

IMPROVING ENERGY RESOLUTION FOR THE STAR FORWARD CALORIMETER SYSTEM

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for the STAR Collaboration

Fall Meeting of APS Division of Nuclear Physics

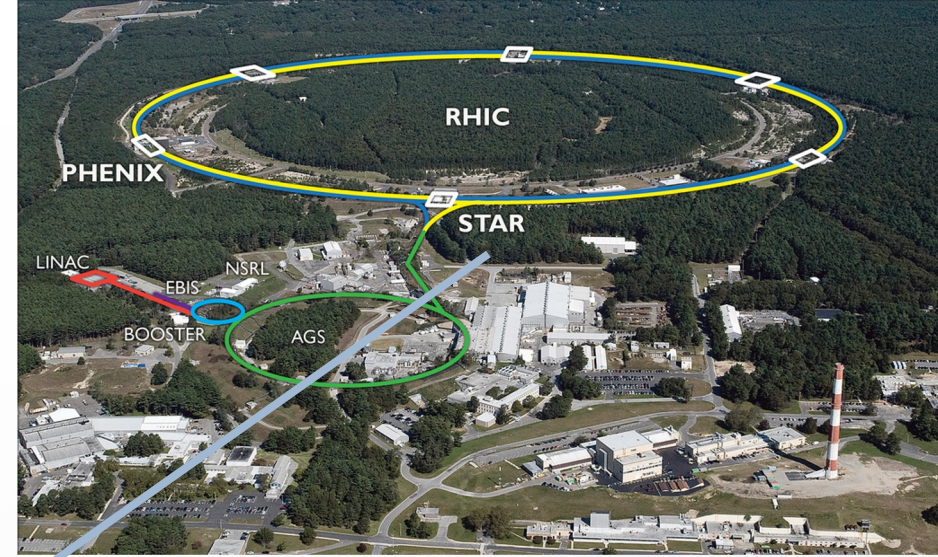
October 2022, New Orleans, LA



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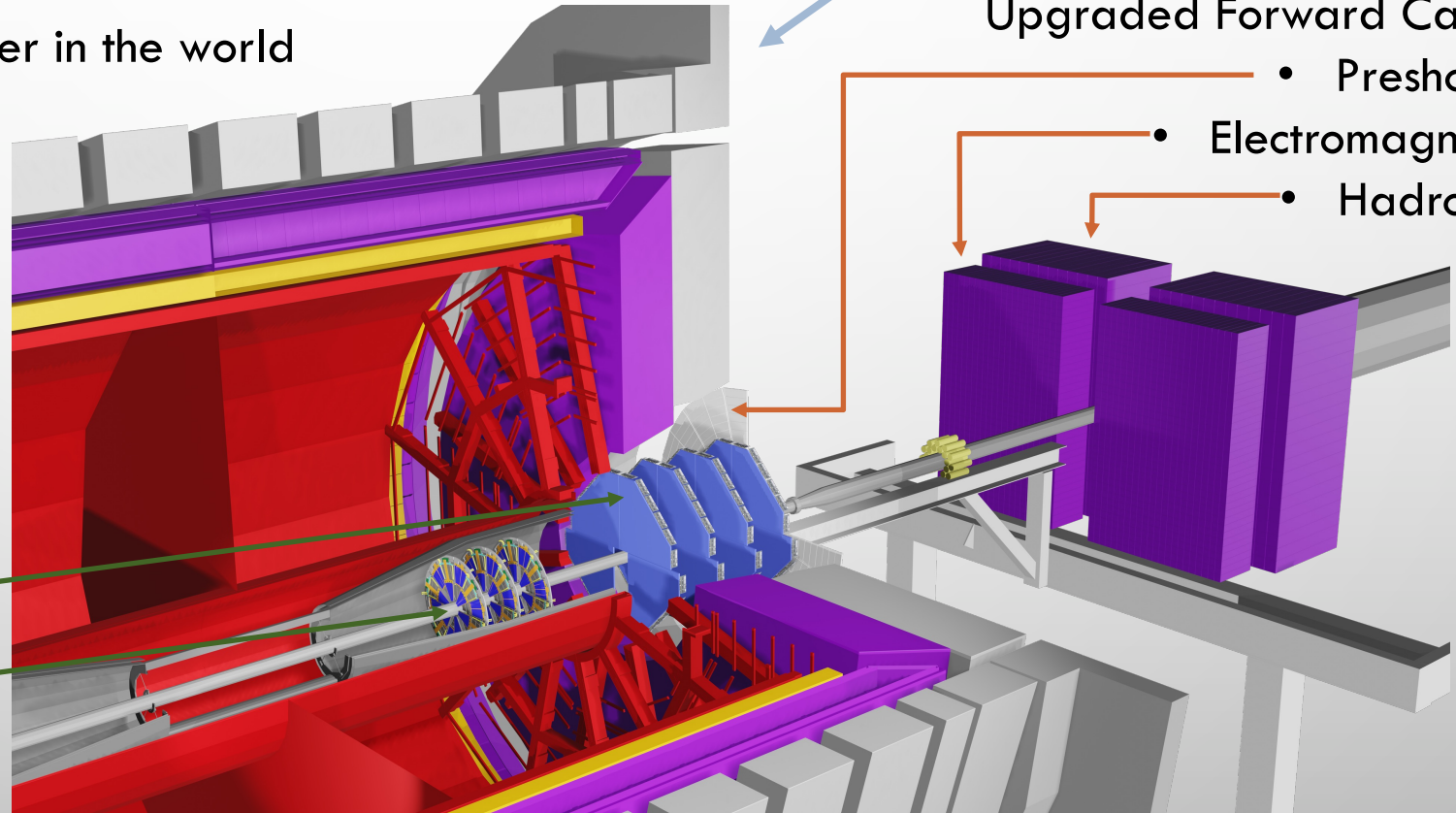
STAR FORWARD UPGRADE



- STAR detector is located at the Relativistic Heavy Ion Collider (RHIC) at Brookhaven National Lab on Long Island NY
- RHIC is the only polarized pp collider in the world
- Upgrades STAR capabilities in $2.5 < \eta < 4$
- Almost full 2π coverage in azimuth
- Tracking and Calorimetry will provide excellent PID

Tracking

- small strip Thin Gap Chambers
- Silicon Disks

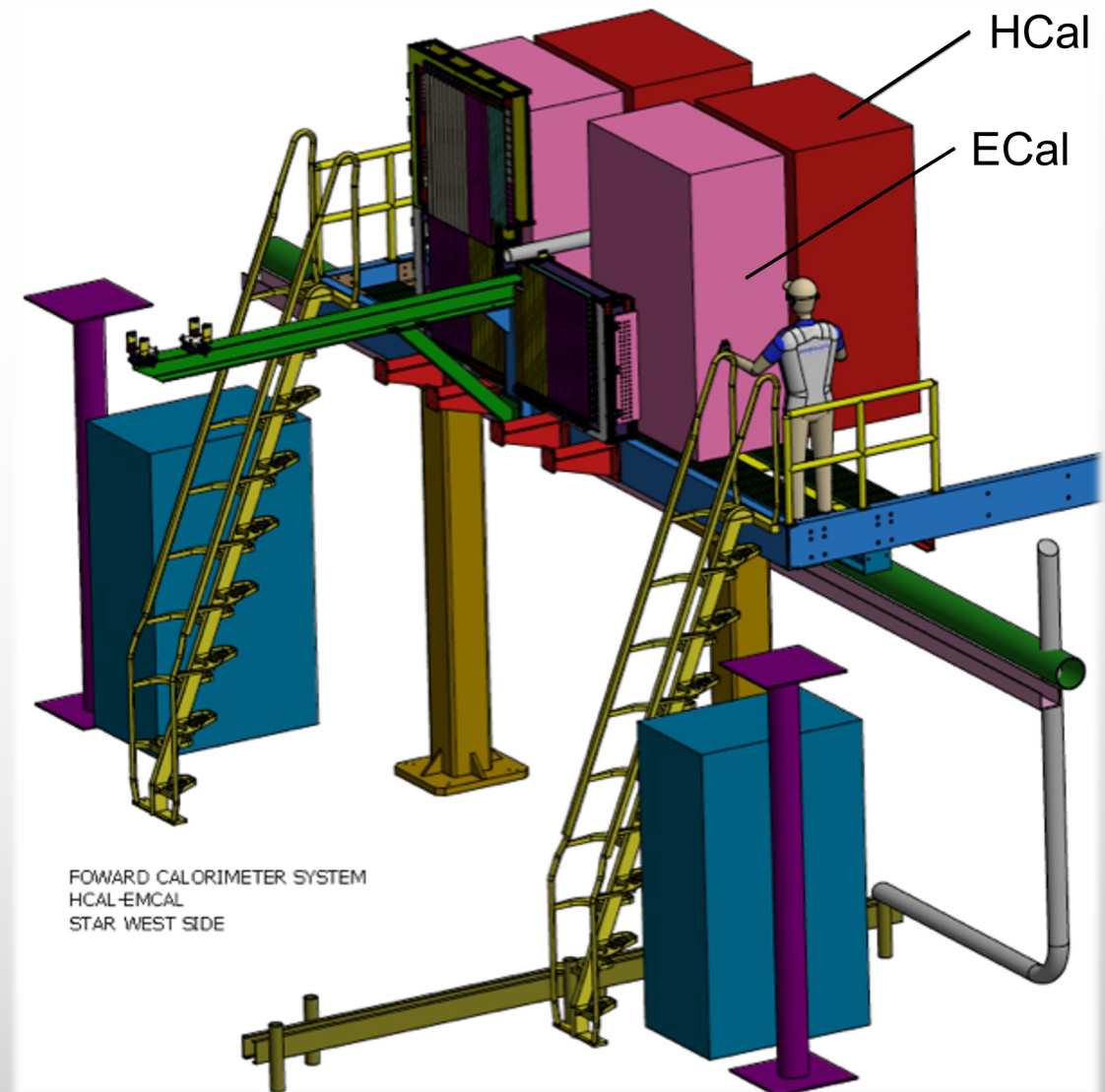


Upgraded Forward Calorimeters

- Preshower (EPD)
- Electromagnetic (EMCal)
- Hadronic (HCal)

FORWARD CALORIMETER SYSTEM (FCS)

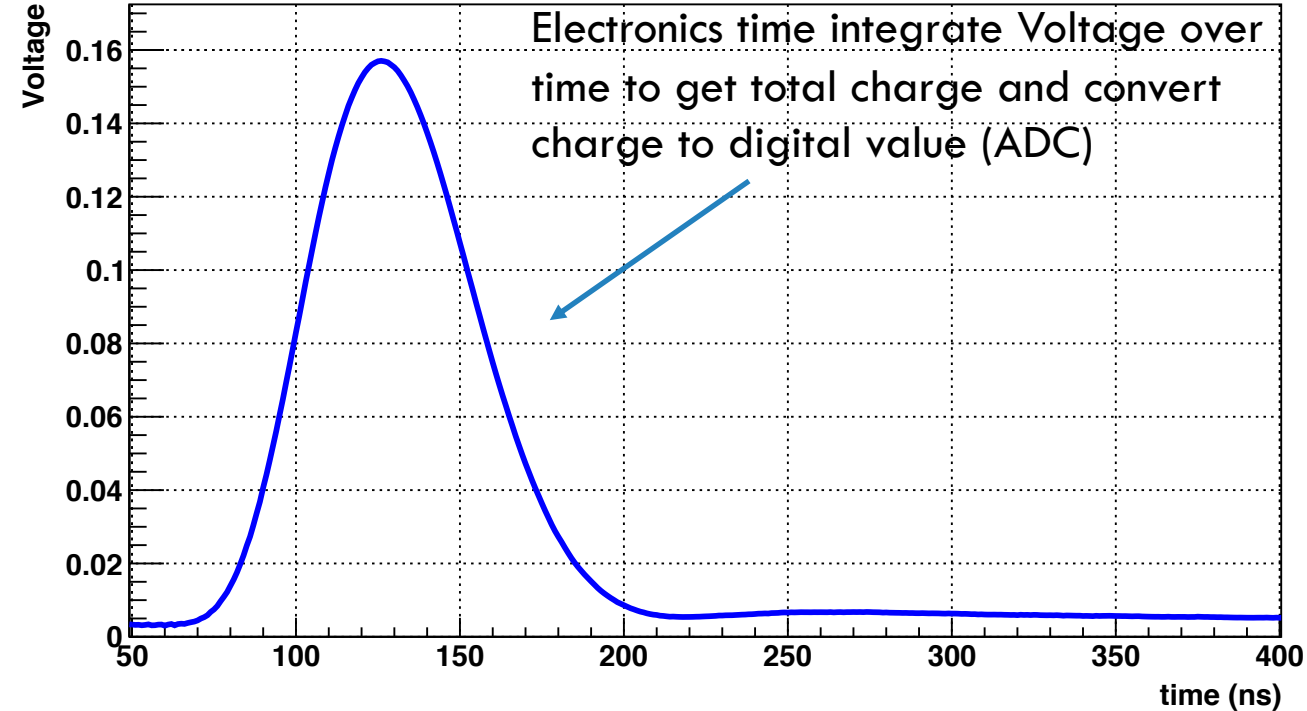
- Located $\sim 7\text{m}$ from STAR IP on STAR west platform
- Both Ecal and Hcal split into two halves (North/South)
- Use SiPMs for readout
- ECal is Pb/Sc sandwich (from PHENIX)
 - $\sim 10\%/\sqrt{E}$
 - $5 \times 5 \text{ cm}^2$ lateral size, $\sim 18 X_0$
- HCal is Steel/Sc sandwich built from scratch
 - $\sim 60\%/\sqrt{E}$
 - $10 \times 10 \text{ cm}^2$ lateral size, $\sim 4.5 \lambda$
- Preshower is existing EPD detector (not shown)



GETTING THE ENERGY

- Signals from a detector are digitized by time integrating the voltage (ADC) of the signal over the whole time of a trigger window
- STAR trigger windows are defined by the time between RHIC bunches (bunch crossing)
- The energy deposited in the detector $(E) = \text{ADC} * \text{Gain}$
 - Gain needs to be calibrated
 - See Xilin Liang's talk next

Sample LED pulse from FCS SiPM

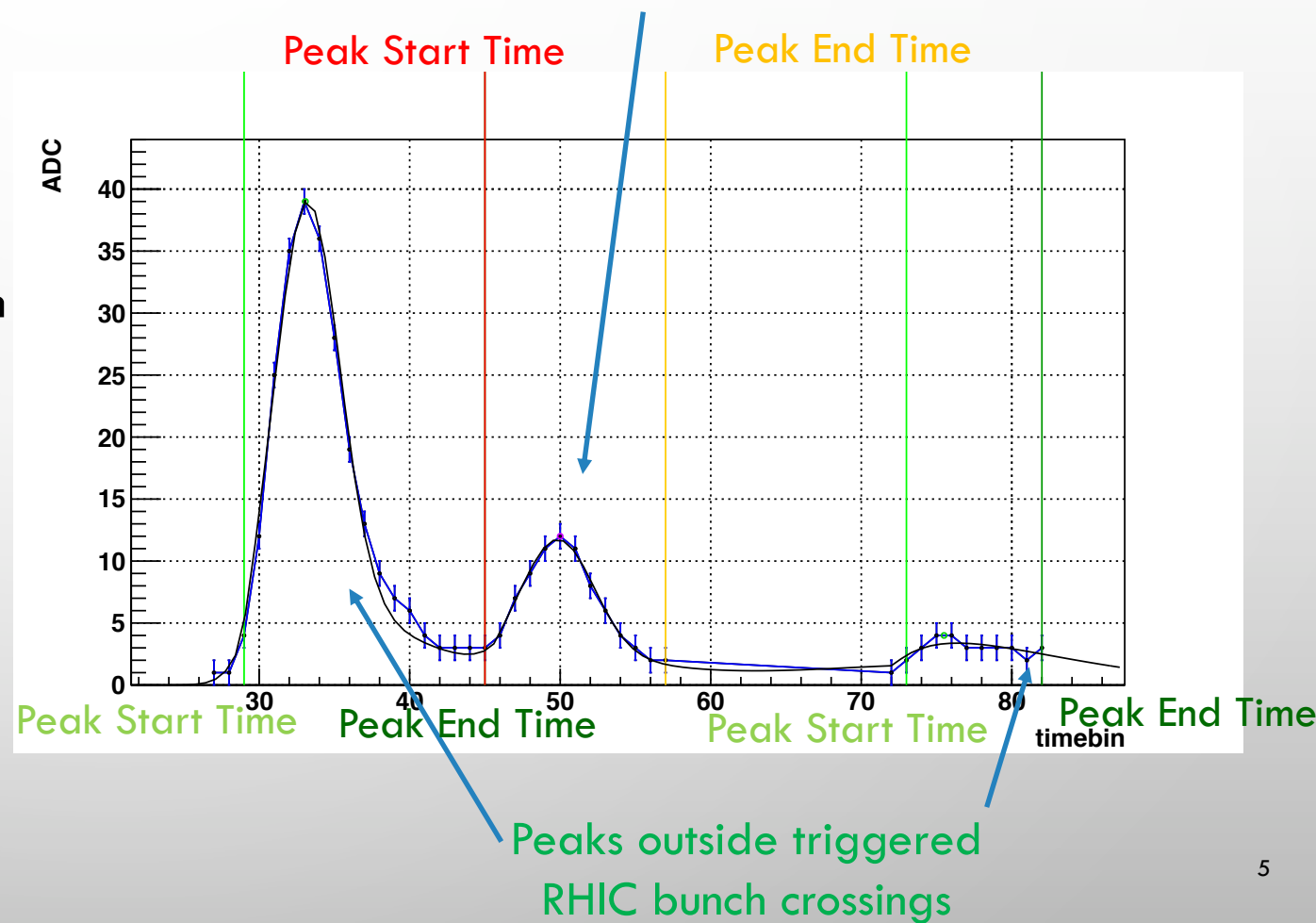


DIGITIZATION OF SIGNAL

- DEP boards digitize signal every ~ 13.5 ns
- This comprises 1 timebin (Tb)
- There are 8 Tb in 1 RHIC bunch crossing
- There is up to 100 Tb of data for every channel in every event
- Energy = (Fitted signal integral)*Gain
 - Each signal/peak is fitted to a Gaussian
 - Fitted signal to all peaks shown in black
 - Amplitude of Gaussian is proportional to integral
- Peaks found using discrete second derivative test

Sample Signal from DEP showing multiple RHIC bunch crossings

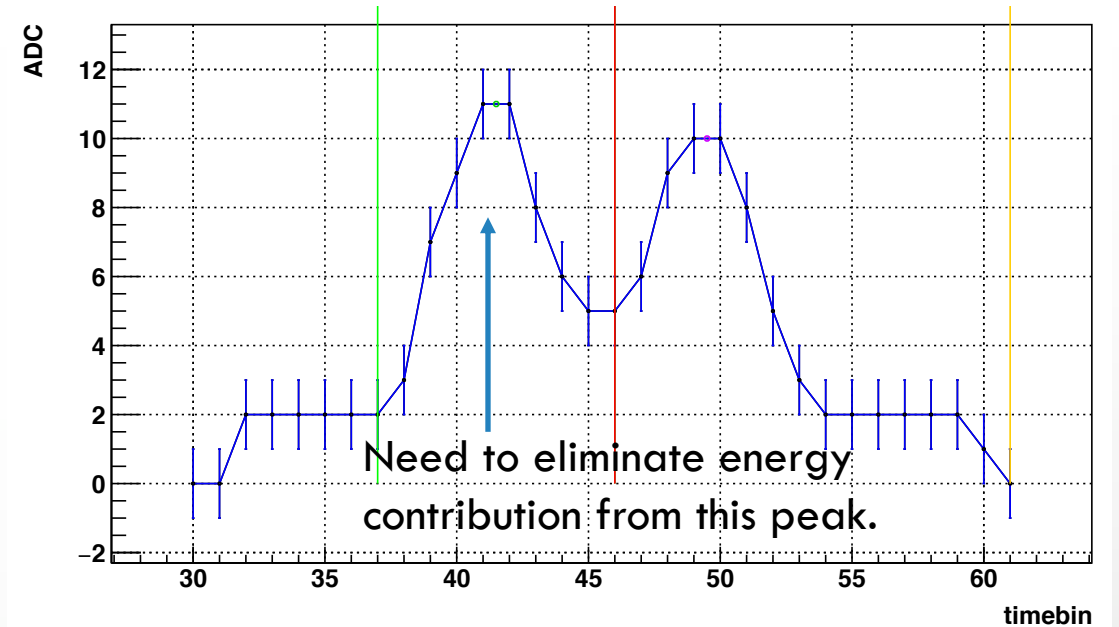
Triggered RHIC crossing peak at Tb=50



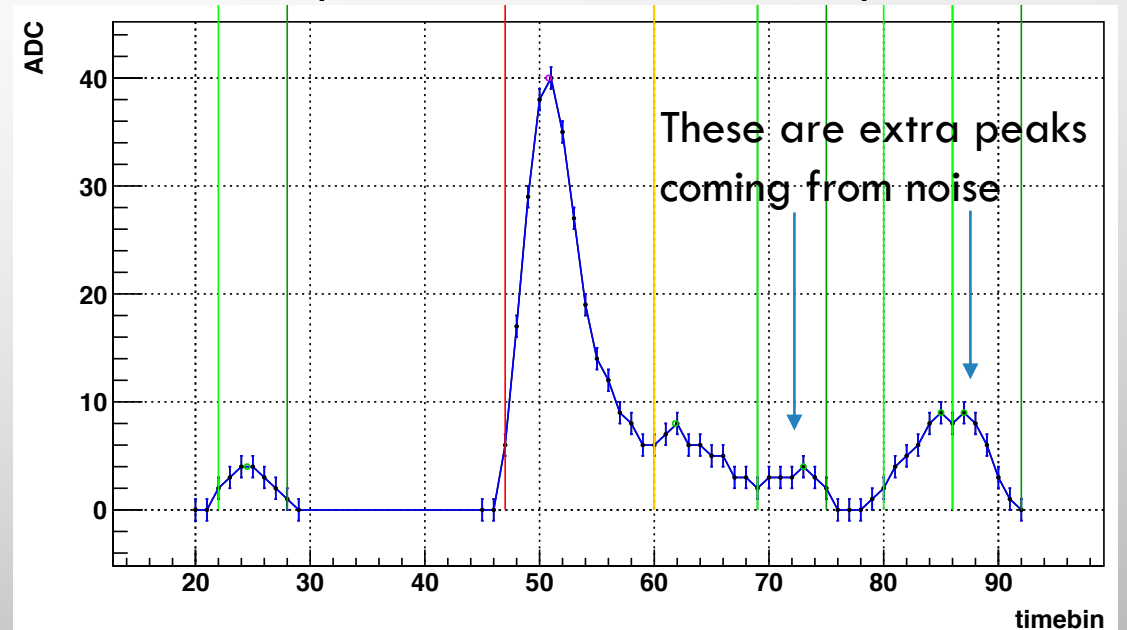
CHALLENGES

1. Fitting is time consuming and want to avoid fitting in cases where it is not needed
 - If only 1 peak in the triggered crossing adding up the ADCs is sufficient
2. Overlapping signals from untriggered RHIC crossing
 - Want to eliminate energy contribution coming from surrounding peaks
 - Fitting will help but need accurate count of peaks
3. Second derivative method can identify noise as peaks
 - Don't want to fit a signal to noise
 - More peaks \rightarrow more fitting parameters \rightarrow computation time

Example of Overlapping signals

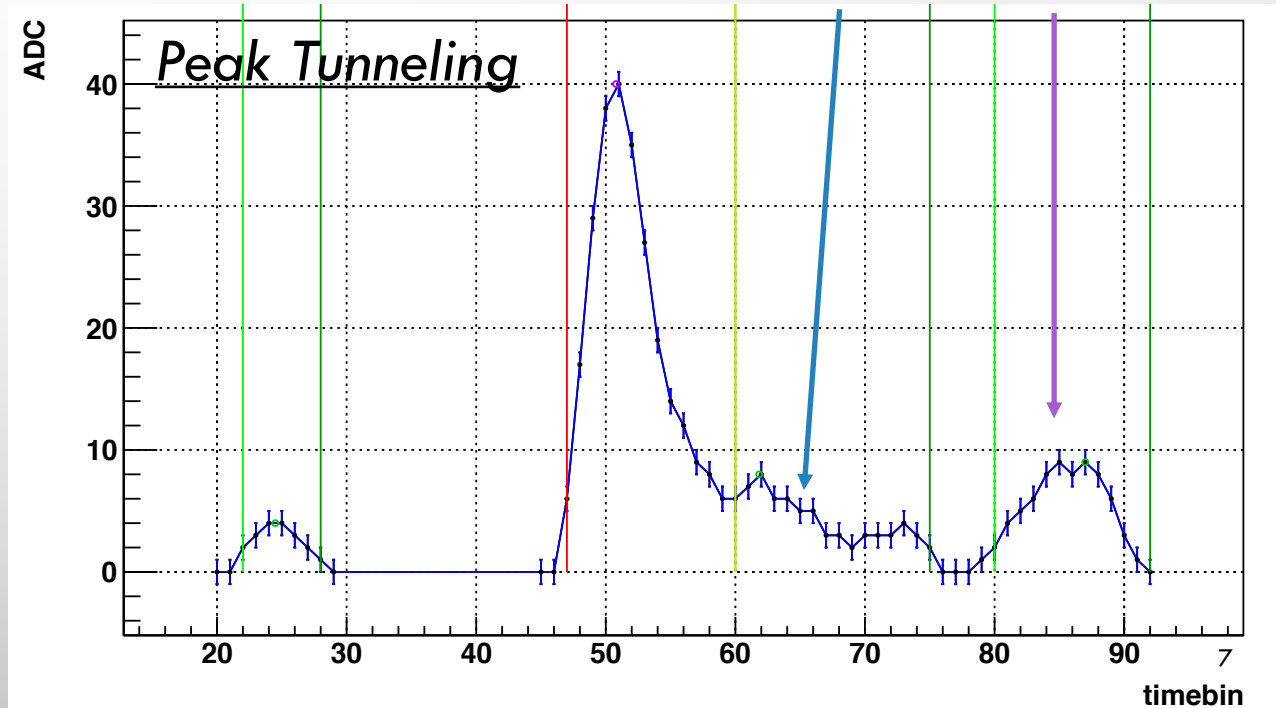
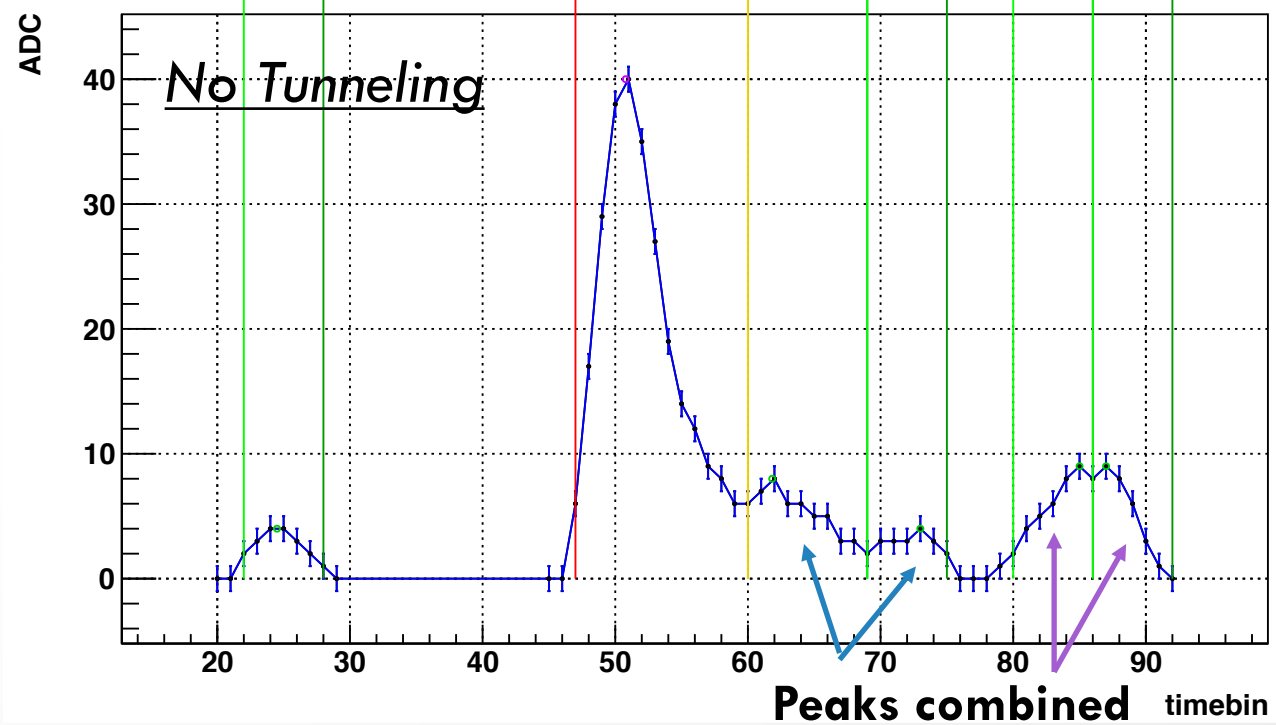


Example of noise identified as peaks



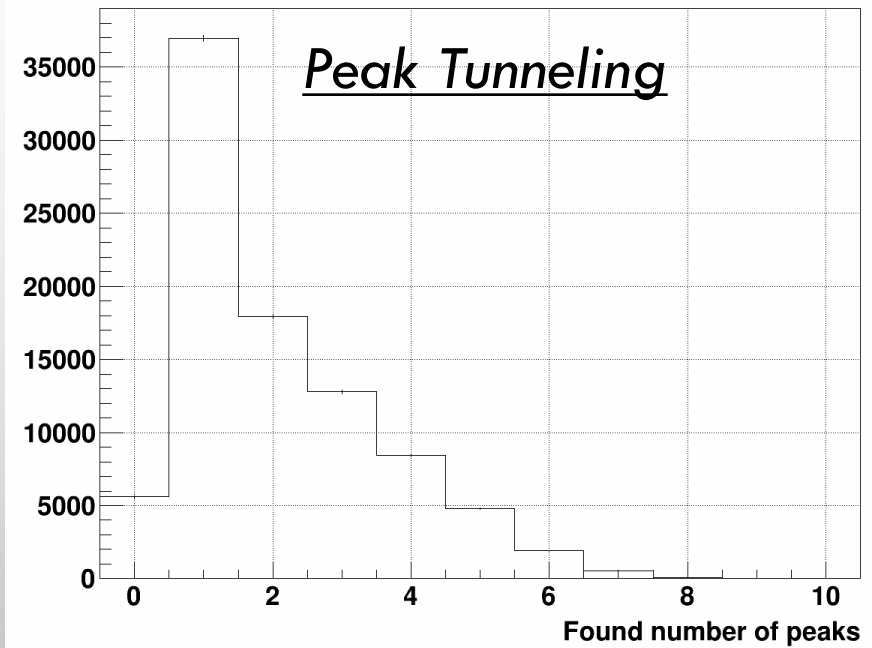
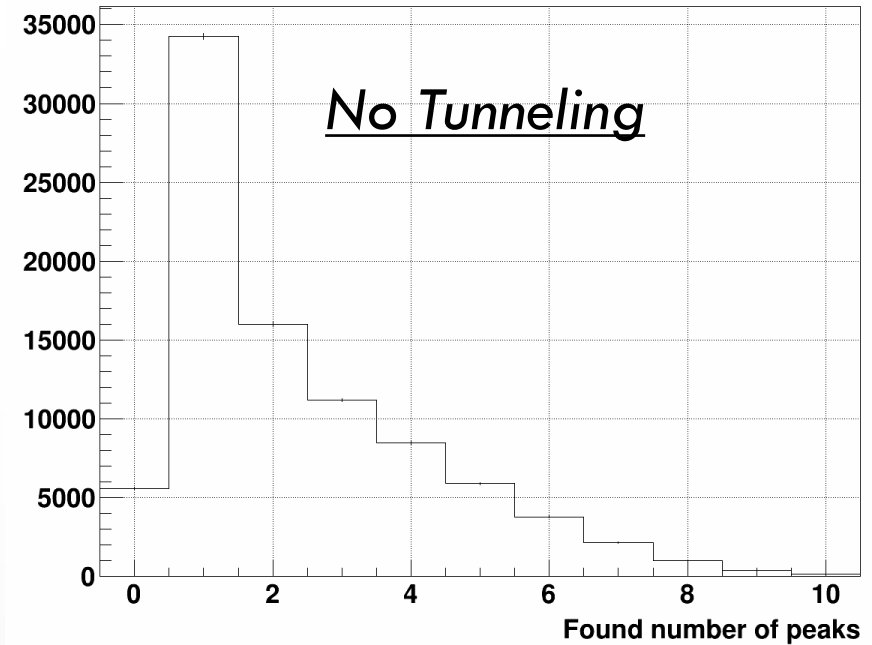
ELIMINATING PEAKS COMING FROM NOISE

- Compute a probability a peak is a real peak or noise
 - Call this method peak tunnelling as it was inspired by quantum tunneling
- Assumes noise follows a Normal distribution
 - Parameters need to be tuned to data set
- Merge peaks that have low probabilities together
 - Higher probability peak is identified as peak position when merging



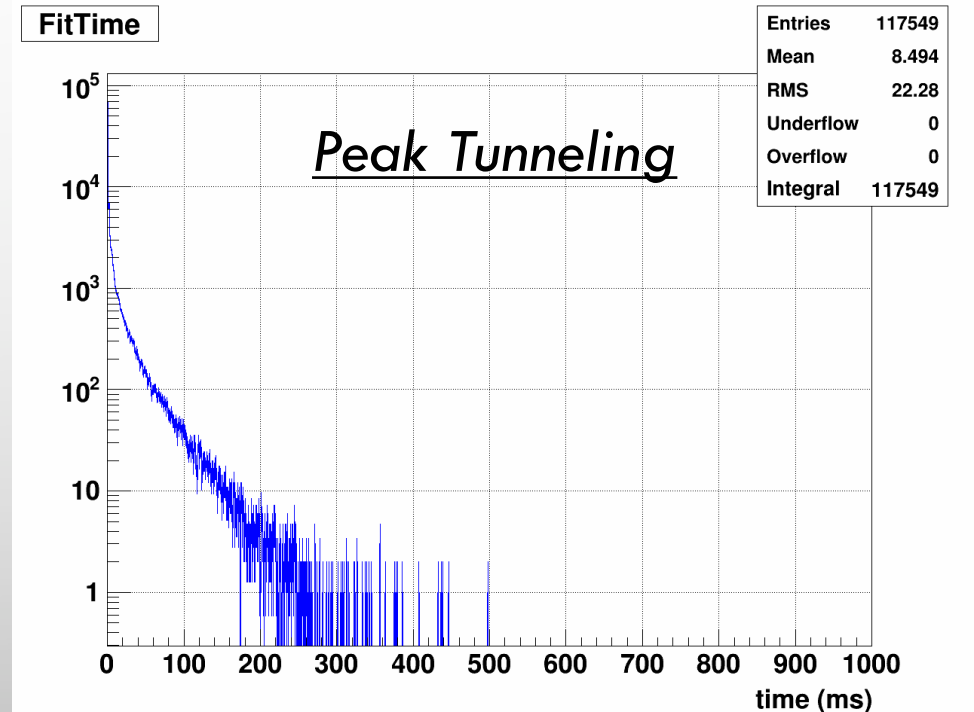
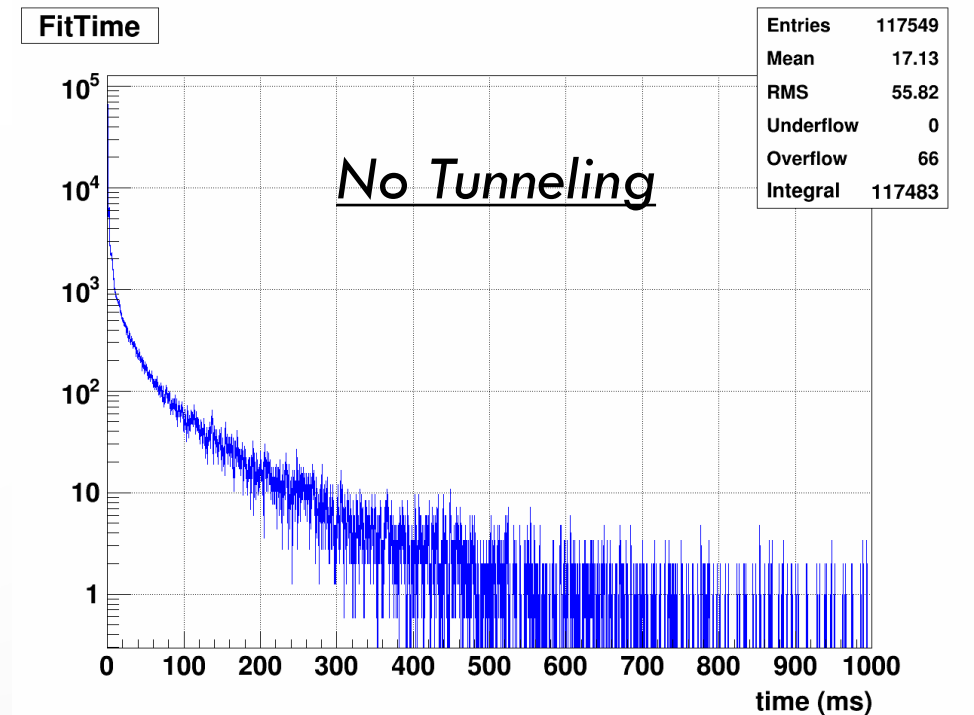
PERFORMANCE: NUMBER OF PEAKS

- Ran 100 events on RHIC run 22 pp data
- Plots show number of peaks with and without peak tunneling method
- Peak tunneling shows much lower number of peaks on average



PERFORMANCE: TIMING

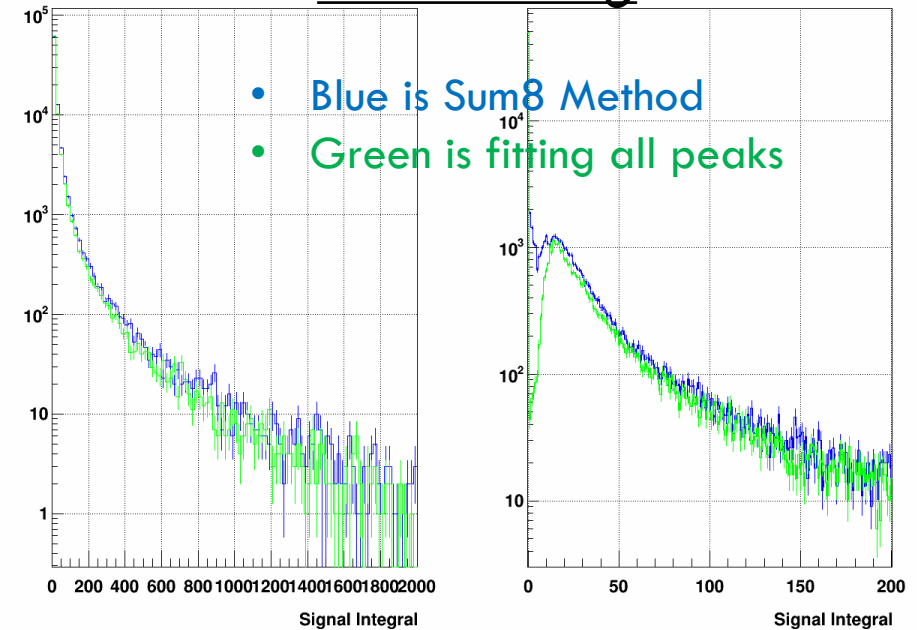
- Fitting all signals to known signal shape with a fixed number of peaks
- Timing with peak tunneling is twice as fast
- This is due to less peaks so fewer fitting parameters



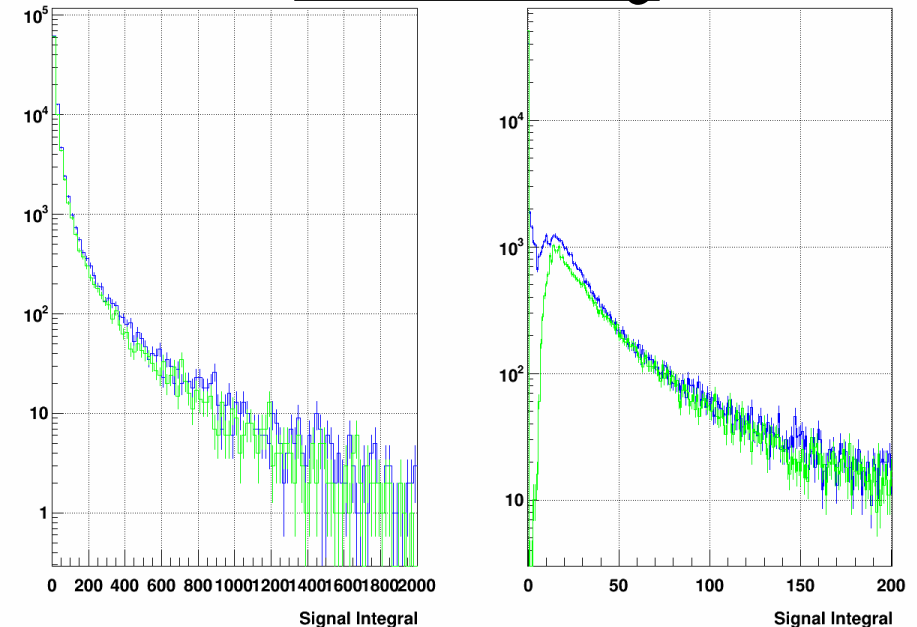
PERFORMANCE: ADC SUM

- Plots are comparing fitted signal integral with just summing ADCs with and without tunneling
- Sum8 method adds ADCs in 8 Tb of triggered RHIC crossing
 - Fastest method
 - Shown in Blue
- Sum8 method to fitting overestimates low side of energy compared to fitting
 - Shown in Green
- Very little difference in integral sum between with and without tunneling

No Tunneling



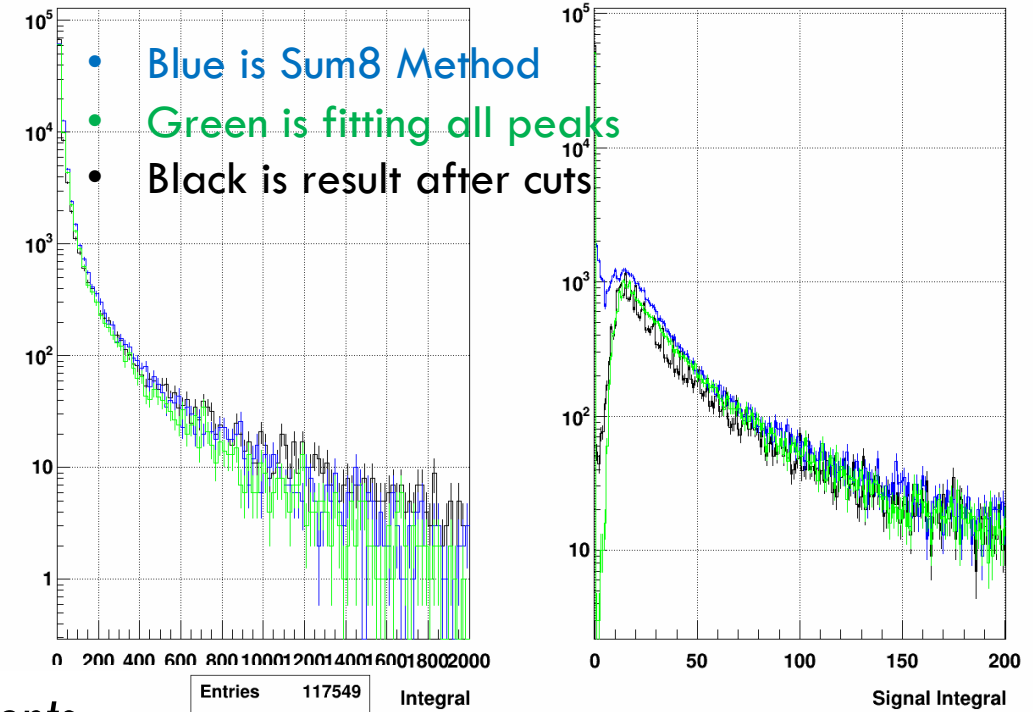
Peak Tunneling



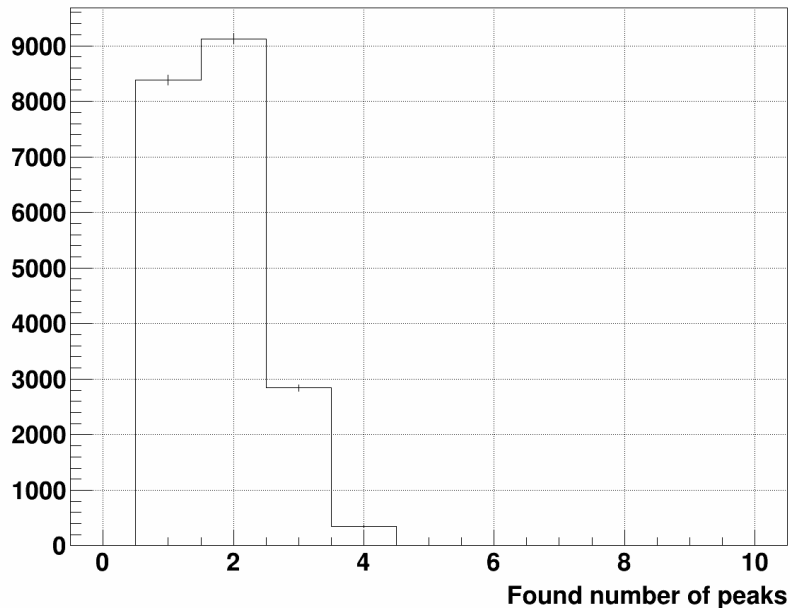
PUTTING IT ALL TOGETHER

- Final algorithm uses peak tunneling to determine number of peaks and their locations
- Cut out all peaks not within 3 RHIC bunch crossings
 - This ensures only overlapping peaks are fit
- Only fit peaks in those cases all other use Sum8 method

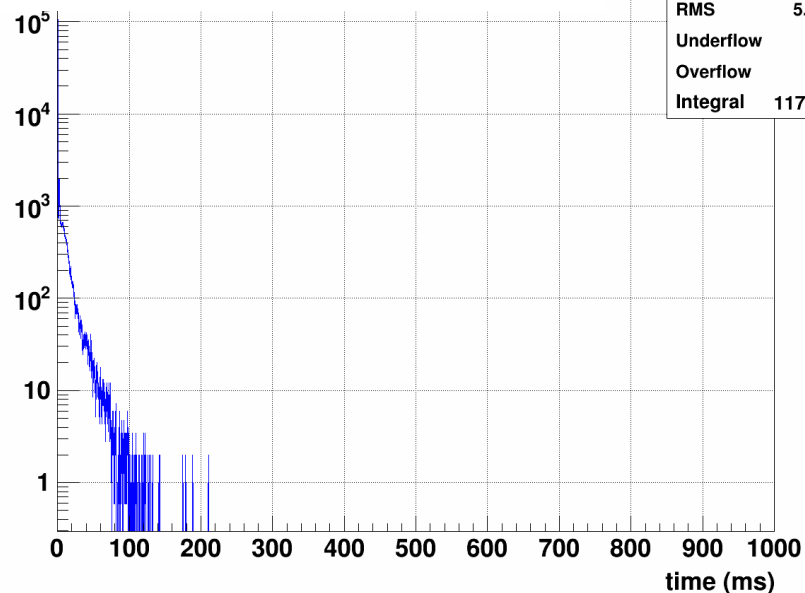
Final ADC sum



Final number of peaks



Final Timing for 100 events



- Accurate determination of peaks leads to faster computation time
- Aids in fitting overlapping peaks to correctly determine the energy

CONCLUSIONS

- STAR Forward upgrade greatly enhances capabilities of STAR in forward region
 - Upgrade will allow much better particle ID at STAR forward region
- The FCS uses latest technology to help improve energy determination
 - DEP boards have much finer digitization to record more details about the raw signal
 - More information means more analysis, which means more computation time to determine energy
- Developed robust algorithm to find peaks and determine energy quickly
 - Employed peak tunneling to reduce noise on found peaks
 - Cut out fitting peaks that don't overlap

BACKUP

FITTED SIGNAL SHAPE

- Signal shape $C + \frac{A_1}{\tau_1^2} (x - x_1)^{p_1} + e^{-(x-x_1)/\tau_1} + \frac{A_2}{\tau_2^2} (x - x_2)^{p_2} + e^{-(x-x_2)/\tau_2} + Ae^{-\frac{(x-x_0)^2}{2\sigma^2}}$
- C is a constant offset usually fixed to 0 unless there is a pedestal (baseline)
- $A_1, A_2, \tau_1, \tau_2, p_1, p_2, x_1, x_2$ are fixed constants to characterize size of tail
- A is the magnitude of the Gaussian and is also a fitting parameter
- x_0 is mean of Gaussian and is also a fitting parameter
- σ is the standard deviation of the Gaussian and is also a fitting parameter
- Multiple peaks just linearly sum this formula

PEAK TUNNEL METHOD

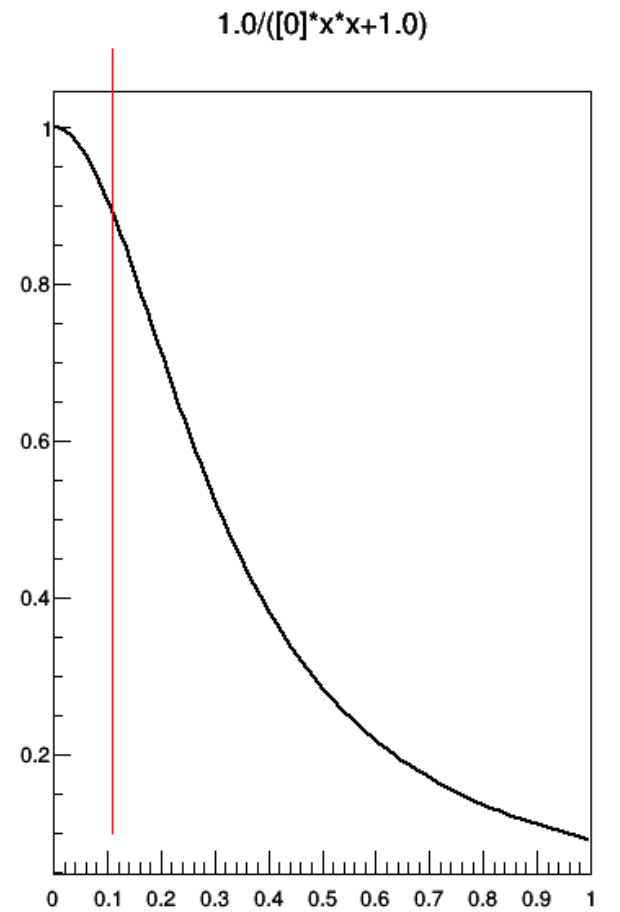
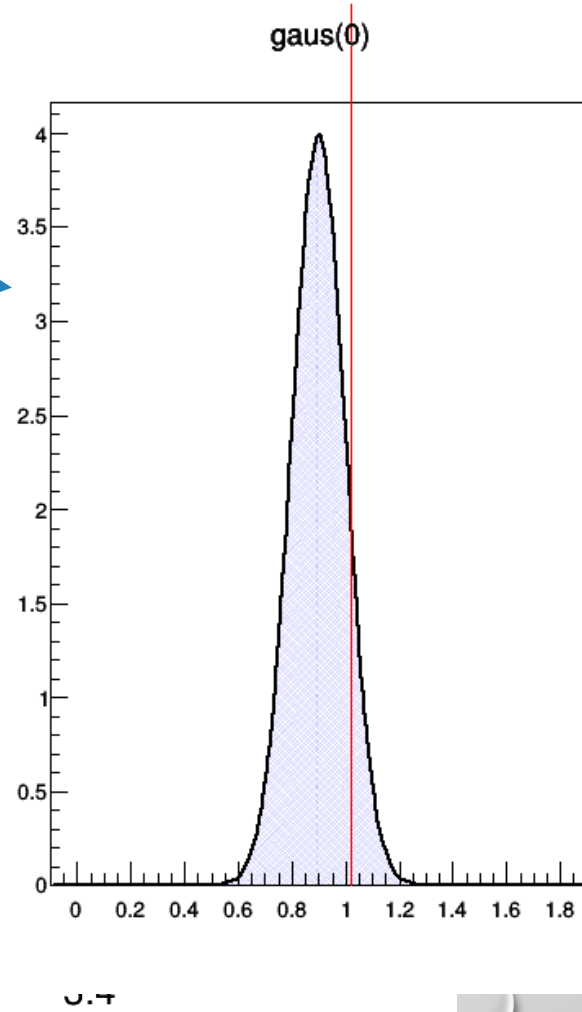
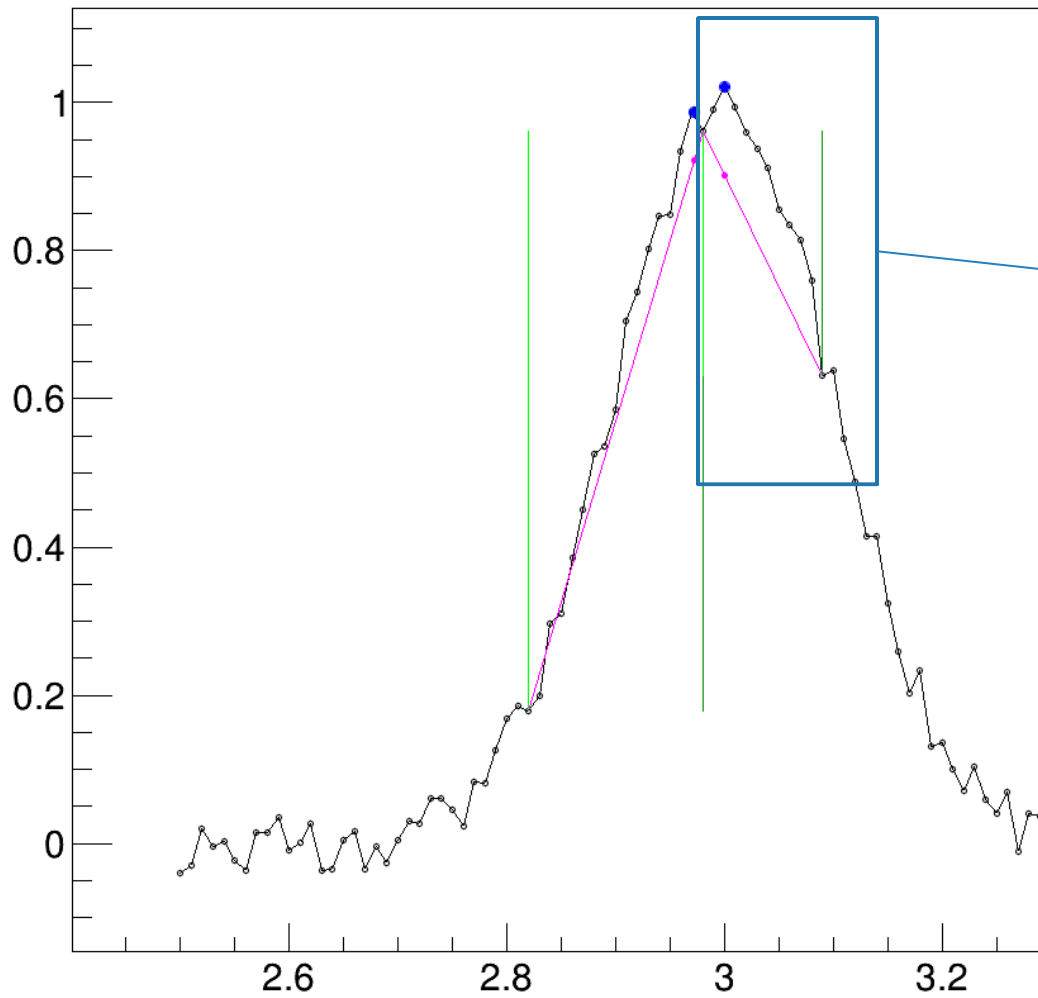
- Compute a probability that a given peak window contains a real peak or noise
- Probability considers two things: width of the window, size of the peak relative to the slope line

$$1. \quad P_{width}(x_s, x_e) = \frac{1}{a*(x_e - x_s)^2 + 1}$$

$$2. \quad P_{height} = \text{Erfc}\left(\frac{\text{heightdiff}}{\sqrt{2}\sigma}\right)$$

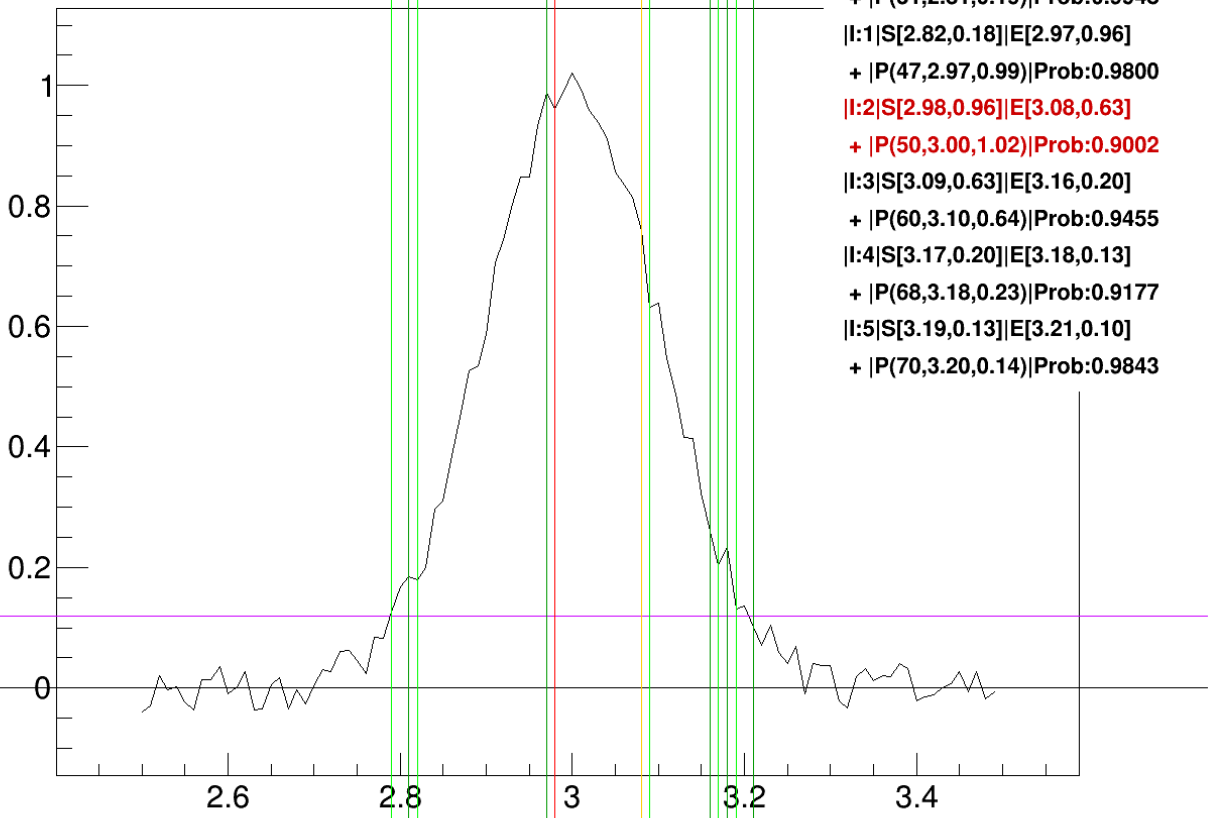
- $\text{heightdiff} = p_y - \text{yline}(p_x)$
- yline is the line formed by the points x_s, y_s, x_e, y_e
- x_s, y_s, x_e, y_e are the x, y points of the start and end of the found peak respectively
- a and σ are two scale parameters that need to be set.
- The total probability is $P_{width} * P_{height}$, with lower probability means more likely it's a peak
- If a peak has a probability higher than some threshold, then it is merged with the adjacent found peak with the lower probability peak dominating.

TUNNEL METHOD ILLUSTRATED



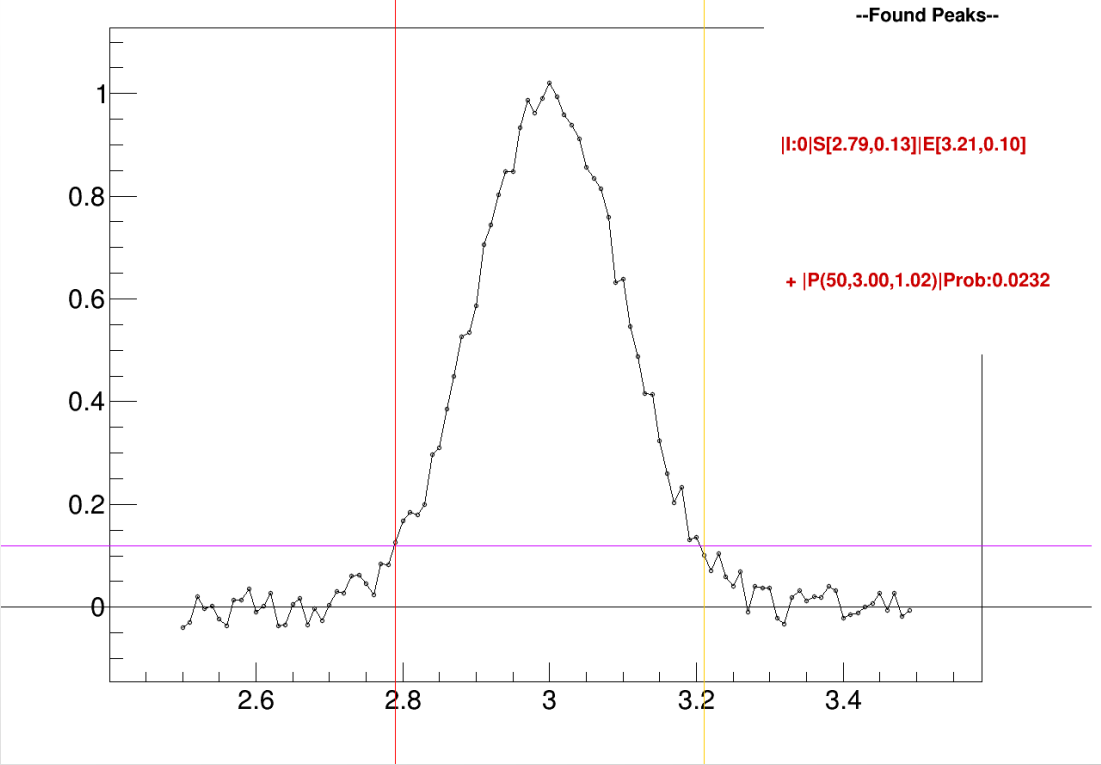
TUNNEL METHOD AT WORK

Test Noisy Gauss



--Found Peaks--
 |I:0|S[2.79,0.13]|E[2.81,0.18]
 + |P(31,2.81,0.19)|Prob:0.9948
 |I:1|S[2.82,0.18]|E[2.97,0.96]
 + |P(47,2.97,0.99)|Prob:0.9800
|I:2|S[2.98,0.96]|E[3.08,0.63]
+ |P(50,3.00,1.02)|Prob:0.9002
 |I:3|S[3.09,0.63]|E[3.16,0.20]
 + |P(60,3.10,0.64)|Prob:0.9455
 |I:4|S[3.17,0.20]|E[3.18,0.13]
 + |P(68,3.18,0.23)|Prob:0.9177
 |I:5|S[3.19,0.13]|E[3.21,0.10]
 + |P(70,3.20,0.14)|Prob:0.9843

Test Noisy Gauss with Tunr



--Found Peaks--
|I:0|S[2.79,0.13]|E[3.21,0.10]
+ |P(50,3.00,1.02)|Prob:0.0232