



# **Precision Measurements of Kinematic Scan** for Fluctuations of (Net-)proton Multiplicity Distributions in Au+Au Collisions from RHIC-STAR

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### Outline

### 1. Motivation

### 2.STAR Experiment

### **3. Selected Results**

- 1) Net-proton Number Distributions
- 2) Rapidity /  $p_T$  Scan of (Net-)Proton (Factorial) Cumulants
- 3) Energy Dependence and Significance
- 4) Finite-Size Scaling Study with  $C_2$  and  $C_4$

### 4. Summary

utions Proton (Factorial) Cumulants gnificance ith  $C_2$  and  $C_4$ 



### Motivation



function of  $\mu_R$  and T by varying the collision energy

[1]X.Luo, et al.: Particles 3(2020)2,278-307



- Model predictions: First-order phase transition at high  $\mu_B$  and a critical end point
- Needs confirmation by experimental data!

# Beam energy scan program in STAR maps the proton high moments as a









## Motivation

N: Event-by-event multiplicity  $\delta N = N - \langle N \rangle$ 

### Cumulants

$$\Box C_{1} = \langle N \rangle$$
  
$$\Box C_{2} = \langle \delta N^{2} \rangle$$
  
$$\Box C_{3} = \langle \delta N^{3} \rangle$$
  
$$\Box C_{4} = \langle \delta N^{4} \rangle - 3 \langle \delta N^{2} \rangle^{2}$$

1) Related to the correlation length  $\xi^{[1]}$  $C_2 \sim \xi^2$  and  $C_4 \sim \xi^7$  $\xi$  diverges at the critical point Higher order  $\rightarrow$  More sensitive!

### **Factorial Cumulants**

$$\Box \kappa_1 = C_1$$
$$\Box \kappa_2 = -C_1 + C_2$$
$$\Box \kappa_2 = 2C_1 - 3C_2 + 3C_2$$

- $\Box \kappa_3 = 2C_1 3C_2 + C_3$  $\Box \kappa_4 = -6C_1 + 11C_2 6C_3 + C_4$
- 1) ordinary cumulant;
- 2) multi-particle correlations.

# the search for QCD critical end point

[1]M.A.Stephanov: Phys.Rev.Lett. 107(2011),052301 [2]R.V.Ravai and S. Gupta: Phys.Lett.B 696(2011),459-463 [3]S.Ejiri, F.Karsch, K.Redlich: Phys.Lett.B 633(2006),275–282 [4]A.Bazavov, et al.: Phys.Rev.Lett. 109(2012),192302 [5]A.Borsanyi, et al.: Phys.Rev.Lett. 111(2013),062005

2) Related to the susceptibility  $\chi^q$ 

Directly comparable to the model<sup>[2-5]</sup> calculations!

Can be expressed by linear combinations of

Directly capture and are sensitive to genuine

 $C_4/C_2 (= \kappa \sigma^2)$ baseline

Non-monotonic energy dependence of  $C_4/C_2$  for the conserved baryon number (using protons as a proxy) indicates the existence of a critical region.<sup>[1]</sup>

Higher-order cumulants of conserved charges serve as an important probe in







[1]STAR: arXiv:2504.00817[nucl-ex] [2]B.Ling and M.A.Stephanov: Phys.Rev.C 93(2016)3,034915

### EEMC Magnet MTD BEMC



# eTOF TOF TPC iTPC E

# STAR Detector System

IIT

-

![](_page_5_Picture_5.jpeg)

### **Data Sets**

Energy (GeV)		7.7	9.2	11.5	14.6	17.3
Vz  Cut (< cm)	y  < 0.5	50	50	50	50	50
	y  < 0.6	20	30	30	40	40
Number of Events (M)	y  < 0.5	45	78	116	178	116
	y  < 0.6	17	42	61	133	94
	y  < 0.5 (BES-I)	3	-	6.6	20	-

- purity larger than 95%;
- and detector acceptance.

![](_page_6_Figure_4.jpeg)

1. With appropriate cut selection and PID method, we ensure the proton

2. The selection of vertex-Z ( $V_z$ ) is constrained by the requirements for purity

![](_page_6_Picture_8.jpeg)

# **Net-proton Number Distributions**

![](_page_7_Figure_1.jpeg)

### **Efficiency Uncorrected Net-proton Number Distributions**

- 2.

![](_page_7_Picture_5.jpeg)

**Net-proton Cumulant Ratios:** Rapidity Scan

![](_page_8_Figure_1.jpeg)

- Cumulant ratios decrease smoothly along rapidity window;
- 2. UrQMD<sup>[1]</sup> describes the trend but fails to quantitatively reproduce the measurement, especially at high collision energy and within wide rapidity range.

[1]S.A.Bass, et al.: Prog.Part.Nucl.Phys. 41(1998),255-369

![](_page_8_Picture_7.jpeg)

# **Net-proton Cumulant Ratios:** $p_{T}$ Scan

![](_page_9_Figure_1.jpeg)

- 2. UrQMD deviates from data in high energy and wide  $p_{\rm T}$  region.

1. Cumulant ratios decrease smoothly along  $p_{\rm T}$  window, and saturate at around 1.8 GeV/c;

![](_page_9_Picture_6.jpeg)

Energy Dependence: Net-proton  $C_4/C_2$ 

![](_page_10_Figure_1.jpeg)

Skellam baseline.

The wider the y (or  $p_T$ ) window size, the farther net-proton  $C_4/C_2$  deviates from

![](_page_10_Picture_6.jpeg)

![](_page_11_Figure_1.jpeg)

![](_page_11_Picture_5.jpeg)

![](_page_12_Figure_1.jpeg)

[1]STAR: arXiv:2504.00817[nucl-ex] [2]UrQMD: S.A.Bass, et al.: Prog.Part.Nucl.Phys. 41(1998),255-369 [3]Hydro. EV: Phys.Rev.C 105(2022)1,014904

![](_page_12_Picture_6.jpeg)

![](_page_13_Figure_1.jpeg)

[1]STAR: arXiv:2504.00817[nucl-ex] [2]UrQMD: S.A.Bass, et al.: Prog.Part.Nucl.Phys. 41(1998),255-369 [3]Hydro. EV: Phys.Rev.C 105(2022)1,014904

![](_page_13_Picture_6.jpeg)

![](_page_14_Figure_1.jpeg)

[1]STAR: arXiv:2504.00817[nucl-ex] [2]UrQMD: S.A.Bass, et al.: Prog.Part.Nucl.Phys. 41(1998),255-369 [3]Hydro. EV: Phys.Rev.C 105(2022)1,014904

![](_page_14_Picture_6.jpeg)

### **Proton Factorial Cumulant Ratios:** Rapidity Scan

![](_page_15_Figure_1.jpeg)

• Smaller exponents than expected power-law  $\kappa_n/\kappa_1 \sim (\Delta y)^{n-1}$  are observed.

Near the critical region, factorial cumulants' dependence on  $\Delta y (= 2 \times y^{\max})$  is simpler and are suggested to study<sup>[1]</sup>

Deep red solid curve (—): Fitting to  $y = Ax^{\gamma}$ 

Light gray dashed curve (--): Fitting to  $y = Ax^{n-1}$ 

![](_page_15_Picture_8.jpeg)

![](_page_15_Figure_9.jpeg)

![](_page_15_Figure_10.jpeg)

![](_page_15_Figure_11.jpeg)

![](_page_15_Figure_12.jpeg)

<sup>[1]</sup>B.Ling and M.A.Stephanov: Phys.Rev.C 93(2016)3,034915

# **Proton Factorial Cumulant Ratios:** $p_{\rm T}$ Scan

![](_page_16_Figure_1.jpeg)

1.  $\kappa_2/\kappa_1$  is negative,  $\kappa_3/\kappa_1$  is positive, and their amplitude increases with increasing window size; 2.  $\kappa_4/\kappa_1$  is close to zero and doesn't show significant  $\Delta p_T$  dependence; 3. UrQMD can't quantitatively describe STAR data.

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![](_page_16_Picture_4.jpeg)

### Finite-Size Scaling Study

![](_page_17_Figure_1.jpeg)

•  $\mu_{Bc} = 625 \pm 60$  MeV in Ref<sup>[1]</sup> (w/ BES-I data)

### 2. Consistency observed from the overlap region of $U_4(\mu_R)$ .

[1]A.Sorensen and P.Sorensen: arXiv:2405.10278[nucl-th] [2]A.Andronic, et al.: Nature 561(2018)7723,321-330 [3]J.V.Sengers and J.G.Shanks: Journal of Statistical Physics 137,857(2009) [4]STAR: Phys.Rev.C 107(2023)2,024908 [5]HADES: Phys.Rev.C 102(2020)2,024914

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2) Binder cumulant:  $U_4 = -3C_4/C_2^2$ 

- 3) Rapidity window size: W
- 4) Freeze out parameters<sup>[1][2]</sup>:  $T, \mu, dV/dy$
- 5) Critical exponents<sup>[3]</sup>:  $\gamma$ ,  $\nu$

6) Uncertainty:  $\sigma = \sqrt{\sigma_{\text{stat.}}^2 + \sigma_{\text{sys.}}^2}$ 

![](_page_17_Picture_14.jpeg)

![](_page_17_Figure_15.jpeg)

![](_page_17_Figure_16.jpeg)

## Summary

- (factorial) cumulants and their ratios from STAR BES-II;
- 2. The significance of net-proton  $C_4/C_2$  shows the largest negative
- $\kappa_n/\kappa_1 \sim (\Delta y)^{n-1}$  (up to 3<sup>rd</sup> order);

1. We report the measurements of kinematic range scan of (net-)proton

deviation at  $\sqrt{s_{NN}} = 19.6$  GeV, which is consistent with reported results;

3. Smaller exponents are extracted compared to the critical inspired

4. FSS and Binder cumulant study leads to an interesting region of  $\mu_R \sim 550$  to 650 MeV, which is consistent with Sorensens' work.

![](_page_18_Picture_10.jpeg)

### Acknowledgment

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# Thank You!

![](_page_19_Picture_6.jpeg)