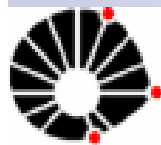




# Thermal Production of particles at RHIC (Test of Chemical Freeze-out at RHIC)

Jun Takahashi for the STAR collaboration  
SQM2008, Beijing, China

高橋純



UNICAMP



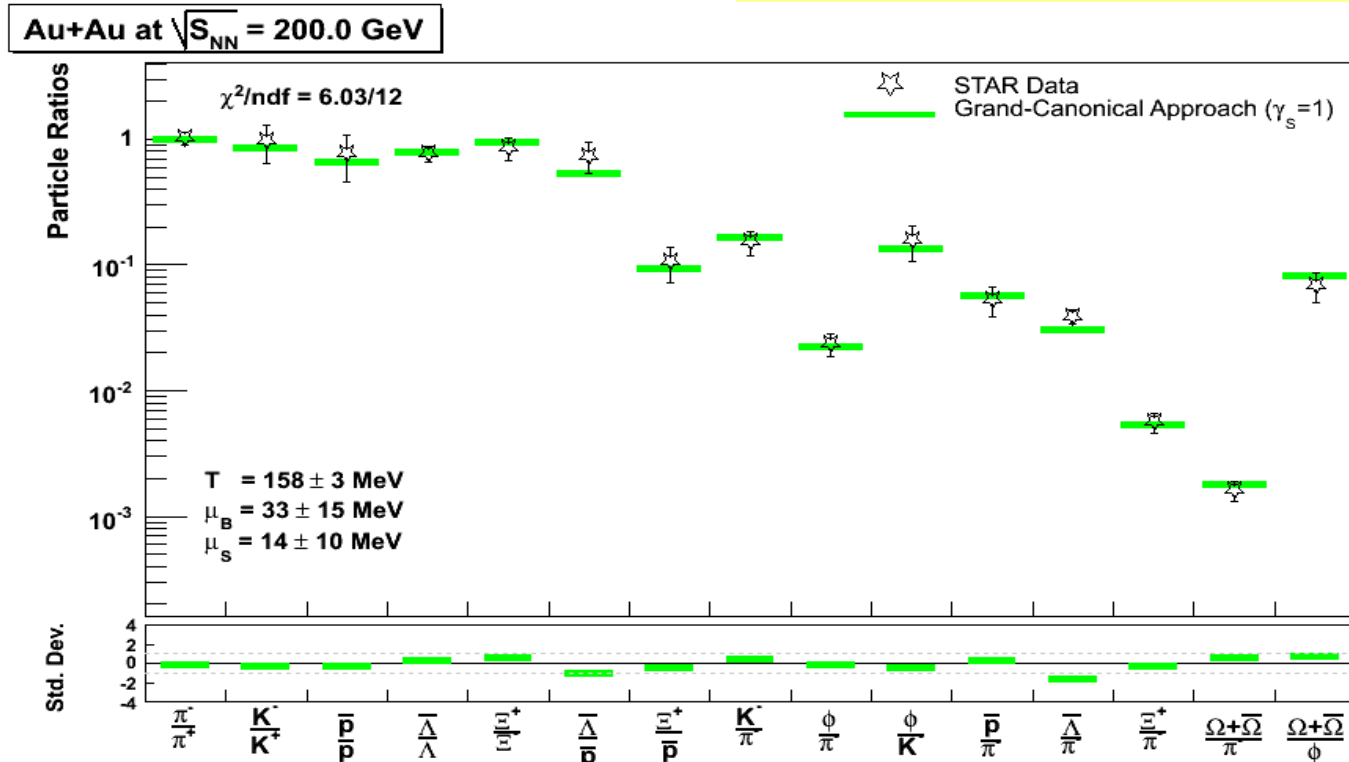
## Introduction:

- Strangeness chemical equilibration is achieved at RHIC, but ...
  - To what extent?
  - Does it vary with system size?
  - Can we disentangle the canonical suppression?
  - What about the rapidity range?
  
- STAR has a large amount of data, with different particles, at different centralities and in different collision systems.

# Statistical Thermal Model

- Statistical Thermal Model (**THERMUS**)<sup>\*</sup> was used fitting  $T_{ch}$ ,  $\mu_B$ ,  $\mu_S$ , and  $\gamma_S$  (strangeness saturation factor).
- Particles used in the fit:  
 $\pi$ ,  $K$ ,  $p$ ,  $\Lambda$ ,  $\Xi$ ,  $\Omega$  and  $\phi$ .

<sup>\*</sup> Thermus, A thermal Model Package for Root  
 S. Wheaton & Cleymans, hep-ph/0407174



# Feed-down corrections:

- Pion yields
  - Weak decay feed-down contributions are subtracted.
- Proton yields
  - $\Lambda$  decay feed-down is subtracted, considering inclusive  $\Lambda$ 's.
  - $\Sigma$  decay feed-down needs to be studied.

## ■ Lambda yields

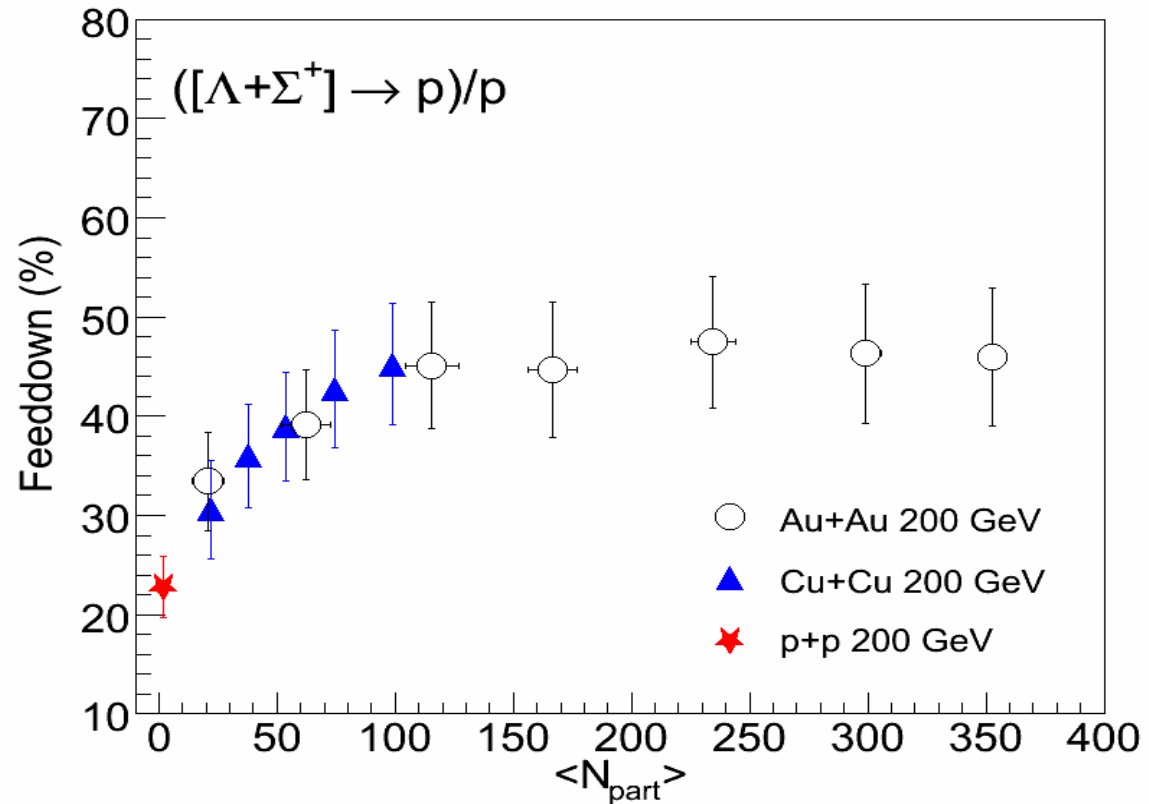
- $\Xi$  decay feed-down is subtracted.
- $\Omega$  decay is negligible. ( $\Omega/\Lambda \sim 0.01$ )

## ■ $\Xi$ yields

- $\Omega$  decay feed-down is negligible. BR 8.6 %

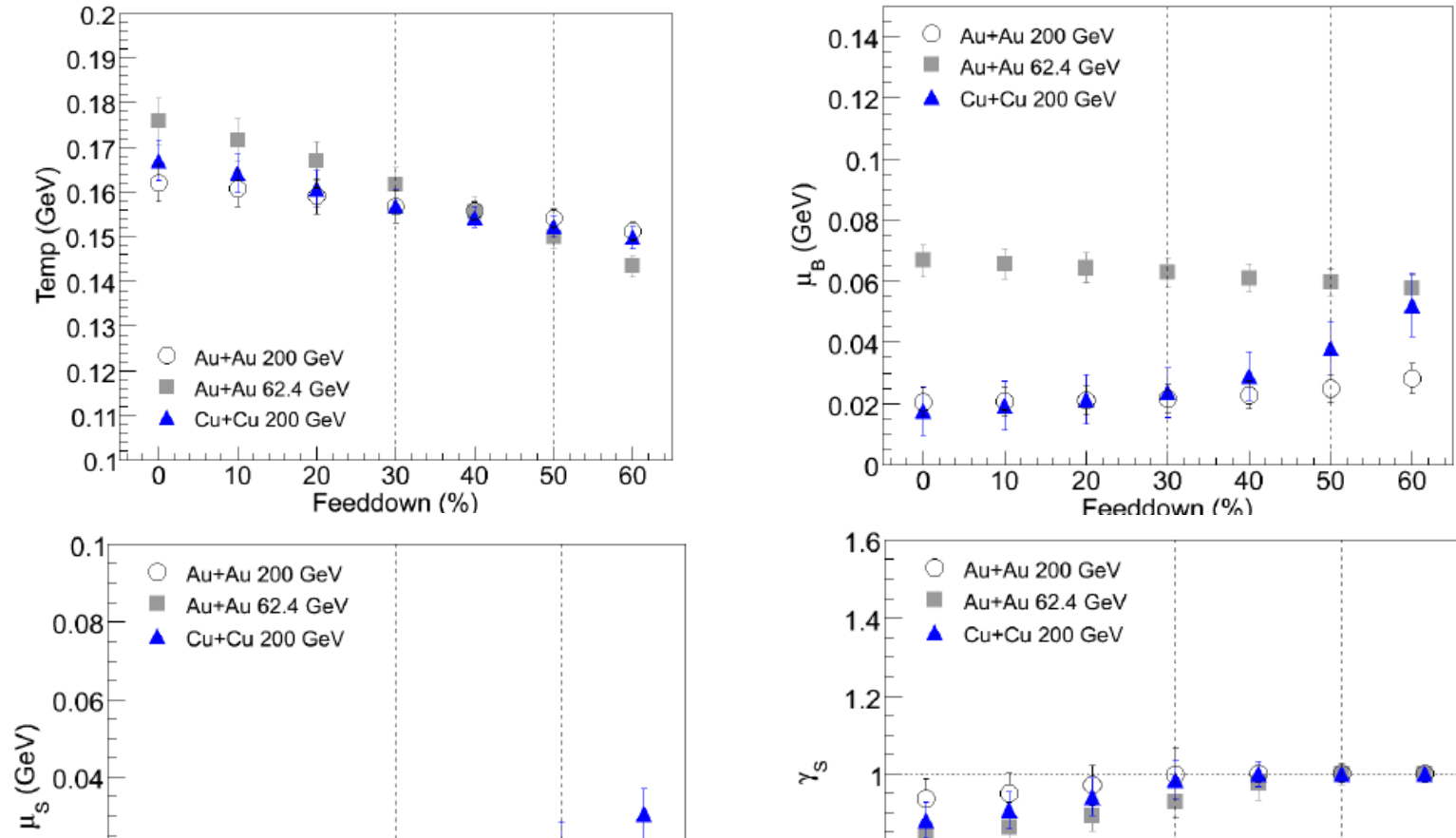
			<u>ng Ratio</u>
$\Xi^-$	$\rightarrow \Lambda + \pi^-$	4,91 cm	$\pm 0,05\)%$
$\Xi^0$	$\rightarrow \Lambda + \pi^0$	8,71 cm	$\pm 0,30\)%$
$\Omega$	$\rightarrow \Lambda + K^-$	2,461 cm	$\pm 0,30\)%$
			$\pm 0,005\)%$

# Feed-down correction of protons from the Sigma



$$\left( \frac{N_{\text{total}}}{N_{\text{total}}} \right)$$
 Total correction to the protons is around 45%, due to Lambdas ( $\Lambda$ ) and Sigmas ( $\Sigma$ ).

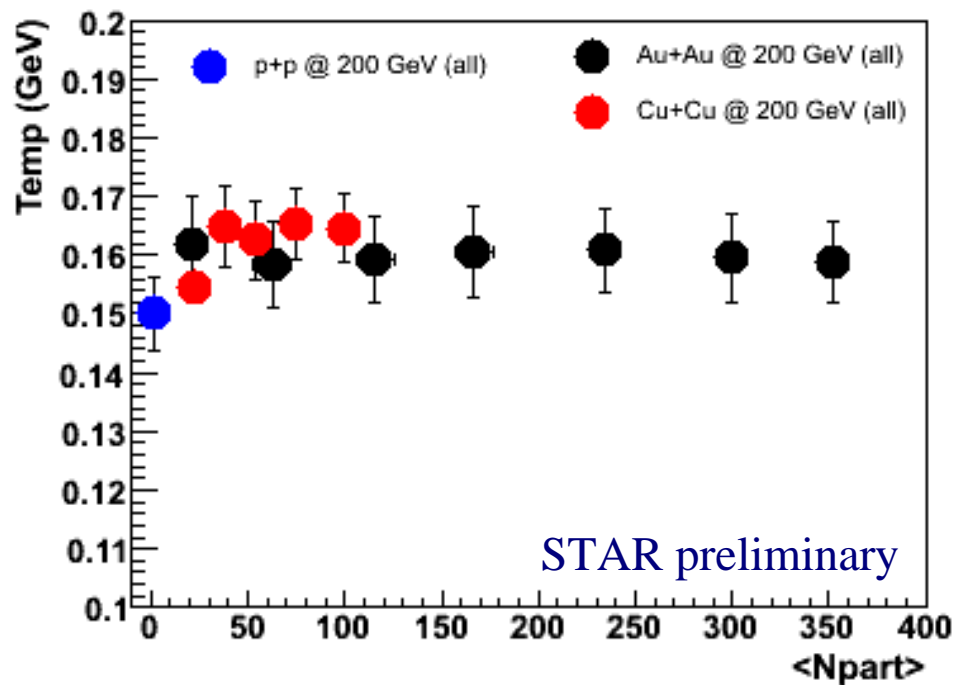
# Effect of feed-down on thermal parameters



Thermal parameters has small variation in the range consistent with the feed-down correction uncertainty of the protons.

$$\Delta T = 5 \text{ MeV (3\%)}, \Delta \mu_B = 0.005, \Delta \mu_S = \text{negligible}, \gamma_S = 1$$

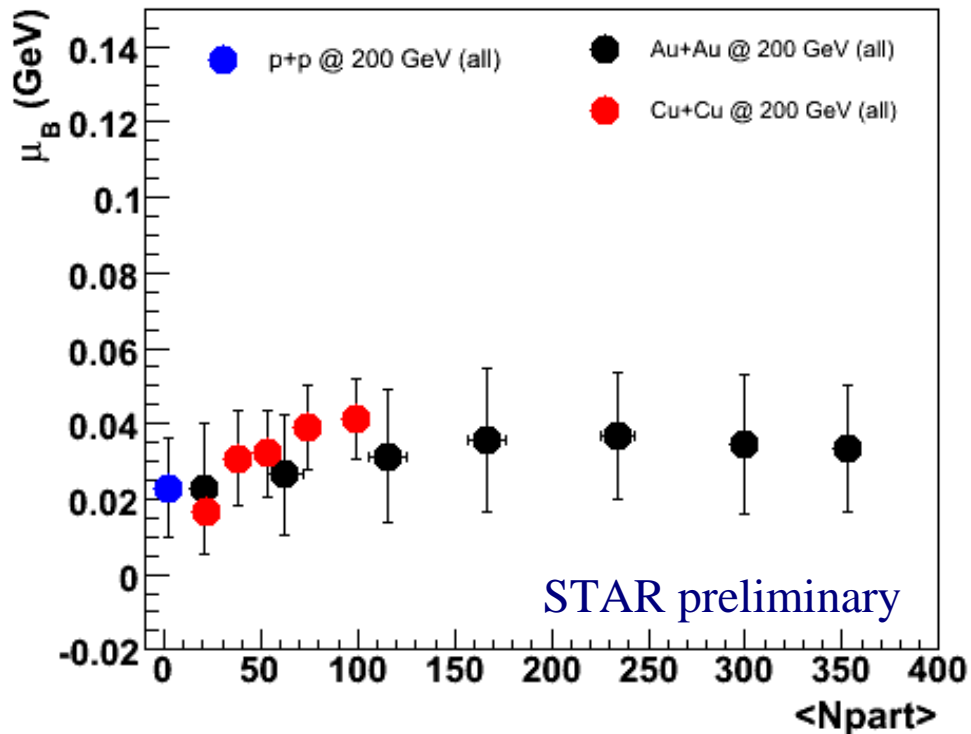
# Chemical freeze-out $T_{ch}$ vs. system size



- Au+Au 200 GeV  
Temperature seems constant with system size.
- Cu+Cu 200 GeV  
Temperature is in agreement with Au+Au 200 GeV (within error bars).

Thermus, a thermal Model Package for Root  
S. Wheaton & Cleymans, hep-ph/0407174

# Baryon chemical potential $\mu_B$ vs. system size

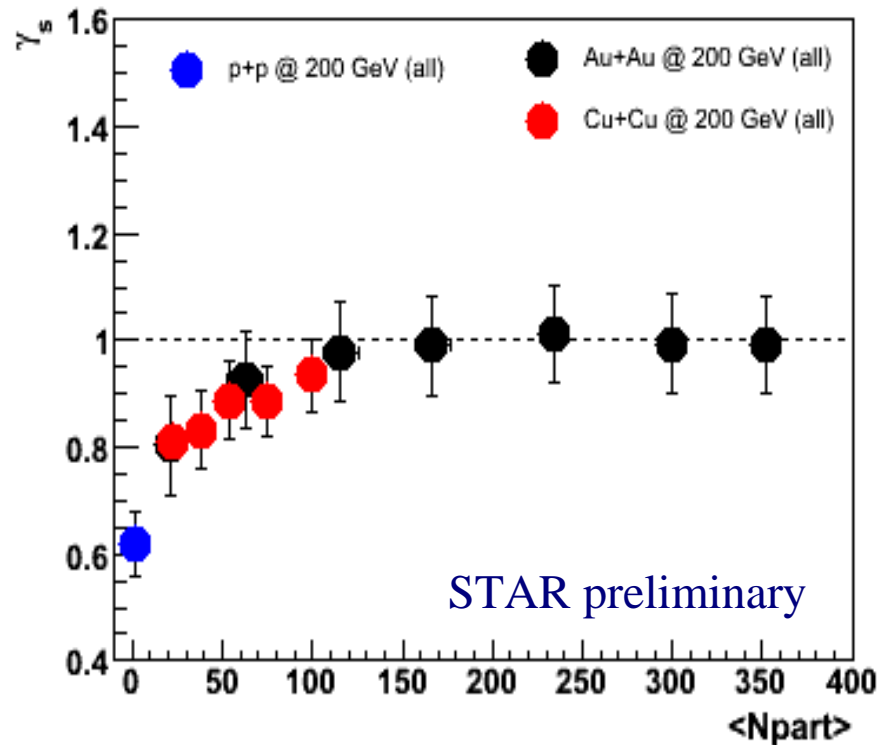


- Baryon chemical potential  $\mu_B$  is small.
- Small variation with system size.
- Cu+Cu 200 GeV baryon chemical potential seems to be in good agreement with Au+Au 200 GeV.

\* Thermus, a thermal Model Package for Root  
 S. Wheaton & Cleymans, hep-ph/0407174



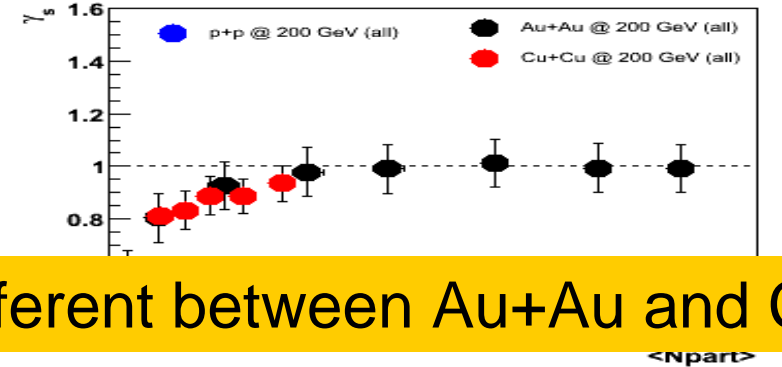
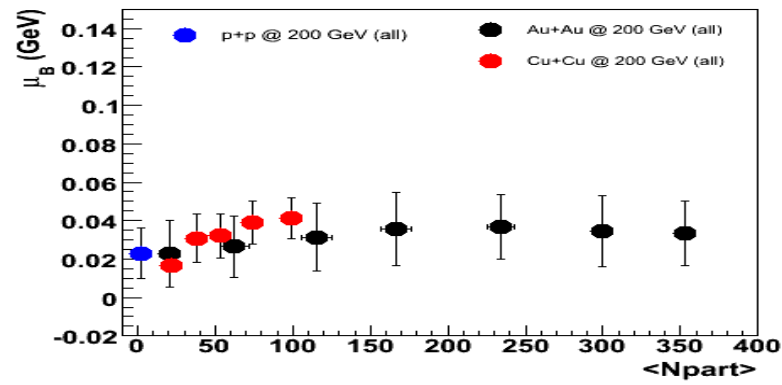
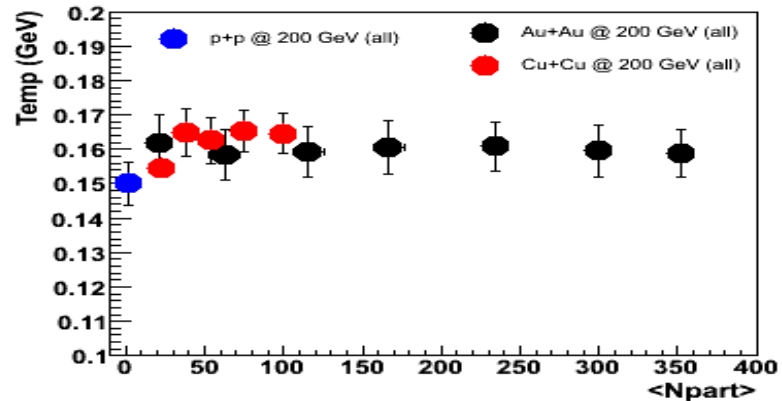
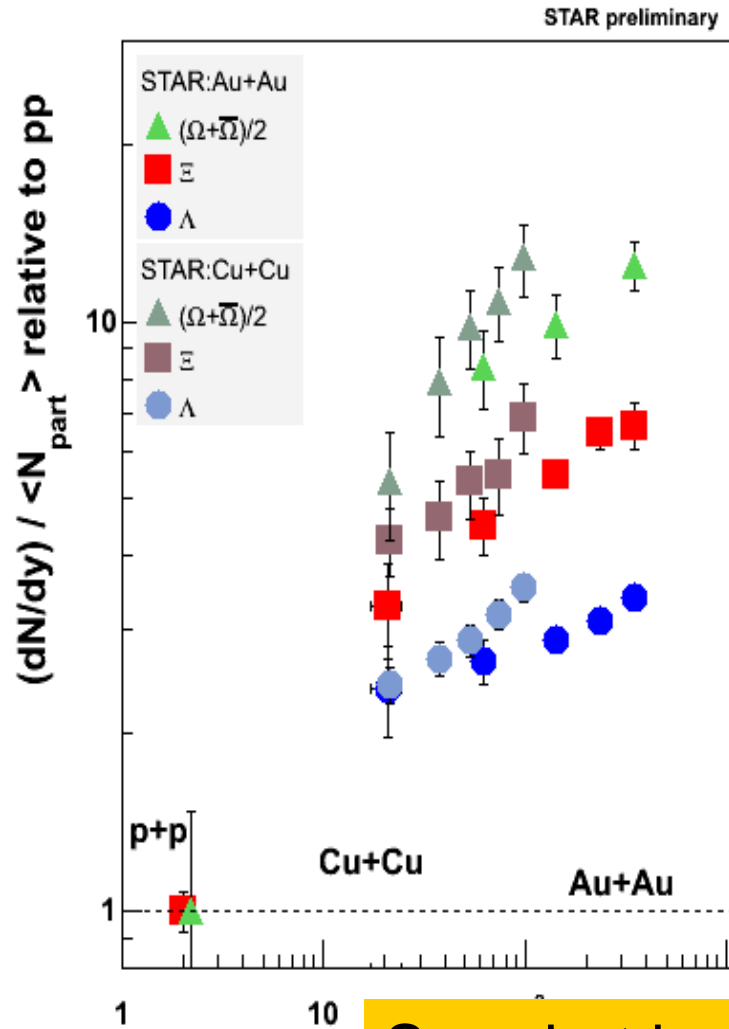
# Strangeness saturation $\gamma_s$ vs. system size



- Strangeness saturation constant, shows an increase with system size, reflecting the increase of strangeness enhancement.
- Cu+Cu 200 GeV data consistent with Au+Au 200 GeV result.
- Deviates from 1 for systems smaller than  $N_{part} < 100$ .

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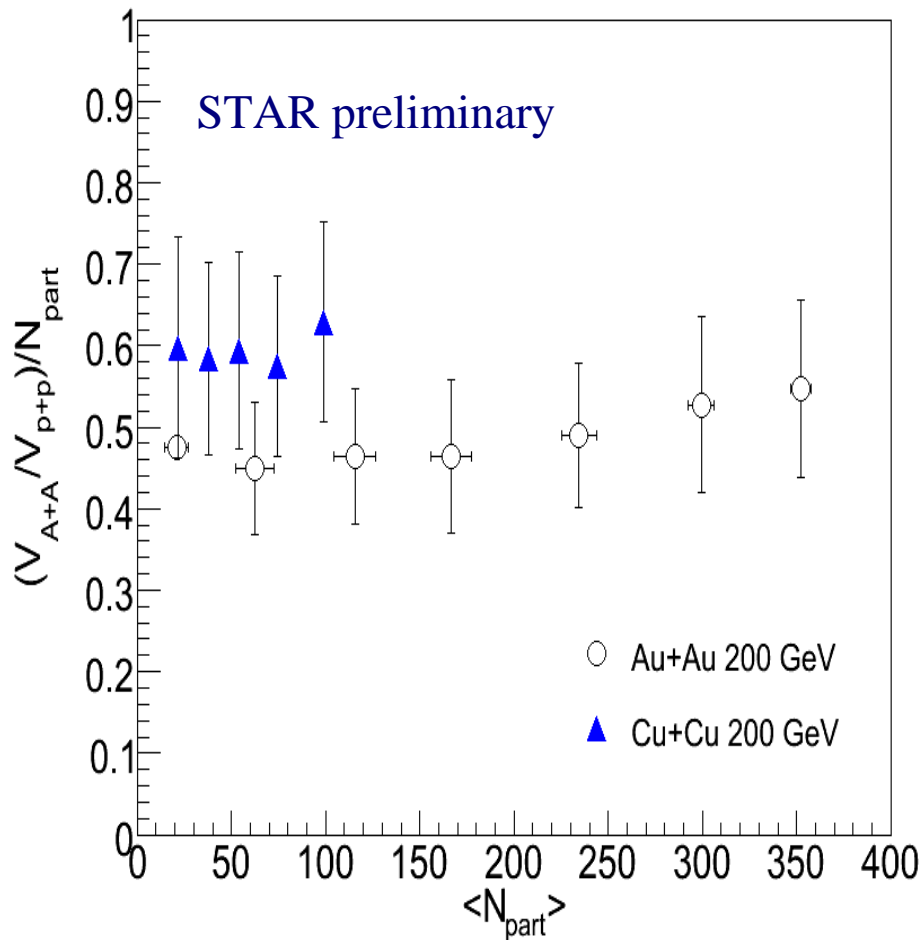
# Cu+Cu vs Au+Au



So, what is different between Au+Au and Cu+Cu?

From A. Timmins talk on monday

# Chemical freeze-out volume vs. system size

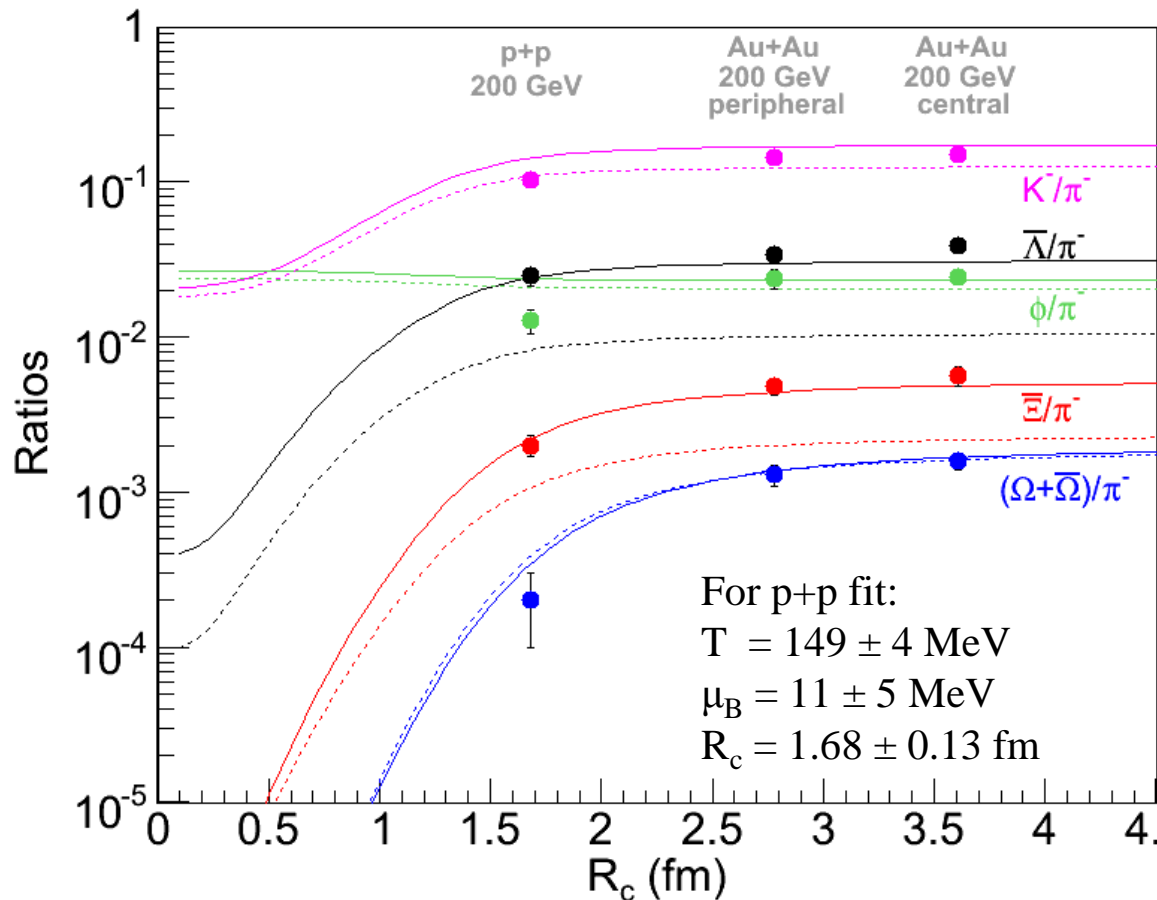


- Volume at Chemical freeze-out was determined using pion yields.
- Relative volume of the fireball at chemical freeze-out in Cu+Cu collision is higher than in Au+Au collision, for the same equivalent  $N_{part}$ .
- Higher strange particle yields observed in Cu+Cu compared to Au+Au (Ant. Timmins talk) is related to the volume at CF, so not to  $N_{part}$ .

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S. Wheaton & Cleymans, hep-ph/0407174

# Canonical Suppression effect

- How does strangeness production get affected by the canonical radius?
- Can we tell we have indication of canonical suppression?

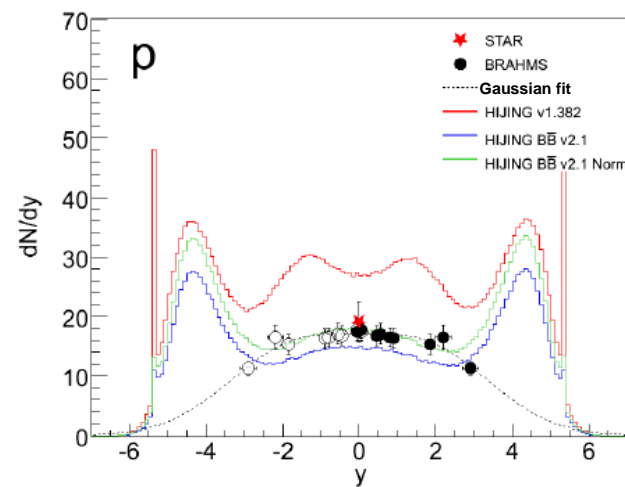
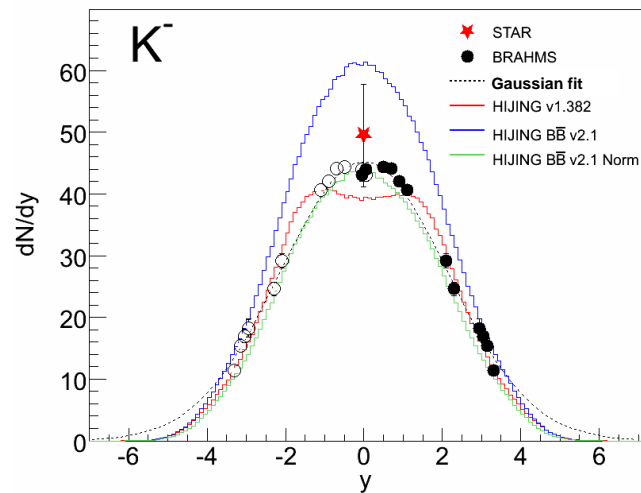
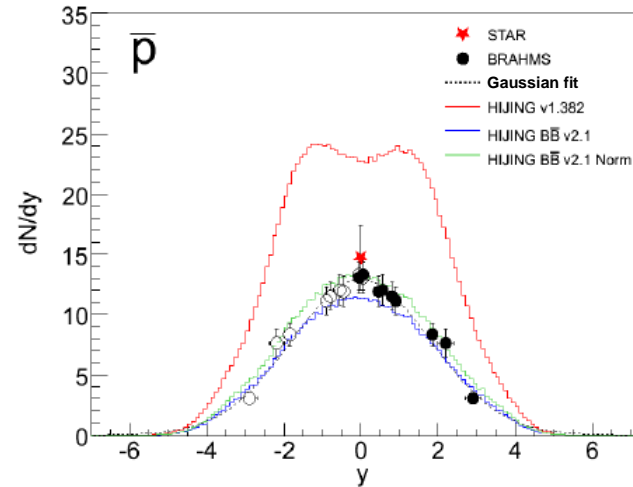
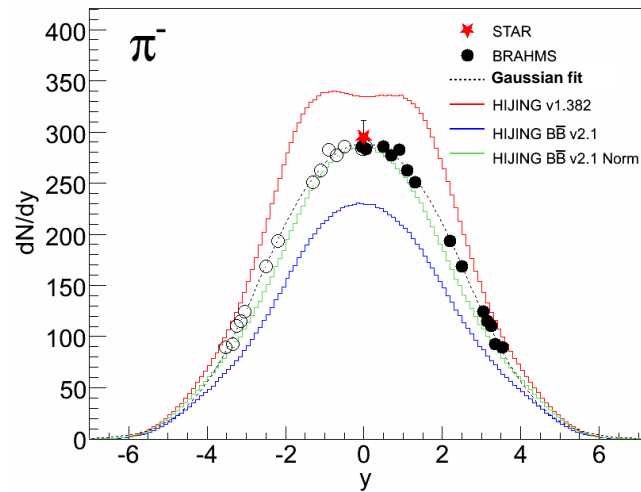


Data of most peripheral Au+Au collision is already in the saturated region, thus, not subject to canonical suppression !!!

It seems that in Strangeness Enhancement scenario, canonical suppression can only account for the initial increase from pp to most peripheral AuAu for open strange particles.

SPS:  $R_c \approx 1.4$  (I. Kraus & K. Redlich [hep-ph/0604242](https://arxiv.org/abs/hep-ph/0604242))

# Rapidity distributions

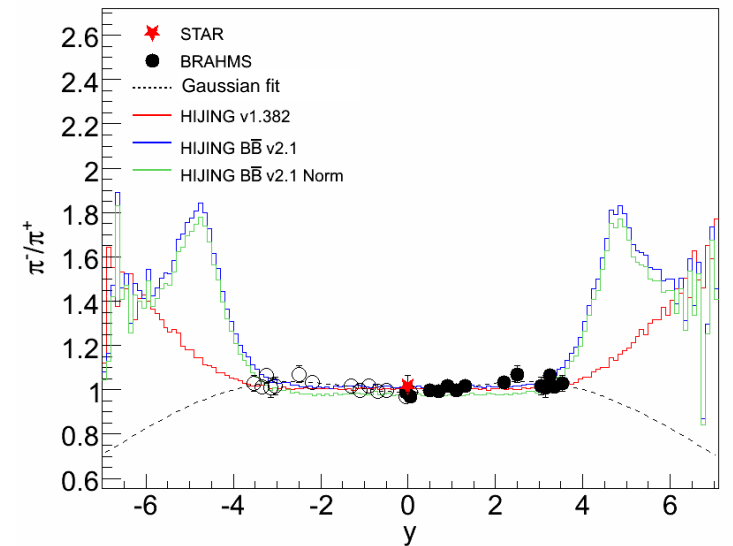
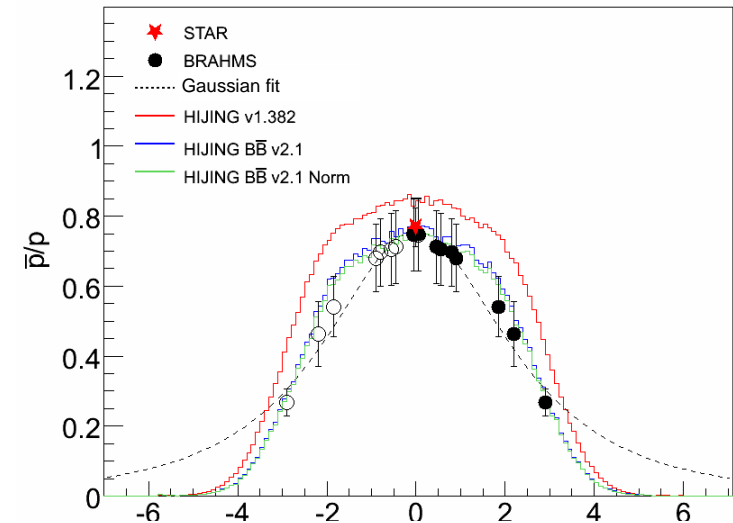
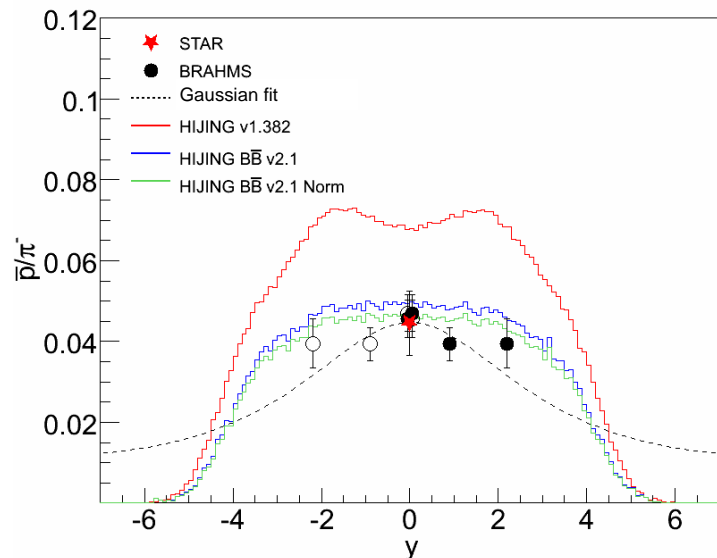


BRAHMS data: PRL94\_162301 (2005) & PRL93\_102301 (2004)

Hijing: X.N. Wang and M. Gyulassy: Comput. Phys. Commun., 83:307, 1994.

# Ratios vs. Rapidity

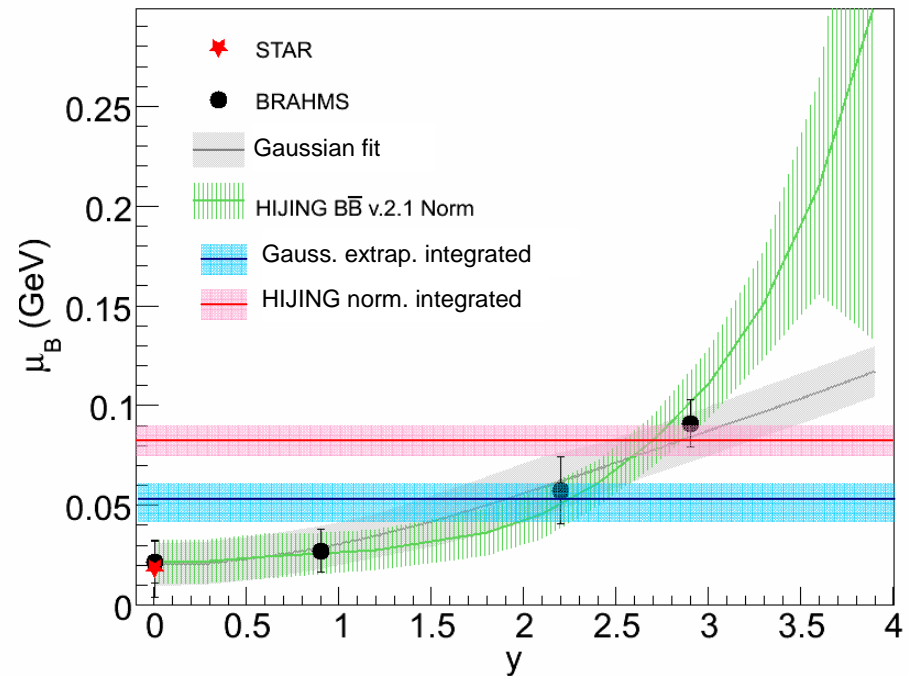
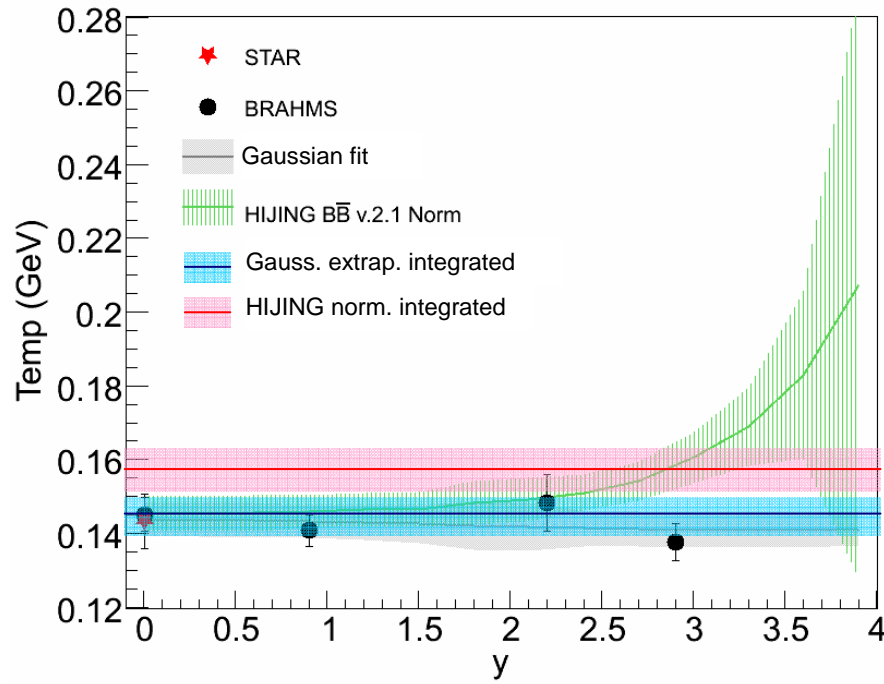
- Hijing and Gaussian distributions describe quite well the particle ratio rapidity dependence for  $|y| < 3.5$ .
- But, different parameterizations disagree for  $|y| > 4$ .



BRAHMS data: PRL94\_162301 (2005) & PRL93\_102301 (2004)

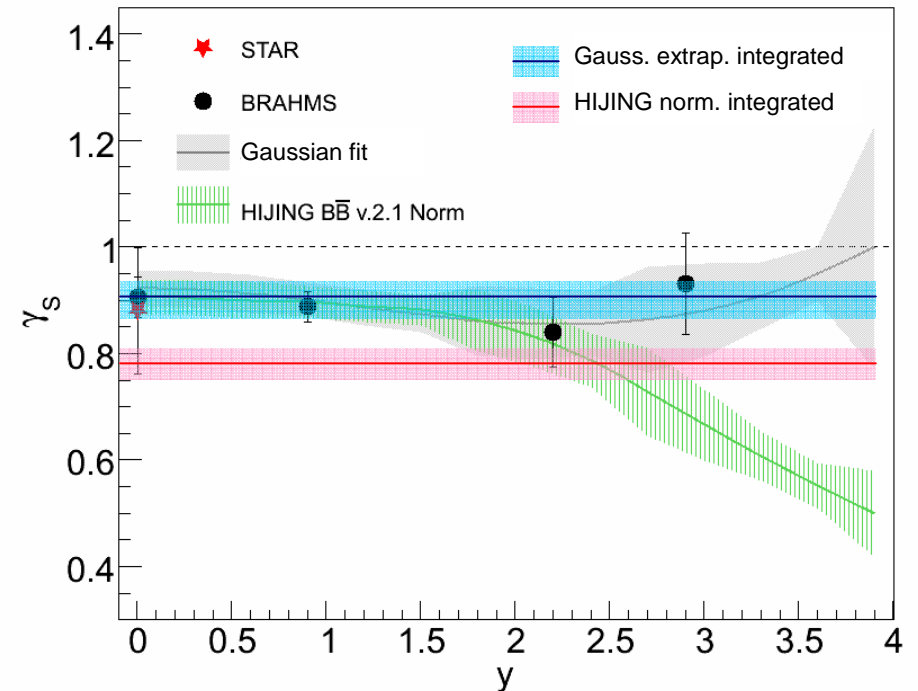
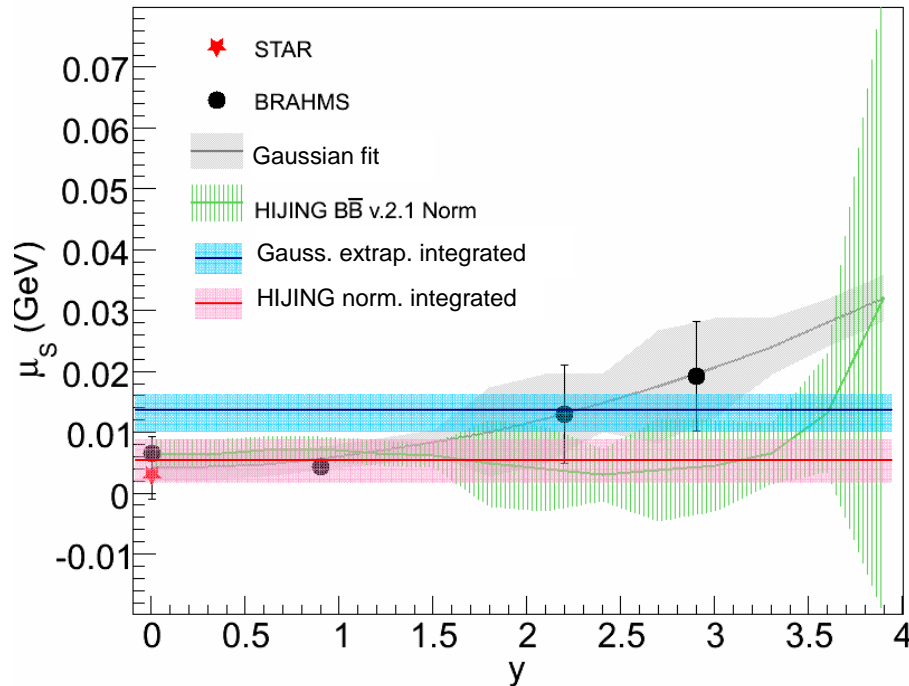
Hijing: X.N. Wang and M. Gyulassy: Comput. Phys. Commun., 83:307, 1994.

# Thermal fit results vs. Rapidity



- Temperature constant up to  $y=3$ .
- $\mu_B$  shows increase with rapidity.
- For rapidity integr. yields using Gauss:  $T=144\pm 8$  MeV,  $\mu_B=53\pm 16$  MeV
- For rapidity integr. yields using Hijing:  $T=156\pm 8$  MeV,  $\mu_B=79\pm 19$  MeV

# Thermal fit results vs. Rapidity



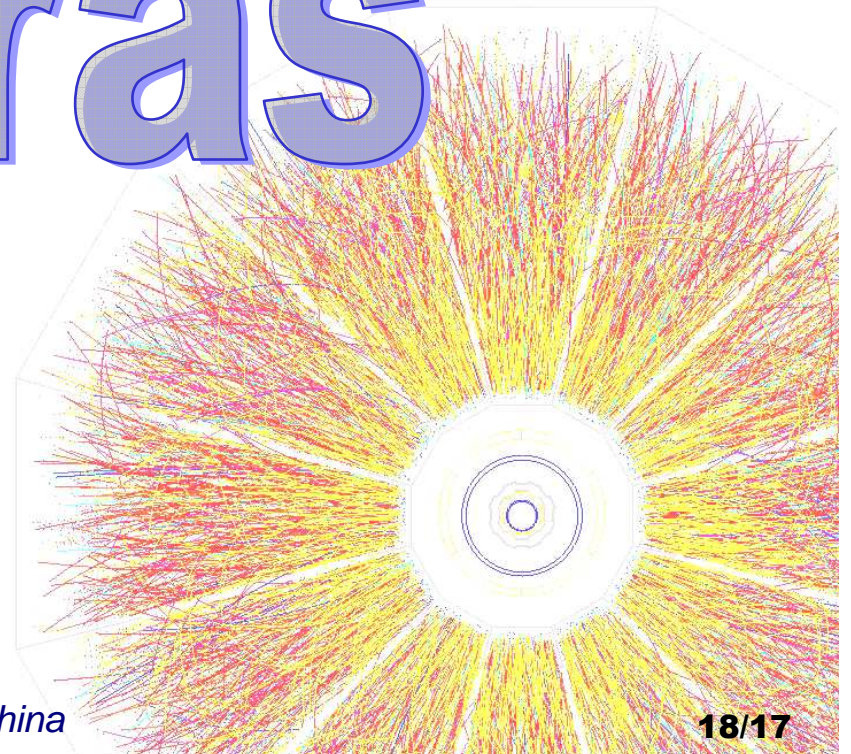
- $\mu_S$  shows increase with rapidity, Hijing extrapolation shows no variation with rapidity in the studied range.
- $\gamma_S$  constant with rapidity, up to  $y=3$ .
- For rapidity integr. yields using Gauss:  $\mu_S=14\pm 4$  MeV,  $\gamma_S=0.91 \pm 0.12$
- For rapidity integr. yields using Hijing:  $\mu_S=7\pm 4$  MeV,  $\gamma_S=0.79 \pm 0.12$



## Conclusions:

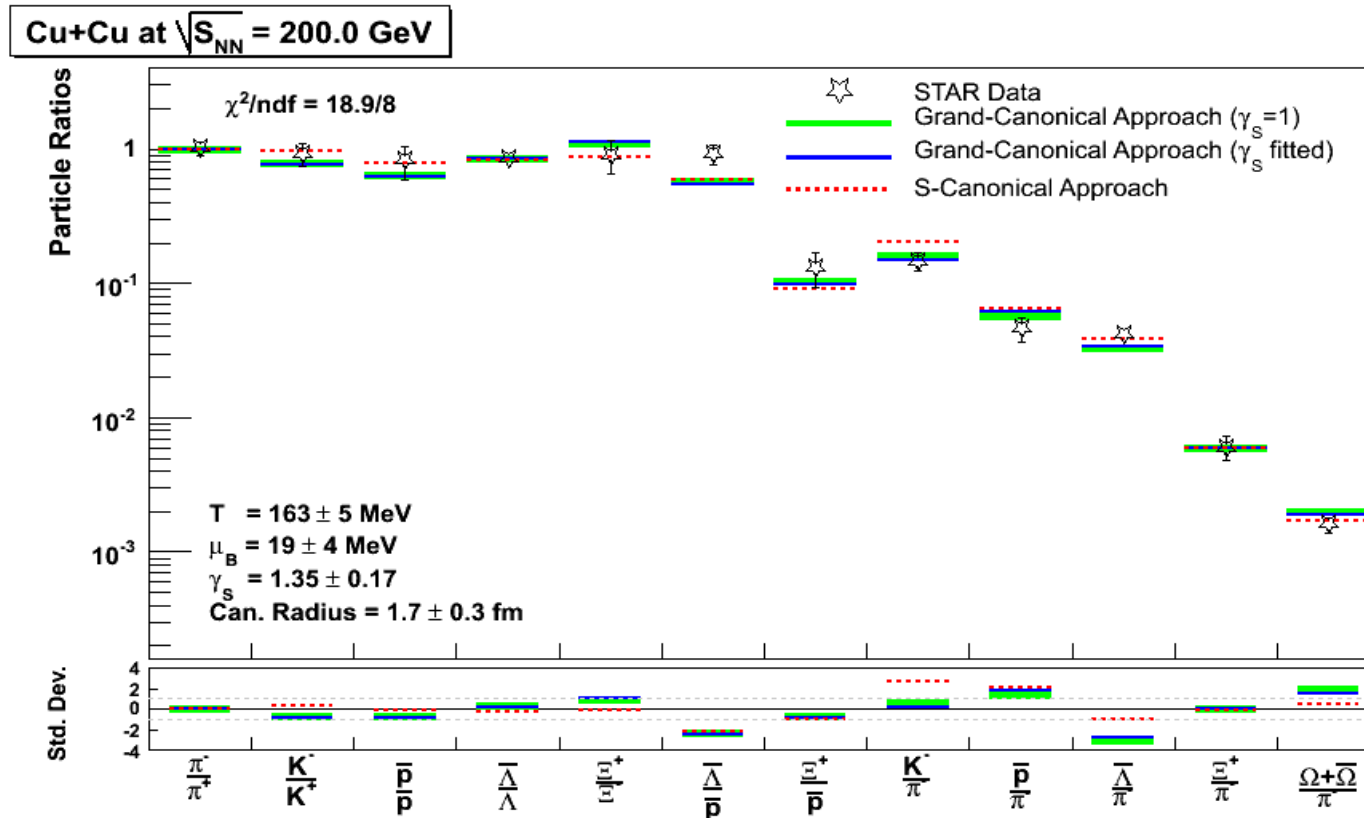
- **Strangeness is chemically equilibrated at RHIC energies.**
- **Thermal parameters are constant for a wide range:**
  - **Centrality:** System with  $N_{\text{part}} > 100$  is well described using GC approach with  $\gamma_s = 1$ .
  - **Rapidity:** at RHIC energies, thermal fit Temperature constant up to  $Y=3$ ;  $\mu_B$  shows increase with rapidity.
- **Canonical** approach shows strong increase of strange particles with  $R_c$ , except for the phi-mesons. **Fits to Au+Au data results in  $R_c$  values where canonical suppression is no longer relevant.**
- **Cu+Cu and Au+Au:** Results of thermal model fits indicate that the **freeze-out volume formed in Cu+Cu collision is higher than that in Au+Au collisions** with the same initial condition.

# Extras

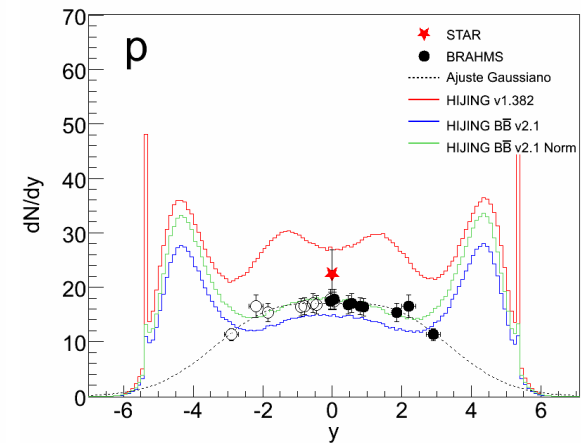
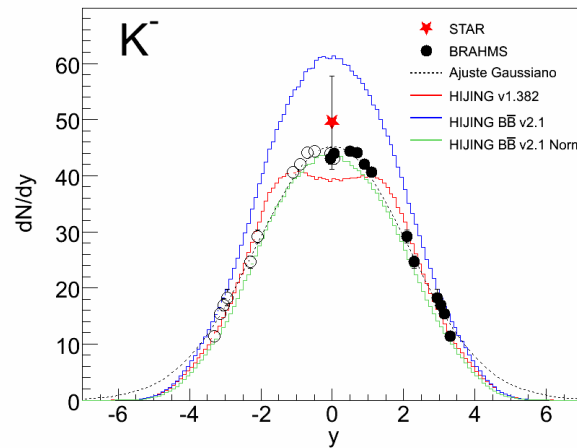
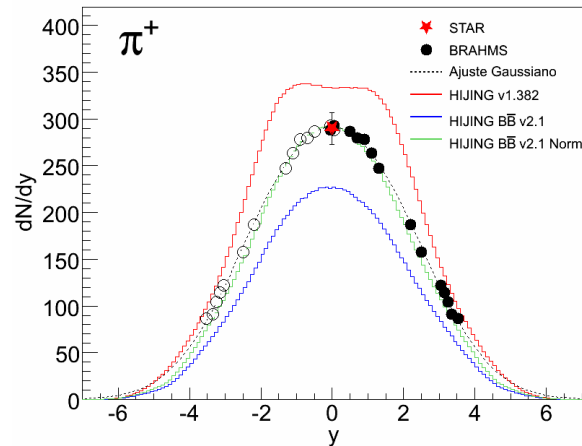
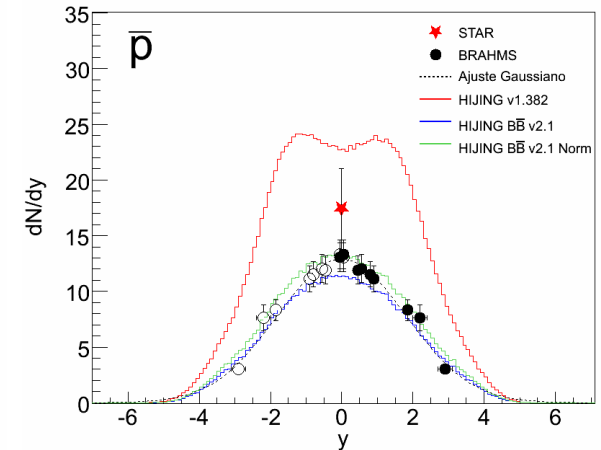
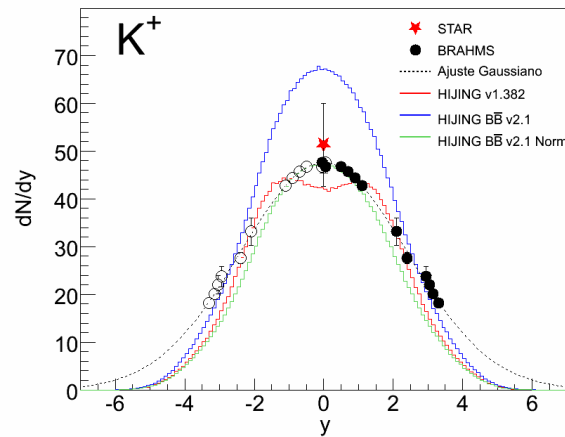
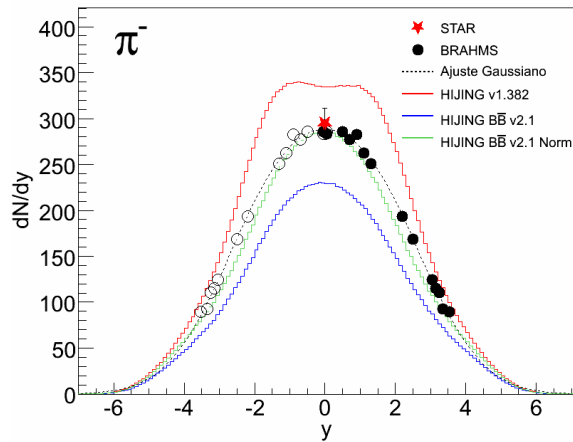


# Comparing different thermal model conditions: Central Cu+Cu collision

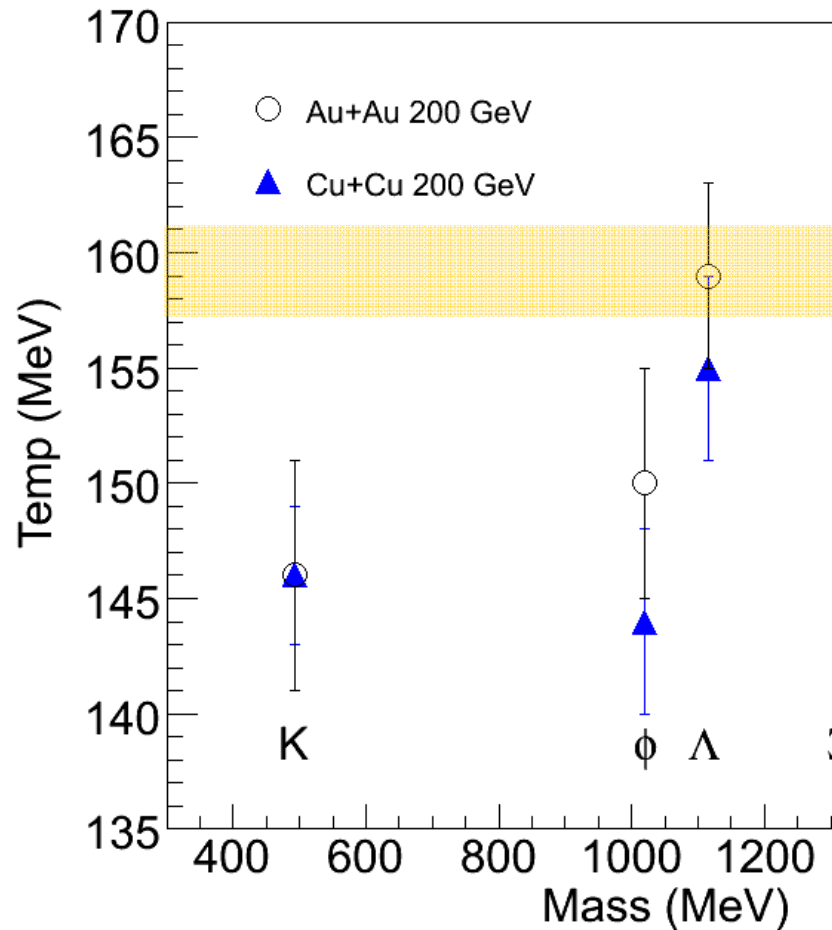
- Particles ratios still well described with GC ensemble.
- Canonical ensemble describes identical particle ratios, Lambdas and Omega better.



# Rapidity distributions

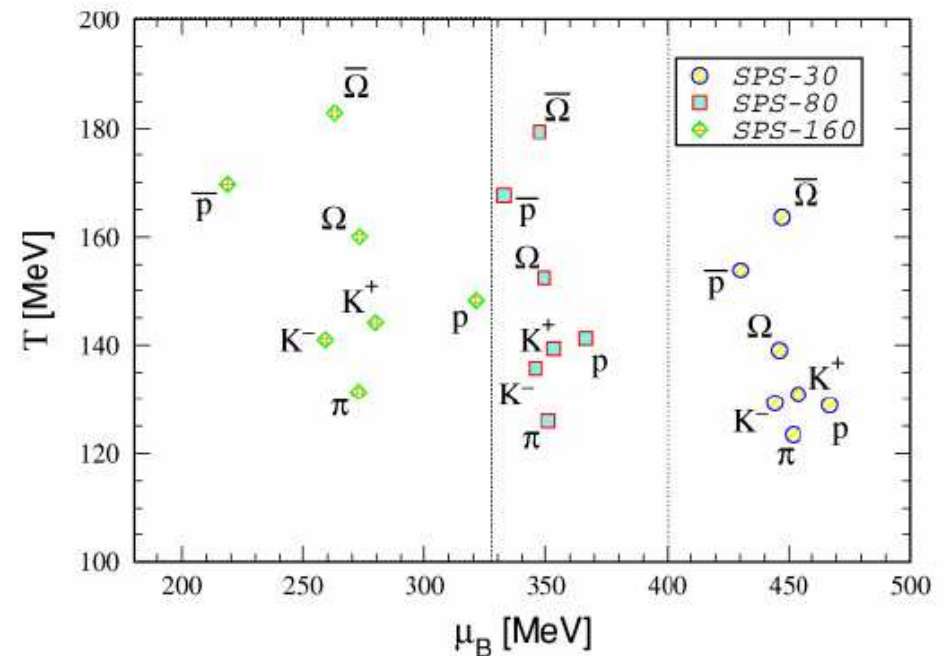


# Varying the strange particle used in the thermal fit



Ratios built with:

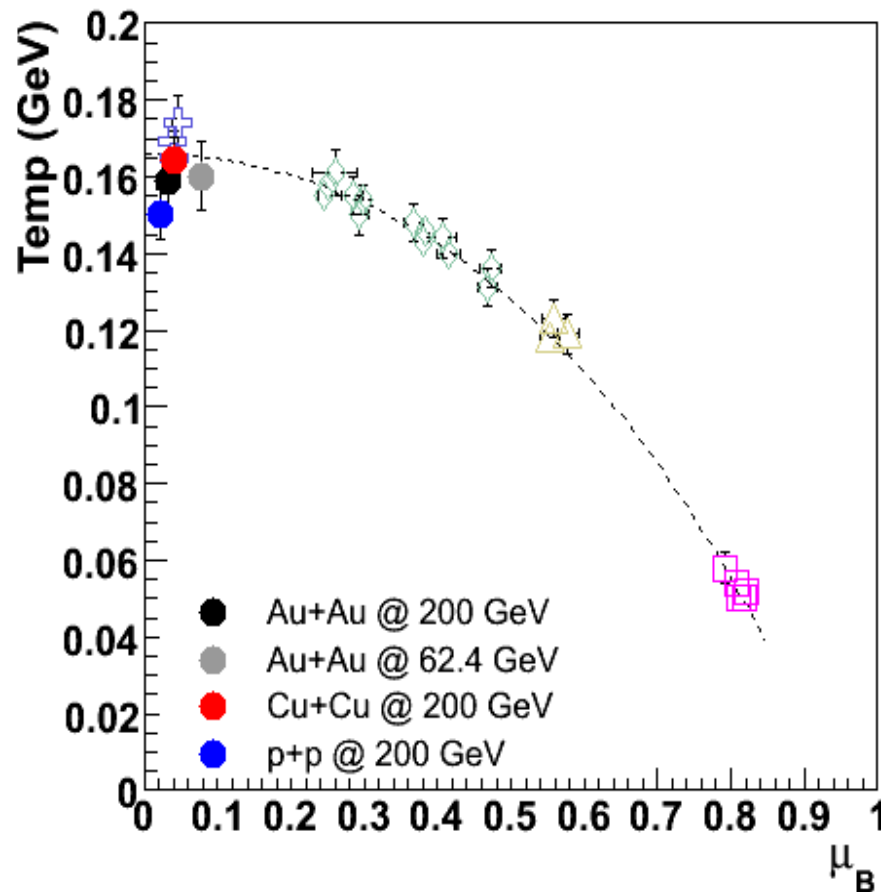
$$p, \bar{p}, \pi^-, \pi^+, S, \bar{S}$$



arXiv:nucl-th/0612033v1: Local and Global strangeness inhomogeneities at freeze-out conditions.  
 PRC\_73\_024902: Inhomogeneous freeze-out in relativistic heavy ion collisions.

# Summary of thermal model study

- **Statistical Thermal** model fits reasonably well the particle ratios measured in STAR, indicating that data is consistent with a thermalized system.



- Au+Au 200 and 62 GeV: Centrality dependence of thermal fits show increase of  $\gamma_s$  parameter, consistent with strangeness enhancement seen in the data yields.
- Cu+Cu 200 GeV: yields the same temperature and baryon chemical potential values obtained from the fit to Au+Au data.
- Strangeness Canonical approach seems to yield better agreement with our data.