

Strange Hadron Production in 0+0 Collisions at $\sqrt{s_{NN}} = 200$ GeV at STAR



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1) Introduction

- At high temperatures QCD matter becomes a new state of matter called the Quark-Gluon plasma (QGP). The QGP behaves as deconfined strongly coupled fluid.
- •Its existence was predicted in 1975 and experimentally discovered in the early 2000s.
- The QGP is predicted to have existed in the early universe in the first μ s after the Big Bang.

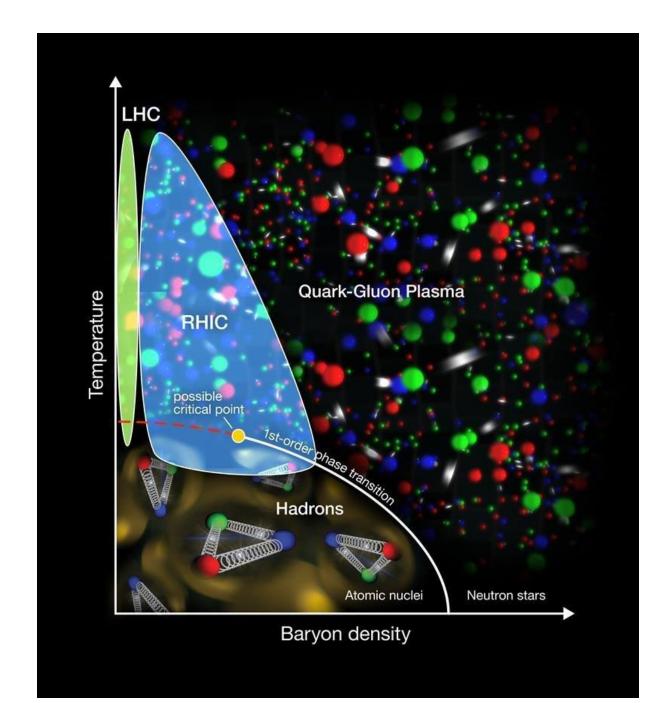
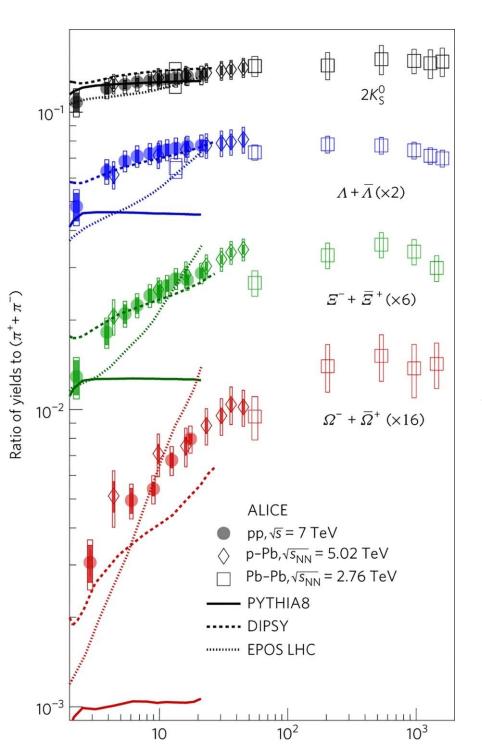


Figure 1: The QGP phase diagram [1]

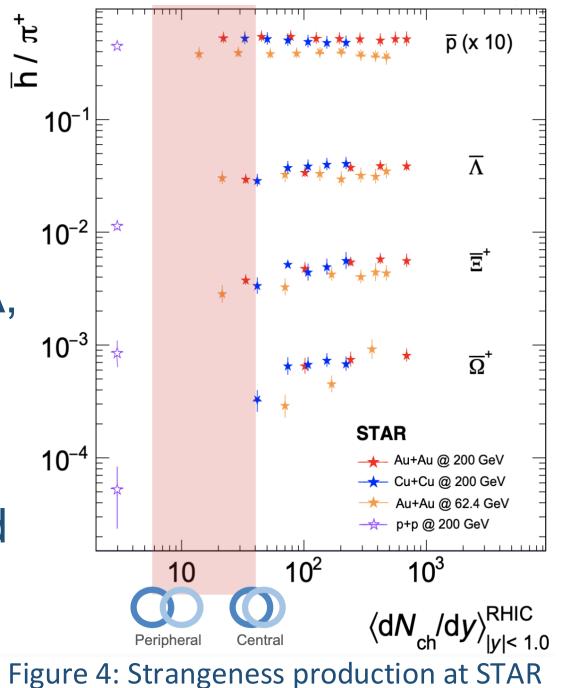
- Figure 2: Lowest Order Feynman diagram for ss production [2]
- Strangeness enhancement was one of the first observables predicted as a signature of the QGP^[2]
- •The thermal production of s-s quark pairs is favorable in the QGP since the s-s masses are close to the QGP transition temperature ~157 MeV.
 - 2 x m_s ~192 MeV
 - There are abundant thermal gluons in the QGP medium.

2) Motivation



 A smooth increase in the ratio of strange hadron production to the pion yield as a function of multiplicity has been found in various collision systems (p+p, p+A, A+A) at TeV collision energies.

 STAR potentially observes a similar trend at $\sqrt{s_{NN}}$ = 200 GeV, but needs more data at low multiplicity.



with the O+O multiplicity are to be filled-in highlighted in red

Au+Au

Figure 5: Different collision systems at RHIC

Figure 3: Strangeness production at ALICE^[3] 2a) STAR Detector

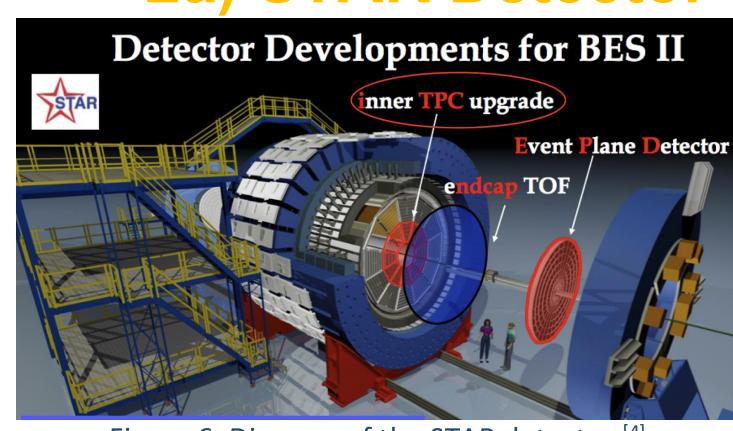
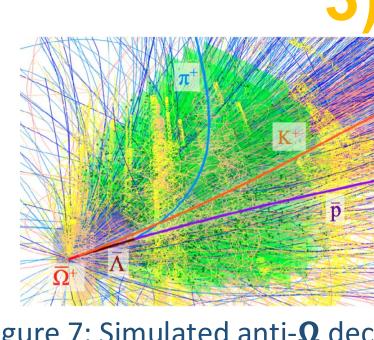


Figure 6: Diagram of the STAR detector [4]

- •From 2018 on, STAR had two detector upgrades: iTPC and eTOF
 - Improved coverage: From $|\eta| < 1.0 \longrightarrow |\eta| < 1.5$
 - Lower p_⊤ coverage: 125 MeV/c → 60 MeV/c
 - Extended PID with eTOF
- •There are ~650M O+O minimum bias events total at $\sqrt{s_{NN}} = 200 \text{ GeV}.$
- ½ of the O+O run was taken with the magnetic field reversed.
- Testing calibration and TPC distortions

3) Particle Reconstruction



 Using Kalman Filter Particle (KF Particle) reconstruction algorithm.

Figure 7: Simulated anti- Ω decay^[5]

- The signal (without background) subtraction) region is $[\mu-3\sigma,\mu+3\sigma]$, and the background region is $[1.095 \text{ to } \mu\text{-}3\sigma,$ $\mu + 3\sigma$ to 1.135 GeV/c²].
- •Fitting function: 2nd poly (for background) + double Gauss function (signal).

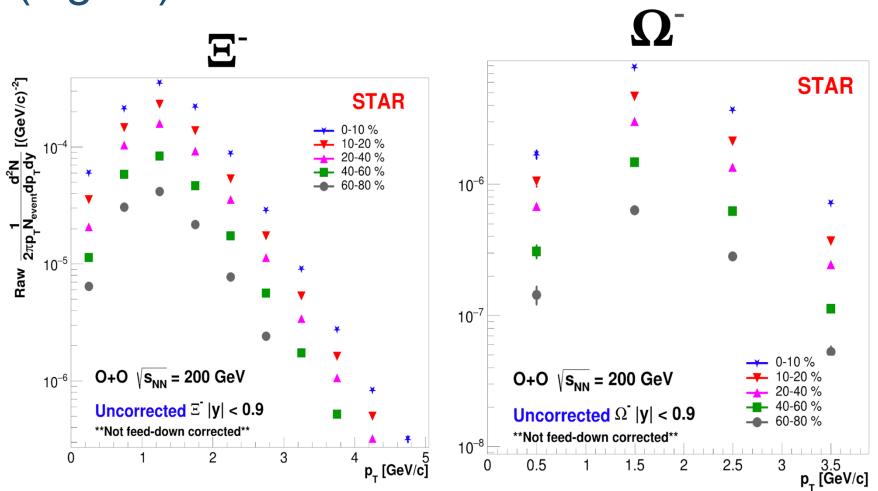


Figure 9: Raw Transverse Momentum Distributions for O+O at $\sqrt{s_{NN}}$ = 200 GeV

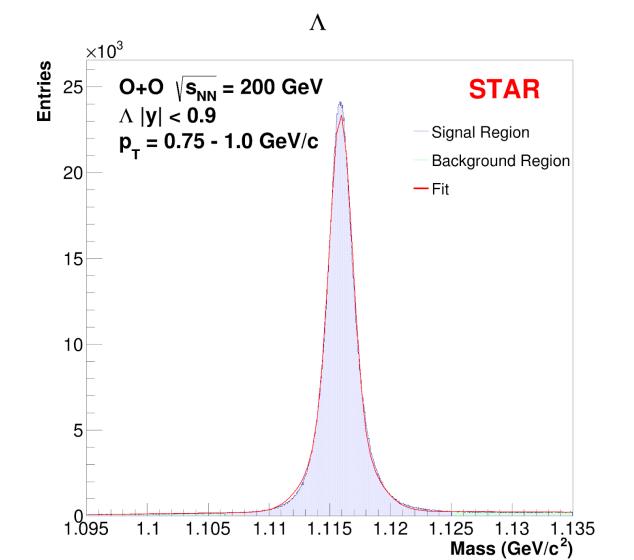


Figure 8: Invariant λ mass peak The blue region is the signal w.o

background subtraction. The green region is the background region (very small).

There is good coverage through 0 - 80% centralities for multistrange hadrons.

4) p_T Spectrum and Particle Yields

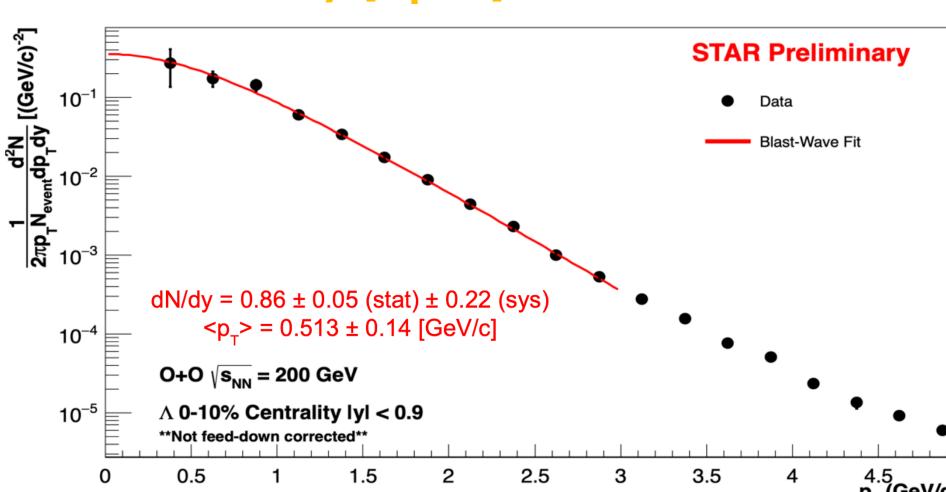


Figure 10: Corrected p_T spectrum for Λ 's in Central O+O Collisions

O+O $\sqrt{s_{NN}} = 200 \text{ GeV}$

Figure 11: Yields as a function of average N_{nat}

O+O Not feed-down corrected

Cu+Cu $\sqrt{s_{NN}} = 200 \text{ GeV (PRL 2012)}$

 $Au + Au \sqrt{s_{NN}} = 200 \text{ GeV (PRL 2012)}$

The p_T spectrum is calculated from the Λ 's invariant mass distributions in different momentum ranges.

The Λ p_T spectrum is the P_T (GeV/c)⁵ average of both magnetic field configurations.

> Most central O+O collisions have a similar < N_{part}> as peripheral Au+Au and Cu+Cu collisions.

$$\frac{dN}{dy} = \int_0^\infty p_t$$

$$\frac{dN}{dy} = 0.86 \pm 0.05 \pm 0.22$$

**O+O yield is not feed-down corrected.

5) Summary and Outlook

 The O+O dataset can fill in the gaps in the low-multiplicity regions of the ratio of strange hadron production to the pion yield for the STAR data.

Mid-rapidity Λ Yield

STAR Preliminary

- We presented the first yield calculation for \(\Lambda\)'s in the 0-10% centrality region for O+O. The O+O yield agrees with previous published STAR ∧ yields at similar N_{part} values.
- Extend the analysis to other hyperons.
 - x) The raw p_T spectra are pending the corrections.
- Use thermal model for freeze-out parameter (e.g. μ_B , T_{ch}) extraction. [1] Brookhaven National Laboratory. (2023, February 24). Clear sign that QGP production 'turns off' at low energy. [2] P. Koch, et al. Phys. Rep. 142, 167 (1986) [3] ALICE Collaboration. Nat. Phys., 13, 535 (2017) [4] Picture: Alex & Maria Schmah [5] Maksym Thesis (2016))





