*)}{\sqrt{\pi}R} F1(q_{\text{inv}}R) - \frac{2\Im f(k^*)}{\sqrt{\pi}R} F2(q_{\text{inv}}R) \right] \right), \quad (4)$$

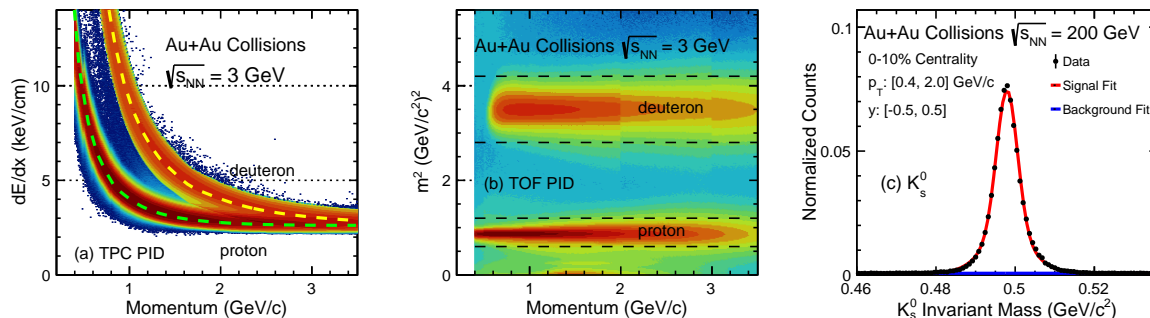


Fig. 1. (a) The dE/dx distribution of protons and deuterons versus total momentum in Au+Au collisions at $\sqrt{s_{\text{NN}}} = 3$ GeV. The dashed curves are the corresponding Bichsel expectations. (b) Particle mass square (m^2) distribution versus total momentum in Au+Au collisions at $\sqrt{s_{\text{NN}}} = 3$ GeV. The dashed lines represent the m^2 ranges used in the analysis. (c) The K_s^0 invariant mass distribution in 0-10% centrality at $\sqrt{s_{\text{NN}}} = 200$ GeV. The red line represents the fit to the data including signal and background, and blue line to the background alone.

50 where R is the invariant radius, λ is correlation strength, $f(k^*)$ is the s-wave
 51 $K^0\bar{K}^0$ scattering amplitude. In this work, we use the resonance masses and
 52 couplings from Ref. [4–7]. The experimental data can be well described by
 53 L-L model, while the Gaussian function failed to describe the dip structure.
 54 This implies that the strong final state interaction has a significant effect
 55 between $K_s^0-K_s^0$ pairs. Figure 3 shows the centrality dependence and energy
 56 dependence of the extracted radii. The results suggest a decreasing trend
 57 from central to peripheral collisions and an increasing trend from low to high
 58 energy. Also, a significant difference in radii between QS and L-L model is
 59 found.

60 3.2. Light Nuclei Correlation Functions at $\sqrt{s_{\text{NN}}} = 3$ GeV

61 The production mechanisms of light nuclei are still under debate. Some
 62 experimental methods to resolve the problem have been proposed [8]. Fem-
 63 toscopy of light nuclei may provide a unique tool to obtain detailed infor-
 64 mation about the production mechanism. At low beam energies ($\sqrt{s_{\text{NN}}} <$
 65 20 GeV), an enhancement in the light nuclei yield is expected due to the high
 66 baryon density [9]. This provides a great opportunity for precise femtoscopy
 67 that involves light nuclei [10].

68 Figure 4 shows the first measurements of $p-d$ and $d-d$ correlation func-
 69 tions at mid-rapidity with four centrality classes in Au+Au collisions at
 70 $\sqrt{s_{\text{NN}}} = 3$ GeV. Clear depletions below unity due to final state interaction

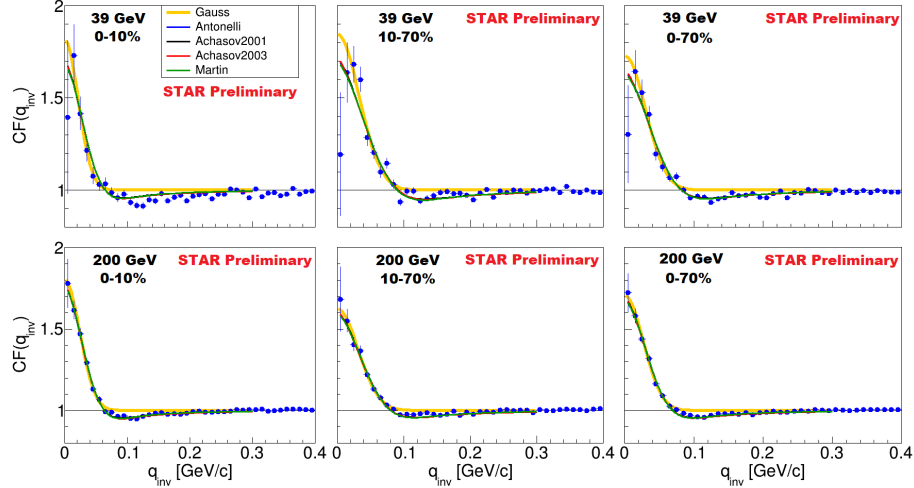


Fig. 2. The K_s^0 - K_s^0 correlation functions at $\sqrt{s_{NN}} = 39$ GeV (upper row) and 200 GeV (lower row) with three centrality classes (0-10%, 10-70% and 0-70%) as a function of q_{inv} . The statistical errors are shown as vertical lines. The colored lines represent the Gaussian fit and L-L model fit results.

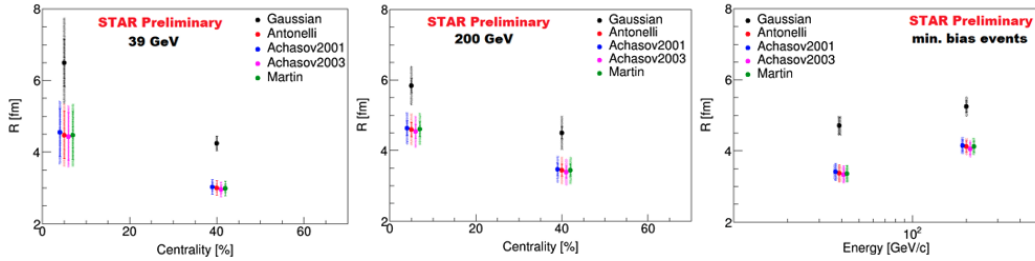


Fig. 3. Centrality and energy dependence of R extracted by fitting the Gaussian function and L-L model to K_s^0 - K_s^0 correlation function at $\sqrt{s_{NN}} = 39$ GeV and 200 GeV in Au+Au collisions. The statistics and systematics errors are shown as the vertical lines and shadowed bands, respectively.

71 at small k^* are observed in the measured correlations. Also no significant
72 centrality dependence is found.

73 To further understand light nucleus formation, a transport model, Simu-
74 lating Many Accelerated Strongly-interacting Hadrons Model (SMASH) [11],
75 is used to simulate particle production in $\sqrt{s_{NN}} = 3$ GeV Au+Au collisions.
76 We use the cascade mode in SMASH model where the particles are propa-
77 gated as in vacuum between collisions with other particles. In this work, two

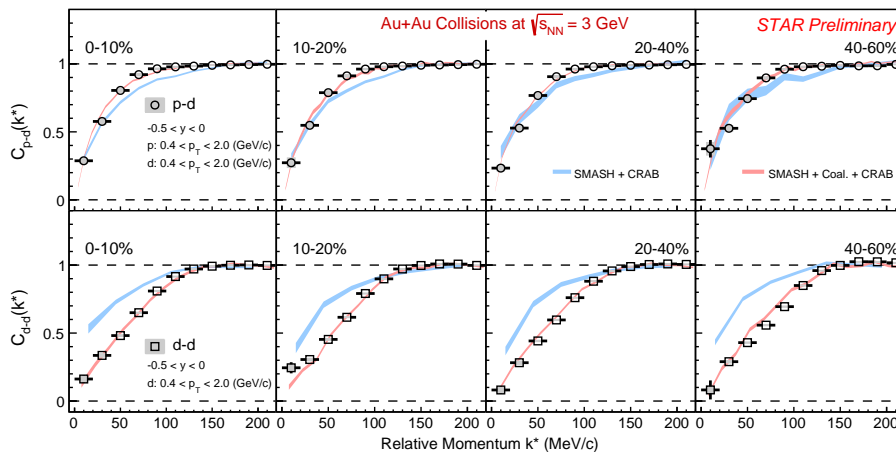


Fig. 4. The $p-d$ and $d-d$ correlation functions in different collision centralities in Au+Au collisions at $\sqrt{s_{\text{NN}}} = 3$ GeV. The statistical and systematic errors are shown as vertical lines and grey bands, respectively. The colored bands represent the $p-d$ and $d-d$ correlations obtained with the deuteron from nucleon coalescence (red) in SMASH and directly produced from SMASH via hadronic scattering (blue), respectively.

78 different versions of SMASH model are chosen to generate deuterons. In the
 79 first version, SMASH can generate deuterons via hadron scattering, such as
 80 $p + n + \pi \leftrightarrow d + \pi$. This is so-called directly produced deuterons. The
 81 second version of SMASH model produces nucleons and the deuterons
 82 are formed from nucleon coalescence with an afterburner package based on
 83 Wigner function [12]. The SMASH model itself does not contain femto-
 84 scopic correlation between two particles after kinetic freeze-out. The in-
 85 teraction (Coulomb potential, strong interaction, and quantum statistics)
 86 between two particles is introduced by Correlation Afterburner (CRAB)
 87 package [13], and the input potentials of $p-d$ and $d-d$ pairs are taken from
 88 Ref. [14].

89 The resulting $p-d$ and $d-d$ correlation functions from the models are
 90 shown as color bands in Fig. 4, respectively. It is found that the experi-
 91 mental data are well reproduced by the SMASH plus coalescence calcula-
 92 tions. On the other hand, the correlation functions with directly produced
 93 deuterons from SMASH can only qualitatively reproduce the overall trends,
 94 but over/under estimate the data depending on the particle pair. Those
 95 observations imply that at $\sqrt{s_{\text{NN}}} = 3$ GeV Au+Au collisions, deuterons are
 96 likely formed via the nucleon coalescence processes.

4. Summary

In summary, we report the femtoscopy results of protons, light nuclei, and strange hadrons in Au+Au collisions from STAR. For $K_s^0-K_s^0$ correlations, by fitted with L-L model, it is found that the FSI plays an important role and extracted radii of particle emitting source clear centrality and energy dependence. As for light nuclei correlations, the SMASH model with coalescence can well reproduce data, which suggests that coalescence mechanism may dominate the deuteron production at $\sqrt{s_{NN}} = 3$ GeV Au+Au collisions.

5. Acknowledgments

We thank Prof. S. Mrowczynski for insightful discussions about production mechanism of light nuclei. This work was supported by the National Key Research and Development Program of China (2020YFE0202002 and 2018YFE0205201), the National Natural Science Foundation of China (12122505, 11890711) and the Fundamental Research Funds for the Central Universities(CCNU220N003).

REFERENCES

- [1] M. A. Lisa, et al., *Ann. Rev. Nucl. Part. Sci.* 55 (2005) 357–402.
- [2] R. Lednicky, *Nucl. Phys. A* 774 (2006) 189–198.
- [3] S. Pratt, *Phys. Rev. D* 33 (1986) 1314–1327.
- [4] E. N. Martin, Alan D., E. J. Squires, *Nucl. Phys. B* 121 (1977) 514–530.
- [5] V. Baru, et al., *Eur. Phys. J. A* 23 (2005) 523–533.
- [6] A. Antonelli, eConf C020620 (2002) THAT06.
- [7] N. N. Achasov, V. V. Gubin, *Phys. Rev. D* 63 (2001) 094007.
- [8] S. Mrowczynski, *Eur. Phys. J. ST* 229 (22-23) (2020) 3559–3583.
- [9] J. Adam, et al., *Phys. Rev. C* 99 (2019) 064905.
- [10] S. Mrówczyński, P. Słoń, *Phys. Rev. C* 104 (2) (2021) 024909.
- [11] J. Weil, et al., *Phys. Rev. C* 94 (5) (2016) 054905.
- [12] W. Zhao, et al., *Phys. Rev. C* 98 (5) (2018) 054905.
- [13] <https://web.pa.msu.edu/people/pratts/freecodes/crab/home.html>.
- [14] B. K. Jennings, et al., *Phys. Rev. C* 33 (1986) 1303–1306.