Strangeness production in Au+Au collisions at $\sqrt{s_{NN}} = 7.7$ -19.6 GeV with the STAR experiment

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One of main motivations of the Beam Energy Scan (BES) program at RHIC is to search 4 for the QCD critical point and the onset of deconfinement. Strangeness production has been 5 suggested as a sensitive probe to the early dynamics of the deconfined matter created in 6 heavy-ion collisions. The rapidity density of (anti-)strange baryons may give insight into 7 the baryon stopping mechanism. Ratios of particle yields involving strange particles are 8 often utilized to understand the hadronization mechanism. Furthermore, the collision energy g threshold of QGP production in heavy-ion collisions can be explored by measuring the colliding 10 energy dependence of baryon-to-meson enhancement. In particular, results from BES-I indicate 11 potential changes in the Ω/ϕ ratio at and below 11.5 GeV, which may be indicative of a change 12 in the underlying strange quark dynamics. However, the limited statistics prohibit a strong 13 conclusion. 14

During BES-II, STAR has accumulated high statistics data in Au+Au collisions, which 15 can help reduce the uncertainties in the strange hadron measurements, particularly for the 16 multi-strange hadrons. Benefiting from the iTPC upgrade, the strangeness measurements are 17 now extended from the previous rapidity window of |y| < 0.5 to |y| < 1.5. In this talk, we will 18 present new STAR measurements of strange hadron $(K_s^0, \Lambda, \bar{\Lambda}, \Xi, \bar{\Xi}, \Omega, \bar{\Omega}, \phi)$ production in 19 Au+Au collisions at $\sqrt{s_{NN}} = 7.7, 9.2, 11.5, 14.6, 17.3$ and 19.6 GeV from BES-II. We will report 20 transverse-momentum and rapidity spectra, nuclear modification factors, antibaryon-to-baryon 21 ratios and baryon-to-meson ratios. In particular, precise measurements of the energy and 22 centrality dependence of Ω/ϕ ratios down to $\sqrt{s_{NN}} = 7.7$ GeV will be presented for the first 23 time. These results will be compared to theoretical calculations and physics implications will 24 be discussed. 25