

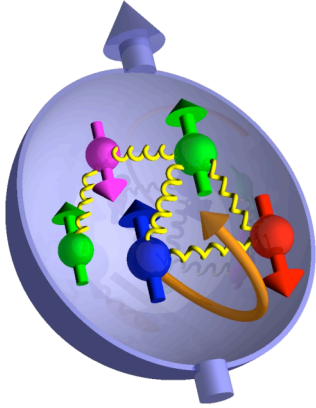
The STAR W Physics Program : New Results and Future Measurements

Joe Seele (MIT) for the

 Collaboration

2009 APS DNP/JPS Meeting

The Spin Puzzle



The proton is viewed as being a "bag" of bound quarks and gluons interacting via QCD

Spins + orbital angular momentum need to give the observed spin 1/2 of proton

$$\frac{1}{2} = \frac{1}{2} \Delta\Sigma + L_q^z + \Delta G + L_g^z$$

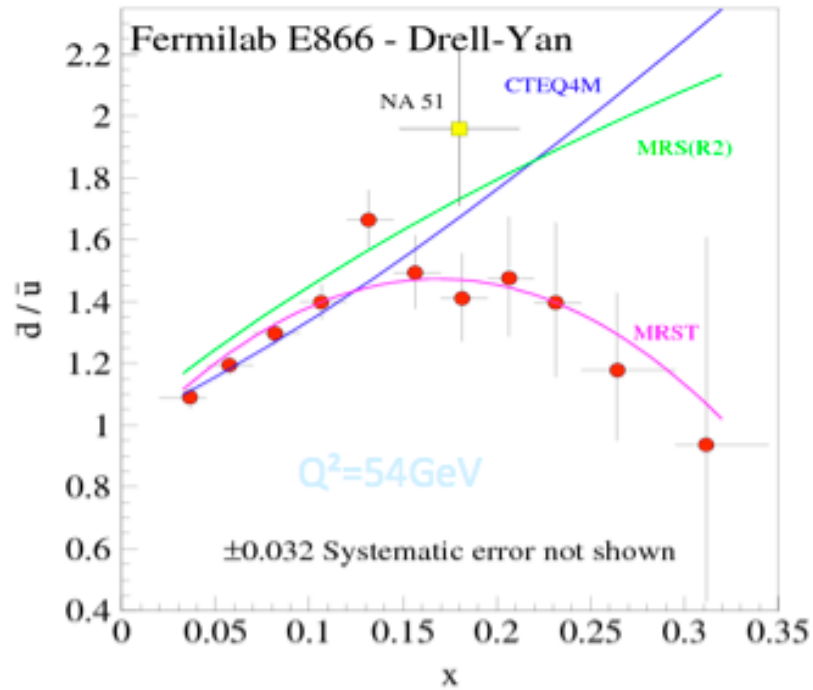
Being measured
at RHIC

Fairly well measured
only ~30% of spin

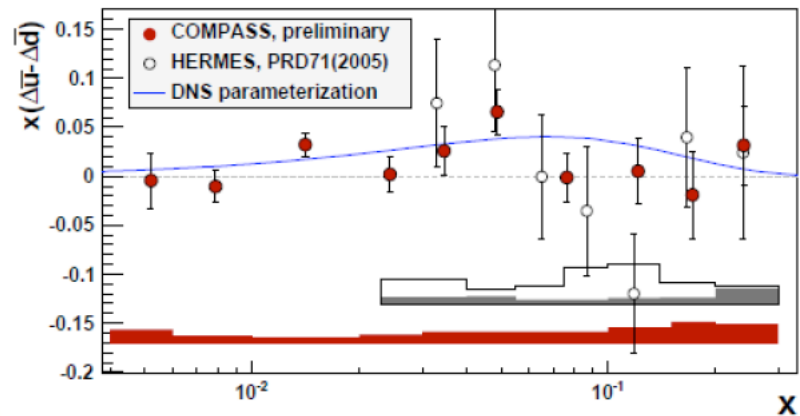
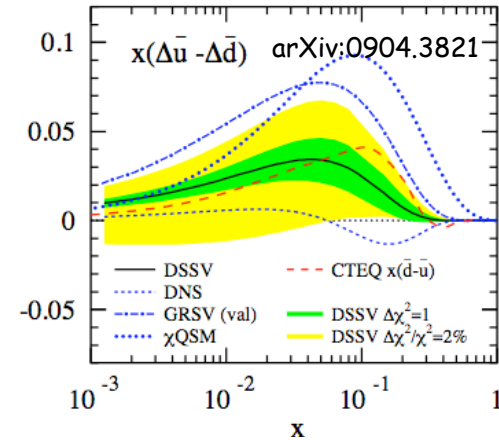
Its decomposition is not well understood,
especially the sea... needs data

$$\Delta\Sigma = \int (\Delta u + \Delta d + \Delta s + \Delta\bar{u} + \Delta\bar{d} + \Delta\bar{s} + \dots) dx$$

Flavor Asymmetry in the Sea

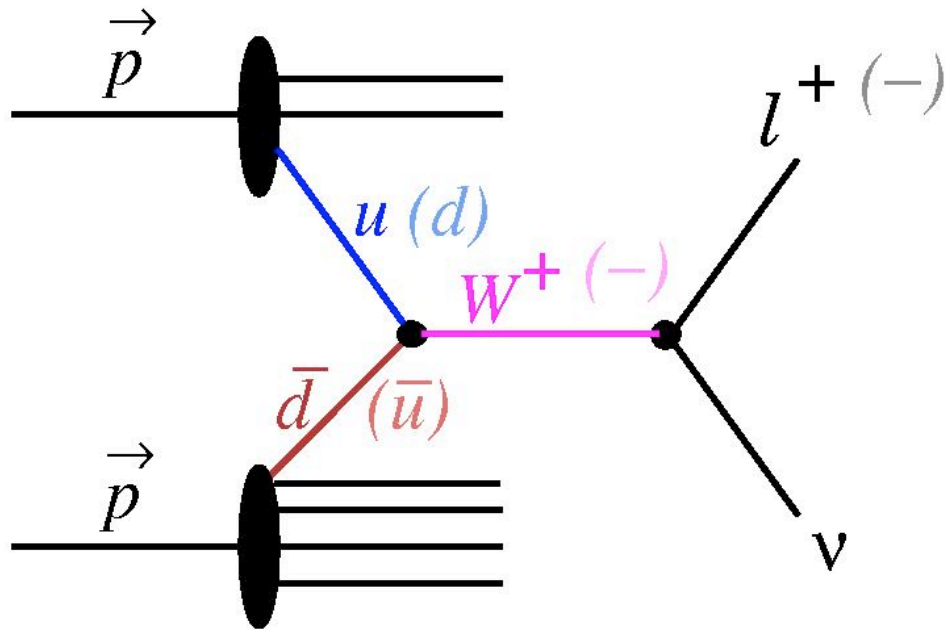


- E866 results are qualitatively consistent with pion cloud models, instanton models, chiral quark soliton models, etc.



- Pauli blocking should contribute to the observed signal, but how much is currently debated
- Non-perturbative processes may be needed in generating the sea

Probing the Sea through Ws



$$u + \bar{d} \rightarrow W^+ \rightarrow e^+ + \nu$$

$$\bar{u} + d \rightarrow W^- \rightarrow e^- + \bar{\nu}$$

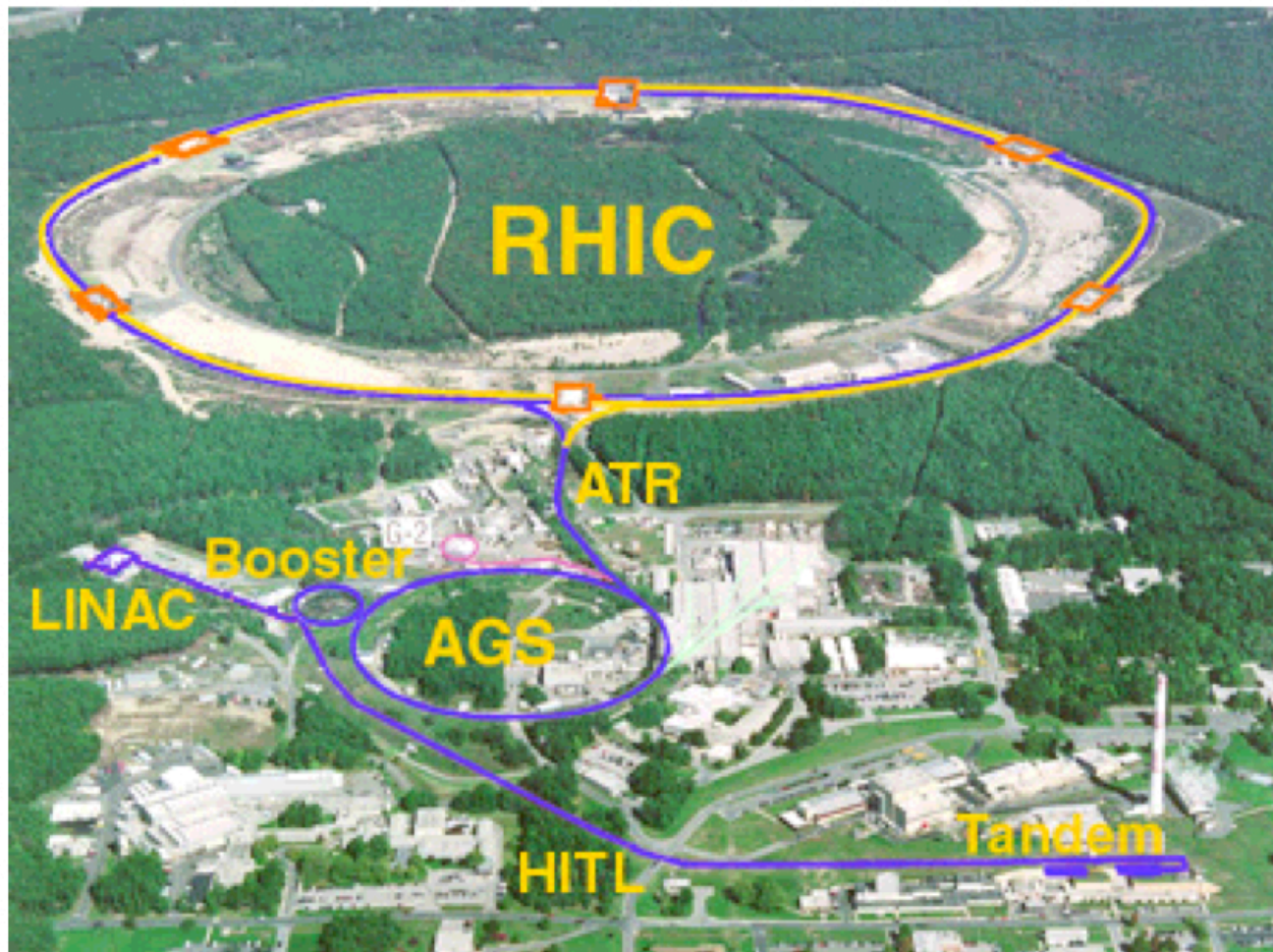
- Detect Ws through e^+ and e^- decay channels
- V-A coupling leads to perfect spin separation
- Neutrino helicity gives preferred direction in decay

Measure parity violating single helicity asymmetry A_L
 (Helicity flip in one beam while averaging over the other)

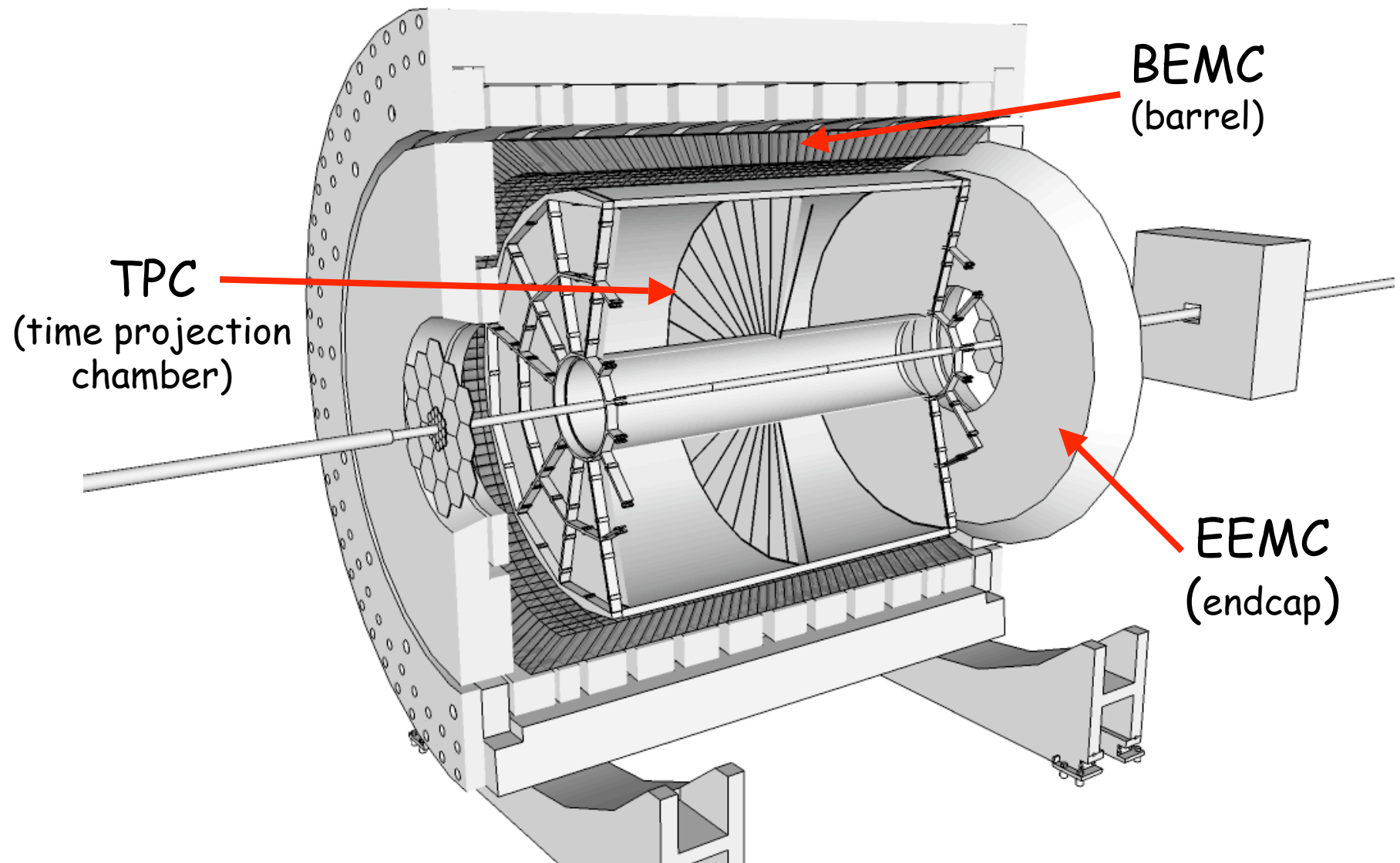
$$A_L^{W^-} \propto -\Delta d(x_1)\bar{u}(x_2) + \Delta\bar{u}(x_1)d(x_2) \quad A_L^{W^+} \propto -\Delta u(x_1)\bar{d}(x_2) + \Delta\bar{d}(x_1)u(x_2)$$

RHIC

A polarized proton collider to study spin in QCD

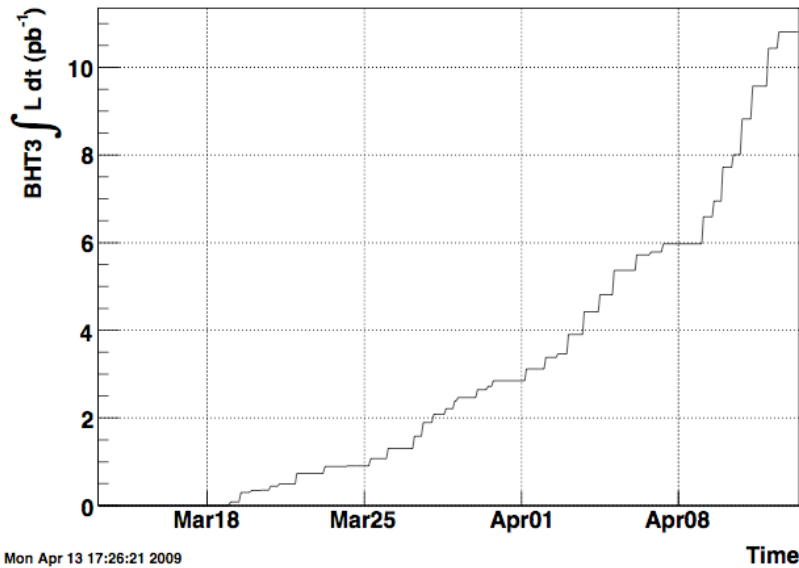


STAR



2009 500 GeV Data Set

2009 STAR 500 GeV pp LongPol BHT3 Recorded Luminosity

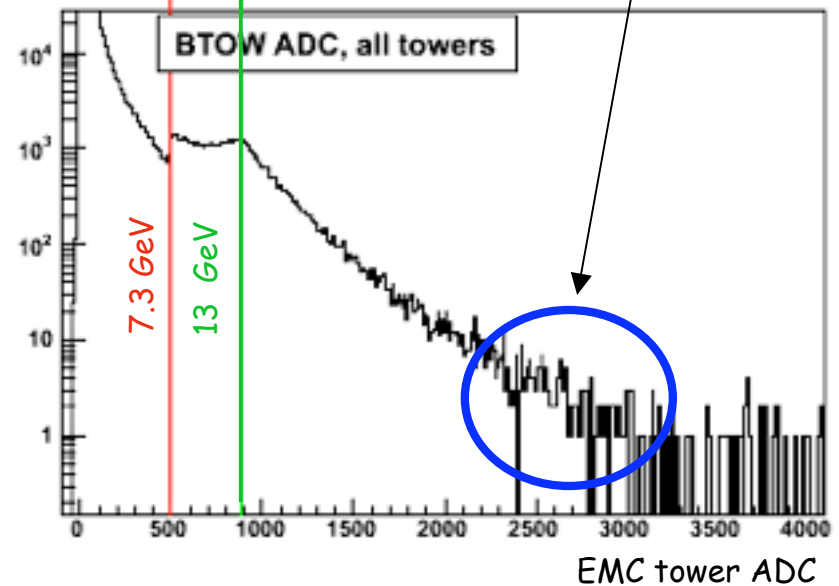


Mon Apr 13 17:26:21 2009

STAR recorded $\sim 10\text{pb}^{-1}$ in the 500 GeV

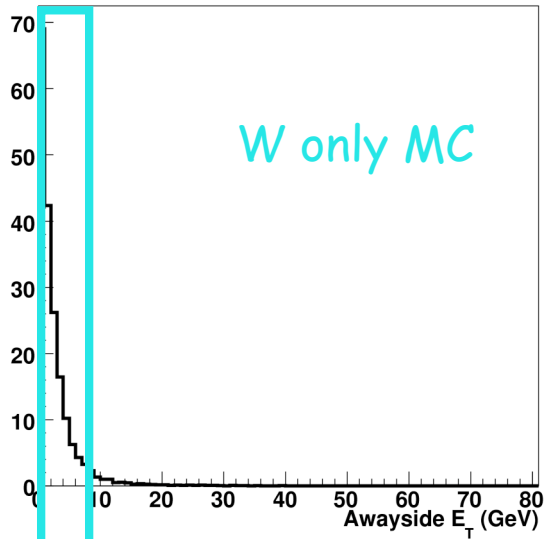
Required a **high tower trigger** ($E_T > 7.3 \text{ GeV}$) and a **high E_T 2x2 clusters** ($E_T > 13 \text{ GeV}$)

leptons from W s should appear here

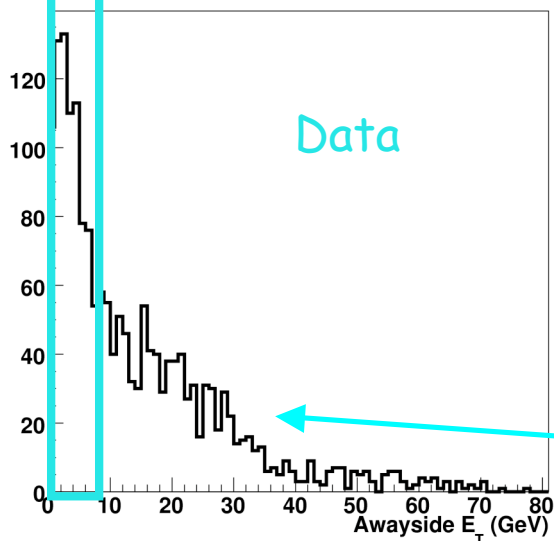


W Algorithm

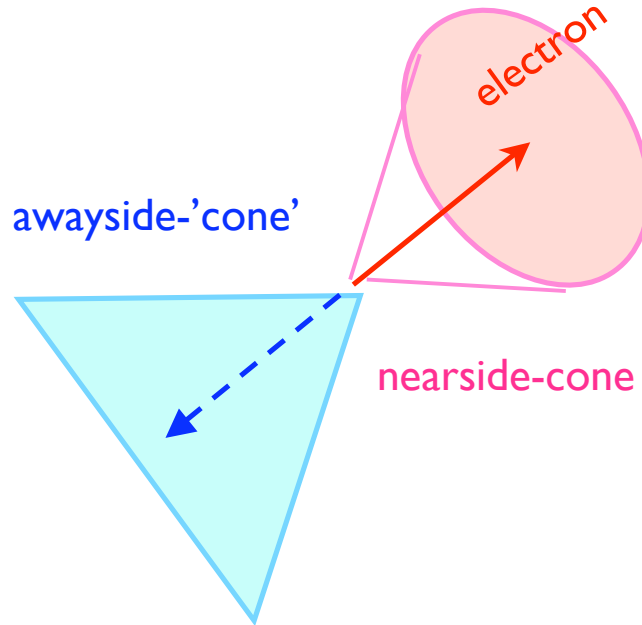
Awayside Sum E_T



Awayside Sum E_T



$W \rightarrow lepton + neutrino$

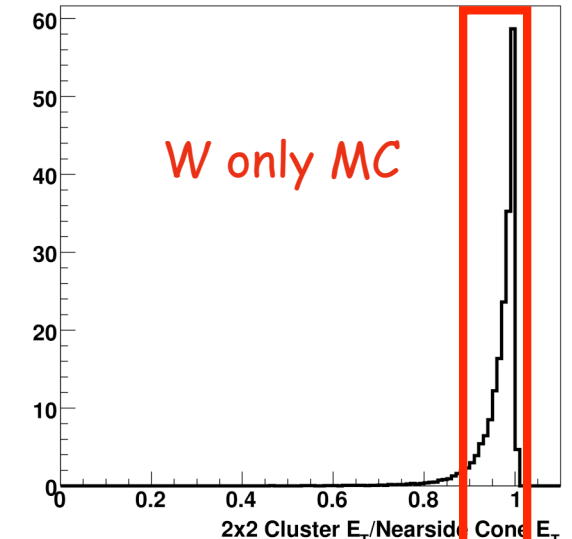


Expect a large QCD background

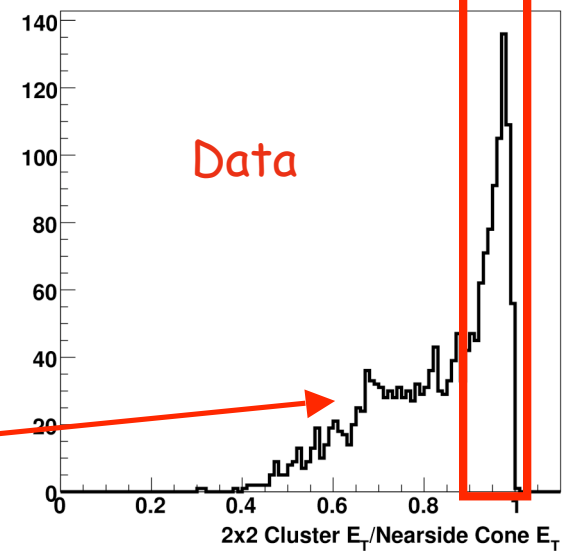
(data shown after some cuts are applied)

J. Seele - 2009 DNP

2x2 Cluster E_T /Nearside Cone E_T



2x2 Cluster E_T /Nearside Cone E_T

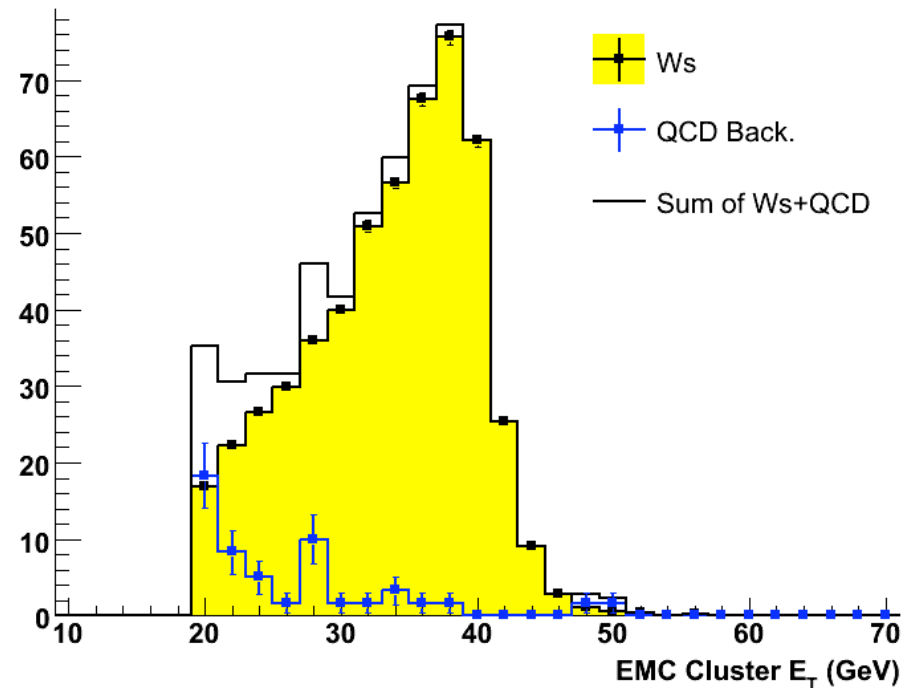


W Algorithm - MC

Full PYTHIA + GEANT simulations were done to assess the performance of the algorithm

Spectra are scaled to 10pb^{-1}

MC Performance of W-Algo



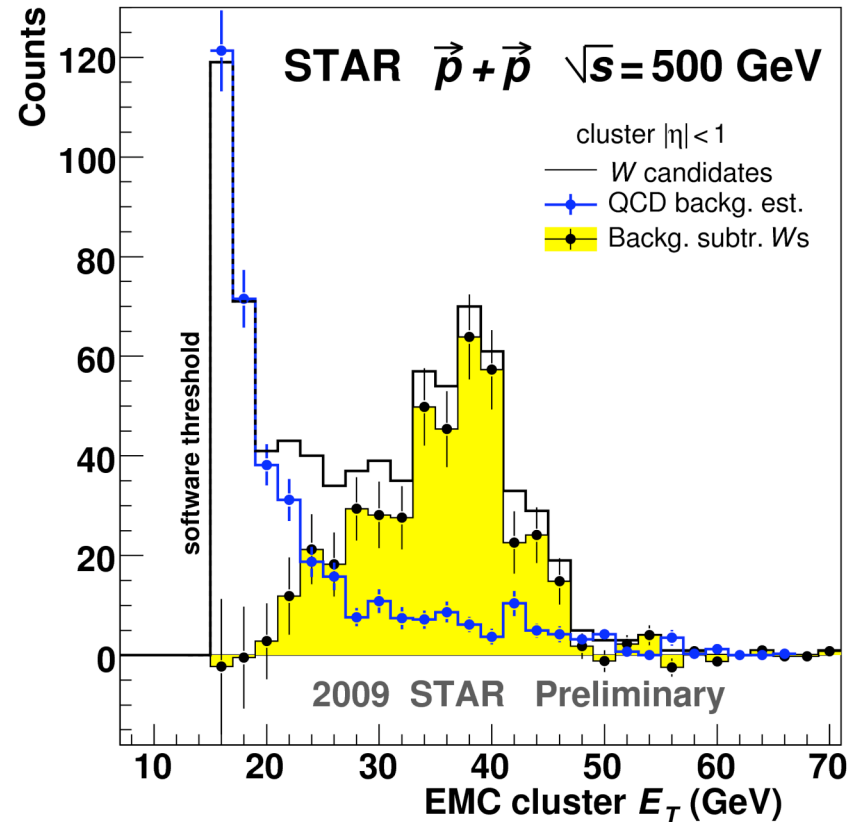
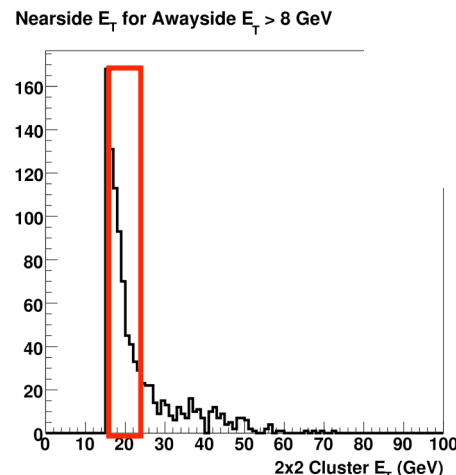
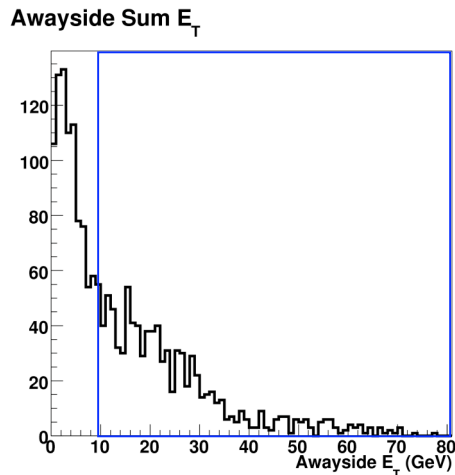
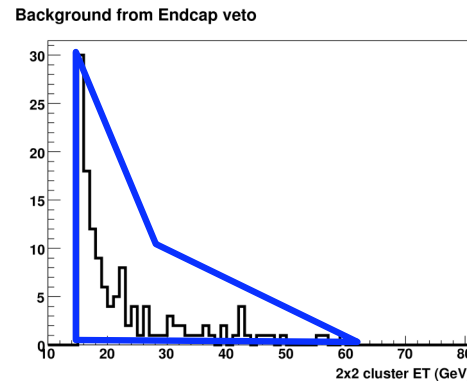
Disclaimer : PYTHIA has been known to not accurately predict the number of high-z fragmenting jets

First STAR W Results

Using the algorithm a clear Jacobian peak can be seen in the data

Estimate QCD background in a fully data-driven manner

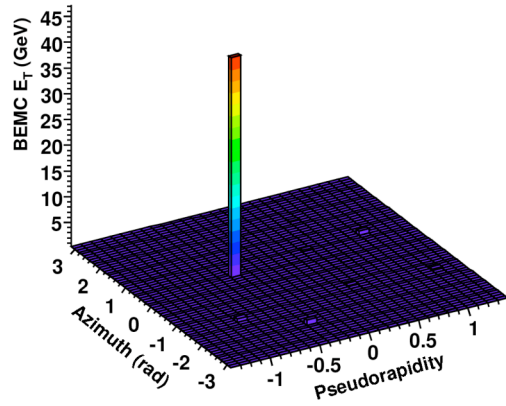
Background from "missing" endcap



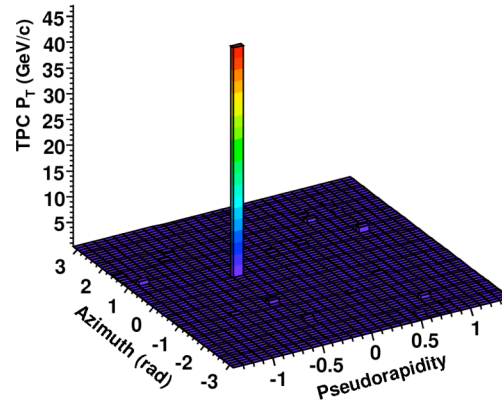
The normalized background is subtracted from the E_T spectrum to obtain a W signal.

Example Lego Plots

BEMC E_T Distribution (GeV)

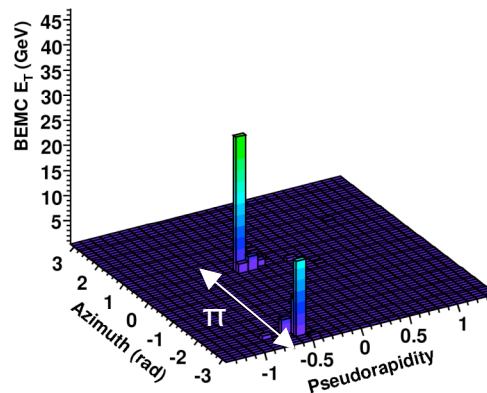


TPC p_T Distribution (GeV/c)

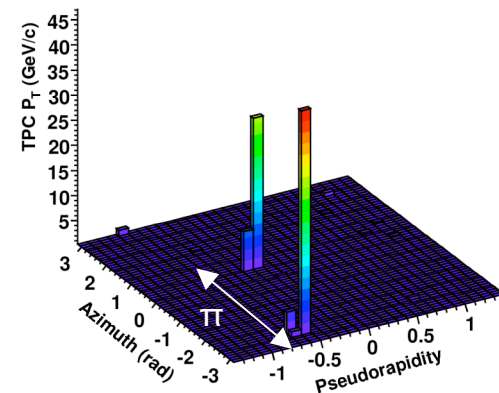


← W event

BEMC E_T Distribution (GeV)



TPC p_T Distribution (GeV/c)



Dijet event →

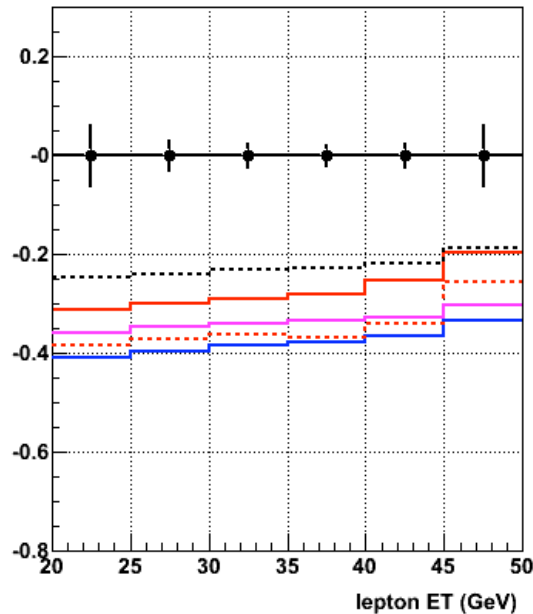


Future W Measurements

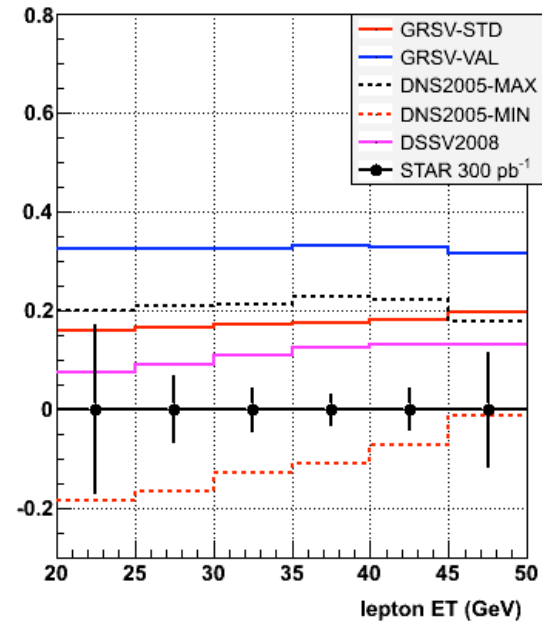
Future W s at mid-rapidity

STAR projections for $LT=300 \text{ pb}^{-1}$, $\text{Pol}=0.7$, $\text{effi}=70\%$, including QCD background, 2 beams, no vertex cut

$A_L(W^+)$ for positron $|\eta| < 1$



$A_L(W^-)$ for electron $|\eta| < 1$



STAR has shown the capability to detect the W at mid-rapidity.

With the expected 300 pb^{-1} for the 500 GeV program STAR will provide strong constraints on the polarized sea pdfs using the mid-rapidity data

Future W s at forward rapidity

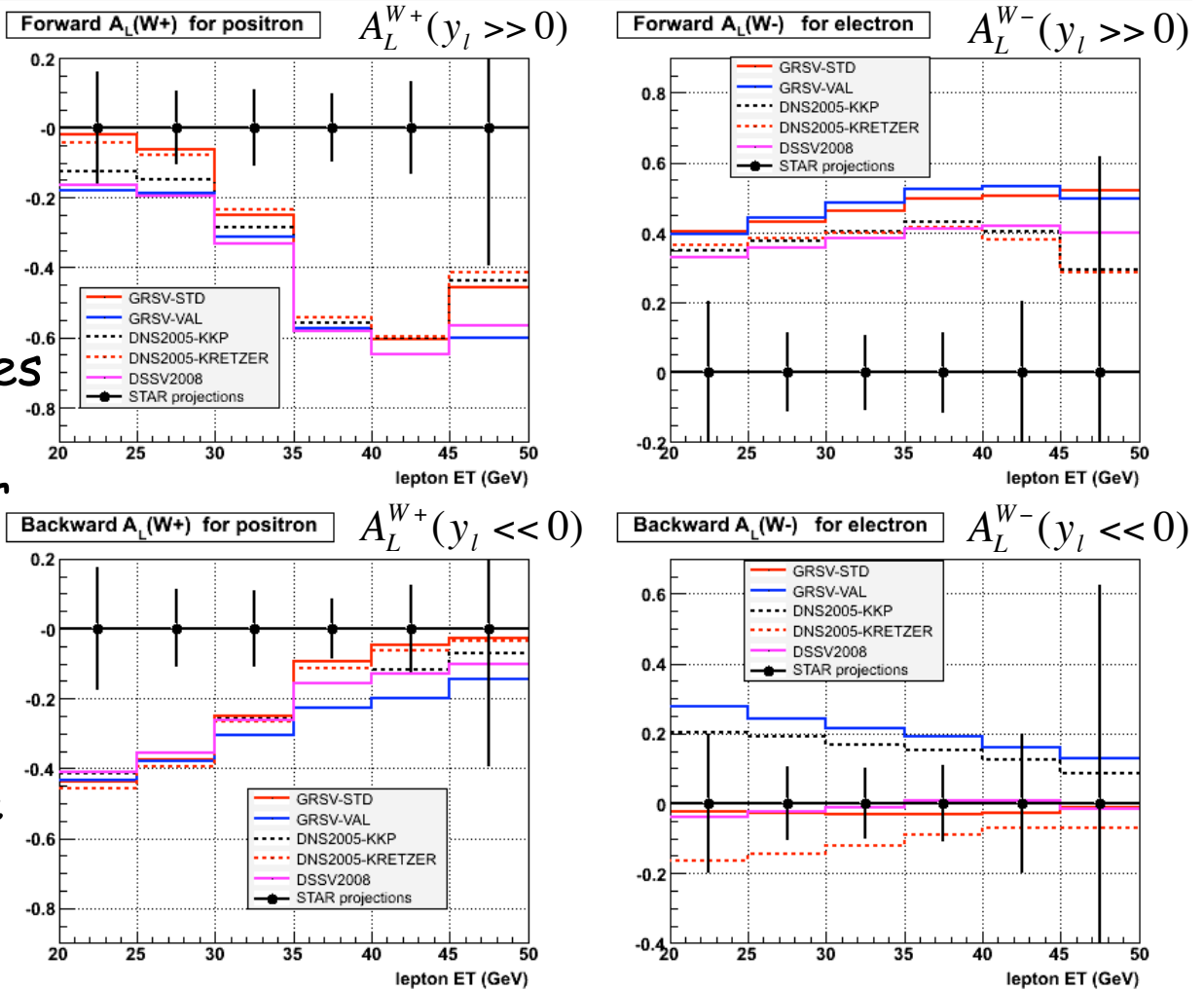
At forward/backward rapidity the A_L s become more sensitive to a single quark flavor

The expected uncertainties for the 500 GeV program are shown to the right for the endcap acceptance

$$1 < \eta < 2$$

Measurements rely on the planned Forward GEM Tracker upgrade (see B. Surrow's talk on Sat.)

STAR projections for $LT=300 \text{ pb}^{-1}$, $\text{Pol}=0.7$, including QCD background and detector effects



Conclusions

- Measurements of the W in polarized p+p collisions provide needed information about the polarized sea in the proton.
- STAR has shown a first extraction of a Jacobian peak of a W signal in p+p collisions at $\sqrt{s}=500$ GeV confirming our expectations.
- Planned STAR measurements will provide strong constraints on the polarized sea of the proton.

Backup Slides