

1 Thermal dielectron measurements in Au+Au collisions at 2 $\sqrt{s_{NN}} = 7.7, 14.6, \text{ and } 19.6 \text{ GeV}$ with the STAR experiment

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5 **Abstract.** Dielectrons emitted during the evolution of the hot and dense QCD
6 medium created in relativistic heavy-ion collisions offer an effective way to
7 probe the medium properties, as they do not interact via the strong force. The
8 rate of the dielectron emission is proportional to the medium's electromagnetic
9 spectral function. In the dielectron invariant mass range from $400 \text{ MeV}/c^2$ to
10 $800 \text{ MeV}/c^2$, the spectral function probes the in-medium ρ meson propagator
11 which is sensitive to the medium's properties including the total baryon density
12 and the temperature. Meanwhile, the low energy range of the spectral function
13 provides information about the medium's electrical conductivity. Therefore, by
14 measuring thermal dielectron production, we can study the microscopic inter-
15 actions between the electromagnetic current and the medium.

16 The STAR experiment has recorded large datasets of Au+Au collisions dur-
17 ing the Beam Energy Scan Phase-II (BES-II) program, spanning center-of-mass
18 energies ($\sqrt{s_{NN}}$) from 3.0 to 19.6 GeV with detector upgrades that benefit the
19 dielectron measurement via extended transverse momentum and rapidity cover-
20 ages as well as enhanced particle identification capability. In these proceedings,
21 we will report on the measurements of thermal dielectrons produced in Au+Au
22 collisions at $\sqrt{s_{NN}} = 7.7, 14.6, \text{ and } 19.6 \text{ GeV}$ using the STAR experiment.

23 1 Introduction

24 Heavy-ion collisions emit thermal radiation at each stage, encompassing both the Quark-
25 Gluon Plasma (QGP) phase and the hadronic phase. This radiation allows us to investigate
26 various properties of the QCD medium formed during these collisions. Key among these are
27 the interactions between electromagnetic (EM) currents and the medium, as well as the evo-
28 lution of the fireball. Experimentally, this thermal radiation can be detected through dilepton
29 production.

30 The rate of dilepton production in thermal radiation is determined by a confluence of
31 factors: the lepton phase-space factor, the thermal Bose-Einstein distribution, and the elec-
32 tromagnetic spectral function [1, 2]. Model calculations show that for dileptons in the in-
33 termediate mass range (with $M_\phi < M_{ll} < M_{J/\psi}$), the EM current predominantly exhibits
34 partonic behavior [3]. Conversely, in the low mass range (with $M_{ll} < 1 \text{ GeV}/c^2$), the current
35 is dominated by vector mesons [9]. Essentially, the EM spectral function is a blend of the
36 imaginary components of vector meson propagators, predominantly influenced by the isospin
37 effect, manifesting as ρ -dominance [4].

38 To examine how the medium temperature and total baryon density impact the EM spectral
39 function, the STAR experiment is conducting thermal dielectron measurements. These are

40 being carried out at lower collision energies (between $\sqrt{s_{NN}} = 19.6$ and 3 GeV) from the
 41 Beam Energy Scan Phase-II (BES-II). This phase features advanced detector capabilities and
 42 enhanced statistics, facilitating more accurate observations.

43 2 Experiment and analysis

44 The data presented in these proceedings were obtained using the STAR detector at the RHIC
 45 during the years 2019 and 2021. These data are taken from Au+Au collisions at collision
 46 energies of $\sqrt{s_{NN}} = 19.6, 14.6,$ and 7.7 GeV. The number of events, meeting stringent quality
 47 criteria, amounted to 200 million at 19.6 GeV, 120 million at 14.6 GeV, and 37 million at
 48 7.7 GeV. The 2019 19.6 GeV data has an approximate factor of 10 increase in sample size
 49 compared to the data collected during the Beam Energy Scan Phase-I (BES-I) in 2010-2014.

50 Electron identification within the STAR framework primarily utilizes the combination
 51 of two detector subsystems: the Time Projection Chamber (TPC) which is instrumental in
 52 determining track momentum and ionization energy loss (dE/dx), and the Time of Flight
 53 (TOF) system which is vital for measuring the velocity of charged particles.

54 Dielectron signal processing begins with the construction of a e^+e^- continuum, formed
 55 by pairing all possible combinations of electrons and positrons from the same event. We
 56 estimate the combinatorial and correlated background by employing the acceptance-corrected
 57 geometric mean of like-sign pairs. The acceptance discrepancy between unlike- and like-sign
 58 pairs is estimated using the mixed-event technique.

59 Subtracting the background spectrum from the dielectron e^+e^- continuum yields the
 60 raw signal. The raw spectrum is then adjusted for detector inefficiencies. A comprehen-
 61 sive exposition of the dielectron signal reconstruction process is detailed in [5]. The phys-
 62 ical background (cocktail including hadronic decays and Drell-Yan process) for the ther-
 63 mal radiation dielectron production includes the following channels: $DY \rightarrow e^-e^+, D\bar{D} \rightarrow$
 64 $e^-e^+ \nu_e \bar{\nu}_e X, J/\psi \rightarrow e^-e^+, \pi^0(\eta, \eta') \rightarrow \gamma e^-e^+, \omega \rightarrow \pi^0 e^-e^+, \omega \rightarrow e^-e^+, \phi \rightarrow \eta e^-e^+, \phi \rightarrow e^-e^+.$

65 The efficiency correction for the efficiency-correct spectrum encompasses both single-
 66 track electron reconstruction efficiency and the dielectron loss at $M_{ee} < 0.2$ GeV/ c^2 due to
 67 photon conversion removal. The methodology for efficiency calculation is the same as the
 68 approach described in [6].

69 3 Results and discussion

70 The efficiency-corrected invariant mass spectra, presented in Fig. 1, correspond to the 0-80%
 71 most-central Au+Au collisions at $\sqrt{s_{NN}} = 19.4, 14.6,$ and 7.7 GeV. In Fig. 1, the blue points
 72 represent the invariant mass spectra within the STAR acceptance at mid-rapidity ($|\eta^e| < 1,$
 73 $p_T^e > 0.2$ GeV/ $c, |y^{ee}| < 1$), where the vertical blue lines indicate statistical uncertainty, and
 74 the blue boxes denote the systematic uncertainty for each data point. The experimental data
 75 are juxtaposed with the physical background, depicted as a solid black line with an accom-
 76 panying grey band representing uncertainty, following the same methodology as described in
 77 [5–8].

78 Figure 2 displays the detector acceptance-corrected excess yield, calculated as, Excess
 79 yield = (Eff corrected data - physical background)/Acc, Acc is the detector-acceptance cor-
 80 rection. The acceptance correction employs a virtual photon method described in [7]. Panel
 81 (a) of Fig. 2 contrasts the excess yield invariant mass spectra from RHIC BES-I (triangles)
 82 and BES-II (circles) at $\sqrt{s_{NN}} = 19.6$ GeV, affirming consistency between the two measure-
 83 ments. Additionally, due to enhanced statistics and detector upgrades in BES-II, the total
 84 error is reduced by a factor of 4 compared to BES-I. Panel (b) of Fig. 2 depicts the excess

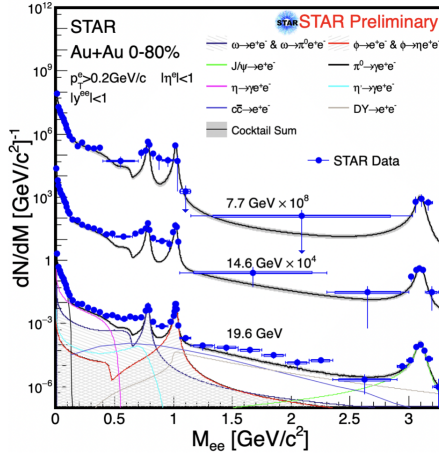


Figure 1. Efficiency-corrected dielectron invariant mass spectrum within the STAR acceptance for Au+Au at $\sqrt{s_{NN}} = 19.6, 14.6,$ and 7.7 GeV. Experimental results with statistical and systematic uncertainties are shown as points. The physical background is illustrated with solid lines.

85 yield invariant mass spectra from BES-II across various collision energies (19.6, 14.6, and
 86 7.7 GeV, from top to bottom), with the 19.6 GeV excess yield closely aligning with the Rapp
 87 model's predictions [9] for the same system and collision energy. The decreasing yield with
 88 lower collision energies at low mass range underscores medium effects on the EM spectral
 89 function, including increased total baryon density and reduced average medium temperature
 90 at lower collision energies.

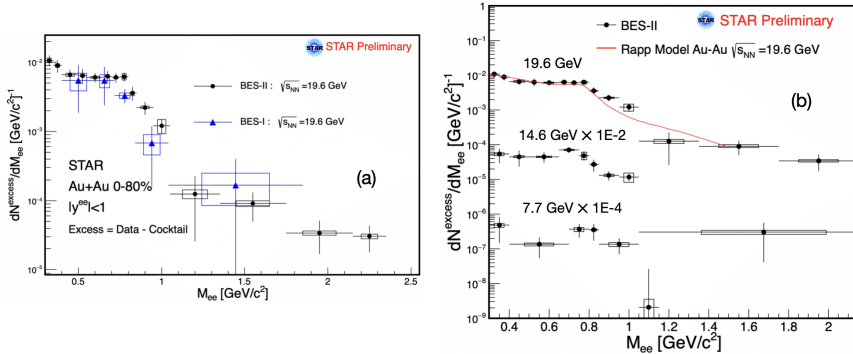


Figure 2. (a) Acceptance-corrected excess yield invariant mass spectra for Au+Au $\sqrt{s_{NN}} = 19.6$ GeV, with statistical and systematic uncertainties. For comparison, BES-I data are shown as triangles [7], and BES-II data as circles. (b) BES-II acceptance-corrected excess yield invariant mass spectra at Au+Au $\sqrt{s_{NN}} = 19.6, 14.6,$ and 7.7 GeV (points), compared with the Rapp model at Au+Au $\sqrt{s_{NN}} = 19.6$ GeV (solid line).

91 Figure 3 features the integrated excess yield, for $0.4 < M_{ll} < 0.75$ GeV/c², normalized
 92 by the π^0 yield (estimated as $N_{\pi^0} = \frac{N_{\pi^+} + N_{\pi^-}}{2}$). This figure combines results from multiple ex-
 93 periments: STAR BES-II (red stars) and BES-I (blue stars) [7, 8] in 0-80% centrality Au+Au
 94 collisions, alongside data from NA60 (black crosses) [10] and HADES (black squares) [11]

95 experiments. The experimental results are compared to the Rapp model's predictions (pink
 96 open circles). STAR BES-II's new findings bridge the gap between the BES-I and HADES
 97 results, suggesting a downward trend in the normalized integrated excess yield from high to
 98 low collision energies, which may provide constraints for models describing medium inter-
 99 actions.

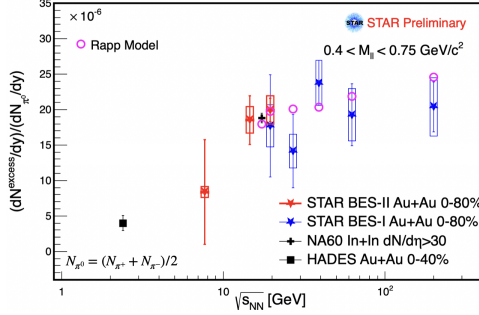


Figure 3. Integrated excess yield from different experiments for $0.4 < M_{II} < 0.75 \text{ GeV}/c^2$, normalized by the π^0 yield. The open circles represent the Rapp model's calculations.

100 4 Summary

101 In this study, we presented thermal dielectron invariant mass spectra obtained from STAR
 102 BES Phase-II Au+Au collisions at $\sqrt{s_{NN}} = 7.7, 14.6, \text{ and } 19.6 \text{ GeV}$. The data from the 19.6
 103 GeV collisions, benefiting from high statistical quality, align well with predictions based on
 104 the hadronic many-body approach for the electromagnetic spectral function. The varying
 105 invariant mass spectra observed at 14.6 and 7.7 GeV illuminate the influence of medium
 106 conditions, particularly the total baryon density and medium temperature, on the EM spectral
 107 function.

108 In addition, STAR's new observations of the integrated excess yield hint a decreasing
 109 trend from higher to lower $\sqrt{s_{NN}}$ values. This trend not only provides new insights into the
 110 dynamics of the medium created in heavy-ion collisions but also offers potential constraints
 111 for theoretical models aimed at describing the complex interactions within this medium.

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