

UPC 2025: The Second International Workshop on the Physics of Ultra Peripheral Collisions

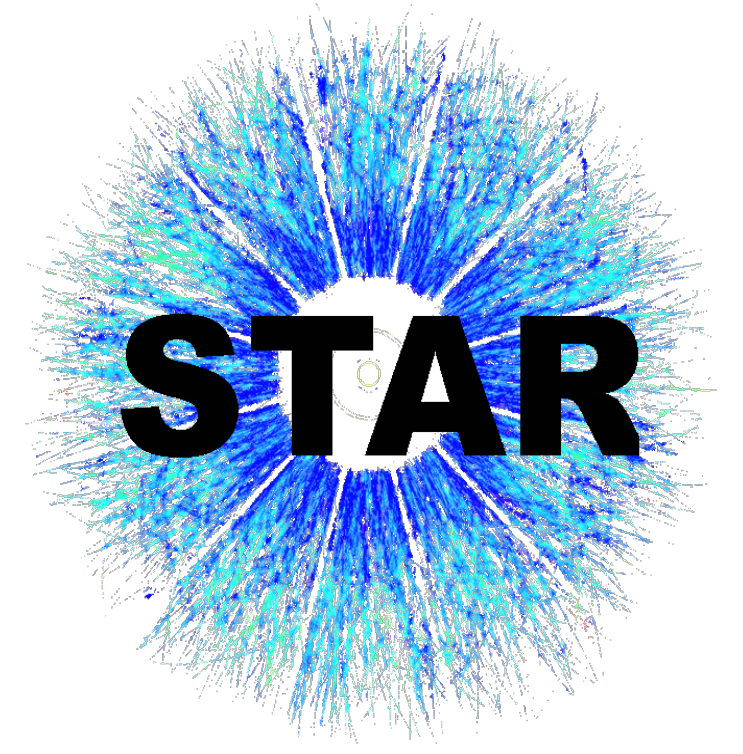
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DAVID TLUSTY (CREIGHTON UNIVERSITY)
FOR THE STAR COLLABORATION

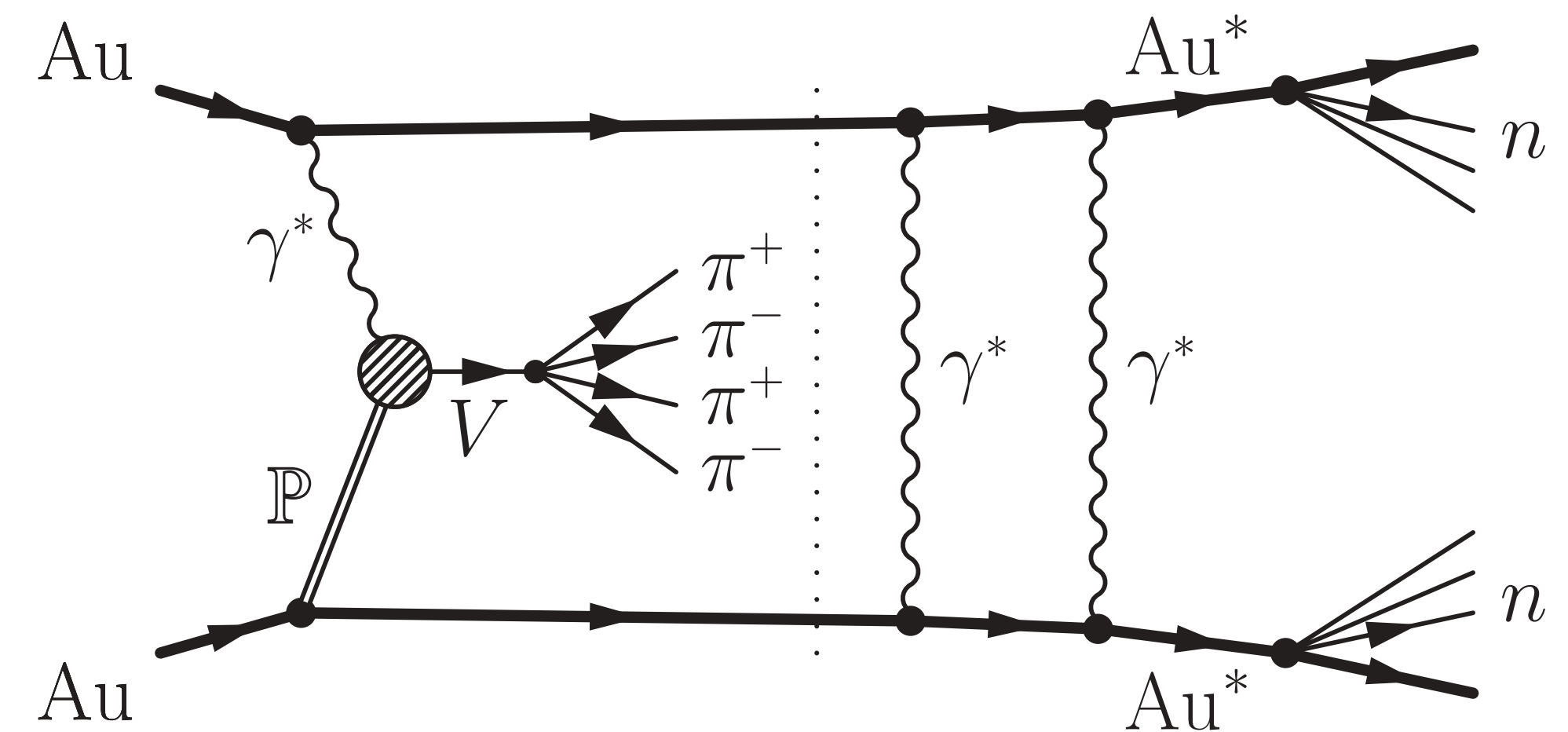
OBSERVATION OF $\pi^+\pi^-\pi^+\pi^-$ AND $\pi^+\pi^-$
FINAL STATE PHOTO PRODUCTION IN UPC AT
 $\sqrt{s_{NN}} = 200$ GEV AT THE STAR DETECTOR

MOTIVATION

- ▶ The first radial excitation 2^3S_1 of ρ_0 is considered to be the ρ_{1450} [PRD 110 030001], but decays suggest it is a hybrid state [PRD 56 1584]
- ▶ ρ_{1700} is assigned to 1^3D_1 state - there is need for precise measurement of mass and width to clarify its nature [PRD 110 030001]
- ▶ Questions of the ρ_{1450} relation to the ρ_{1700} have been raised
- ▶ The relativistic quark model [PRD 32 189] predicts 2^3D^1 state $J^{PC} = 1^{--}$ at 2.15 GeV which can be identified with the $\rho(2150)$

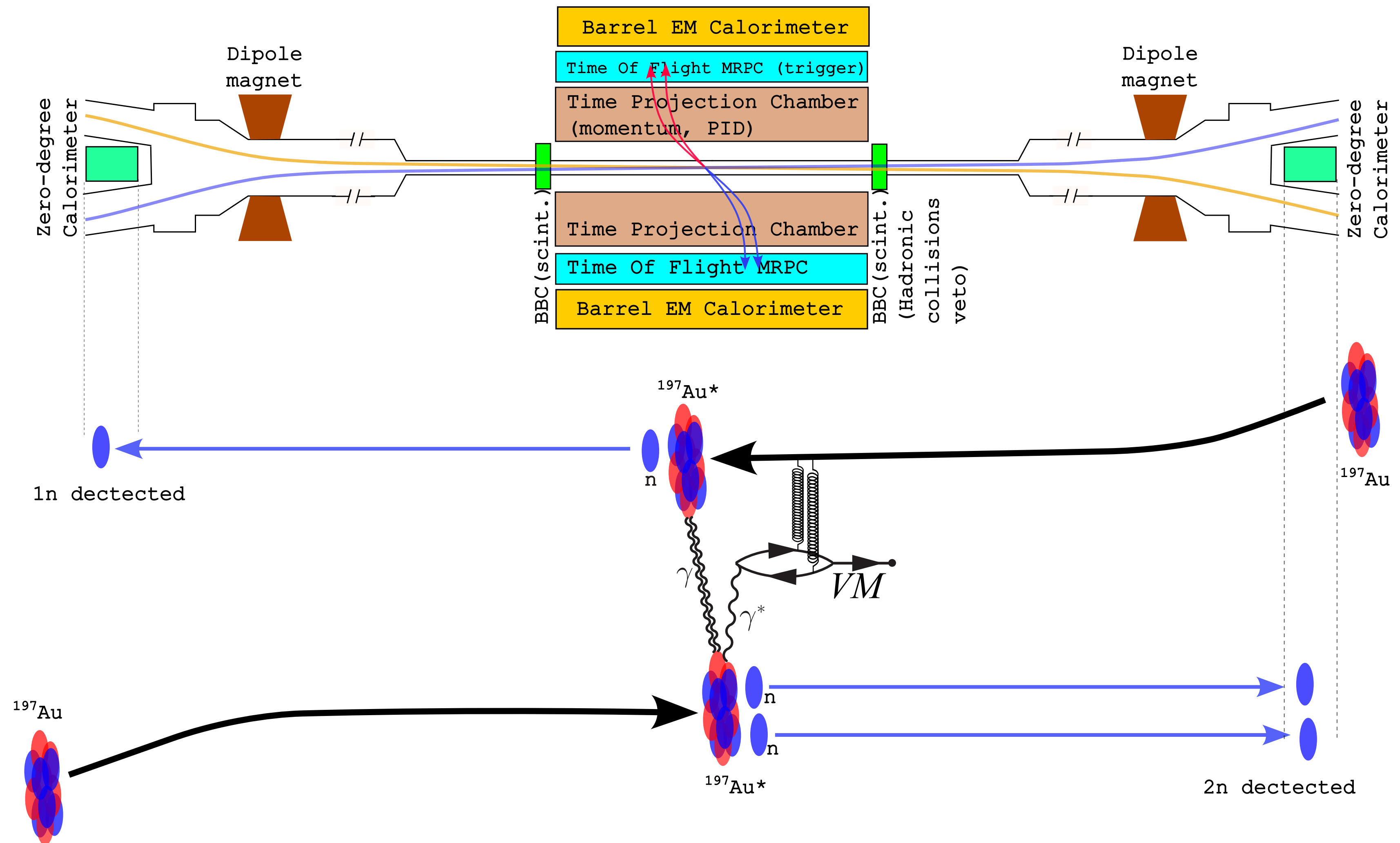
UPC AS A GREAT PRODUCTION TOOL

- ▶ Heavy Ion Collisions - large charge \Rightarrow large photon flux \Rightarrow large production cross section, accompanied by Coulomb excitation of the beam particles which emit neutrons \Rightarrow easy to trigger
- ▶ coherent (on nucleus) and incoherent (on nucleons)
- ▶ coherent photo production
 - ▶ final state is exclusive
 - ▶ easy to separate the signal from background



STAR EXPERIMENTAL SETUP (UPC RELEVANT DETECTORS ONLY)

- ▶ Solenoidal Tracker At RHIC
- ▶ central rapidity coverage
 - ▶ $(-1,1) \xrightarrow{2019} (-1.5,1.5)$
- ▶ neutron tagging
- ▶ charged hadrons PID
 - ▶ plus electron calorimetry including decay topology
- ▶ veto particles in the UPCs rapidity gap regions



DATASETS, LUMINOSITIES AND EVENT SELECTION

▶ Online Event Selection ("UPC_main" trigger)

▶ number of neutrons on each side

▶ 1 - 4.5 (Run 10)

▶ 1 - 3.5 (Run 11)

▶ 1 - 11 (Run 14)

▶ $2 \leq \text{Track Multiplicity} \leq 6$

▶ UPC Rapidity Gap Veto

▶ Offline Event Selection (analysis)

▶ | Z-Pos. of collision vertex | < 130 cm from acceptance center

▶ Track DCA to the vertex < 3cm

▶ TPC PID using dE/dx: normalized $|\sigma_\pi| < 3$

▶ #TPC track hits > 15

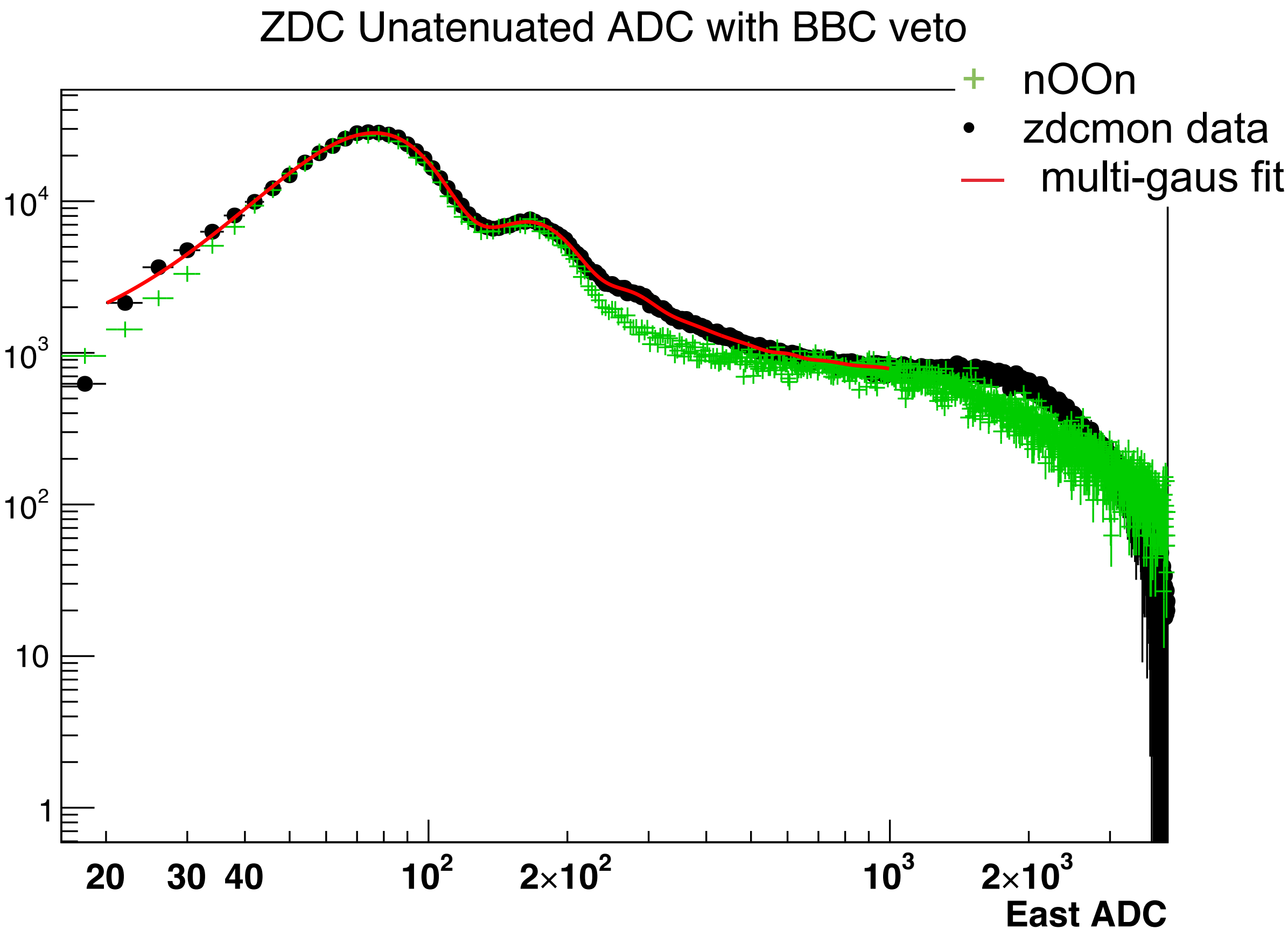
▶ $p_T(\pi^+\pi^-) < 0.15 \text{ GeV}/c$ or $p_T(\pi^+\pi^-\pi^+\pi^-) < 0.15 \text{ GeV}/c$

	Run14	Run11	Run10
$L^{-1} [\mu b^{-1}]$	787	523	926
L^{-1} fraction in $ v_z < 130$	0.664	0.813	0.764

$\sigma_{\text{TRIGGER}} \rightarrow \sigma_{\text{NN}}$

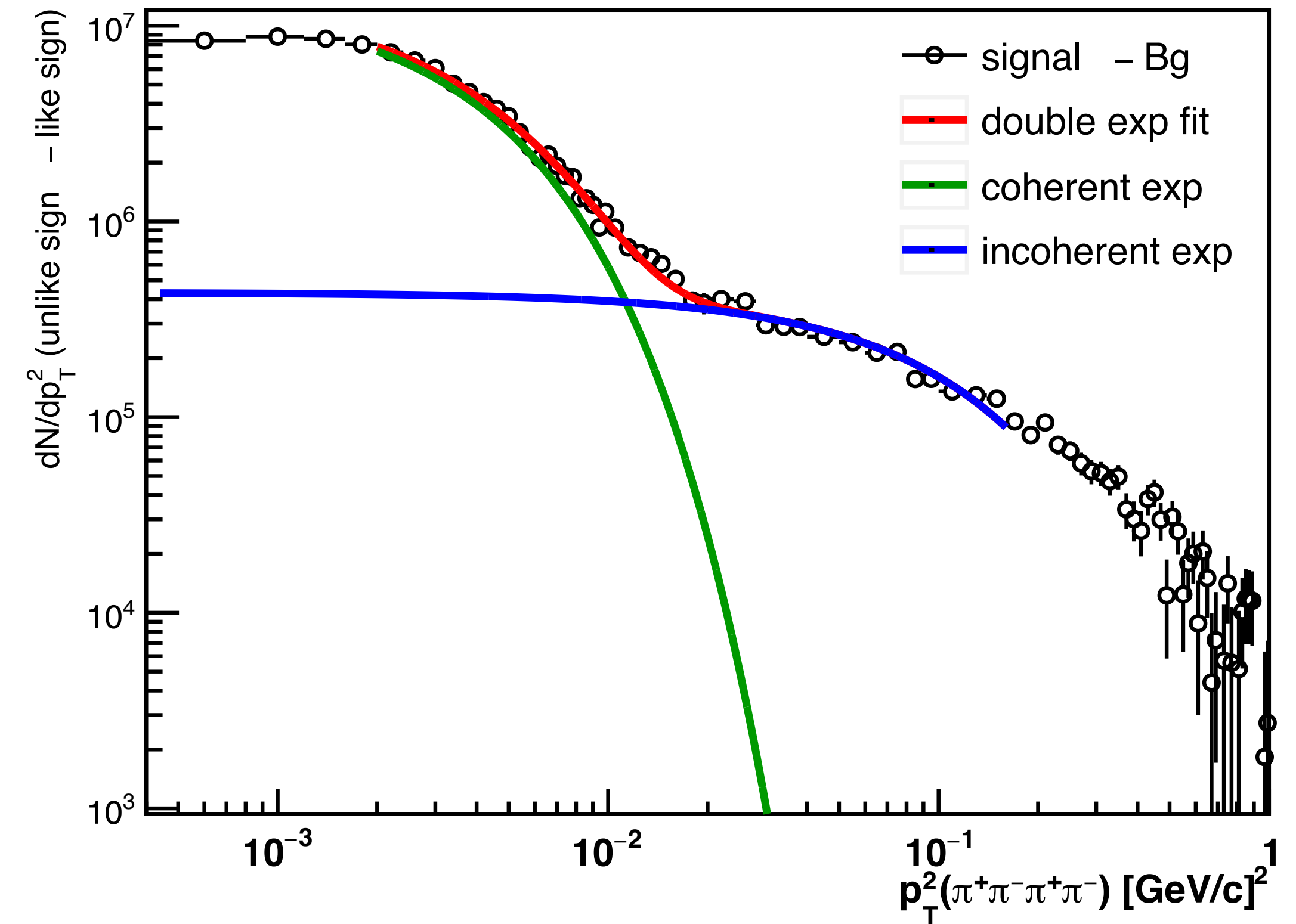
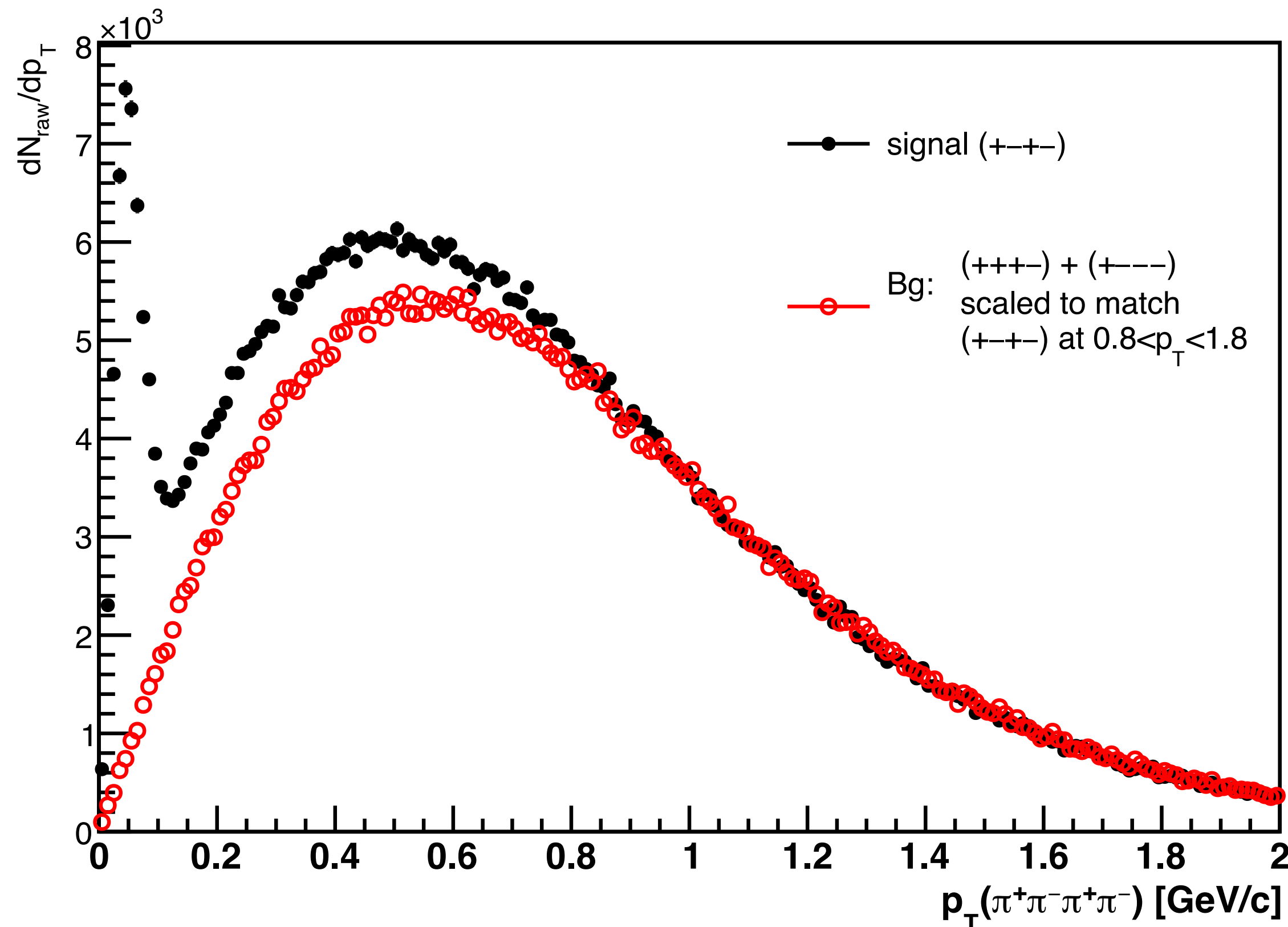
- ▶ UPC_main trigger does not see whole σ_{nn}
- ▶ STAR added a special trigger in Run14 called "zdcmon" that was just ZDC coincidence (no cut on ADC, no hadron veto)
 - ▶ we analyzed these data and compare with UPC_main to what fraction of σ_{nn} the UPC_main trigger "see" in each year.
- ▶ nOOn model [CPC 253 107181] of neutron production can predict neutron distribution in heavy ion collisions

	Run14	Run11	Run10
▶ UPC_main trigger	1-4.5n	1-3.5n	1-11n
▶ fraction from data	56.74%	37.72%	41.58%
▶ fraction from nOOn	63.16%	39.52%	43.52%



we used average of
data and nOOn

INCOHERENT CONTRIBUTION IN $P_T < 0.15$ RANGE



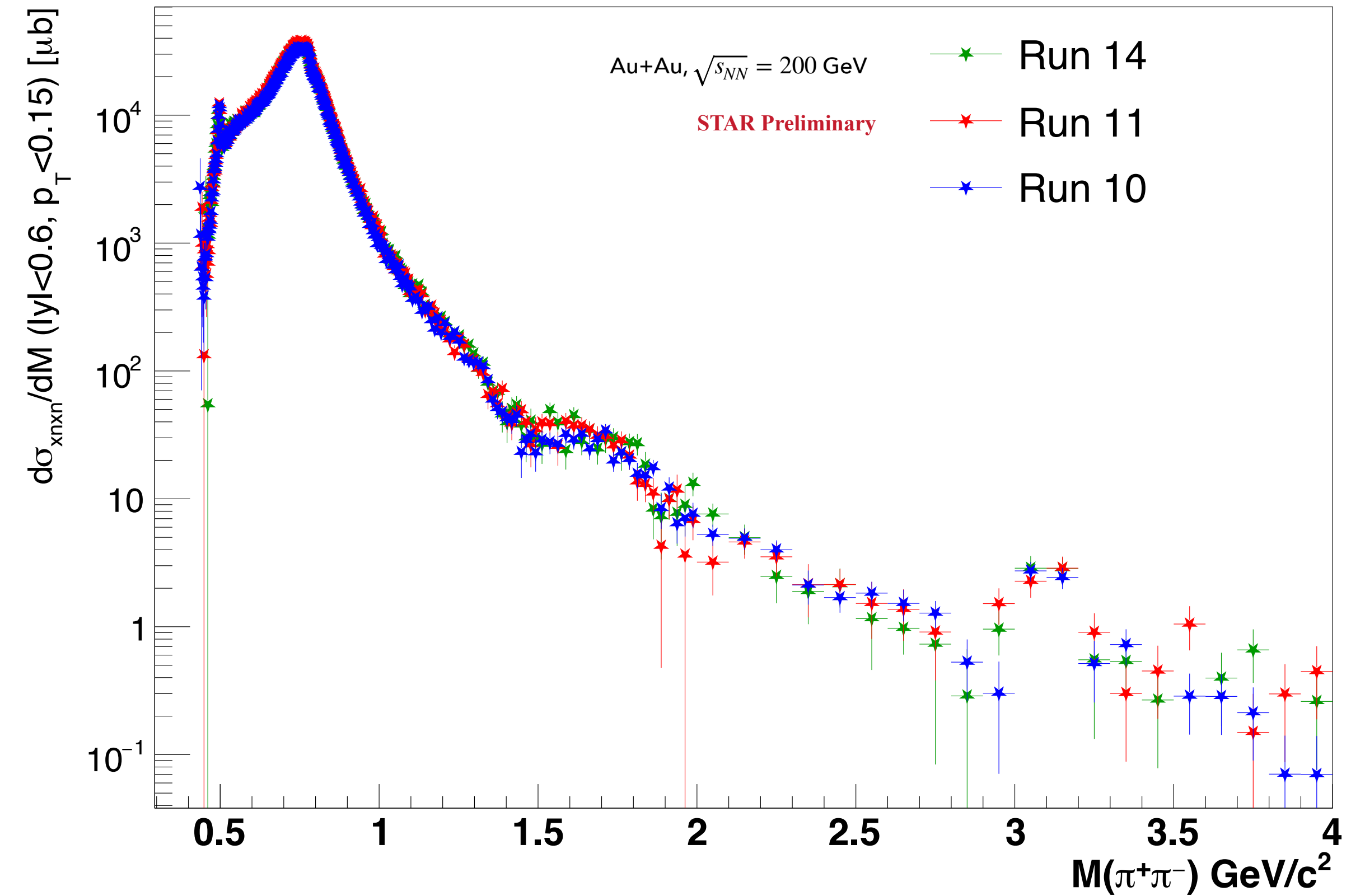
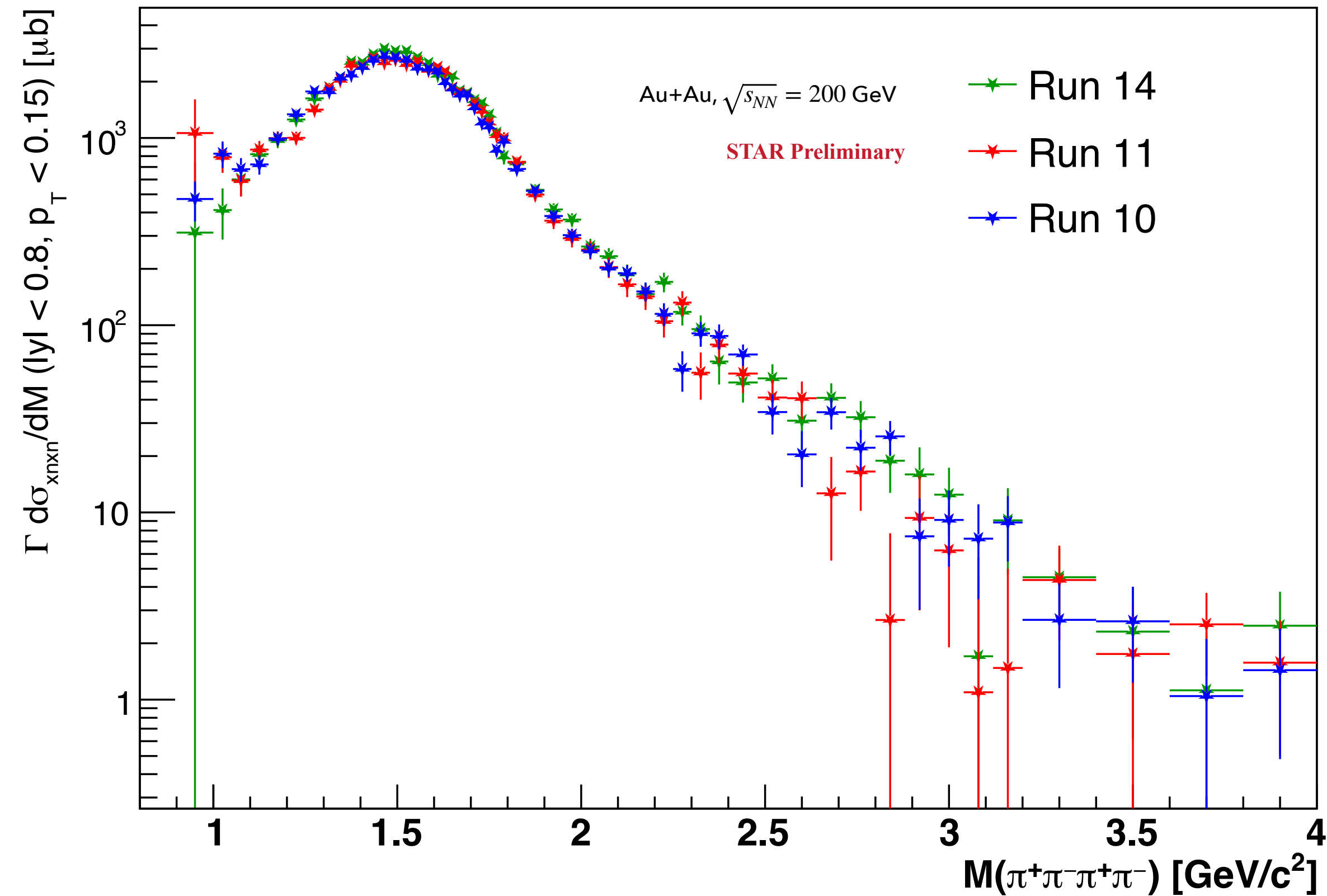
Incoherent fraction for $0.9 < M(4\pi) < 4.0$ GeV : 0.181

Incoherent fraction for $1.5 < M(4\pi) < 2.5$ GeV : 0.160

Incoherent fraction for $0.6 < M(2\pi) < 2.8$ GeV : 0.064

Incoherent fraction for $1.5 < M(2\pi) < 2.5$ GeV : 0.252

MASS SPECTRA OF $\pi^+\pi^-$ AND $\pi^+\pi^-\pi^+\pi^-$ (BOTH $p_T < 0.15$ GEV/C)

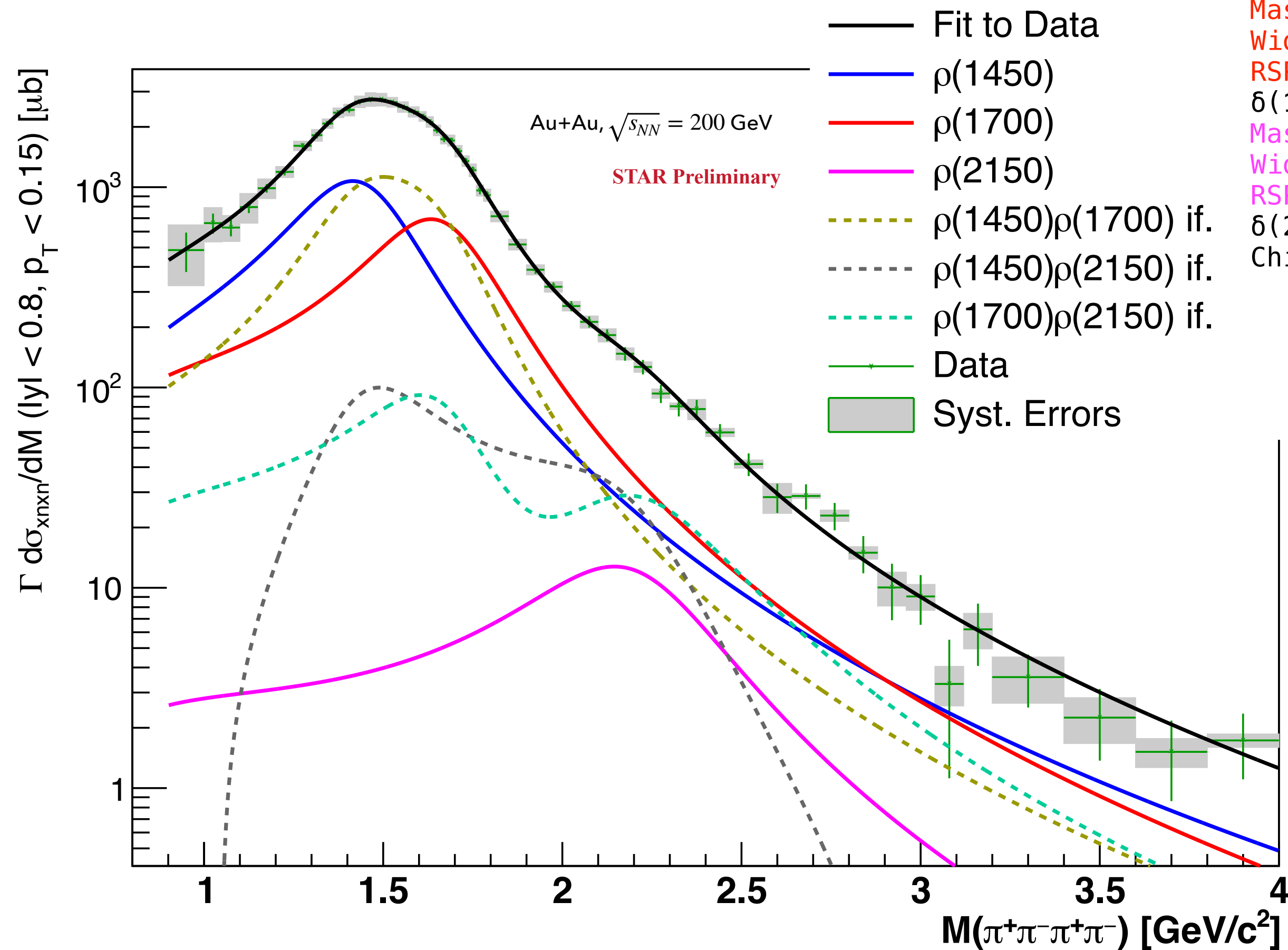


- Main “sanity check” - cross section consistent between datasets

MASS($\pi^+\pi^-\pi^+\pi^-$) AND RAPIDITY COMBINED FROM ALL RUNS

$$\Gamma \frac{d\sigma}{dM} = A^2 |BW(\rho_{1450})|^2 + B^2 |BW(\rho_{1700})|^2 + C^2 |BW(\rho_{2150})|^2 + 2\sqrt{AB} \Re [BW^*(\rho_{1450})BW(\rho_{1700})e^{i\delta(1700)}] + 2\sqrt{BC} \Re [BW^*(\rho_{1700})BW(\rho_{2150})e^{i\delta(2150)}] + 2\sqrt{AC} \Re [BW^*(\rho_{1450})BW(\rho_{2150})e^{i(\delta(1700)+\delta(2150))}]$$

$$BW(\rho) \equiv \left(\frac{M_\rho}{M}\right)^{\text{RSP}} \frac{\sqrt{WM_\rho}}{M^2 - M_\rho^2 + iM_\rho W}, \quad W = W_\rho \frac{M_\rho}{M} \left(\frac{M^2 - 16m_\pi^2}{M_\rho^2 - 16m_\pi^2}\right)^{3/2}$$



PARAMETER	VALUE	ERROR
Mass $\rho(1450)$	1486	20.3
Width $\rho(1450)$	400.3	30
RSP1	2.17	0.27
Mass $\rho(1700)$	1701	15.4
Width $\rho(1700)$	399.6	34.5
RSP2	2.39	0.37
$\delta(1700)$	1.22	0.38
Mass $\rho(2150)$	2247	91.2
Width $\rho(2150)$	570	fixed
RSP3	2.36	2.50
$\delta(2150)$	0.50	0.48
Chi2/ndf	33.0882/41	

$$\sigma_{4\pi,xn,xn}^{\text{coh}}(|y| < 0.8) = 1348 \pm 16_{\text{stat.}} \pm 20_{\text{syst.}} \pm 162_{\text{norm.}}$$

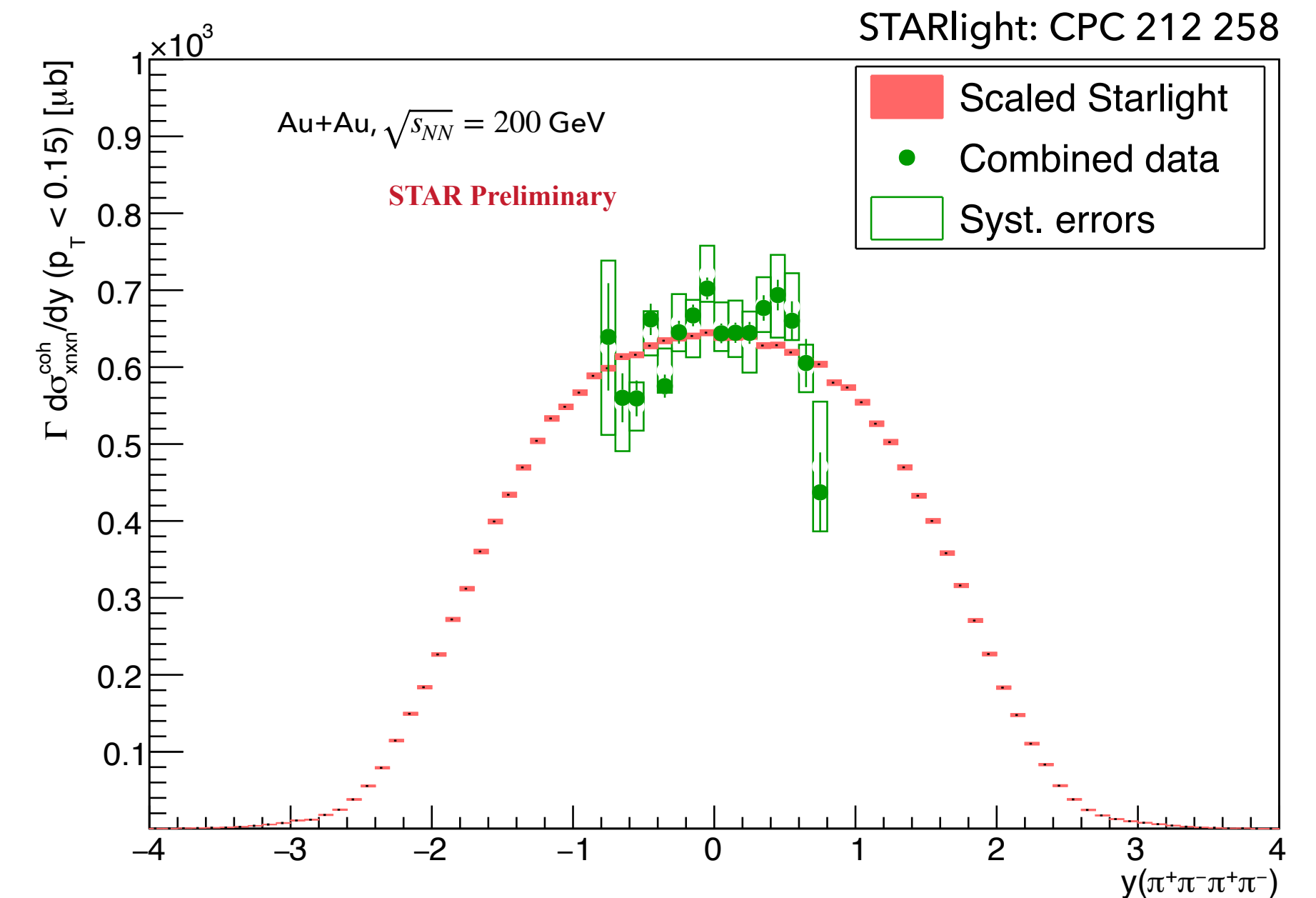
$$\sigma_{4\pi,xn,xn}^{\text{coh}} = 2941 \pm 34_{\text{stat.}} \pm 45_{\text{syst.}} \pm 352_{\text{norm.}}$$

$$\Gamma \sigma_{\rho_{1450},xn,xn}^{\text{coh}}(|y| < 0.8) = 426 \pm 162_{\text{stat.}} \pm 202_{\text{syst.}} \pm 51_{\text{norm.}}$$

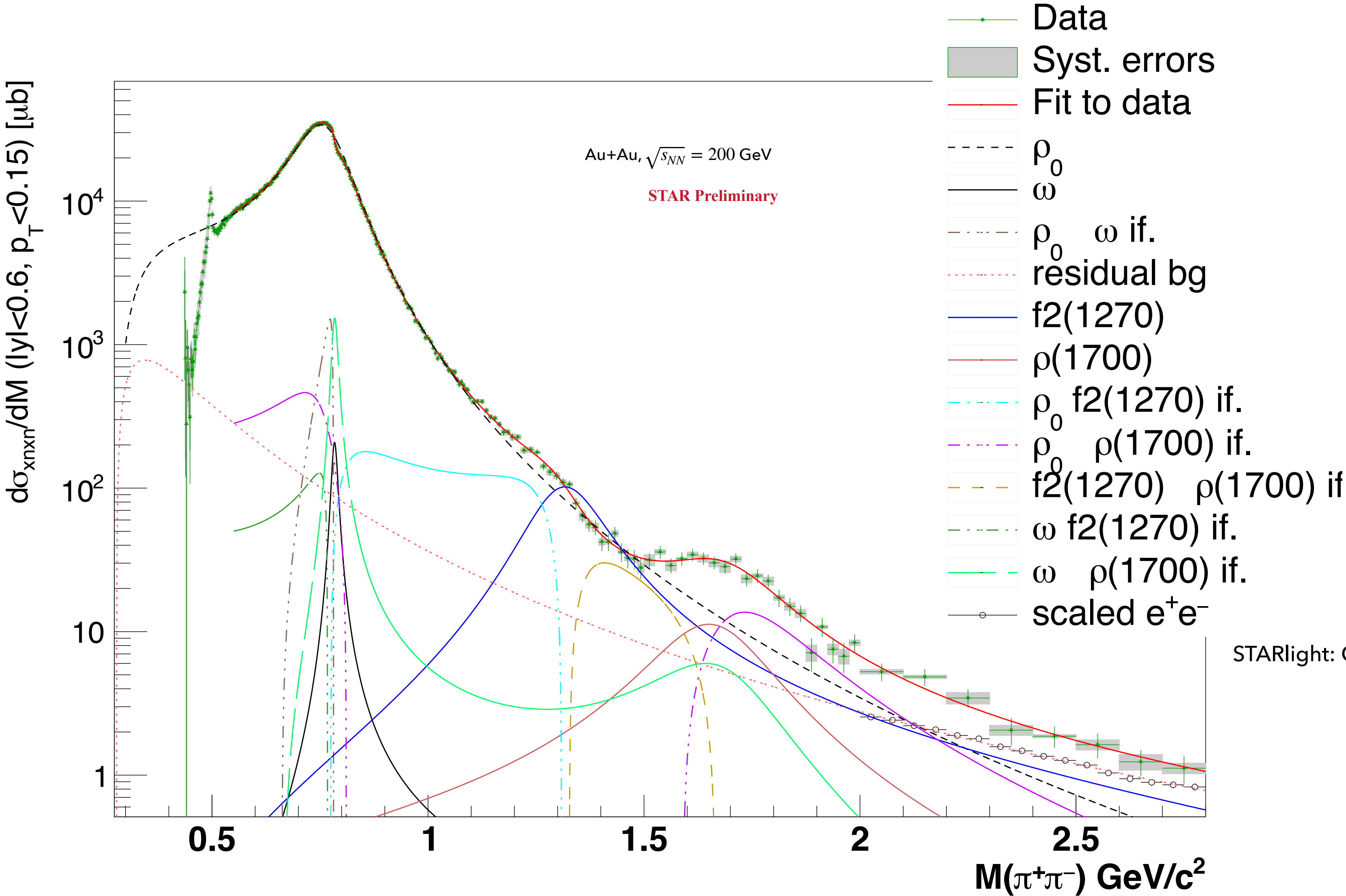
$$\Gamma \sigma_{\rho_{1700},xn,xn}^{\text{coh}}(|y| < 0.8) = 310 \pm 148_{\text{stat.}} \pm 169_{\text{syst.}} \pm 37_{\text{norm.}}$$

$$\Gamma \sigma_{\rho_{2150},xn,xn}^{\text{coh}}(|y| < 0.8) = 8.9 \pm 6.1_{\text{stat.}} \pm 1.3_{\text{syst.}} \pm 1.1_{\text{norm.}}$$

to extrapolate to full rapidity - multiply by 2.18

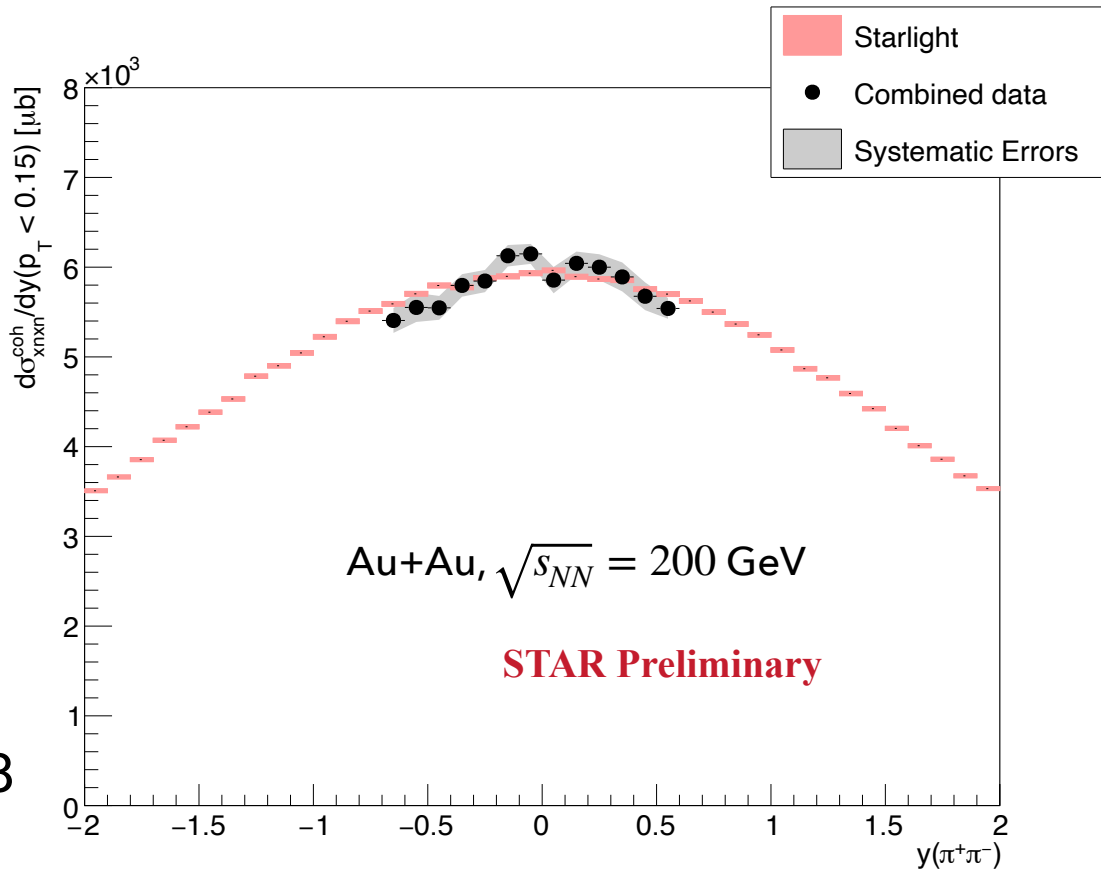


MASS OF $\pi^+\pi^-$ COMBINED FROM ALL RUNS



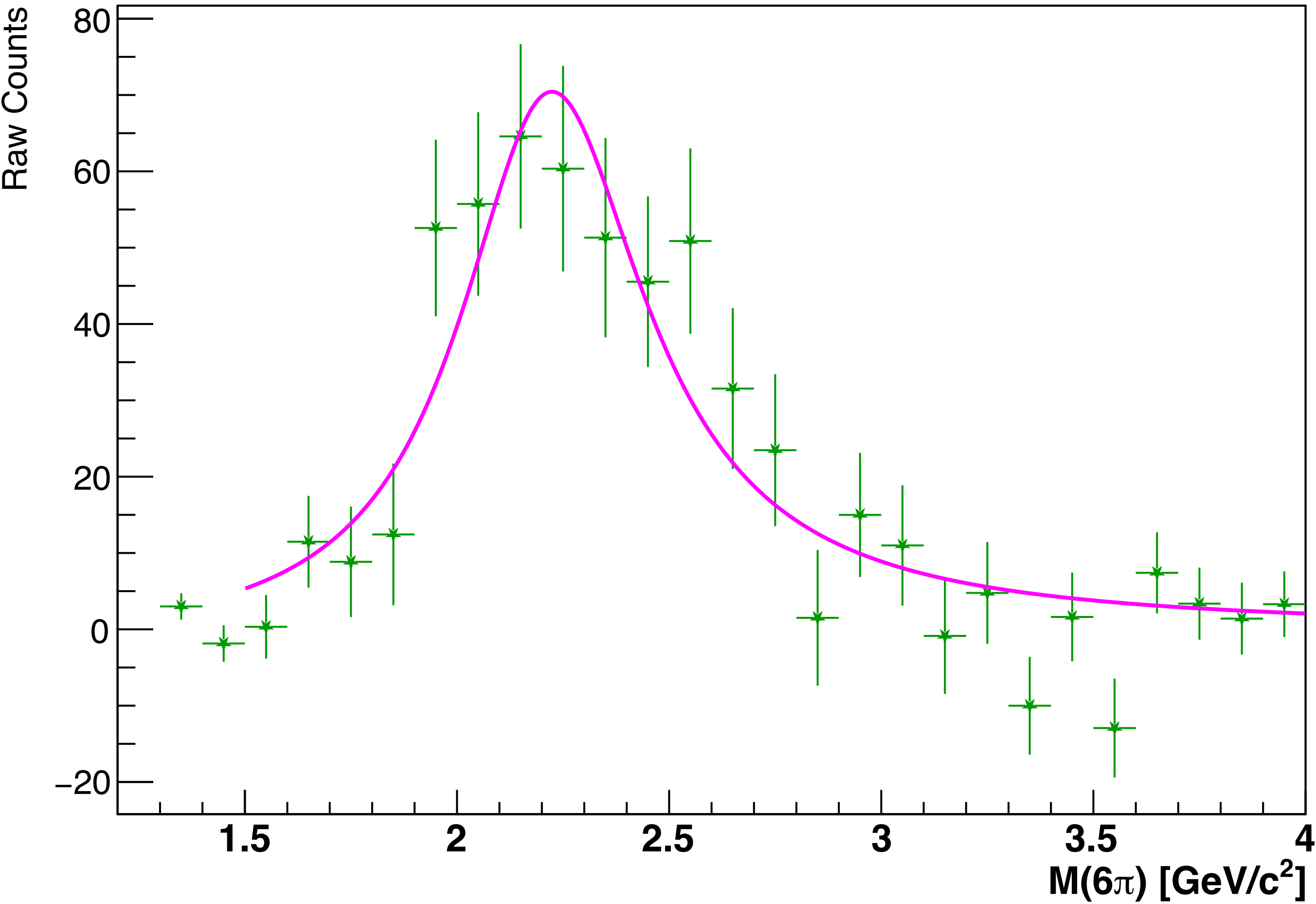
PARAMETER	VALUE	ERROR
Mass ρ_0	774.83	0.30
Width ρ_0	146.39	0.57
RSP ρ_0	2.76	0.02
Mass ω	783.82	0.50
Width ω	16.13	0.85
$\delta(\omega)$	1.97	0.17
Mass f2	1339.26	19.63
Width f2	210.35	38.30
Mass $\rho(1700)$	1700.65	22.71
Width $\rho(1700)$	317.79	46.75
$\delta(f2)$	3.12	0.18
$\delta(1700)$	0.38	0.18
Chi2/ndf=282.248/260		

STARlight: CPC 212 258



$\sigma_{\rho_0, xn, xn}^{\text{coh}}(|y| < 0.6) = 7219 \pm 28_{\text{stat.}} \pm 70_{\text{syst.}} \pm 866_{\text{norm.}}$
 $\Gamma \sigma_{\rho(1700), xn, xn}^{\text{coh}}(|y| < 0.6) = 3.43 \pm 1.05_{\text{stat.}} \pm 1.47_{\text{syst.}} \pm 0.41_{\text{norm.}}$
to extrapolate to full rapidity - multiply by 3.66

6π PHOTO PRODUCTION RESULTS

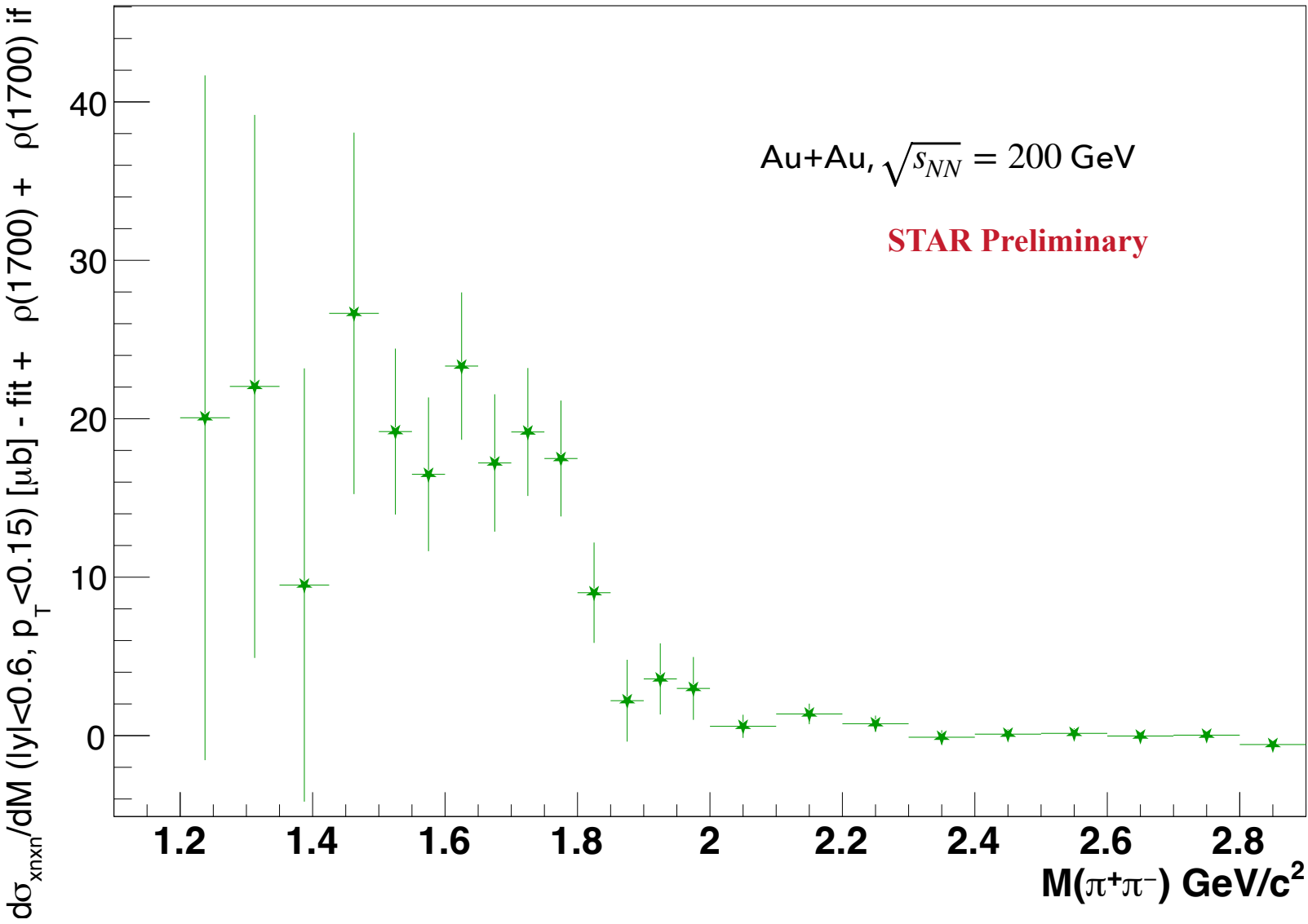


PARAMETER	VALUE	ERROR
Mass $\rho(2150)$	2276	57.8
Width $\rho(2150)$	573	84.5
RSP $\rho(2150)$	0.82	0.42
Chi2/ndf	27.9946/21	

First observation at STAR

RATIO OF THE BRANCHING FRACTIONS OF THE RHO1700 TO 2PI AND 4PI

- comparison of yields directly from Breit-Wigner functions (let's do it from full rapidity and mass)
 - $\sigma_{\rho 1700, xn, xn}^{coh} = 4.53 \pm 1.38_{stat.} \pm 1.93_{syst.}$ in $|y| < 0.8$ and $12.6 \pm 3.8_{stat.} \pm 5.4_{syst.}$ in full rapidity from $\pi^+ \pi^-$
 - $\sigma_{\rho 1700, xn, xn}^{coh} = 328 \pm 161_{stat.} \pm 178_{syst.}$ in $|y| < 0.8$ and $716 \pm 351_{stat.} \pm 388_{syst.}$ in full rapidity from $\pi^+ \pi^- \pi^+ \pi^-$
 - $(\rho_{1700} \rightarrow \pi^+ \pi^-) / (\rho_{1700} \rightarrow \pi^+ \pi^- \pi^+ \pi^-) = 1.38 \pm 0.80_{stat.} \pm 0.95_{syst.} \% [(1.76 \pm 1.01 \pm 1.20)\% \text{ in full rapidity}]$
- an alternative method using an excessive yield in $\pi^+ \pi^-$ and yield in $\pi^+ \pi^- \pi^+ \pi^-$ in the mass window from 1.5 to 2.5 GeV/c² - a good proxy for ρ_{1700}
 - the excessive yield in $\pi^+ \pi^-$ can be calculated as $\pi^+ \pi^-$ data – components of the fit function excluding ρ_{1700} Breit-Wigner and its interference



Excess in $1.5 < \text{Mass} < 2.5 \quad |y| < 0.8 = 6.72 \pm 0.61_{stat.} \pm 0.33_{syst.}$
 Excess in $1.5 < \text{Mass} < 2.5 \text{ full } y = 18.7 \pm 1.7_{stat.} \pm 0.91_{syst.}$

$$\sigma_{4\pi, xn, xn}^{coh} (|y| < 0.8, 1.5 < M < 2.5) = 612 \pm 8_{stat.} \pm 17_{syst.}$$

$$\sigma_{4\pi, xn, xn}^{coh} (1.5 < M < 2.5) = 1336 \pm 18_{stat.} \pm 37_{syst.}$$

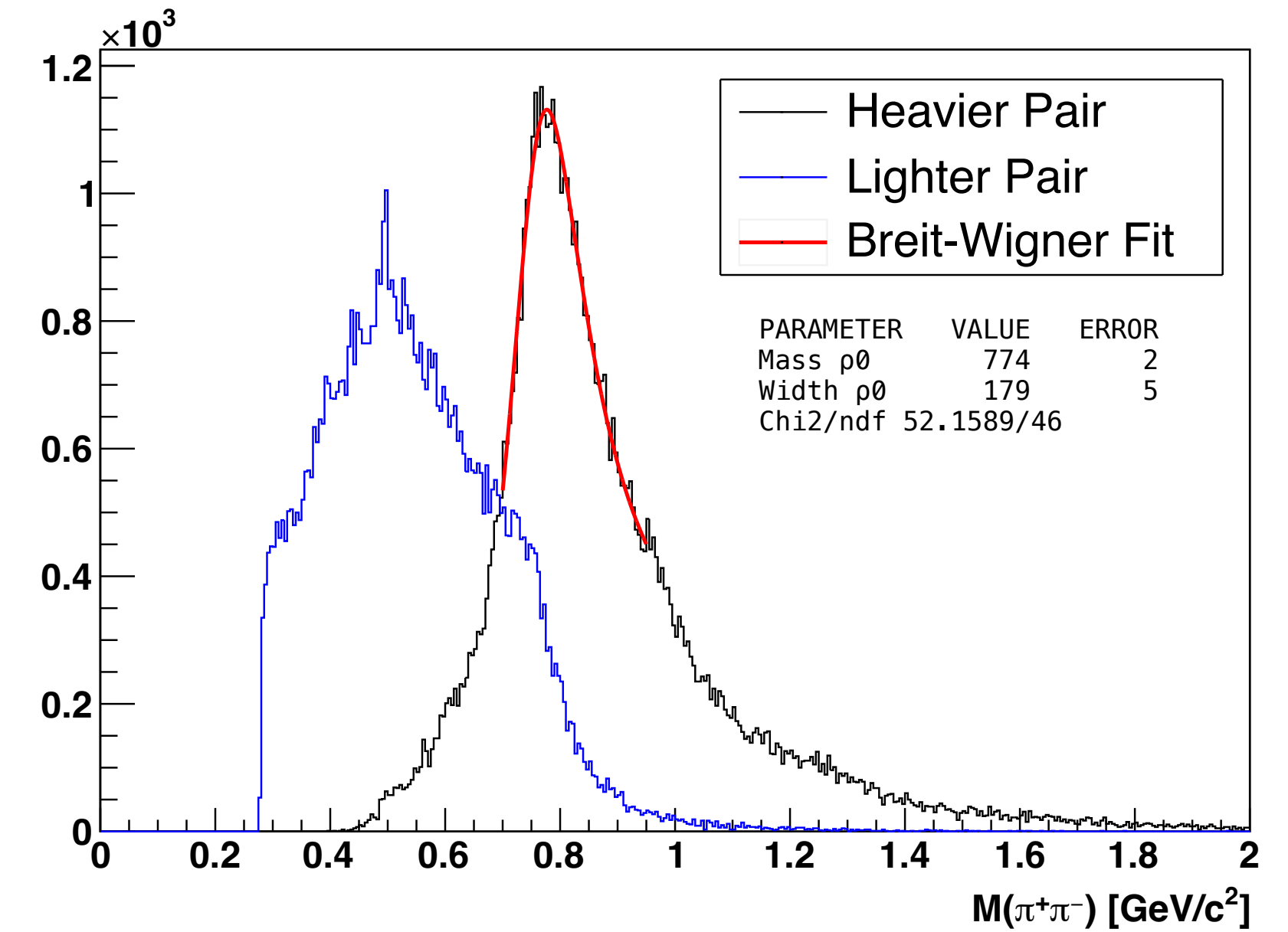
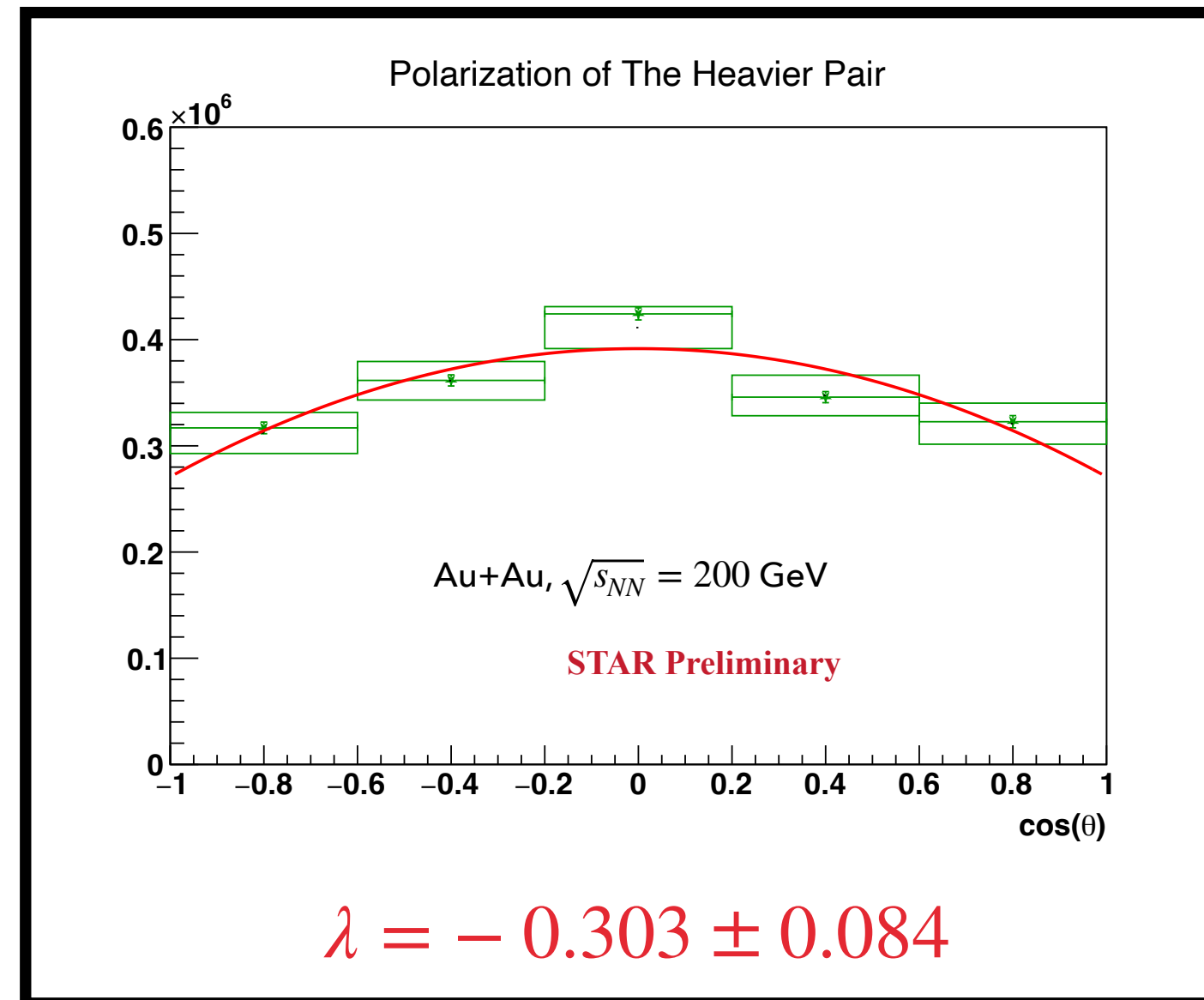
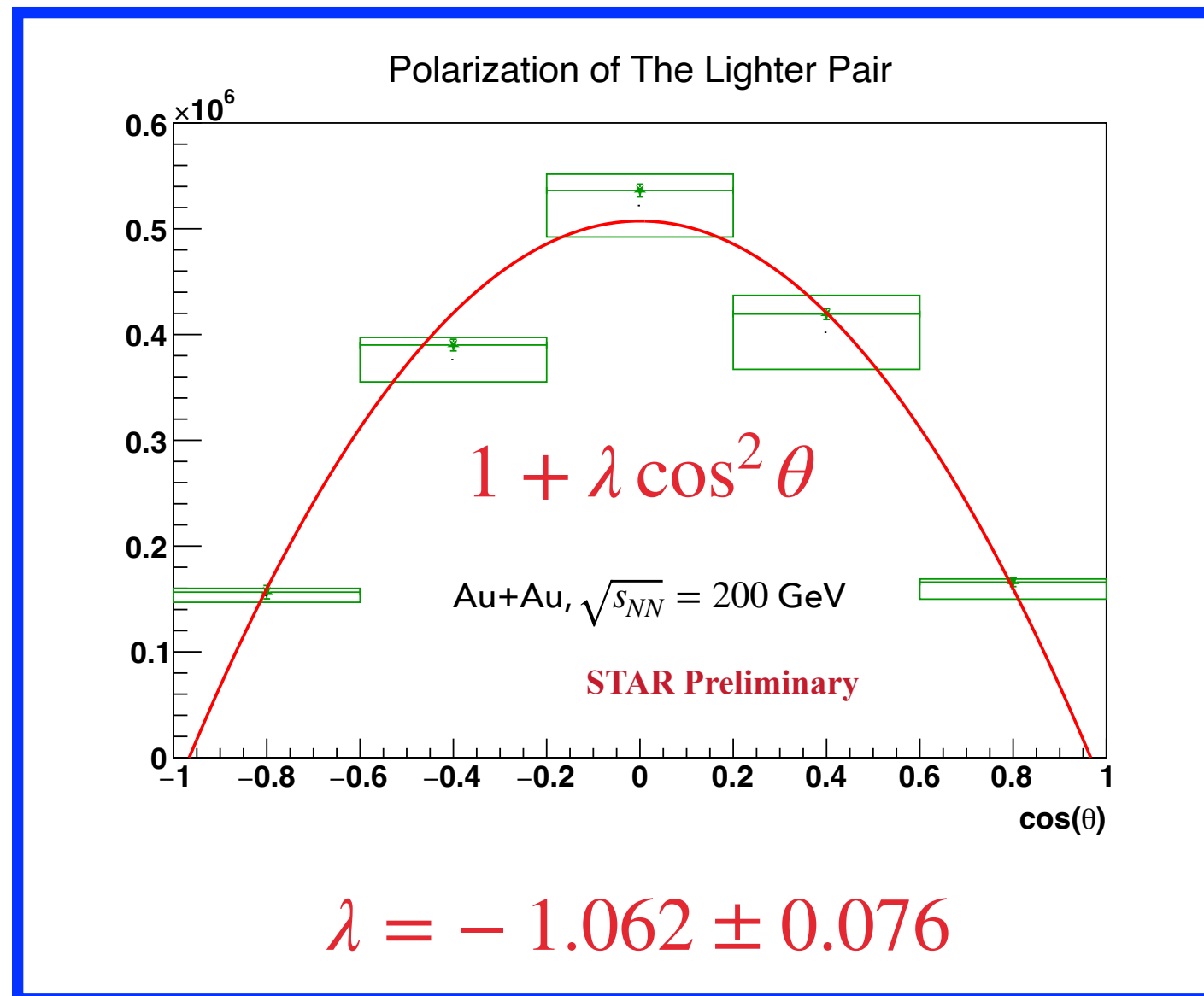
$$(\rho_{1700} \rightarrow \pi^+ \pi^-) / (\rho_{1700} \rightarrow \pi^+ \pi^- \pi^+ \pi^-) = 1.10 \pm 0.10_{stat.} \pm 0.06_{syst.} \%$$

$$(\rho_{1700} \rightarrow \pi^+ \pi^-) / (\rho_{1700} \rightarrow \pi^+ \pi^- \pi^+ \pi^-) = 1.40 \pm 0.13_{stat.} \pm 0.08_{syst.} \% \text{ in full rapidity}$$

PRC 81 044901	
$\sigma_{4\pi, xn, xn}^{coh} / \sigma_{\rho 0, xn, xn}^{coh} (y < 0.8) = 14.1 \pm 0.4_{stat.} \pm 0.5_{syst.} \%$	$16.4 \pm 1.0_{stat} \pm 5.2_{syst} \%$
$\sigma_{4\pi, xn, xn}^{coh} / \sigma_{\rho 0, xn, xn}^{coh} = 11.1 \pm 0.3_{stat.} \pm 0.4_{syst.} \%$	$13.4 \pm 0.8_{stat} \pm 4.4_{syst} \%$

J=1 TRANSFER TO $\cos\theta$ ANISOTROPY

$$\rho(1450) + \rho(1700) \rightarrow (\pi^+\pi^-)_{\text{heavy}} + (\pi^+\pi^-)_{\text{light}} \rightarrow \pi^+\pi^-\pi^+\pi^-$$



$$\cos \theta = \frac{\vec{P}_{\pi^+\pi^-} \cdot \vec{p}_{\pi^*}}{\|\vec{P}_{\pi^+\pi^-}\| \|\vec{p}_{\pi^*}\|}$$

- Breit-Wigner fit to the heavier pair mass point to ρ_0 meson.
- Heavy pair polarized, but not fully
- Light pair looks like $f_0(500)$ resonance. But this resonance is supposed to be a scalar meson while its polarization indicates vector meson.

SUMMARY

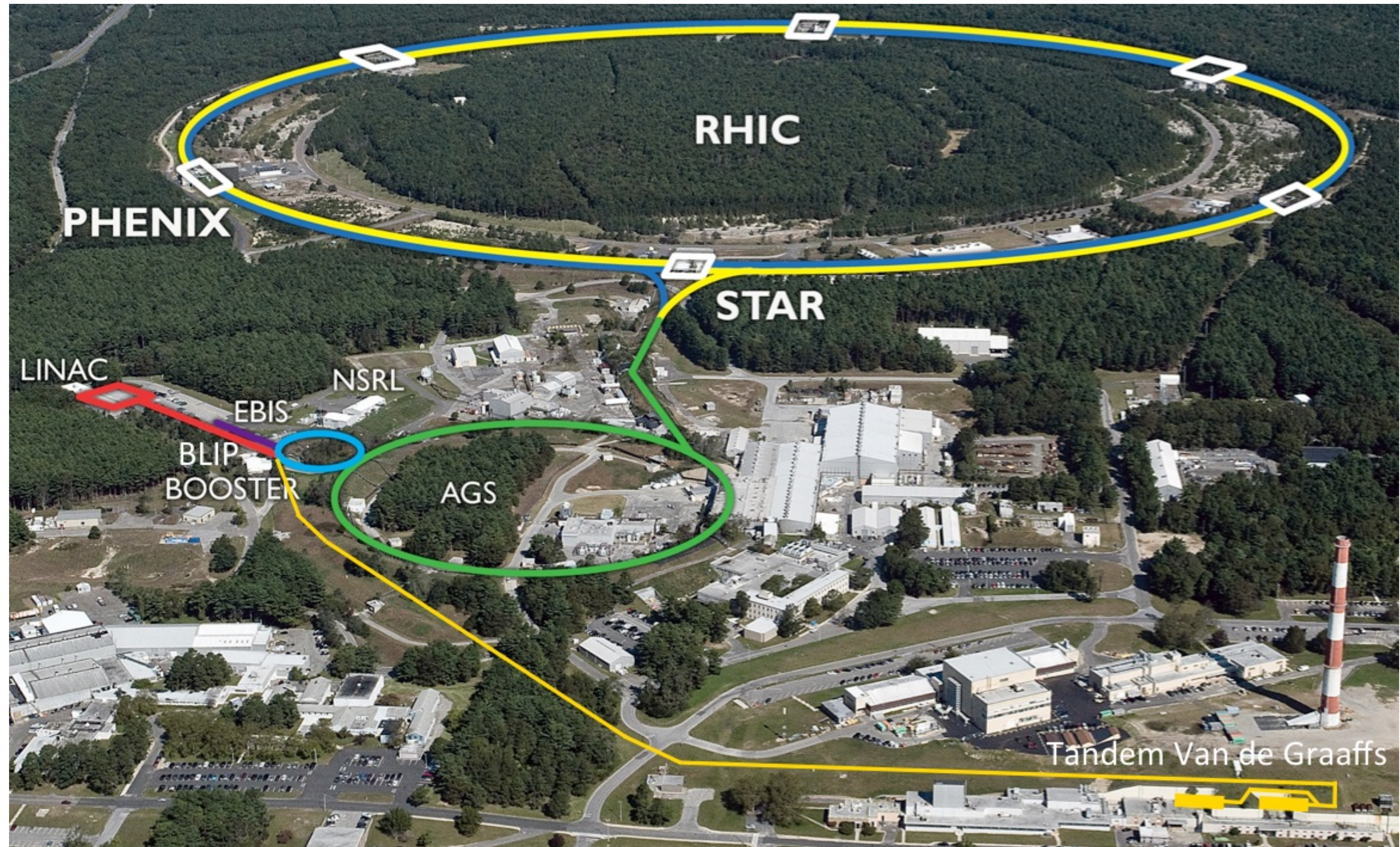
- ▶ STAR presented a precise measurement of $\pi^+\pi^-\pi^+\pi^-$ and $\pi^+\pi^-$ photo-production in Au+Au collisions at $\sqrt{s_{NN}}=200$ GeV
 - ▶ ρ_{1450} (2^3S_1), ρ_{1700} (1^3D_1) clearly and ρ_{2150} (2^3D_1) likely observer in $\pi^+\pi^-\pi^+\pi^-$ mass spectrum
 - ▶ ρ_{1450} mass and width consistent with the world average
 - ▶ ρ_{1700} mass consistent with the world average, but width larger => more decays modes possible, hybrid state indication?
 - ▶ $f_2(1250)$ and ρ_{1700} observer in $\pi^+\pi^-$ mass spectrum
 - ▶ $f_2(1250)$ mass and width too large - might contain ρ_{1450} , but we can't separate these states
 - ▶ ρ_{1700} mass and width consistent with the world average, width lower than in $\pi^+\pi^-\pi^+\pi^-$ => existence of intermediate states in decay
 - ▶ ρ_{2150} (2^3D_1) observed in $\pi^+\pi^-\pi^+\pi^-$ mass spectrum - first time at STAR
 - ▶ $(\rho_{1700} \rightarrow \pi^+\pi^-)/(\rho_{1700} \rightarrow \pi^+\pi^-\pi^+\pi^-) = 1.10 \pm 0.10_{\text{stat.}} \pm 0.06_{\text{syst.}} \%$ in mid rapidity and $1.40 \pm 0.13_{\text{stat.}} \pm 0.08_{\text{syst.}} \%$ in full rapidity
 - ▶ $\sigma_{4\pi, xn, xn}^{\text{coh}}/\sigma_{\rho_0, xn, xn}^{\text{coh}} = 14.1 \pm 0.4_{\text{stat.}} \pm 0.5_{\text{syst.}} \%$ in mid and $11.1 \pm 0.3_{\text{stat.}} \pm 0.4_{\text{syst.}} \%$ in full rapidity
- ▶ $\pi^+\pi^-\pi^+\pi^-$ states' (all supposed to have J=1) decay can be separated to 2 $\pi^+\pi^-$ pairs by their mass
 - ▶ the lighter pair whose mass spectrum resembles $f_0(500)$, a scalar meson, decays like fully polarized particle ($\lambda = -1$)
 - ▶ the heavier pair whose mass spectrum resembles ρ_0 is partially polarized ($\lambda = -0.3$)

THANK YOU

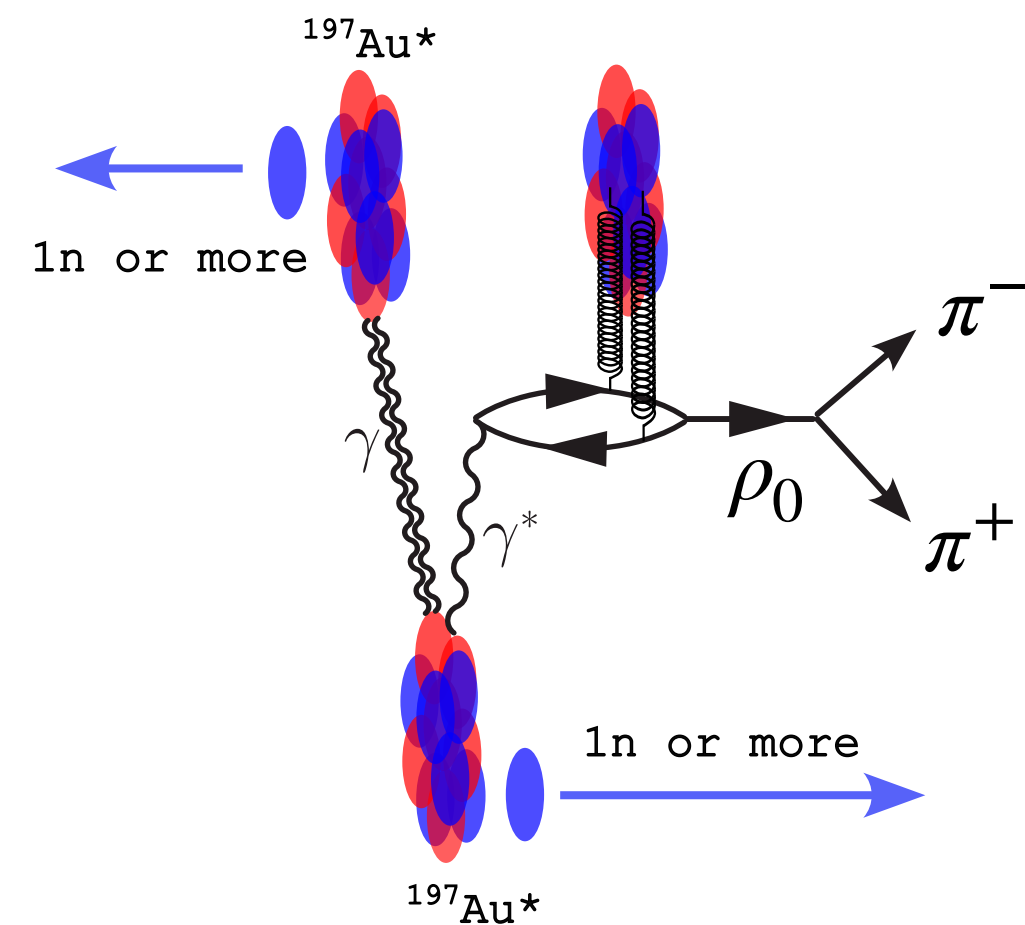
BACKUP SLIDES

ULTRA-PERIPHERAL COLLISIONS AT RHIC

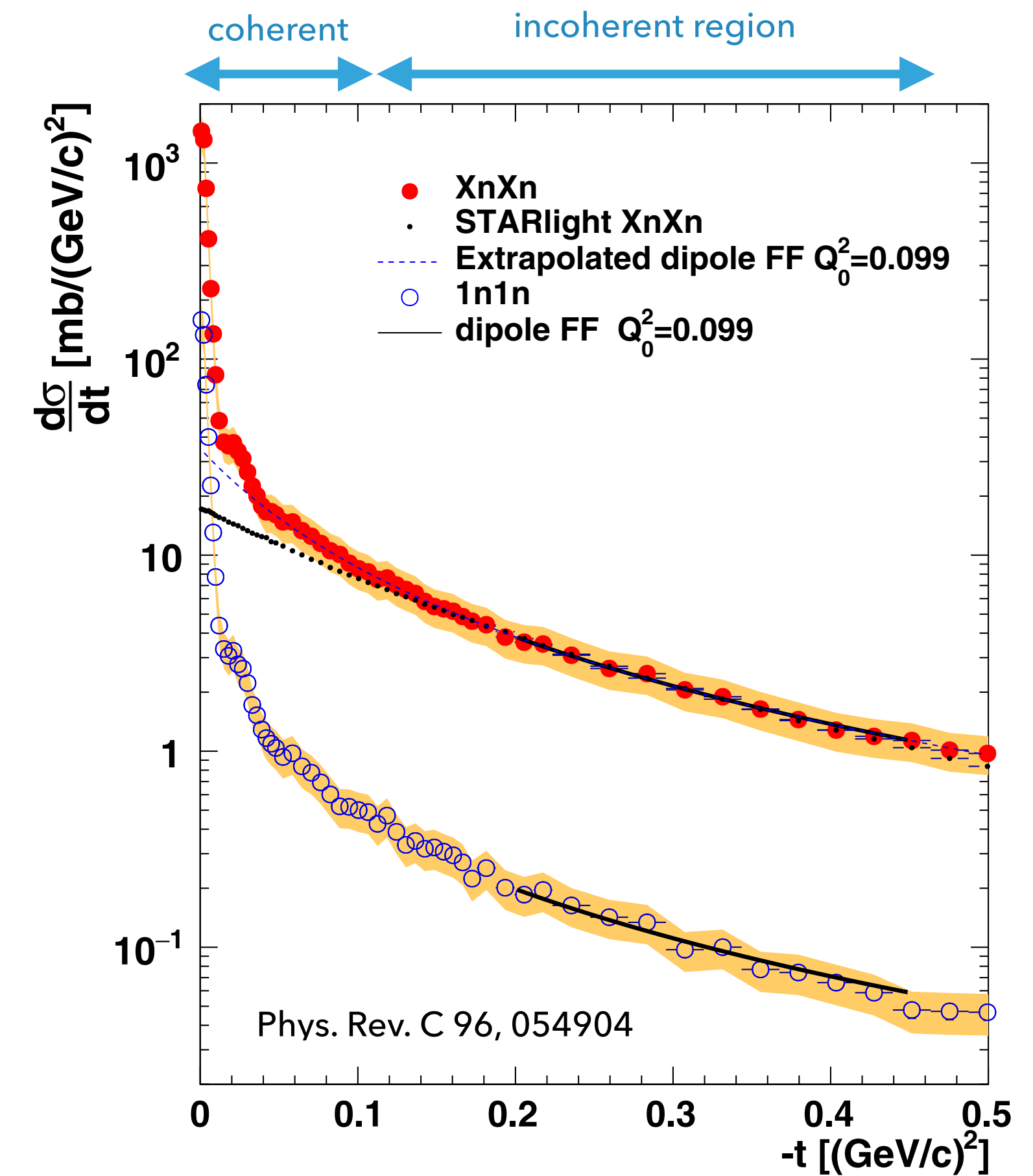
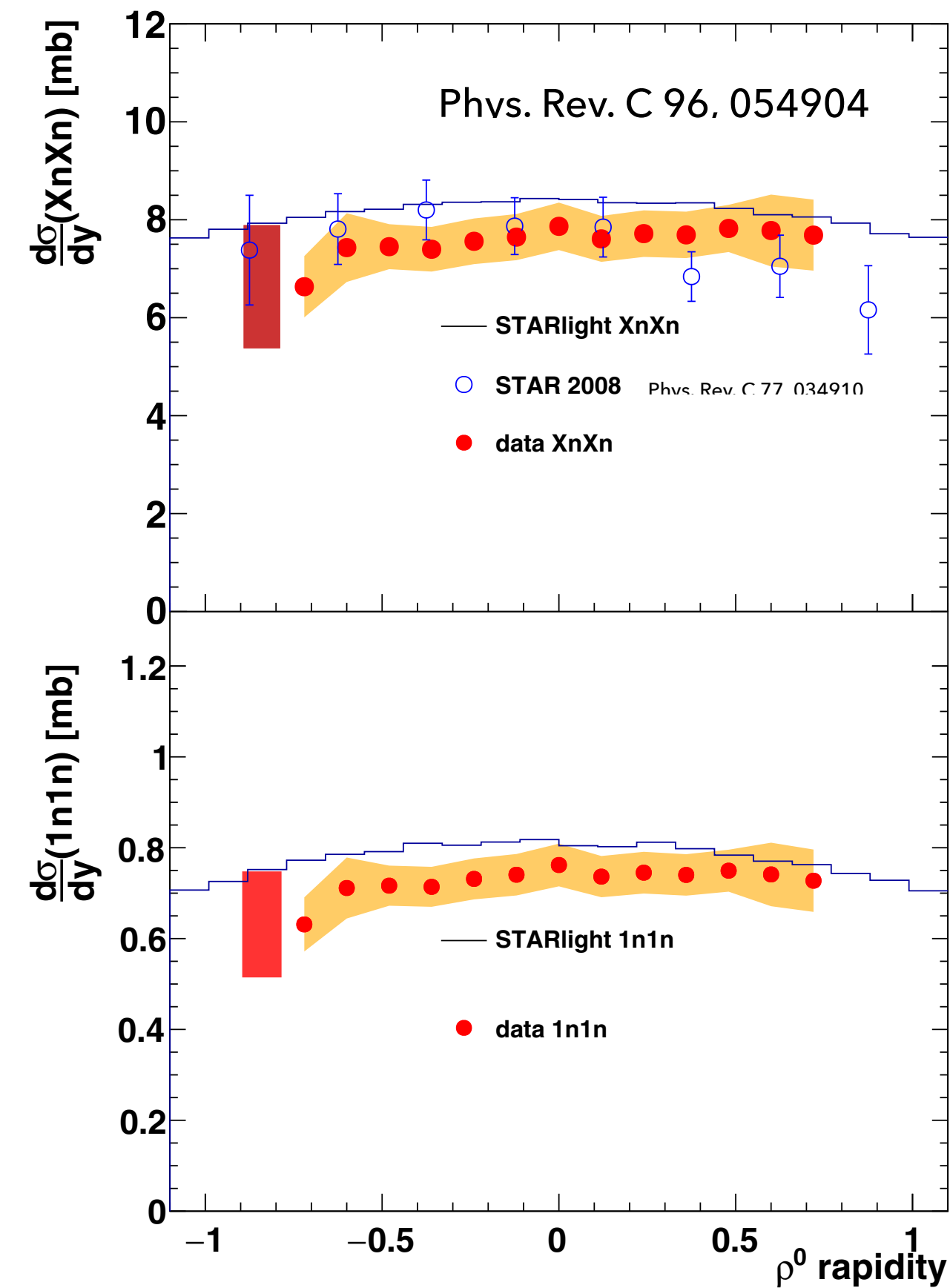
- ▶ Relativistic Heavy Ion Collider
- ▶ located in Brookhaven National Laboratory (Long Island, USA)
- ▶ different species, energy, and proton polarization



U^{238} , Au^{197} , Zr^{96} , Ru^{96} , d^2 at 200 GeV and pp at 510 GeV

ρ_0 CROSS SECTION

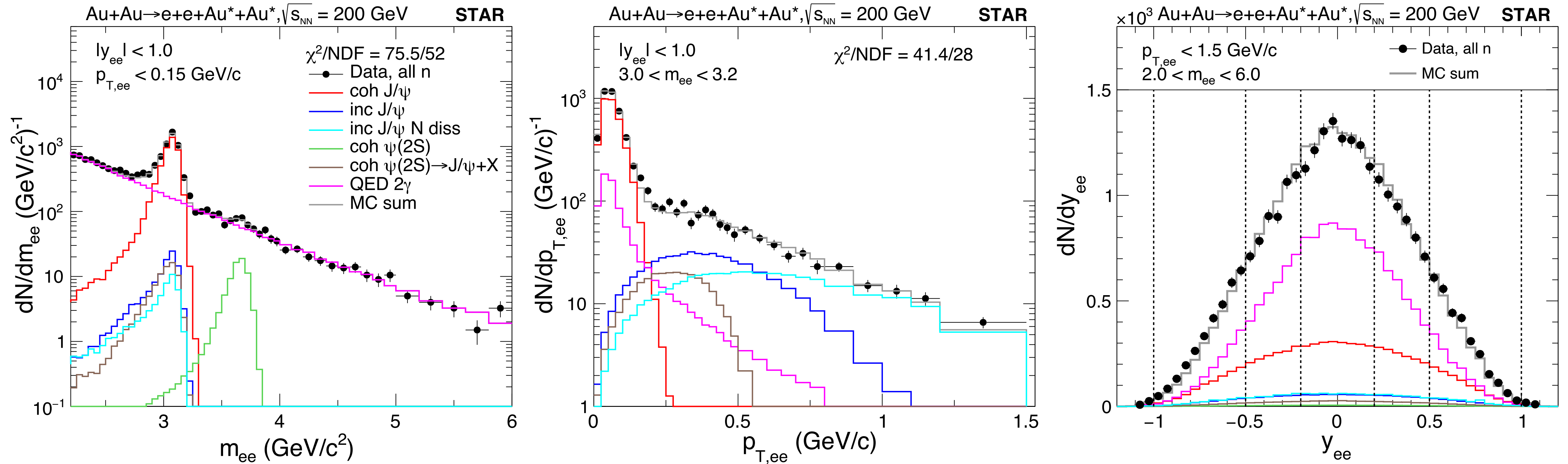
- integrated luminosity of 1100 ± 100 μb^{-1} of data collected in 2010
- XnXn extrapolated from 1n1n using STARlight
- incoherent components in $d\sigma/dt$ are fit in range $-t = (0.2, 0.45)$
 - σ_{incoh} are integrals of the fits



Parameter	XnXn	1n1n
$\sigma_{coh.}$	6.49 ± 0.01 (stat.) ± 1.18 (syst.) mb	0.770 ± 0.004 (stat.) ± 0.140 (syst.) mb
$\sigma_{incoh.}$	2.89 ± 0.02 (stat.) ± 0.54 (syst.) mb	0.162 ± 0.010 (stat.) ± 0.029 (syst.) mb
$\sigma_{incoh.}/\sigma_{coh.}$	0.445 ± 0.015 (stat.) ± 0.005 (syst.)	0.233 ± 0.007 (stat.) ± 0.007 (syst.)

Nuclear excitation and ρ_0 photo production are not completely independent

J/ψ PHOTOPRODUCTION IN AU+AU UPC EVENTS AT 200 GEV



- when $Q^2 \sim 0$, p_T of J/ψ is directly related to momentum transfer ($t \sim p_T^2$)

OUTLINE

- ▶ STAR Experiment
- ▶ Recently Published UPC Results from STAR
 - ▶ coherent ρ_0 and nuclear imaging
 - ▶ coherent and incoherent J/ψ photo-production in d+Au and Au+Au
 - ▶ coherent $\psi(2S)$ photo-production in Au+Au
 - ▶ di-leptons from Breit-Wheeler process
- ▶ Highlights of the newest preliminary results
- ▶ Outlook
- ▶ Summary