



Probing gluon structure with J/ψ photoproduction in isobaric ultra-peripheral collisions at $\sqrt{s_{NN}} = 200$ GeV with the STAR

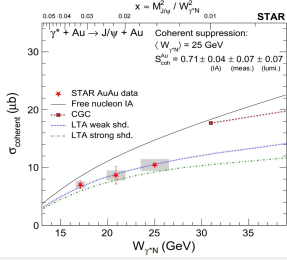
Zengzhi Li (zengzhi-li@m.scnu.edu.cn), for the STAR Collaboration
South China Normal University



Abstract

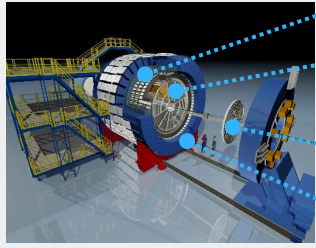
In ultra-peripheral collisions (UPCs), coherent J/ψ photoproduction has been recognized as one of the most sensitive probes of the nuclear gluon distribution. The collision system size in isobaric collisions ($^{96}\text{Ru} + ^{96}\text{Ru}$ and $^{96}\text{Zr} + ^{96}\text{Zr}$) lies between d+Au and Au+Au collisions. Therefore, the measurement of coherent J/ψ photoproduction in isobaric UPCs offers a unique opportunity to study the system size dependence of gluon structure. In this poster, we present the differential cross sections of photoproduced coherent J/ψ as a function of rapidity (y) in isobaric UPCs at $\sqrt{s_{NN}} = 200$ GeV. These data provide crucial constraints on the system size dependence of the gluon structure function within nuclei in the kinematic range x_{parton} , the momentum fraction carried by the gluon, $\sim 0.015-0.03$. The results are compared with STARlight and previous STAR measurements. Physics implications are also discussed.

1. Motivation



- The measurement of the coherent J/ψ cross section in Au+Au collisions provides insight into the gluon structure of heavy nuclei and shows good agreement with the shadowing model[1][2].
- A direct probe of the system-size dependence of the gluon structure can be made using Ru+Ru and Zr+Zr collisions.

2. STAR Detector



- Barrel ElectroMagnetic Calorimeter
 - Trigger and identify electron/positron
- Time Projection Chamber
 - Tracking
 - Particle identification (dE/dx)
- Beam-Beam Counters
 - Trigger UPC events
- Time Of Flight Detector
 - Time of flight
 - Particle identification

3. Dataset and Analysis Details

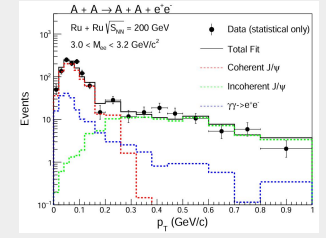
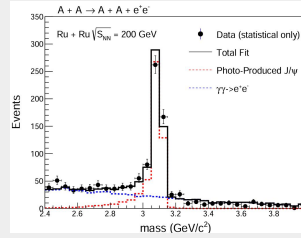
- Dataset: Run18 Ru+Ru and Zr+Zr collisions at $\sqrt{s_{NN}} = 200$ GeV
- Luminosity: 2.6 nb^{-1} for each Ru+Ru and Zr+Zr
- Trigger Requirements for UPC Coherent Events:
 - $2 \leq \text{TOF multiplicity} \leq 6$
 - EMC have back to back high tower hit
 - No signal in BBCs from both sides
- Event Selection:
 - $|V_z| < 100$ cm with UPC trigger requirement
 - BEMC clusters ≤ 6 with UPC trigger requirement
- Track Selection:
 - TPC hits used to reconstruct tracks ≥ 15
 - TPC hits used to calculate $dE/dx \geq 11$
 - $|\eta_e| < 1$
- Electron Identification:
 - $\chi^2_{ee} = n\sigma_{e1}^2 + n\sigma_{e2}^2 < 10$ & $\chi^2_{ee} < \chi^2_{\pi\pi}, n\sigma_e = \frac{1}{R_{dE/dx}} \log \frac{(dE/dx)^{Mea.}}{(dE/dx)^{Th.}}$

7. References

- [1] M. I. Abdulhamid et al. (STAR), Phys. Rev. Lett. 133, 052301 (2024)
- [2] M. I. Abdulhamid et al. (STAR), Phys. Rev. C. 110, 014911 (2024)
- [3] P. Wang et al, Chinese Phys. C. 46074103 (2022)
- [4] W. Zha et al, Phys. Lett. B. 800, 135089 (2020)

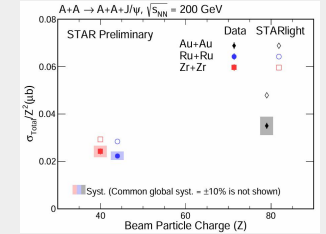
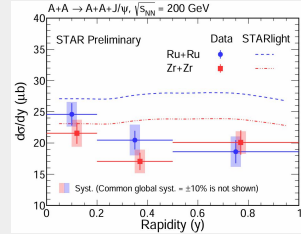
4. Raw Signal

- $J/\psi \rightarrow e^+ + e^-$, Br = 5.92 %.



- The uncorrected p_T and mass spectra were reconstructed from the data.
- A combined fit, incorporating coherent J/ψ , incoherent J/ψ (EPA-VMD) and $\gamma\gamma \rightarrow e^+e^-$ (QED calculation) contributions, was applied to the mass and p_T spectra to extract the coherent J/ψ yields[3][4]. The same fitting method is applied separately to three rapidity bins: $[0, 0.2]$, $[0.2, 0.5]$, and $[0.5, 1.0]$.

5. Physics Results



- Left: coherent J/ψ cross section as function of rapidity, fully corrected for efficiency, in ultra-peripheral Ru+Ru and Zr+Zr collisions.
 - The measured data is lower than the STARlight prediction.
- Right: the J/ψ cross section as a function of Z , scaled by $1/Z^2$, for data and STARlight simulation.
 - Photon flux dependence is roughly cancelled by dividing by Z^2 .
 - The measured data is lower than STARlight predictions, indicating shadowing effects in coherent photoproduction of J/ψ in Ru, Zr, and Au.
 - There is no significant system size dependence.

6. Summary

- The coherent J/ψ cross section is lower than predicted by STARlight, but shows a similar trend with rapidity in isobaric UPCs.
- The J/ψ cross section scaled by $1/Z^2$ also shows the data is lower than STARlight, indicating shadowing effects, but there is no significant system size dependence.

Supported in part by the



Office of
Science



华南师范大学
SOUTH CHINA NORMAL UNIVERSITY



The STAR
Collaboration

