

# Light- and Hyper-Nuclei Collectivity in Au+Au Collisions at RHIC-STAR

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**Abstract.** We report the first results on collision energy and particle mass dependence of directed flow  $v_1$  of light- and hyper-nuclei in mid-central Au+Au collisions at center of mass energies per nucleon pair of 3.0, 3.2, 3.5, and 3.9 GeV. All data have been collected by the STAR experiment in the fixed-target mode during the second phase of the RHIC beam energy scan. The mid-rapidity  $v_1$  slope,  $dv_1/dy|_{y=0}$ , of hyper-nuclei shows a similar energy and particle mass dependence to that of light-nuclei. The results suggest that the coalescence mechanism plays a dominant role in the formation of light- and hyper-nuclei.

## 1 Introduction

The two phases of the Beam Energy Scan program at RHIC, BES-I and BES-II, aim to investigate nuclear matter at different temperatures and baryon chemical potentials by colliding gold nuclei at various collision energies. The goal is to study the nature of phase transition to the deconfined phase and search for QCD critical point within the high baryon density region. In astrophysics, the hyperon puzzle in neutron star research refers to the difficulty to reconcile the measured masses of neutron stars with the presence of the hyperons in their interiors. Extracting the strength of the baryon-baryon interaction in a dense nuclear medium experimentally is essential for understanding the inner structure of compact stars.

In the past few decades, the study of baryon-baryon interaction and the properties of QCD matter using light- and hyper-nuclei production in heavy-ion collisions has been a topic of interest. Thermal model [1] and hadronic transport model with coalescence afterburner [2, 3] calculations have predicted abundant production of light- and hyper-nuclei at high baryon density region. This allows for the study of production mechanism of light- and hyper-nuclei in fixed-target  $\sqrt{s_{NN}} = 3.0, 3.2, 3.5$  and  $3.9$  GeV Au+Au collisions at the STAR experiment. Light-nuclei, in particular, carry information about local baryon density fluctuations at freeze-out, providing insights into the final state nucleon-nucleon ( $N-N$ ) interaction. Hyper-nuclei, on the other hand, offer access to the hyperon-nucleon ( $Y-N$ ) interaction. Collective flow is commonly used to study the properties of matter produced in high energy nuclear collisions, this is primarily due to its sensitivity to the early dynamics of the collision. The first harmonic coefficient of the Fourier expansion of the particle azimuthal distributions in the momentum space is called the directed flow ( $v_1$ ), which is driven by pressure gradients created in such

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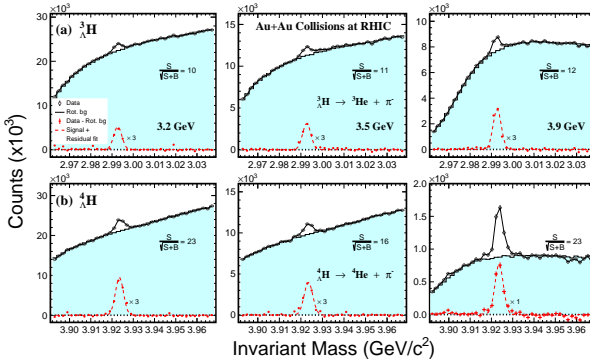
37 collisions. Therefore, measurements of light- and hyper-nuclei collectivity make it possible  
 38 to study  $N - N$  and  $Y - N$  interactions and the Equation of State at high baryon density.

## 39 2 Experiment and data analysis

40 The dataset used in this analysis were collected by the STAR experiment at RHIC with the  
 41 fixed-target setup during 2018-2020 for the center-of-mass energies per nucleon pair of  $\sqrt{s_{NN}}$   
 42 = 3.0, 3.2, 3.5 and 3.9 GeV. A detailed description of the STAR detector system can be found  
 43 in [4]. The Time Projection Chamber (TPC) [5, 6] is used for charged-particle tracking, while  
 44 a combination of TPC and Time-of-Flight (TOF) [7] is used for particle identification. The  
 45 collision centrality is determined by the charged-particle multiplicity distribution measured  
 46 in TPC within the pseudo-rapidity range  $-2.4 < \eta < 0$  combined with Monte Carlo Glauber  
 47 models [8, 9]. The event-plane is reconstructed with the Event-Plane-Detector (EPD) within  
 48 the pseudo-rapidity range  $-5.3 < \eta < -3.3$ . Because of acceptance asymmetry in fixed-target  
 49 mode collision, three sub-events method was used to determined the event plane resolution.

**Table 1.**  $p_T/A - y$  acceptance windows of light- and hyper-nuclei used for collective flow analysis, where the  $A$  is the atomic mass number.

Particle	$p_T/A$ (GeV/c)	$y$	Particle	$p_T/A$ (GeV/c)	$y$
p	(0.4, 0.8)	(-1.0, 0.0)	t	(0.4, 0.8)	(-1.0, -0.3)
$\Lambda$	(0.4, 0.8)	(-1.0, 0.0)	${}^3\text{He}$	(0.4, 0.8)	(-1.0, 0.0)
d	(0.4, 0.8)	(-1.0, -0.2)	${}^3_{\Lambda}\text{H}$	(0.33, 0.83)	(-1.0, 0.0)
			${}^4\text{He}$	(0.4, 0.8)	(-1.0, -0.4)
			${}^4_{\Lambda}\text{H}$	(0.30, 0.75)	(-1.0, 0.0)

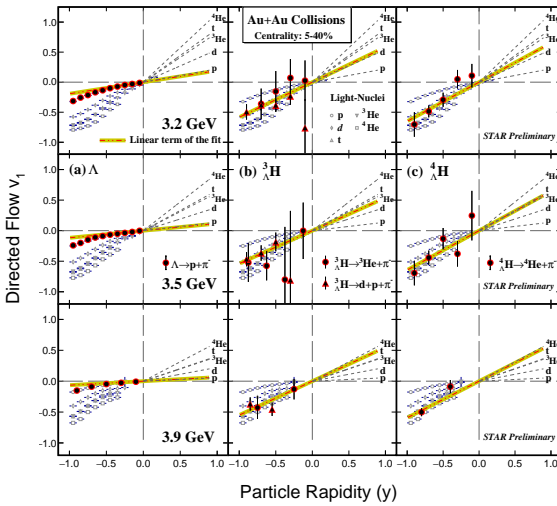


**Figure 1.** Topologically reconstructed (2-body decay)  ${}^3_{\Lambda}\text{H}$  (top panel) and  ${}^4_{\Lambda}\text{H}$  (bottom panel) from 3.2, 3.5 and 3.9 GeV Au+Au collisions. Background subtracted distributions are shown as red symbols. The significances of the mass peaks are also indicated in the figure.

50 For particle identification, the  $\pi^-$ , p, d, t,  ${}^3\text{He}$ , and  ${}^4\text{He}$  are selected based on the ioniza-  
 51 tion energy loss ( $dE/dx$ ) measured in the TPC as a function of rigidity. For the short-lived  
 52 hyper-nuclei reconstruction, the secondary decay topology is reconstructed with the KFParticle  
 53 package based on a Kalman filter method providing a full set of the particle parameters to-  
 54 gether with their uncertainties. Fig.1 shows the reconstructed invariant mass distributions for  
 55  ${}^3_{\Lambda}\text{H}$  and  ${}^4_{\Lambda}\text{H}$  2-body decay using phase space listed in Table 1 for 3.2, 3.5 and 3.9 GeV. Com-  
 56 binatorial backgrounds for hyper-nuclei are constructed by rotating decay daughter particles  
 57 through a random angle between  $10^\circ$  and  $350^\circ$ . The combinatorial background subtracted  
 58 distributions are fitted with a Gaussian plus a linear function for hyper-nuclei to extract the  
 59 signal counts. The signal significance for  ${}^3_{\Lambda}\text{H}$  2-body decay is around 10 and for  ${}^4_{\Lambda}\text{H}$  is around  
 60 20.

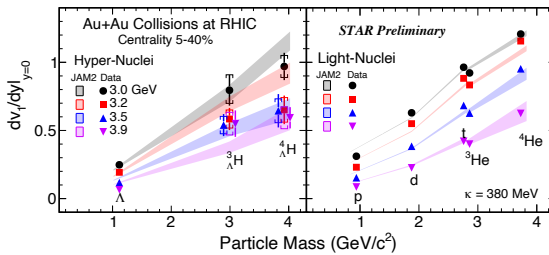
61 The  $v_1(y)$  results of  $\Lambda$  hyperon and hyper-nuclei with the event plane method [10], from  
 62 5-40% mid-central Au + Au collisions at 3.2, 3.5 and 3.9 GeV, are shown in Fig.2. For com-  
 63 parison, the  $v_1(y)$  of p, d, t,  ${}^3\text{He}$  and  ${}^4\text{He}$  from the same dataset are shown as open markers.

64 Since the collective flow depends on the  $p_T$  range of particle species, for a meaningful comparison we use same  $p_T/A$  range when comparing different species and it is listed in Table  
 65 1. For the  $\Lambda$  hyperon and light-nuclei, we use the third-order polynomial  $v_1(y) = p_0y + p_1y^3$   
 66 to fit in order to cover wider rapidity range. The linear terms for the light-nuclei are shown  
 67 by the dash line in the positive rapidity region. Due to limited statistics for  ${}^3_\Lambda\text{H}$  and  ${}^4_\Lambda\text{H}$ , we  
 68 used only the first order polynomial  $v_1(y) = p_0y$  to fit. The linear-terms for hyper-nuclei at the  
 69 mid-rapidity are shown by the red-yellow lines. In all the above cases, the fit is constrained  
 70 to pass through (0,0) taking advantage of the symmetry. Systematic uncertainties of the measured  
 71 collective flow mainly come from the event plane resolution and efficiency corrections,  
 72 identification as well as topological variable cuts. For hyper-nuclei  $v_1$  measurements, the  
 73 dominant source of systematic uncertainty is from topological cuts. For light-nuclei, the major  
 74 contributor to systematic uncertainty is arising from the particle misidentification. The  
 75 total systematic uncertainty is around 13% and 6% for hyper-nuclei and light-nuclei, respectively.  
 76  
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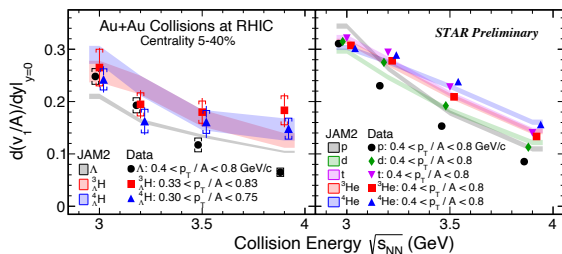
**Figure 2.** Light- and hyper-nuclei directed flow  $v_1$  shown as a function of particle rapidity from 3.2, 3.5 and 3.9 GeV mid-central (5-40%) Au + Au collisions. The rapidity dependence of  $\Lambda$   $v_1$  is shown in left column.  ${}^3_\Lambda\text{H}$ , with both 2-body and 3-body decays, and  ${}^4_\Lambda\text{H}$  rapidity dependence are shown in the middle and right column, respectively. The results of the fit are shown as the red-yellow lines. For comparison, the rapidity dependence for p, d, t,  ${}^3\text{He}$ , and  ${}^4\text{He}$  are shown as open markers and the linear terms of the fitting results for these light-nuclei are displayed as dashed lines in the positive rapidity region.

### 78 3 Results and Discussion



**Figure 3.** Particle mass dependence of the mid-rapidity  $v_1$  slope for hyper-nuclei (left plot) and light-nuclei (right plot) from mid-central 5-40% Au + Au collisions. The results from the 3 GeV data is taken from Ref.[11]. Transport model (JAM [12]) + coalescence results are shown as colored bands.

79 The results of the mid-rapidity  $v_1$  slope as a function of particle mass for hyper-nuclei and  
 80 light-nuclei with centrality 5-40% from the  $\sqrt{s_{NN}} = 3.0$  [11], 3.2, 3.5 and 3.9 GeV are shown  
 81 in Fig. 3. The left panel is for hyper-nuclei mid-rapidity  $v_1$  slope, the right panel is for the  
 82 light-nuclei mid-rapidity  $v_1$  slope. As we can see, at given energy, for both light- and hyper-  
 83 nuclei, the mid-rapidity  $v_1$  slopes are scaled with particle mass, implying that coalescence is  
 84 the dominant process for the light- and hyper-nuclei production. And the feature is also reproduced  
 85 by hadronic transport model JAM2 ( $\kappa = 380$  MeV) [12] with coalescence afterburner  
 86 calculations, shown as colored bands.



**Figure 4.** Collision energy dependence of the mid-rapidity  $v_1$  slope for hyper-nuclei (left plot) and light-nuclei (right plot) from mid-central 5-40% Au + Au collisions. The results from the 3 GeV data is taken from Ref.[11]. Transport model (JAM [12]) + coalescence results are shown as colored bands.

87 Figure 4 shows the collision energy dependence of the mass scaled mid-rapidity  $v_1$  slope  
 88 as a function of collision energy in centrality 5-40% from the  $\sqrt{s_{NN}} = 3.0$  [11], 3.2, 3.5 and  
 89 3.9 GeV, for hyper-nuclei on the left panel, light-nuclei on the right panel. From the figure,  
 90 it can be seen that as the collision energy increases, the mid-rapidity  $v_1$  slope of light- and  
 91 hyper-nuclei decreases. The hadronic transport model (JAM2) with coalescence afterburner  
 92 calculations are consistent with observed energy dependence.

## 93 4 Summary

94 In these proceedings, we report the collision energy and particle mass dependence of the  
 95 directed flow  $v_1$  for both light- and hyper-nuclei in  $\sqrt{s_{NN}} = 3.0, 3.2, 3.5$  and 3.9 GeV Au + Au  
 96 collisions measured by the STAR experiment at RHIC. An approximate atomic mass number  
 97 scaling and energy dependence are observed in the measured  $v_1$  slopes of light- and hyper-  
 98 nuclei at mid-rapidity. Calculations of hadronic transport JAM model with a coalescence  
 99 afterburner can qualitatively reproduce the observed dependences for hyper-nuclei as well  
 100 as light-nuclei collective flow measurements. It suggests that the dominant process in the  
 101 production mechanism for light clusters is the coalescence process. Moreover, results from a  
 102 high statistics 2 billion event samples for the 3.0 GeV Au + Au collisions will significantly  
 103 enhance the precision and help constrain coalescence parameters, as well as constrain  $N - N$   
 104 and  $Y - N$  interactions for both light- and hyper-nuclei at the high density region.

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## 109 References

- 110 [1] A. Andronic, P. Braun-Munzinger, J. Stachel, H. Stocker, Phys. Lett. B **697**, 203 (2011)  
 111 [2] J. Steinheimer et al., Phys. Lett. B **714**, 85 (2012)  
 112 [3] J. Aichelin et al., Phys. Rev. C **101**, 044905 (2020)  
 113 [4] K.H. Ackermann et al. (STAR), Nucl. Instrum. Meth. A **499**, 624 (2003)  
 114 [5] M. Anderson et al., Nucl. Instrum. Meth. A **499**, 659 (2003)  
 115 [6] F. Shen et al., Nucl. Instrum. Meth. A **896**, 90 (2018)  
 116 [7] W.J. Llope (STAR), Nucl. Instrum. Meth. A **661**, S110 (2012)  
 117 [8] M.L. Miller et al., Ann. Rev. Nucl. Part. Sci. **57**, 205 (2007)  
 118 [9] R.L. Ray, M. Daugherty, J. Phys. G **35**, 125106 (2008)  
 119 [10] H. Masui, A. Schmah, A.M. Poskanzer, Nucl. Instrum. Meth. A **833**, 181 (2016)  
 120 [11] B. Aboona et al. (STAR), Phys. Rev. Lett. **130**, 212301 (2023)  
 121 [12] Y. Nara, A. Jinno, K. Murase, A. Ohnishi, Phys. Rev. C **106**, 044902 (2022)