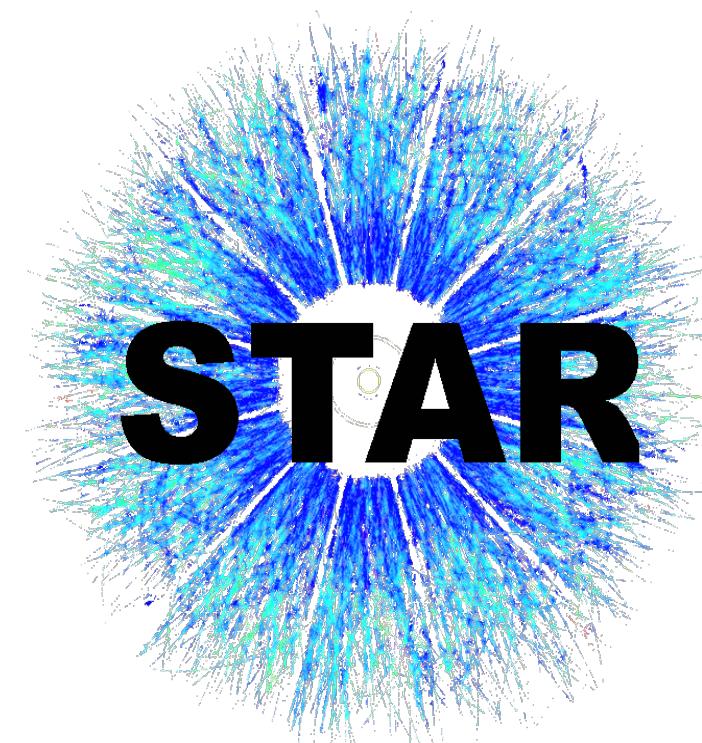


UPC 2025: The Second International Workshop on the Physics of Ultra Peripheral Collisions

Supported in part by the



Creighton  
UNIVERSITY



DAVID TLUSTY (CREIGHTON UNIVERSITY)  
*FOR THE STAR COLLABORATION*

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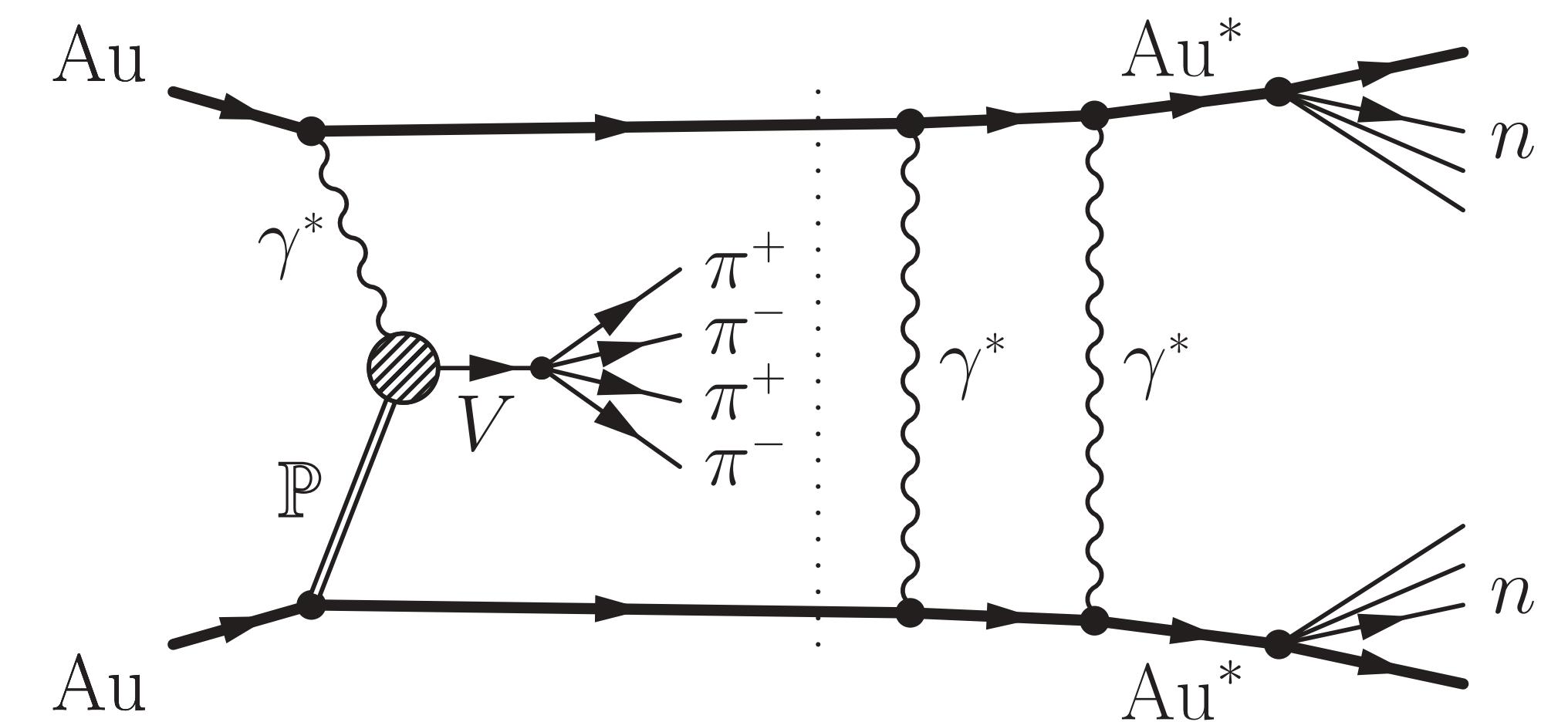
OBSERVATION OF  $\pi^+\pi^-\pi^+\pi^-$  AND  $\pi^+\pi^-$   
FINAL STATE PHOTO PRODUCTION IN UPC AT  
 $\sqrt{s_{NN}} = 200$  GET AT THE STAR DETECTOR

## MOTIVATION

- ▶ The first radial excitation  $2^3S_1$  of  $\rho_0$  is considered to be the  $\rho_{1450}$  [PRD 110 030001], but decays suggest it is a hybrid state [PRD 56 1584]
- ▶  $\rho_{1700}$  is assigned to  $1^3D_1$  state - there is need for precise measurement of mass and width to clarify its nature [PRD 110 030001]
- ▶ Questions of the  $\rho_{1450}$  relation to the  $\rho_{1700}$  have been raised
- ▶ The relativistic quark model [PRD 32 189] predicts  $2^3D^1$  state  $J^{PC} = 1^{--}$  at 2.15 GeV which can be identified with the  $\rho(2150)$

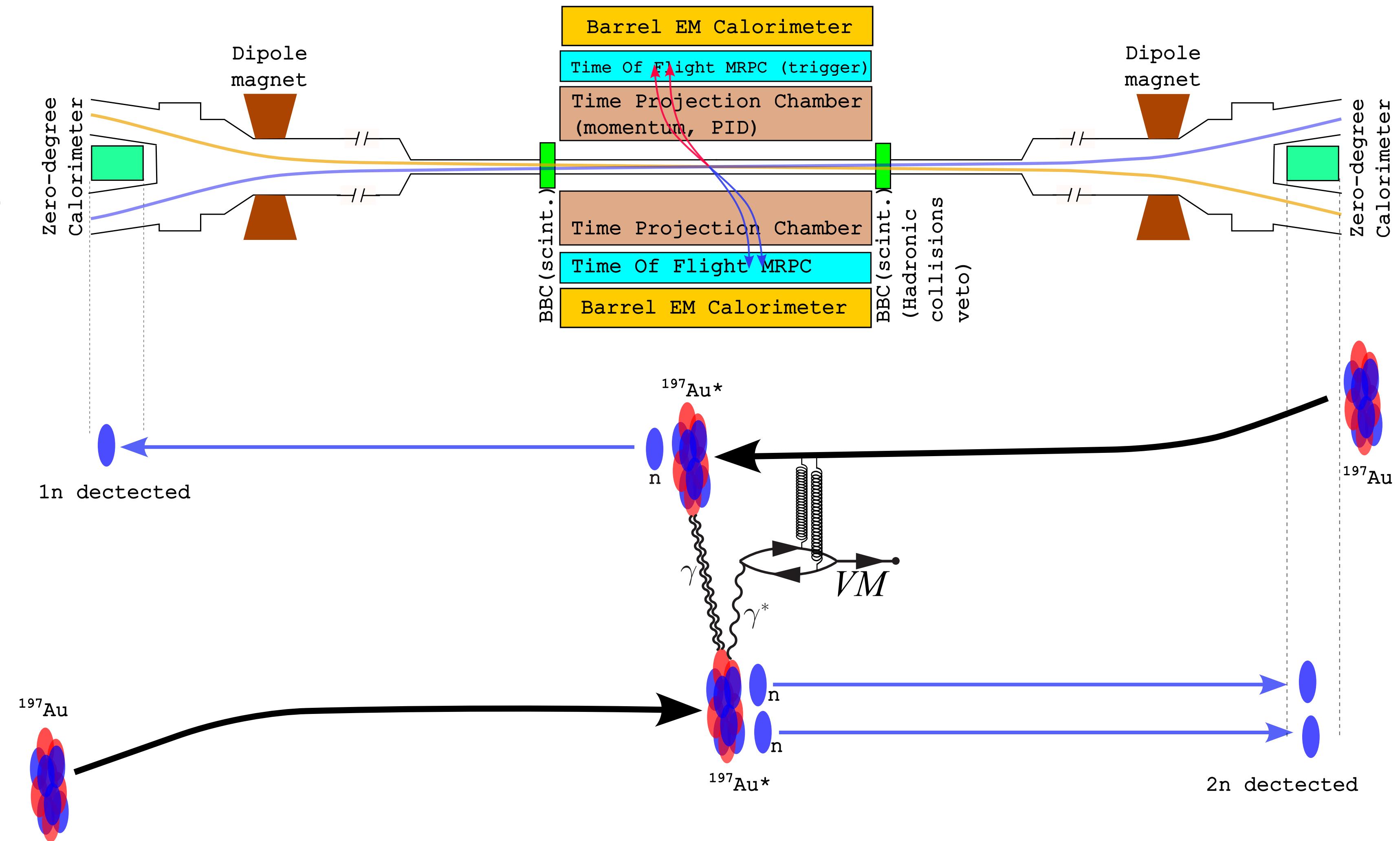
## UPC AS A GREAT PRODUCTION TOOL

- ▶ Heavy Ion Collisions - large charge => large photon flux => large production cross section, accompanied by Coulomb excitation of the beam particles which emit neutrons => easy to trigger
- ▶ coherent (on nucleus) and incoherent (on nucleons)
- ▶ coherent photo production
  - ▶ final state is exclusive
  - ▶ easy to separate the signal from background



# STAR EXPERIMENTAL SETUP (UPC RELEVANT DETECTORS ONLY)

- ▶ Solenoidal Tracker At RHIC
- ▶ central rapidity coverage
  - ▶  $(-1,1) \xrightarrow{2019} (-1.5,1.5)$
- ▶ neutron tagging
- ▶ charged hadrons PID
  - ▶ plus electron calorimetry including decay topology
- ▶ veto particles in the UPCs rapidity gap regions



# DATASETS, LUMINOSITIES AND EVENT SELECTION

## ▶ Online Event Selection ( "UPC\_main" trigger)

- ▶ number of neutrons on each side
  - ▶ 1 - 4.5 (Run 10)
  - ▶ 1 - 3.5 (Run 11)
  - ▶ 1 - 11 (Run 14)
- ▶  $2 \leq \text{Track Multiplicity} \leq 6$
- ▶ UPC Rapidity Gap Veto

	Run14	Run11	Run10
$L^{-1}$ [ $\mu\text{b}^{-1}$ ]	787	523	926
$L^{-1}$ fraction in $ v_z  < 130$	0.664	0.813	0.764

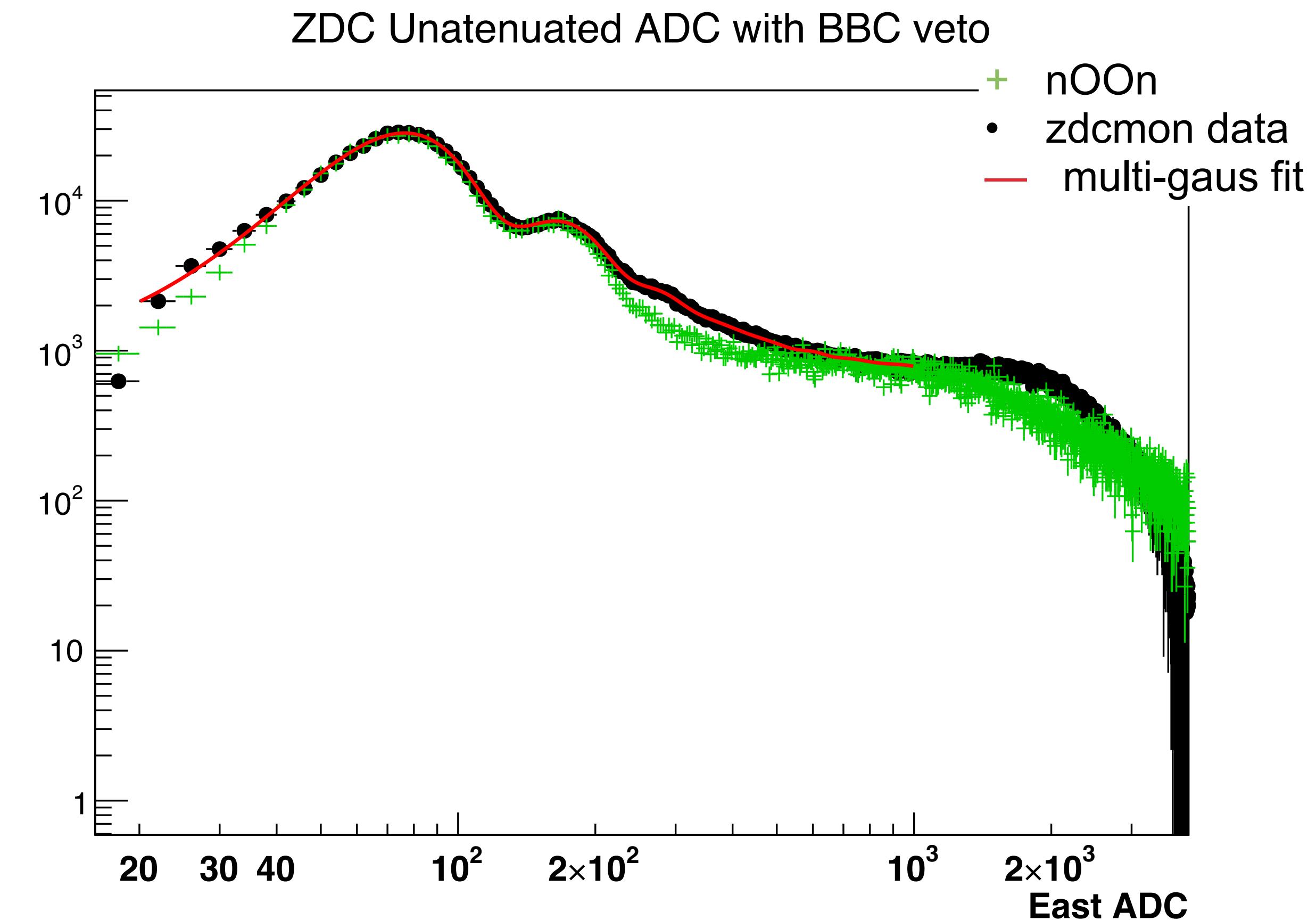
## ▶ Offline Event Selection (analysis)

- ▶  $|\text{Z-Pos. of collision vertex}| < 130 \text{ cm}$  from acceptance center
- ▶ Track DCA to the vertex  $< 3\text{cm}$
- ▶ TPC PID using  $dE/dx$ : normalized  $|\sigma_\pi| < 3$
- ▶ #TPC track hits  $> 15$
- ▶  $p_T(\pi^+\pi^-) < 0.15 \text{ GeV}/c$  or  $p_T(\pi^+\pi^-\pi^+\pi^-) < 0.15 \text{ GeV}/c$

$\sigma_{\text{TRIGGER}} \rightarrow \sigma_{\text{nn}}$

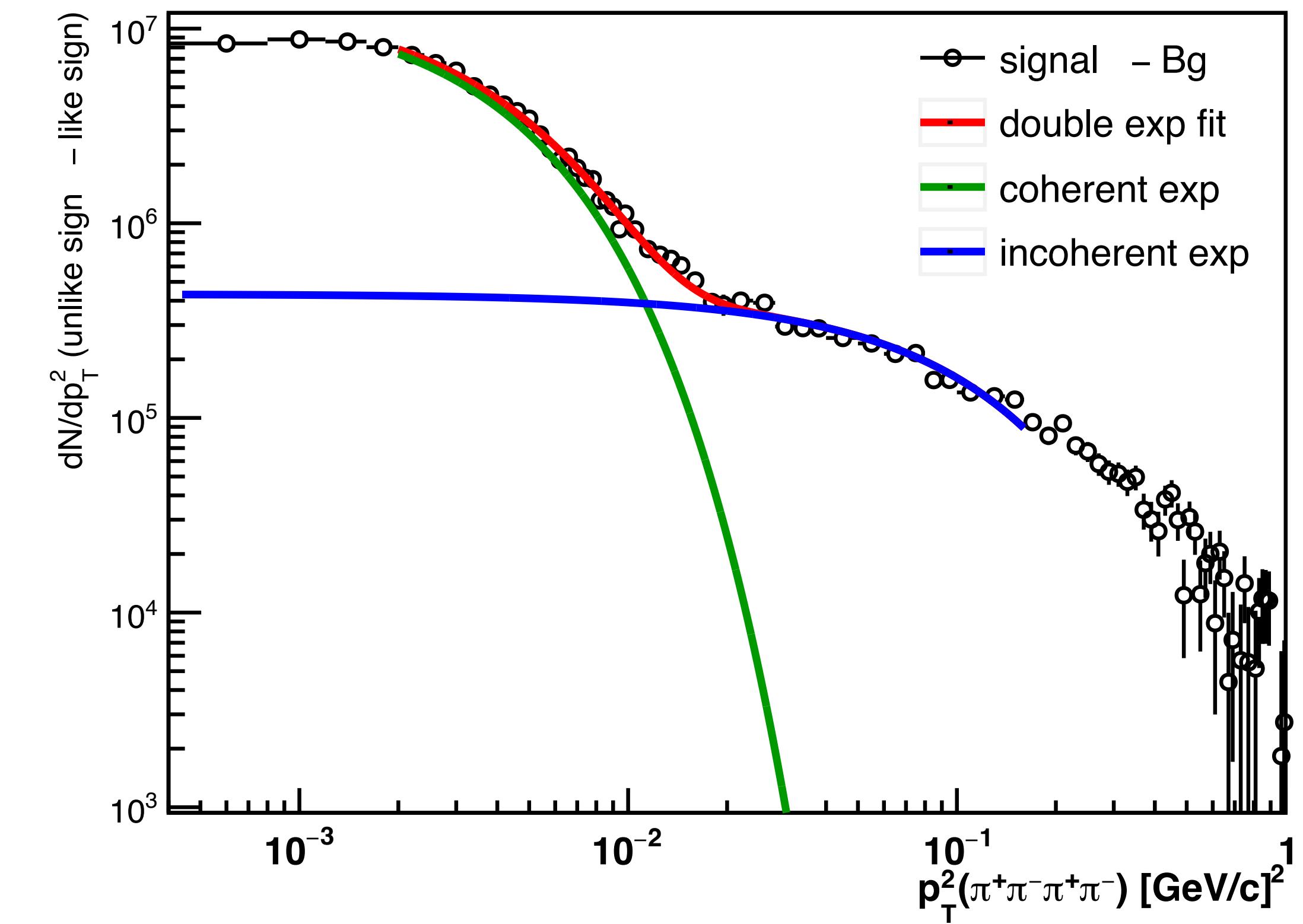
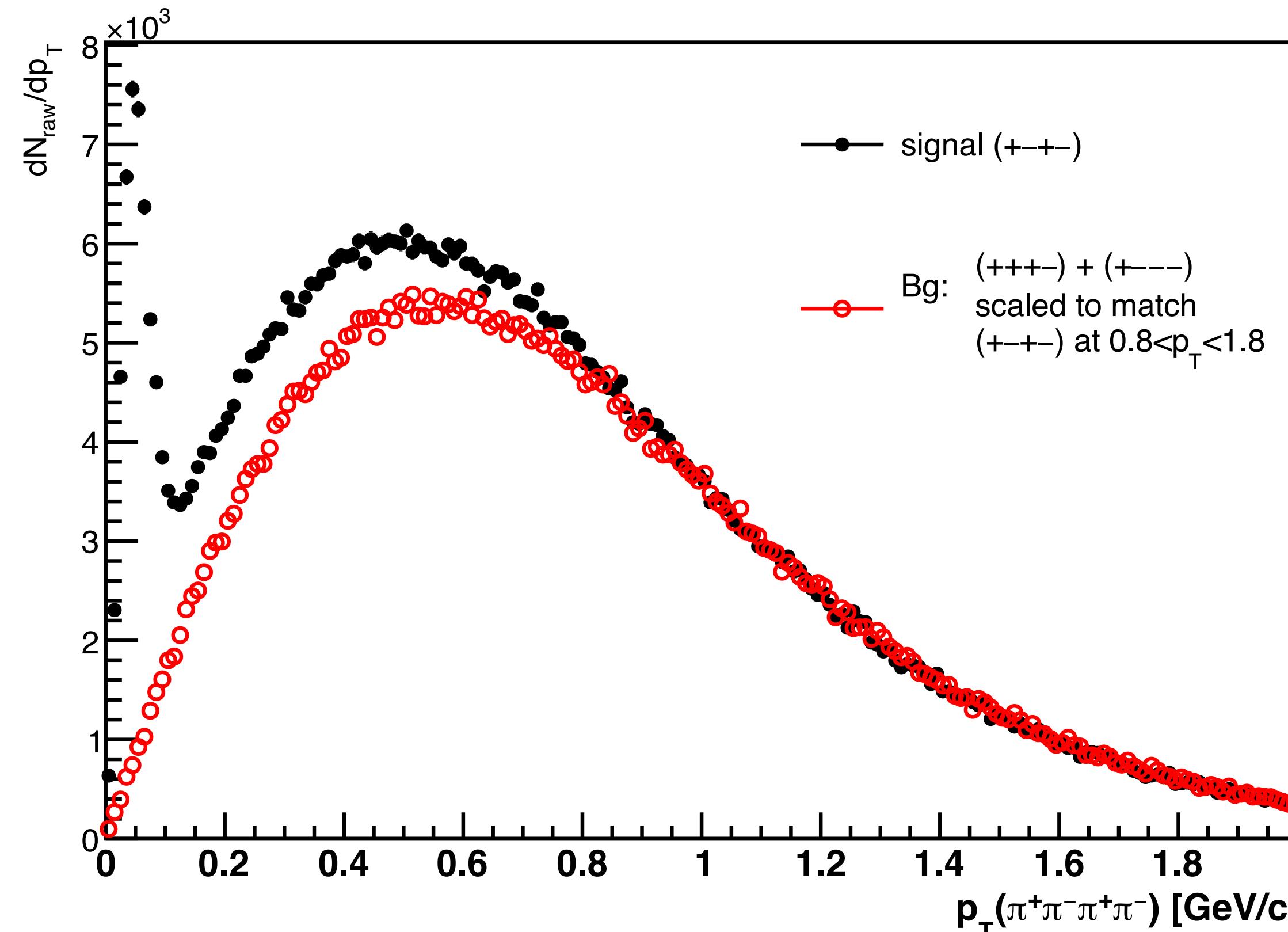
- ▶ UPC\_main trigger does not see whole  $\sigma_{\text{nn}}$
- ▶ STAR added a special trigger in Run14 called "zdcmon" that was just ZDC coincidence (no cut on ADC, no hadron veto)
  - ▶ we analyzed these data and compare with UPC\_main to what fraction of  $\sigma_{\text{nn}}$  the UPC\_main trigger "see" in each year.
- ▶ nOOn model [CPC 253 107181] of neutron production can predict neutron distribution in heavy ion collisions

	Run14	Run11	Run10
▶ UPC_main trigger	1-11n	1-3.5n	1-4.5n
▶ fraction from zdcmon data	56.74%	37.72%	41.58%
▶ fraction from nOOn	63.16%	39.52%	43.52%



we used average of  
zdcmon data and nOOn

# INCOHERENT CONTRIBUTION IN $P_T < 0.15$ RANGE



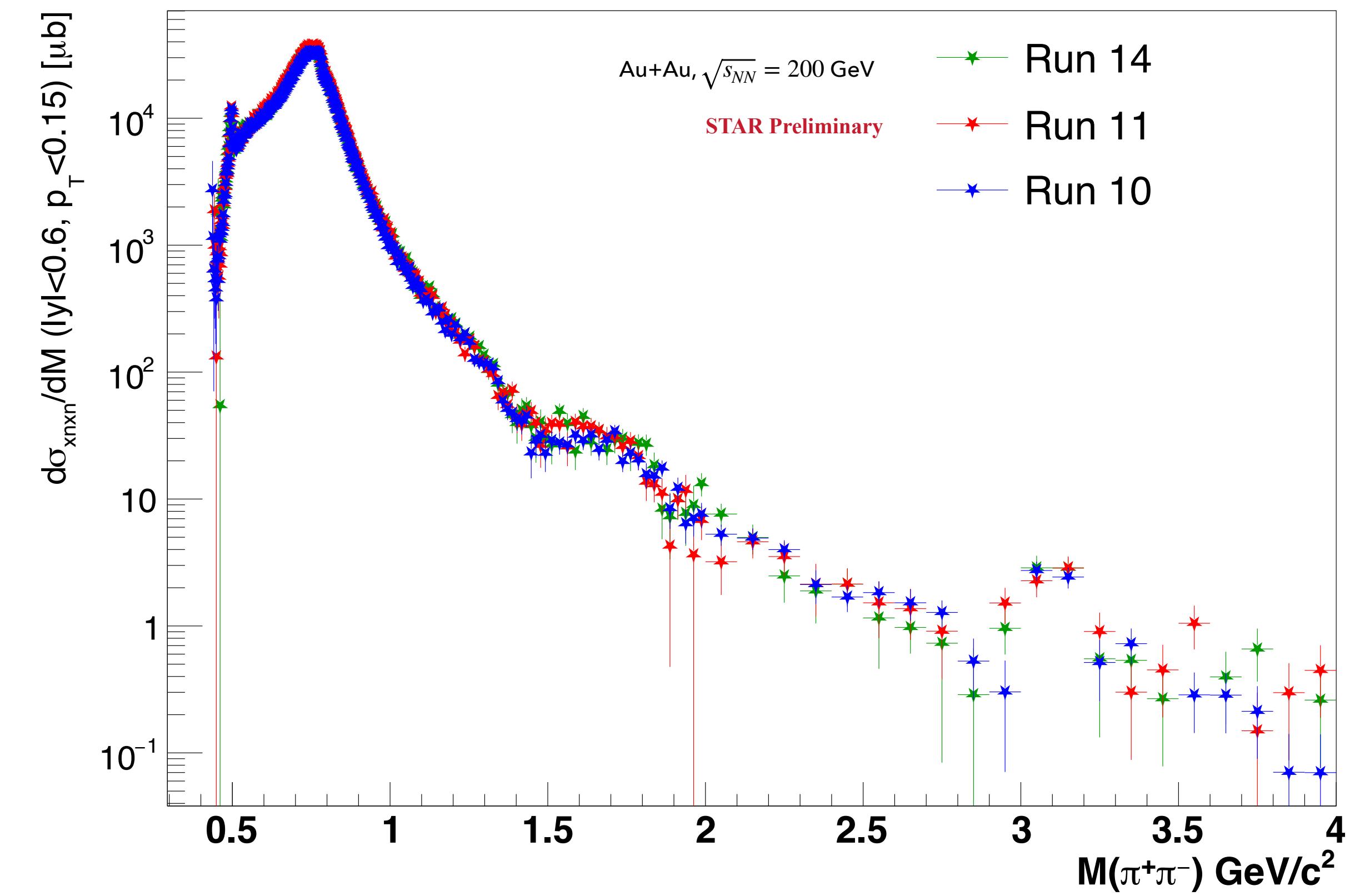
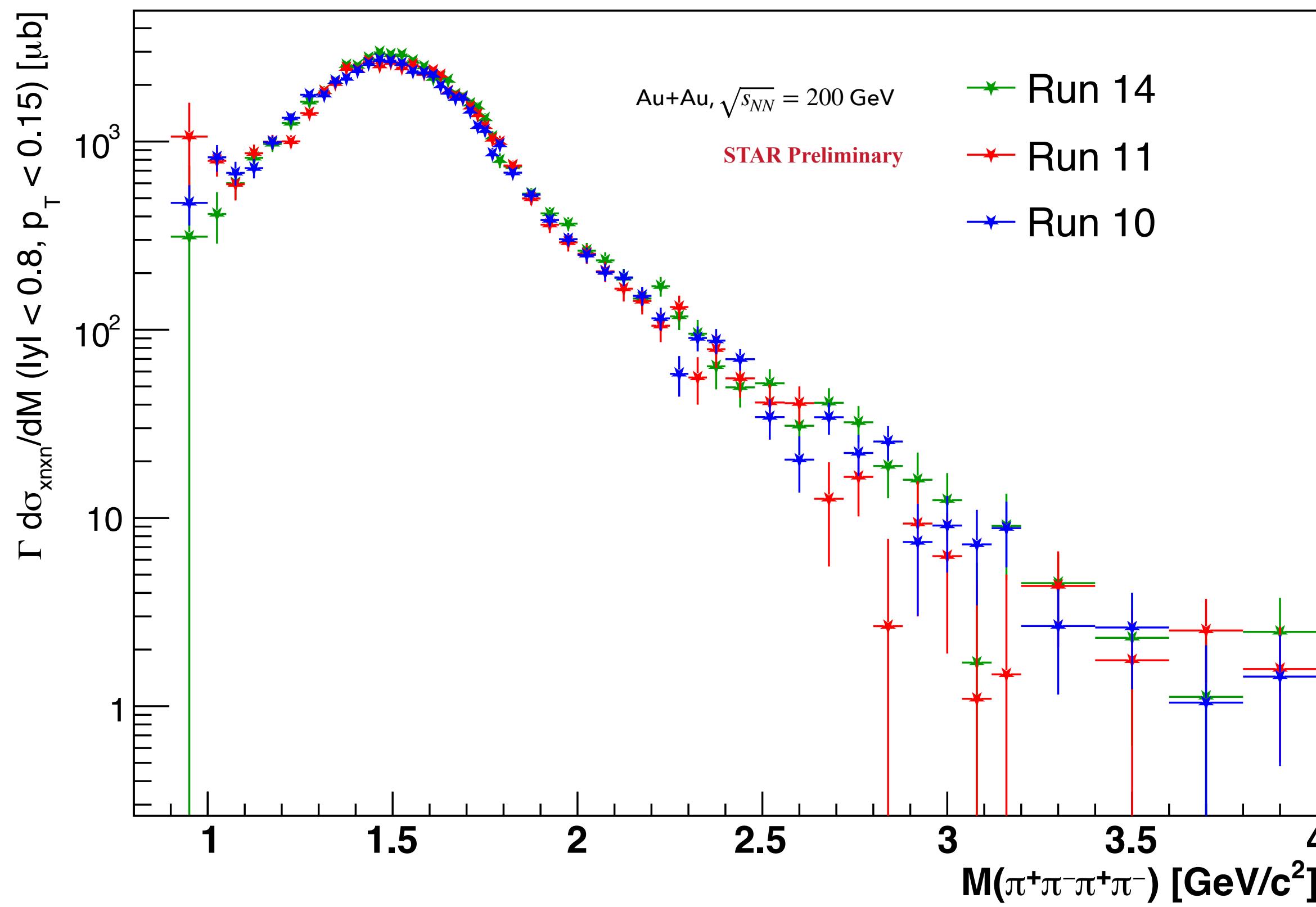
Incoherent fraction for  $0.9 < M(4\pi) < 4.0$  GeV : 0.181

Incoherent fraction for  $1.5 < M(4\pi) < 2.5$  GeV : 0.160

Incoherent fraction for  $0.6 < M(2\pi) < 2.8$  GeV : 0.064

Incoherent fraction for  $1.5 < M(2\pi) < 2.5$  GeV : 0.252

# MASS SPECTRA OF $\pi^+\pi^-$ AND $\pi^+\pi^-\pi^+\pi^-$ (BOTH $P_T < 0.15 \text{ GeV}/c$ )

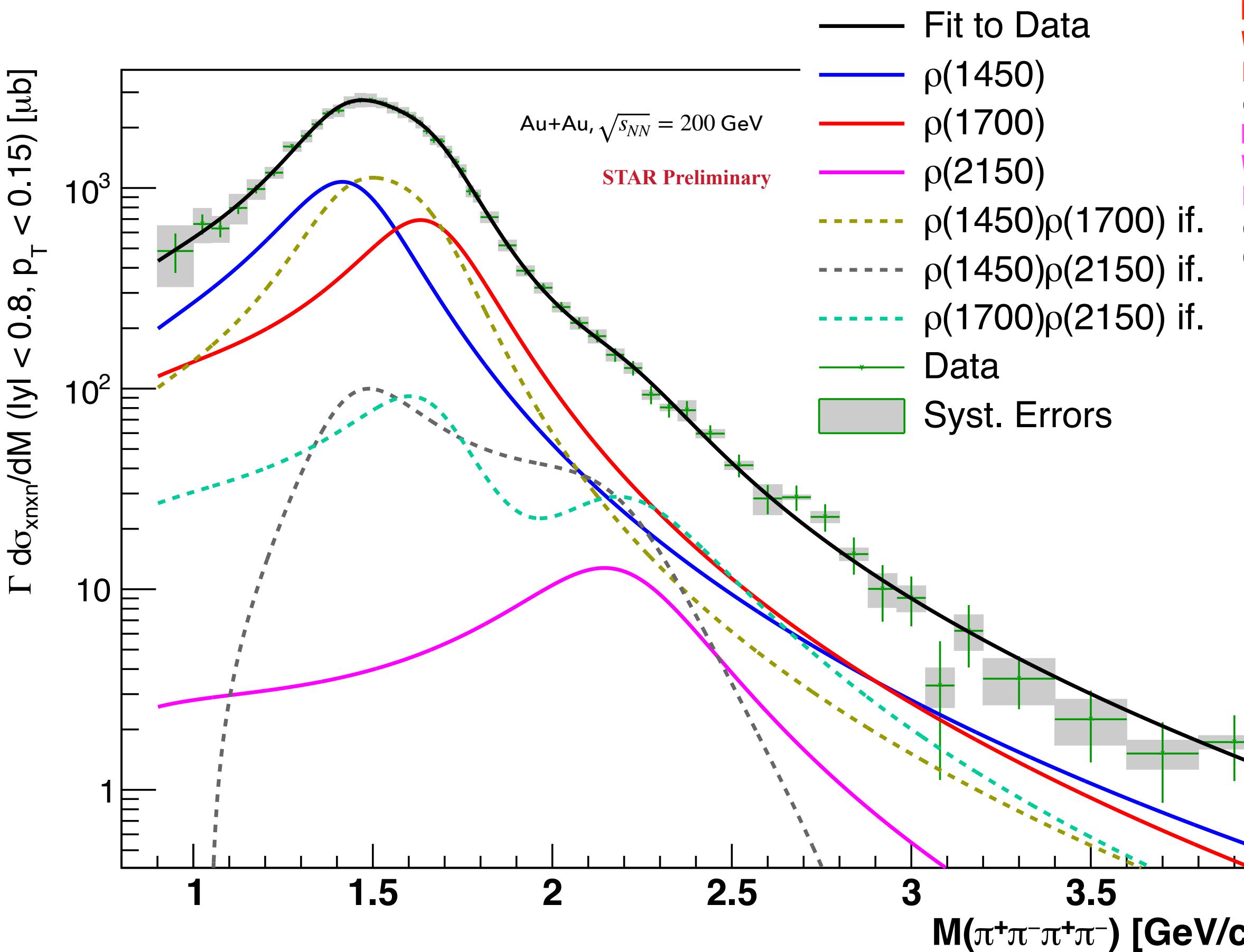


- Main “sanity check” - cross section consistent between datasets

# MASS( $\pi^+\pi^- \pi^+\pi^-$ ) AND RAPIDITY COMBINED FROM ALL RUNS

$$\Gamma \frac{d\sigma}{dM} = A^2 |BW(\rho_{1450})|^2 + B^2 |BW(\rho_{1700})|^2 + C^2 |BW(\rho_{2150})|^2 + 2\sqrt{AB} \Re [BW^*(\rho_{1450})BW(\rho_{1700})e^{i\delta(1700)}] + 2\sqrt{BC} \Re [BW^*(\rho_{1700})BW(\rho_{2150})e^{i\delta(2150)}] + 2\sqrt{AC} \Re [BW^*(\rho_{1450})BW(\rho_{2150})e^{i(\delta(1700)+\delta(2150))}]$$

$$BW(\rho) \equiv \left(\frac{M_\rho}{M}\right)^{\text{RSP}} \frac{\sqrt{WM_\rho}}{M^2 - M_\rho^2 + iM_\rho W}, \quad W = W_\rho \frac{M_\rho}{M} \left(\frac{M^2 - 16m_\pi^2}{M_\rho^2 - 16m_\pi^2}\right)^{3/2}$$



PARAMETER	VALUE	ERROR
Mass $\rho(1450)$	1486	20.3
Width $\rho(1450)$	400.3	30
RSP1	2.17	0.27
Mass $\rho(1700)$	1701	15.4
Width $\rho(1700)$	399.6	34.5
RSP2	2.39	0.37
$\delta(1700)$	1.22	0.38
Mass $\rho(2150)$	2247	91.2
Width $\rho(2150)$	570	fixed
RSP3	2.36	2.50
$\delta(2150)$	0.50	0.48
Chi2/ndf	33.0882/41	

$$\sigma_{4\pi,xn,xn}^{\text{coh}}(|y| < 0.8) = 1348 \pm 16_{\text{stat.}} \pm 20_{\text{syst.}} \pm 162_{\text{norm.}}$$

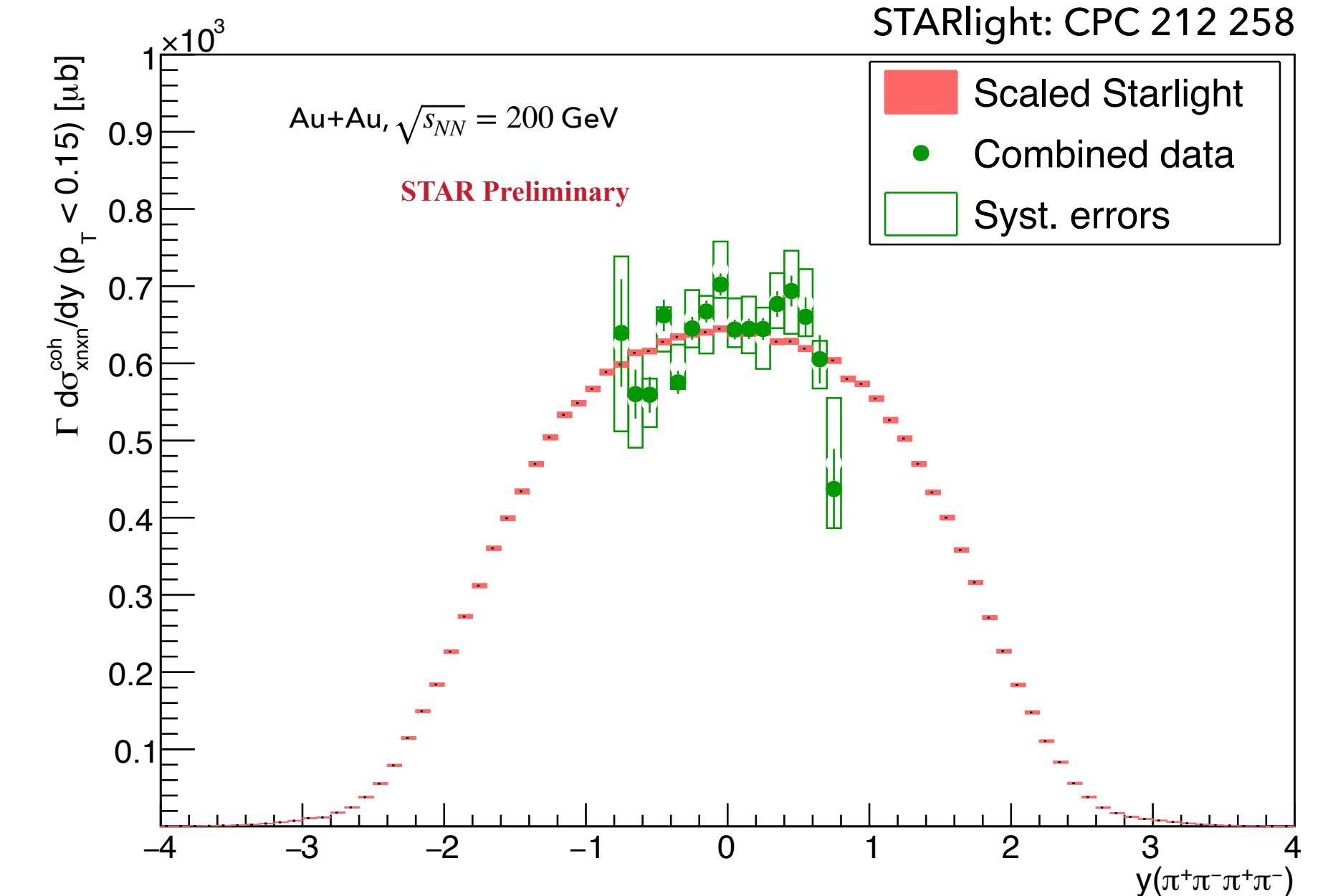
$$\sigma_{4\pi,xn,xn}^{\text{coh}} = 2941 \pm 34_{\text{stat.}} \pm 45_{\text{syst.}} \pm 352_{\text{norm.}}$$

$$\Gamma \sigma_{\rho 1450,xn,xn}^{\text{coh}}(|y| < 0.8) = 426 \pm 162_{\text{stat.}} \pm 202_{\text{syst.}} \pm 51_{\text{norm.}}$$

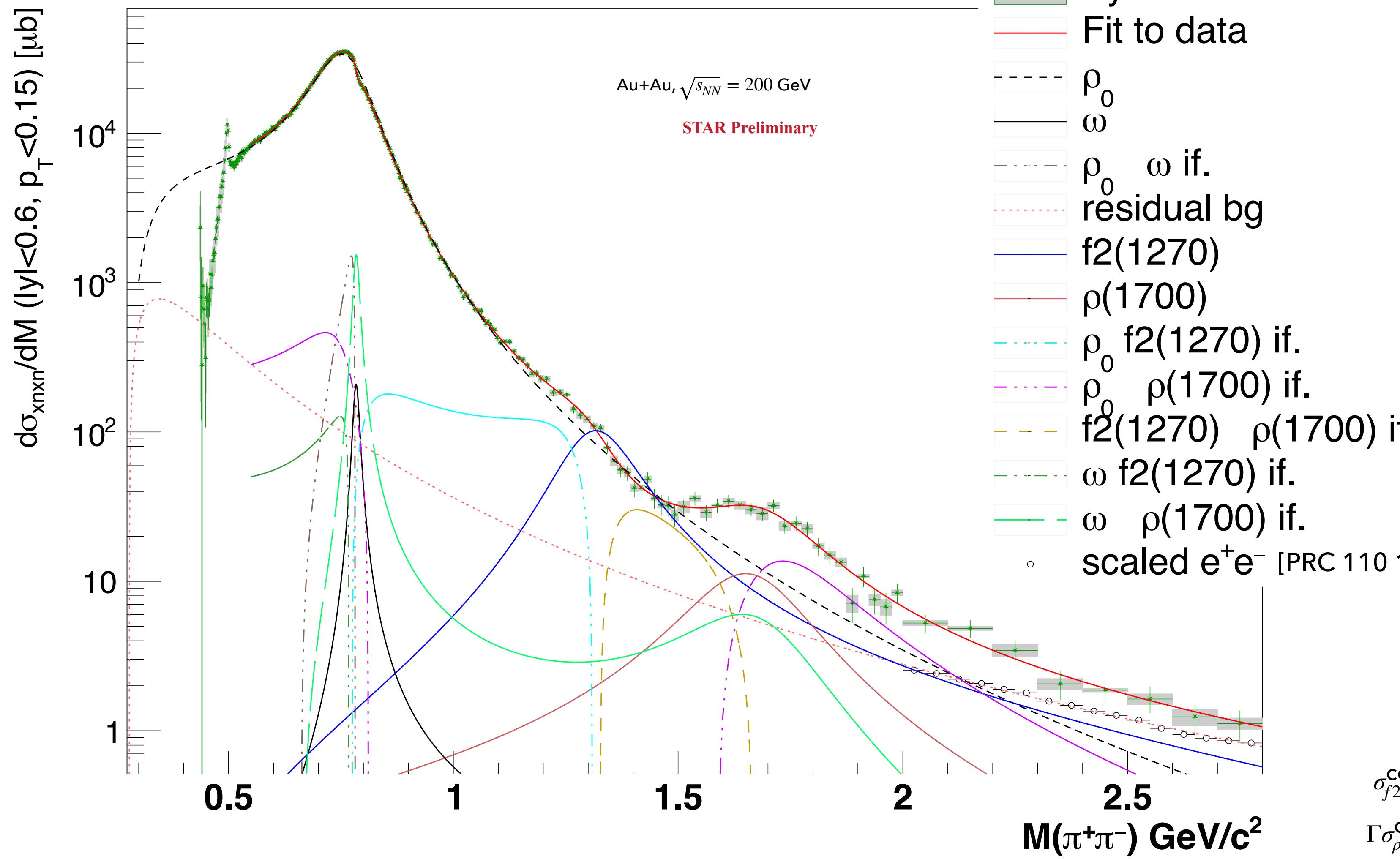
$$\Gamma \sigma_{\rho 1700,xn,xn}^{\text{coh}}(|y| < 0.8) = 310 \pm 148_{\text{stat.}} \pm 169_{\text{syst.}} \pm 37_{\text{norm.}}$$

$$\Gamma \sigma_{\rho 2150,xn,xn}^{\text{coh}}(|y| < 0.8) = 8.9 \pm 6.1_{\text{stat.}} \pm 1.3_{\text{syst.}} \pm 1.1_{\text{norm.}}$$

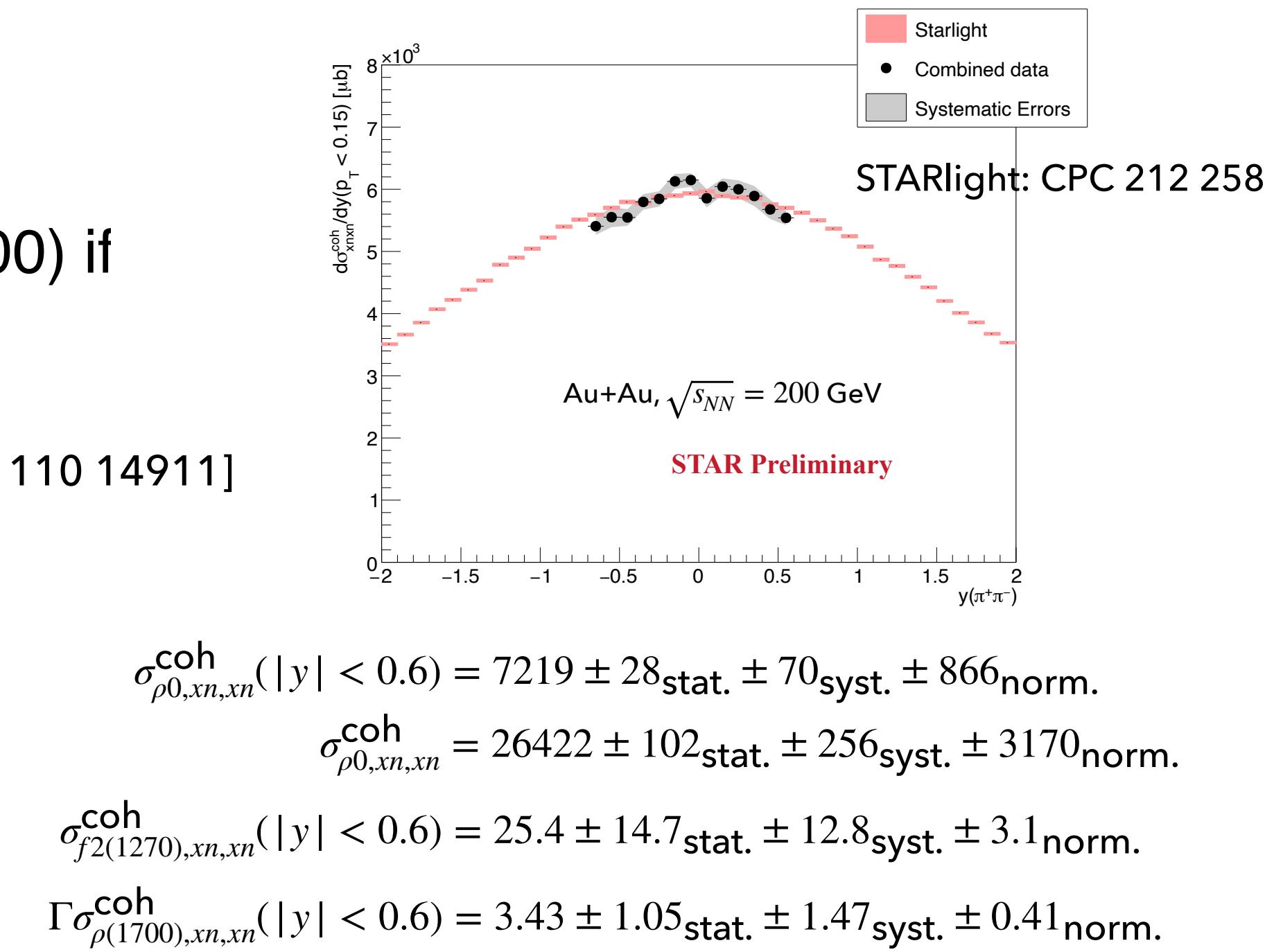
to extrapolate to full rapidity - multiply by 2.18



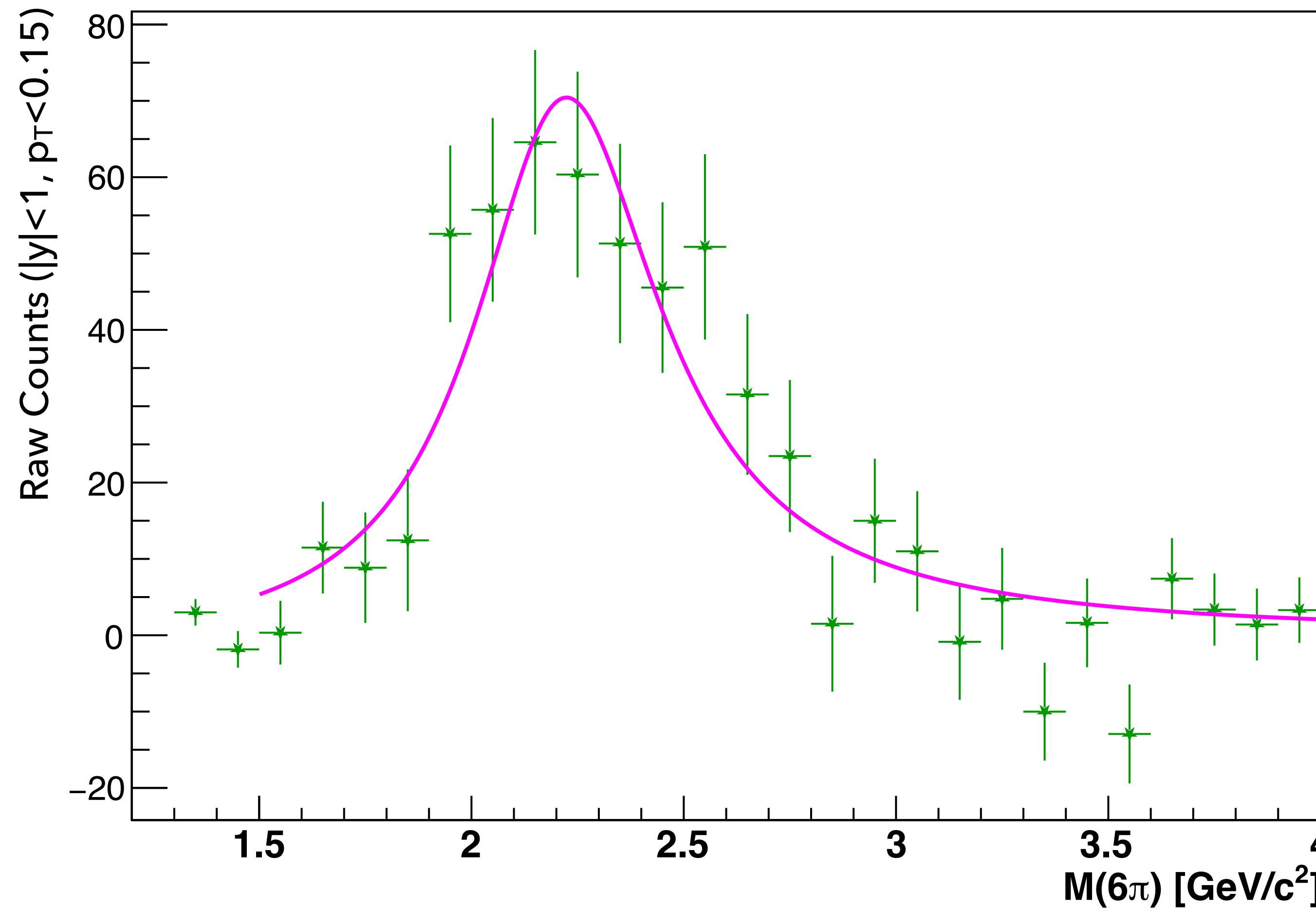
# MASS OF $\pi^+\pi^-$ COMBINED FROM ALL RUNS



PARAMETER	VALUE	ERROR
Mass $\rho_0$	774.83	0.30
Width $\rho_0$	146.39	0.57
RSP $\rho_0$	2.76	0.02
Mass $\omega$	783.82	0.50
Width $\omega$	16.13	0.85
$\delta(\omega)$	1.97	0.17
Mass $f_2$	1339.26	19.63
Width $f_2$	210.35	38.30
Mass $\rho(1700)$	1700.65	22.71
Width $\rho(1700)$	317.79	46.75
$\delta(f_2)$	3.12	0.18
$\delta(1700)$	0.38	0.18
Chi2/ndf	282.248/260	



# 6 $\pi$ PHOTO PRODUCTION RESULTS



PARAMETER	VALUE	ERROR
Mass $\rho(2150)$	2276	57.8
Width $\rho(2150)$	573	84.5
RSP $\rho(2150)$	0.82	0.42
Chi2/ndf	27.9946/21	

First observation at STAR

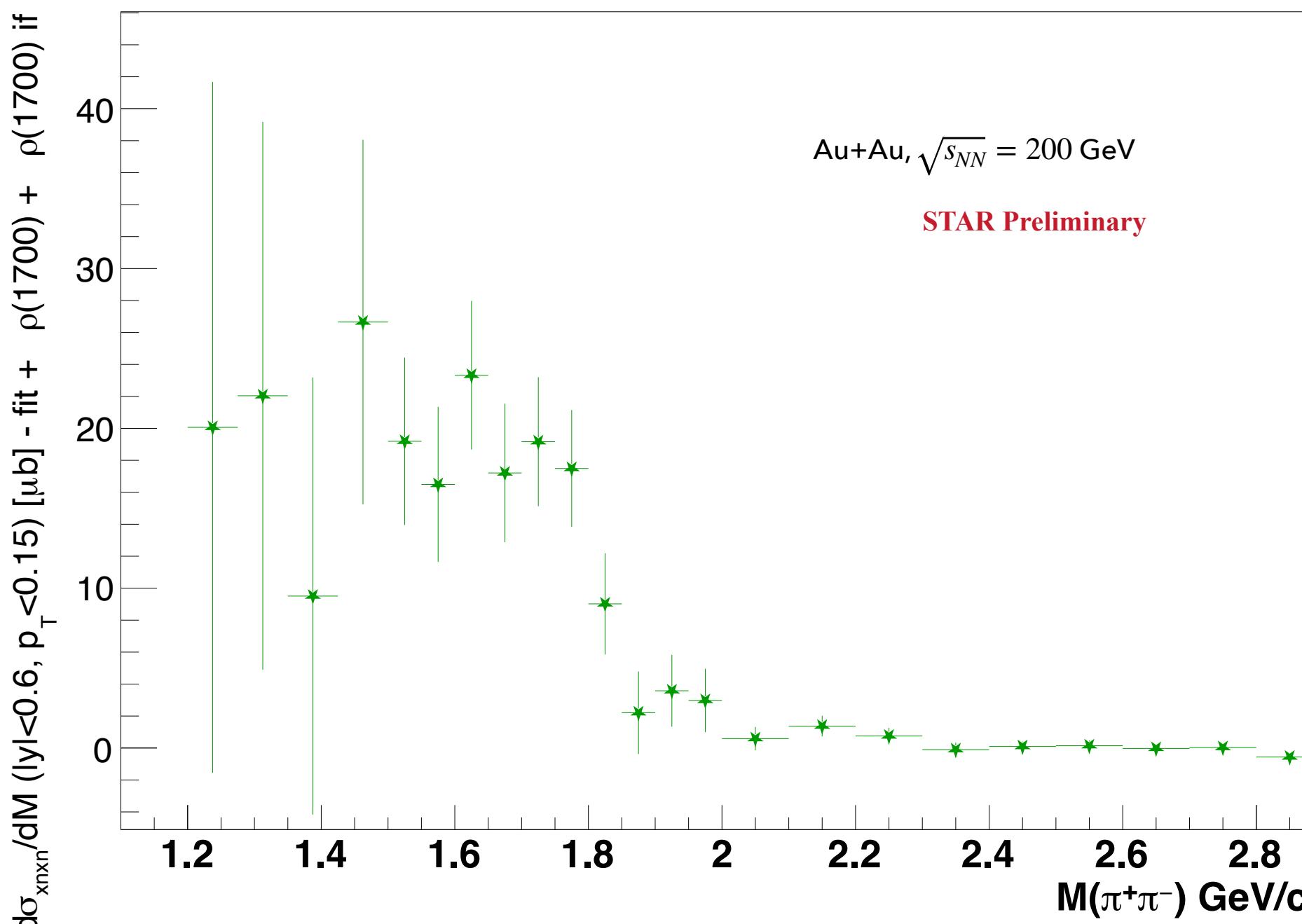
# RATIO OF THE BRANCHING FRACTIONS OF THE RHO1700 TO 2PI AND 4PI

- comparison of yields directly from Breit-Wigner functions

- $\Gamma_{2\pi}\sigma_{\rho_{1700},xn,xn}^{\text{coh}} = 4.53 \pm 1.38_{\text{stat.}} \pm 1.93_{\text{syst.}}$  in  $|y| < 0.8$  from  $\pi^+\pi^-$
- $\Gamma_{4\pi}\sigma_{\rho_{1700},xn,xn}^{\text{coh}} = 328 \pm 161_{\text{stat.}} \pm 178_{\text{syst.}}$  in  $|y| < 0.8$  from  $\pi^+\pi^-\pi^+\pi^-$

$$\Gamma_{2\pi}/\Gamma_{4\pi}(\rho_{1700}) = 1.38 \pm 0.80_{\text{stat.}} \pm 0.95_{\text{syst.}} \%$$

- an alternative method using an excessive yield in  $\pi^+\pi^-$  and yield in  $\pi^+\pi^-\pi^+\pi^-$  in the mass window from 1.5 to 2.5 GeV/c<sup>2</sup> - a good proxy for  $\rho_{1700}$
- the excessive yield in  $\pi^+\pi^-$  can be calculated as  $\pi^+\pi^-$  data – components of the fit function excluding  $\rho_{1700}$  Breit-Wigner and its interference



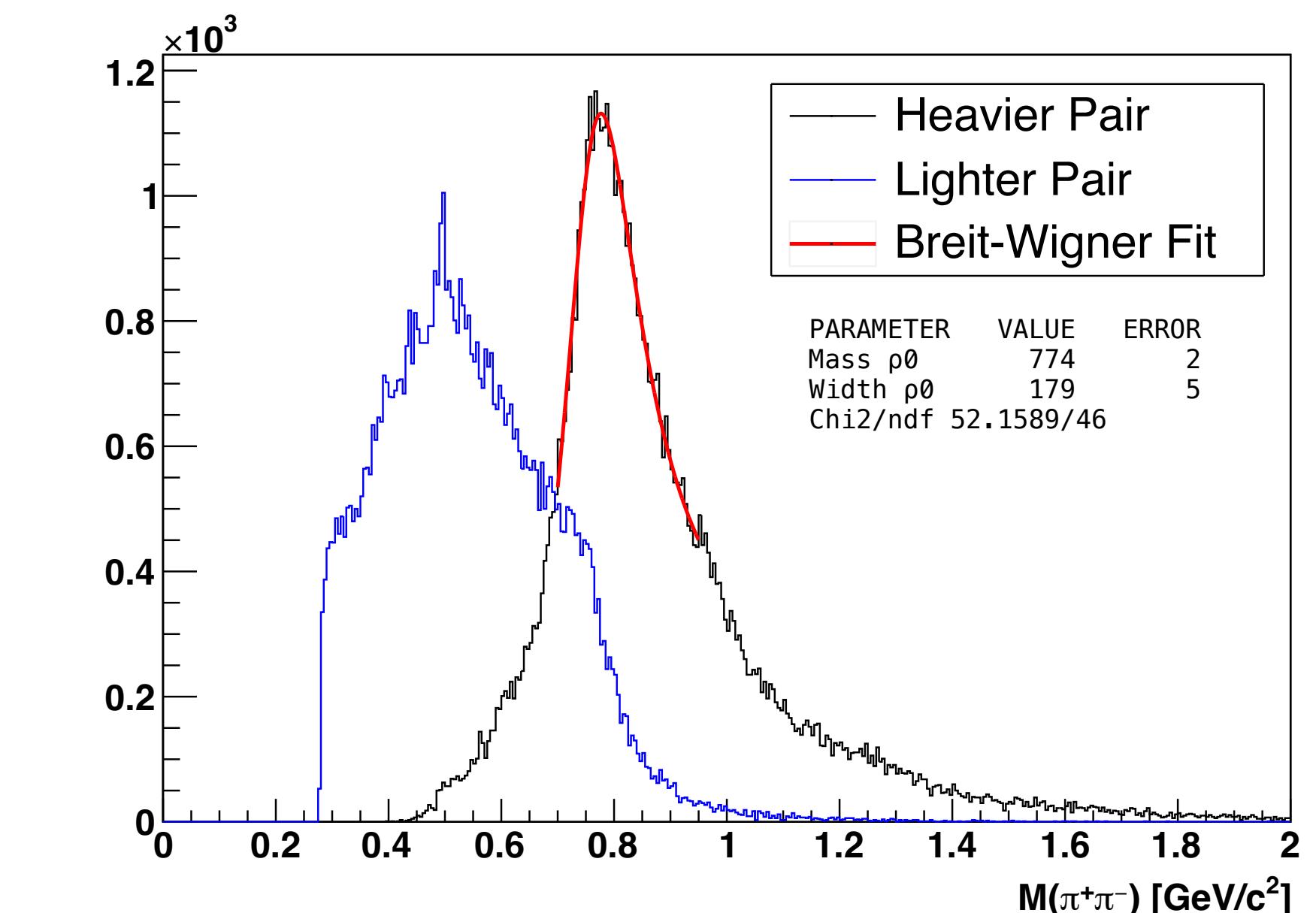
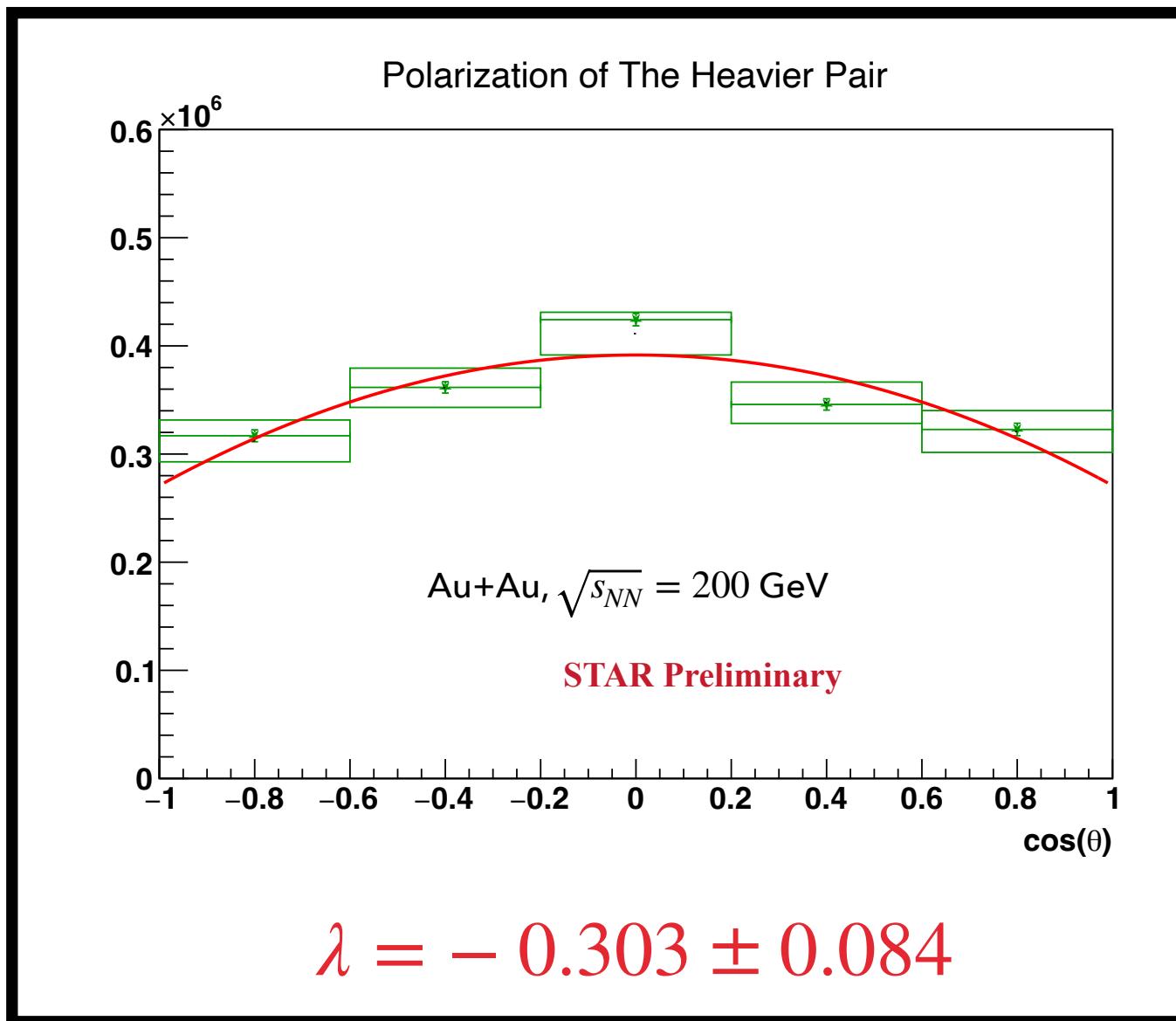
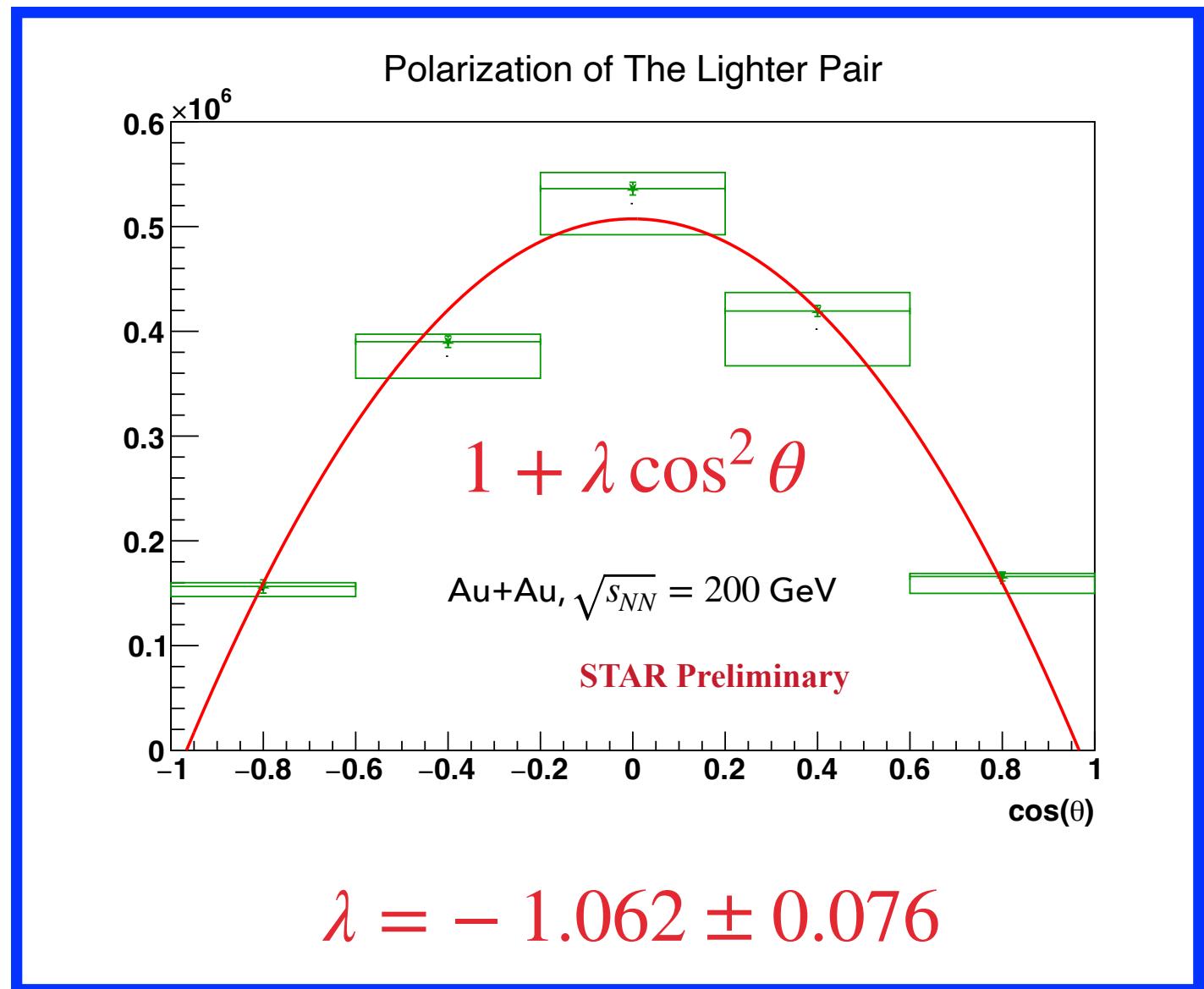
Excess in  $1.5 < \text{Mass} < 2.5 \quad |y| < 0.8 = 6.72 \pm 0.61_{\text{stat.}} \pm 0.33_{\text{syst.}}$

 $\sigma_{4\pi,xn,xn}^{\text{coh}}(|y| < 0.8, 1.5 < M < 2.5) = 612 \pm 8_{\text{stat.}} \pm 17_{\text{syst.}}$ 

$$\Gamma_{2\pi}/\Gamma_{4\pi}(|y| < 0.8, 1.5 < M < 2.5) = 1.10 \pm 0.10_{\text{stat.}} \pm 0.06_{\text{syst.}} \%$$

		PRC 81 044901
$\sigma_{4\pi,xn,xn}^{\text{coh}}/\sigma_{\rho_0,xn,xn}^{\text{coh}}( y  < 0.8) = 14.1 \pm 0.4_{\text{stat.}} \pm 0.5_{\text{sys.}} \%$		$16.4 \pm 1.0_{\text{stat.}} \pm 5.2_{\text{syst.}} \%$
$\sigma_{4\pi,xn,xn}^{\text{coh}}/\sigma_{\rho_0,xn,xn}^{\text{coh}} = 11.1 \pm 0.3_{\text{stat.}} \pm 0.4_{\text{sys.}} \%$		$13.4 \pm 0.8_{\text{stat.}} \pm 4.4_{\text{syst.}} \%$

## J=1 TRANSFER TO COS $\theta$ ANISOTROPY



$$\cos \theta = \frac{\overrightarrow{P_{\pi^+\pi^-}} \cdot \overrightarrow{p_\pi^*}}{\|\overrightarrow{P_{\pi^+\pi^-}}\| \|\overrightarrow{p_\pi^*}\|}$$

- ▶ Breit-Wigner fit to the heavier pair mass point to  $\rho_0$  meson.
- ▶ Heavy pair polarized, but not fully
- ▶ Light pair likely contains  $f_0(500)$  resonance. But this resonance is supposed to be a scalar meson while its polarization indicates vector meson.

## SUMMARY

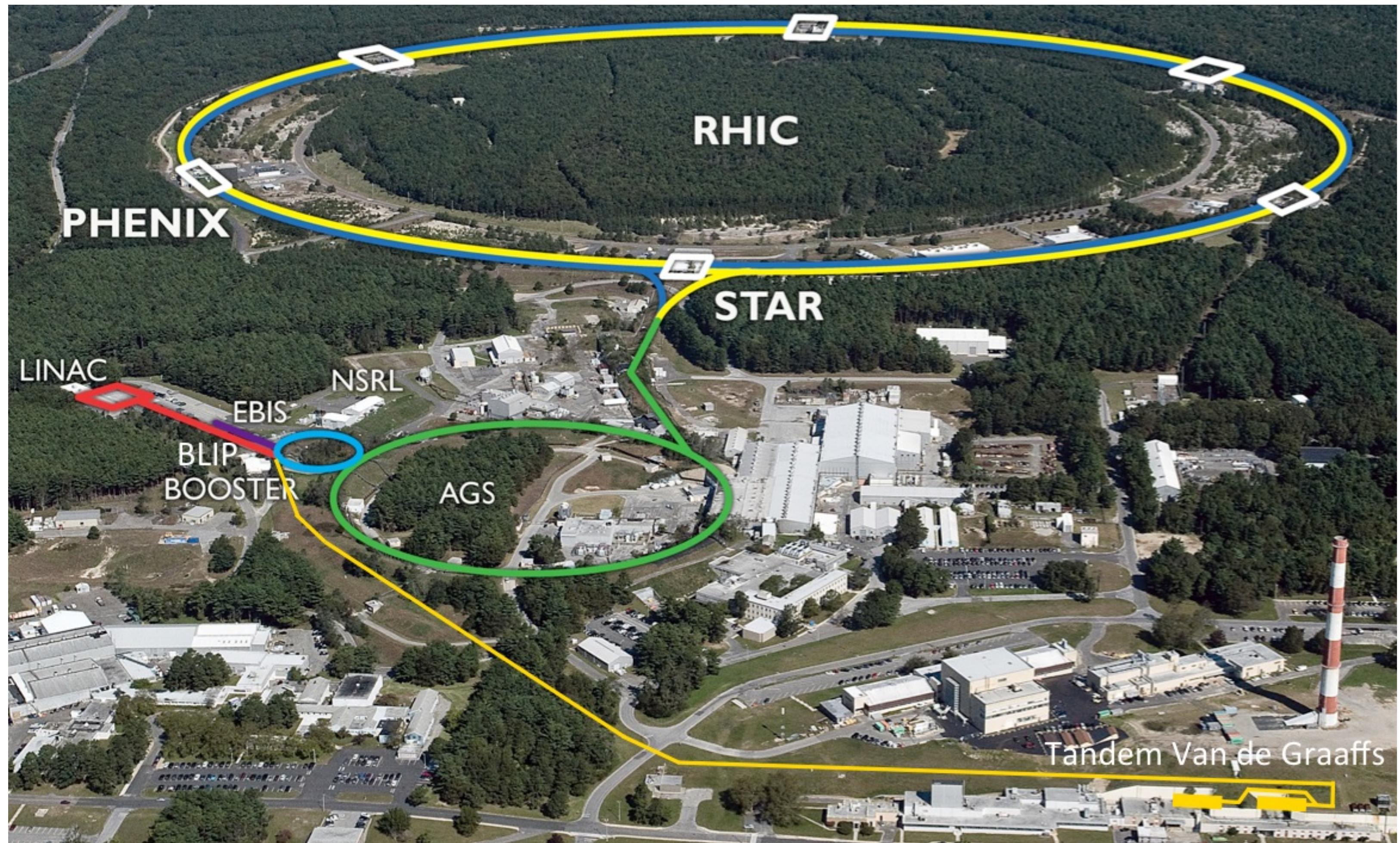
- ▶ STAR presented a precise measurement of  $\pi^+\pi^-\pi^+\pi^-$  and  $\pi^+\pi^-$  photo-production in Au+Au collisions at  $\sqrt{s_{NN}}=200$  GeV
  - ▶  $\rho_{1450}$  ( $2^3S_1$ ),  $\rho_{1700}$  ( $1^3D_1$ ) clearly and  $\rho_{2150}$  ( $2^3D_1$ ) likely observed in  $\pi^+\pi^-\pi^+\pi^-$  mass spectrum
    - ▶  $\rho_{1450}$  mass and width consistent with the world average
    - ▶  $\rho_{1700}$  mass consistent with the world average, but width larger => more decays modes possible, hybrid state indication?
  - ▶  $f_2(1250)$  and  $\rho_{1700}$  observed in  $\pi^+\pi^-$  mass spectrum
    - ▶  $f_2(1250)$  mass larger - might contain  $\rho_{1450}$ , but we can't separate these states
    - ▶  $\rho_{1700}$  mass and width consistent with the world average, width lower than in  $\pi^+\pi^-\pi^+\pi^-$  => existence of intermediate states in decay
  - ▶  $\rho_{2150}$  ( $2^3D_1$ ) observed in  $\pi^+\pi^-\pi^+\pi^-\pi^+\pi^-$  mass spectrum - first time at STAR
  - ▶  $(\rho_{1700} \rightarrow \pi^+\pi^-)/(\rho_{1700} \rightarrow \pi^+\pi^-\pi^+\pi^-) = 1.10 \pm 0.10_{\text{stat.}} \pm 0.06_{\text{syst.}} \%$  in mid rapidity and
  - ▶  $\sigma_{4\pi,xn,xn}^{\text{coh}}/\sigma_{\rho_0,xn,xn}^{\text{coh}} = 14.1 \pm 0.4_{\text{stat.}} \pm 0.5_{\text{sys.}} \%$  in mid and  $11.1 \pm 0.3_{\text{stat.}} \pm 0.4_{\text{sys.}} \%$  in full rapidity
- ▶  $\pi^+\pi^-\pi^+\pi^-$  states' (all supposed to have  $J=1$ ) decay can be separated to 2  $\pi^+\pi^-$  pairs by their mass
  - ▶ the lighter pair which likely contains  $f_0(500)$ , a scalar meson, decays like fully polarized particle ( $\lambda = -1$ )
  - ▶ the heavier pair whose mass spectrum resembles  $\rho_0$  is partially polarized ( $\lambda = -0.3$ )

# THANK YOU

# BACKUP SLIDES

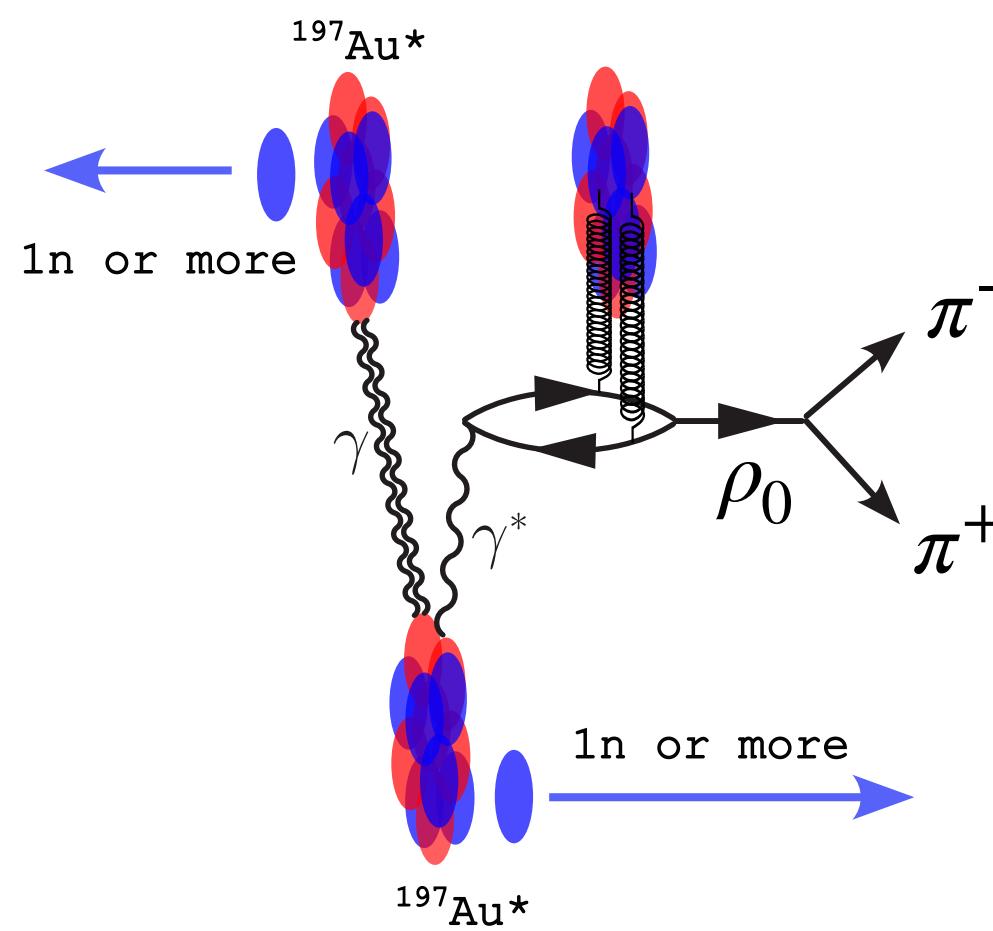
# UTRA-PERIPHERAL COLLISIONS AT RHIC

- ▶ Relativistic Heavy Ion Collider
- ▶ located in Brookhaven National Laboratory (Long Island, USA)
- ▶ different species, energy, and proton polarization

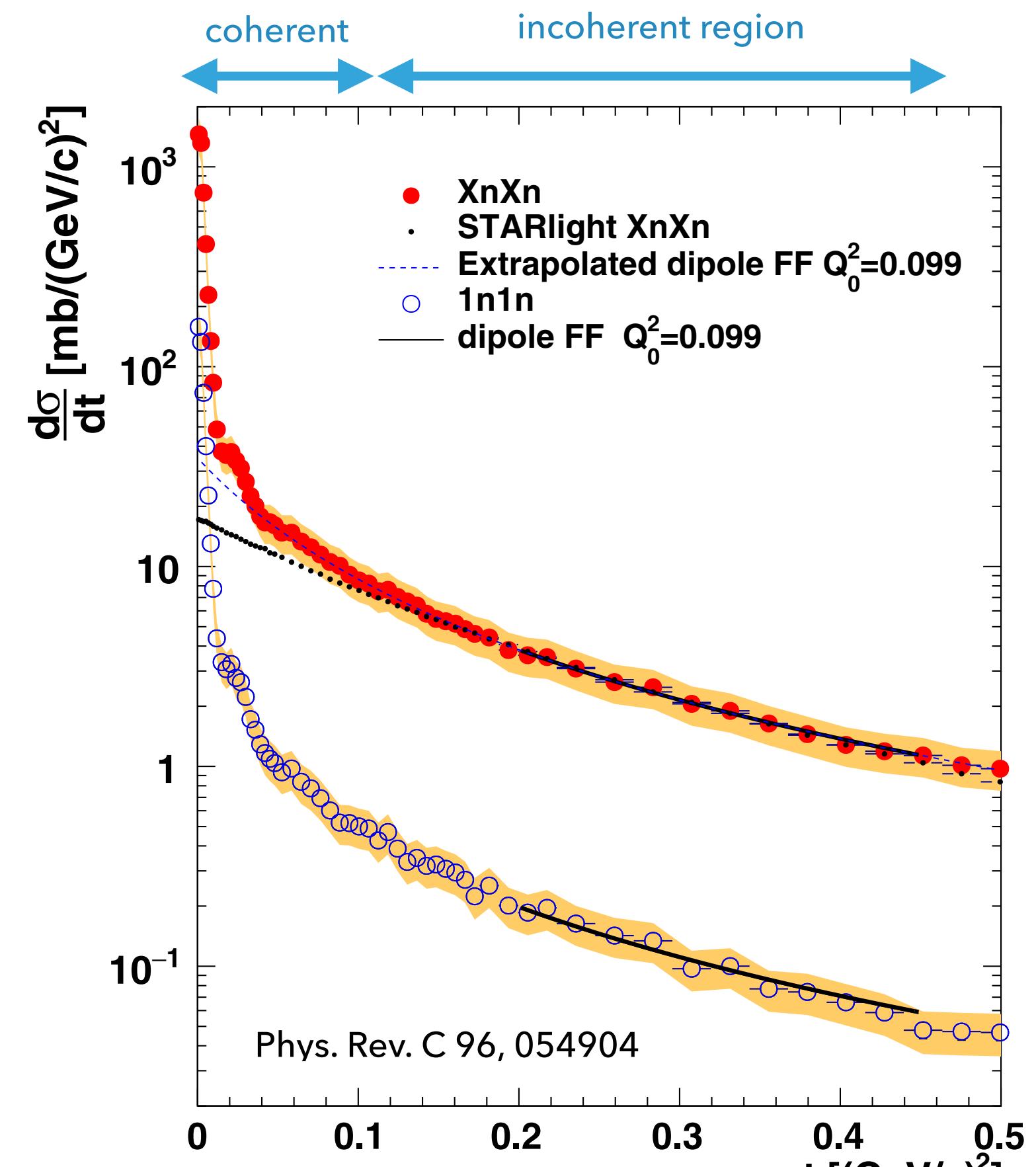
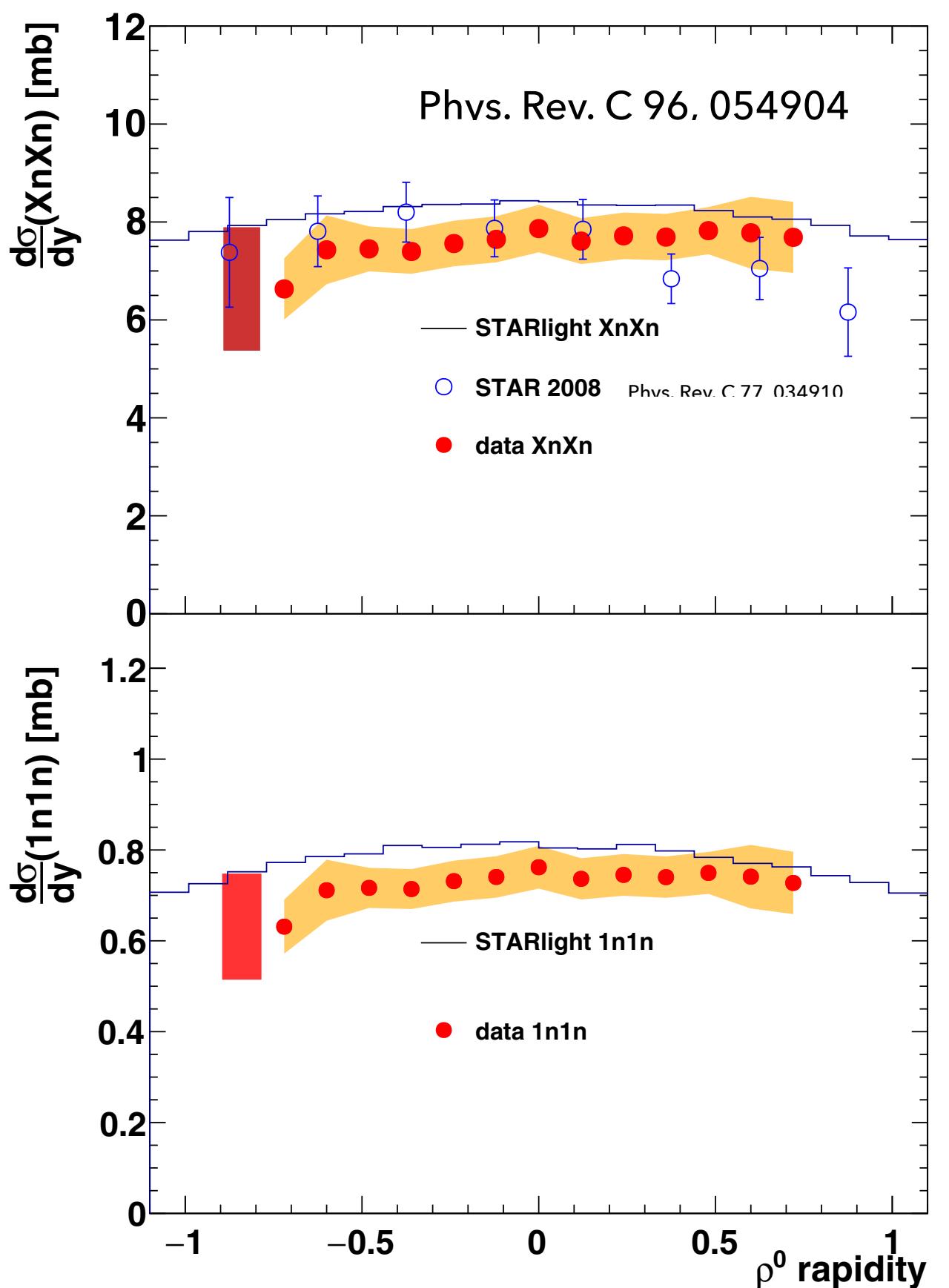


$U^{238}$ ,  $Au^{197}$ ,  $Zr^{96}$ ,  $Ru^{96}$ ,  $d^2$  at 200 GeV and  $pp$  at 510 GeV

## $\rho_0$ CROSS SECTION



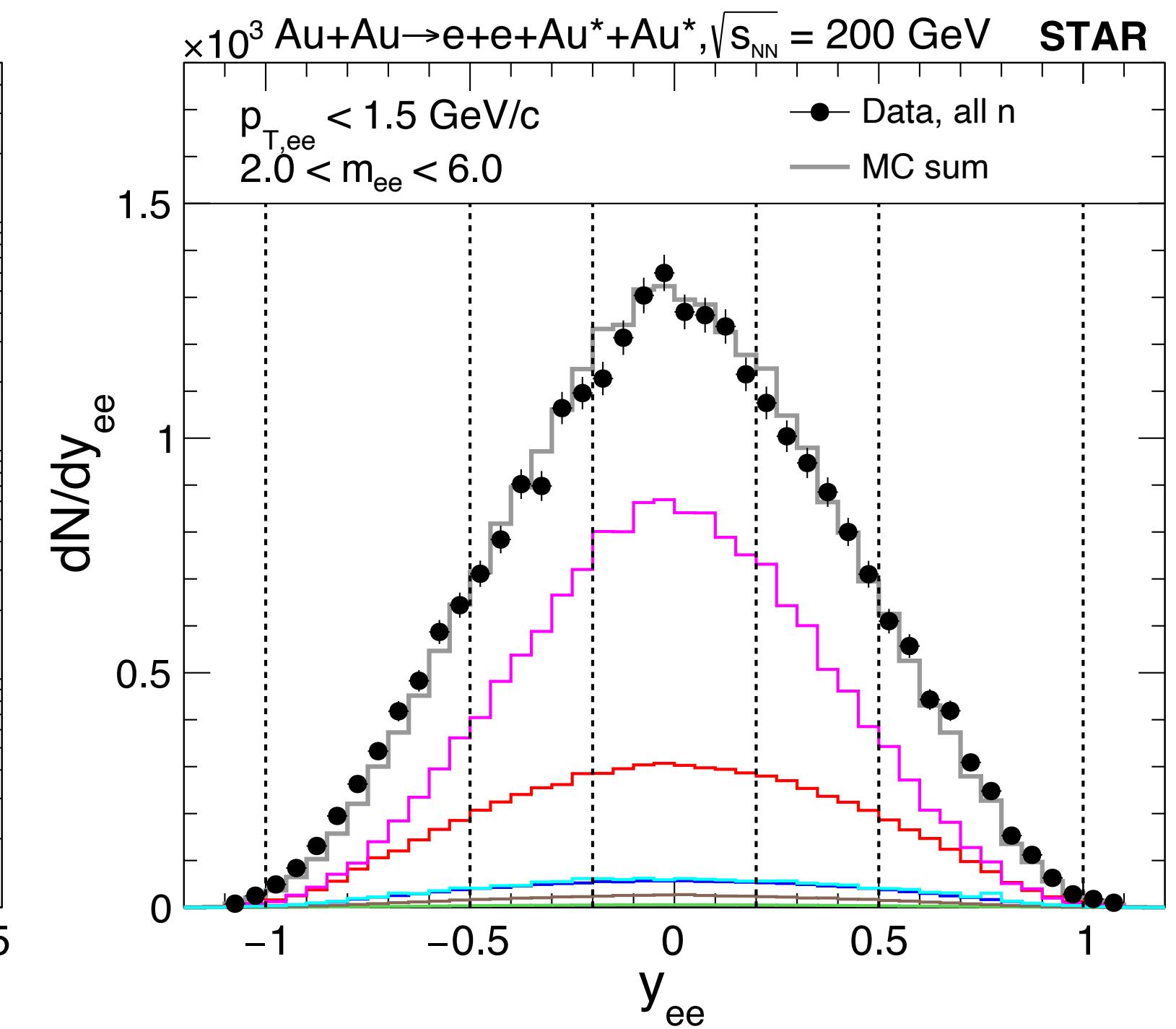
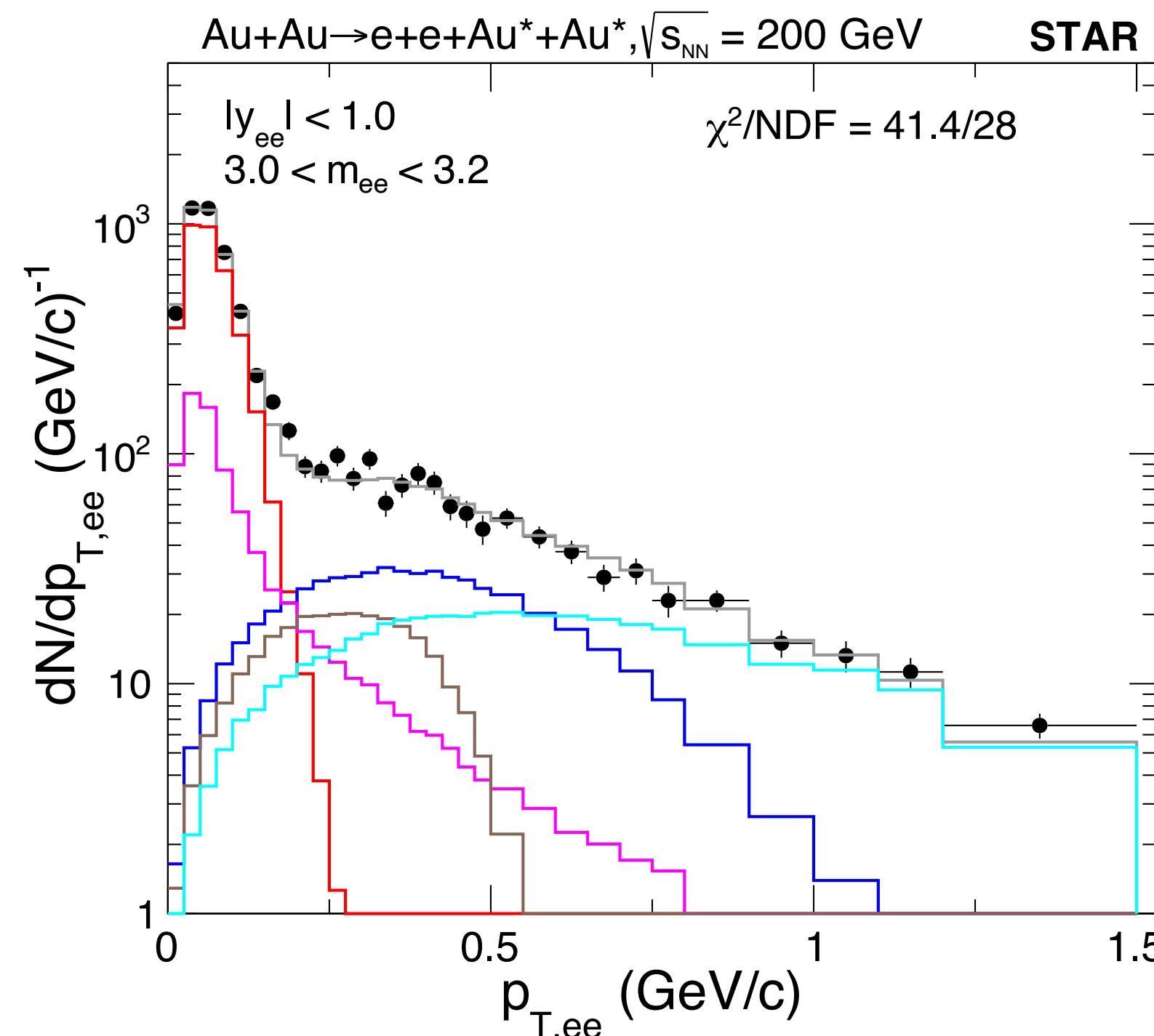
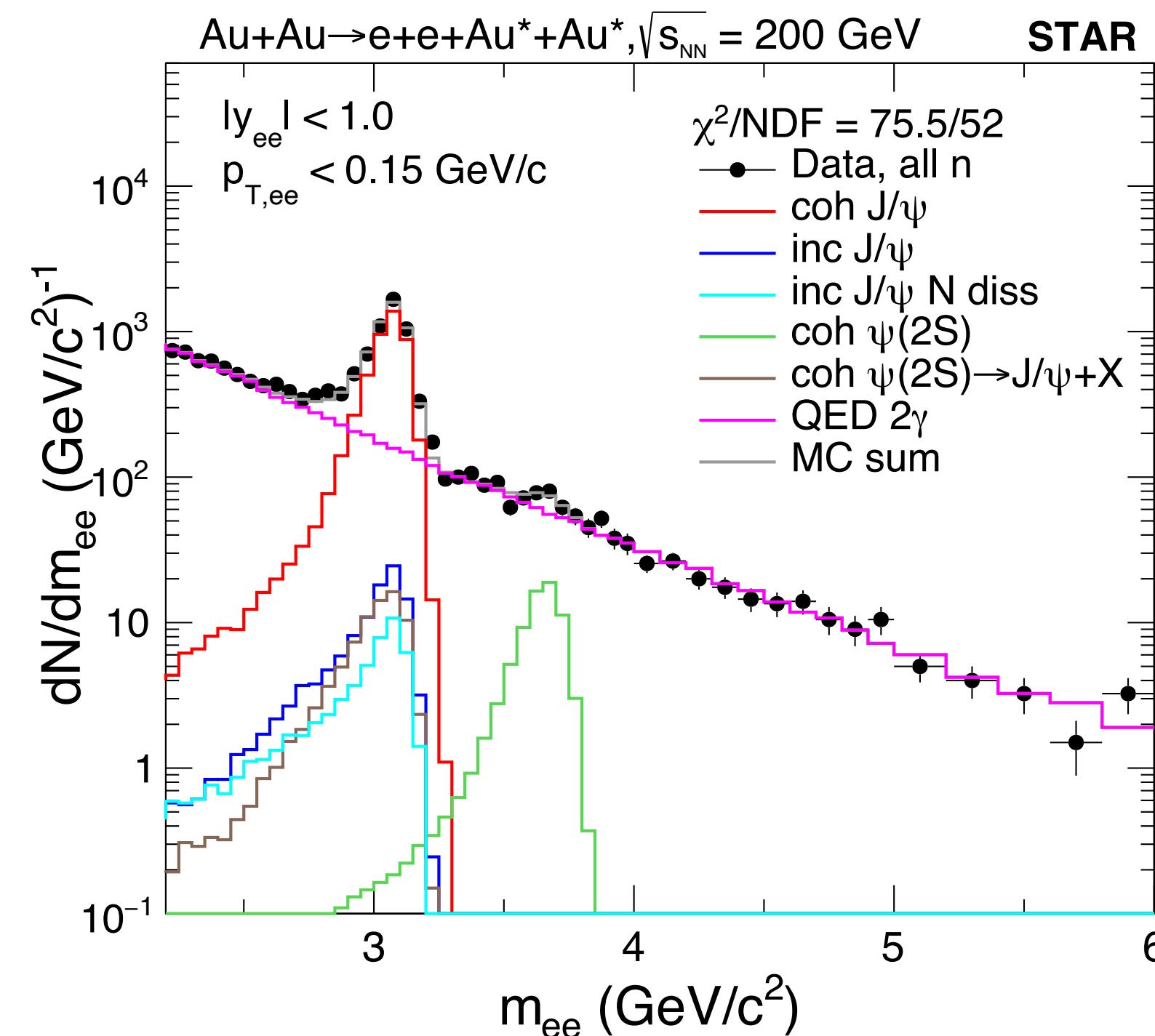
- integrated luminosity of  $1100 \pm 100 \mu\text{b}^{-1}$  of data collected in 2010
- $XnXn$  extrapolated from  $1n1n$  using STARlight
- incoherent components in  $d\sigma/dt$  are fit in range  $-t = (0.2, 0.45)$ 
  - $\sigma_{incoh}$  are integrals of the fits



Parameter	$XnXn$	$1n1n$
$\sigma_{coh.}$	$6.49 \pm 0.01 \text{ (stat.)} \pm 1.18 \text{ (syst.) mb}$	$0.770 \pm 0.004 \text{ (stat.)} \pm 0.140 \text{ (syst.) mb}$
$\sigma_{incoh.}$	$2.89 \pm 0.02 \text{ (stat.)} \pm 0.54 \text{ (syst.) mb}$	$0.162 \pm 0.010 \text{ (stat.)} \pm 0.029 \text{ (syst.) mb}$
$\sigma_{incoh.}/\sigma_{coh.}$	$0.445 \pm 0.015 \text{ (stat.)} \pm 0.005 \text{ (syst.)}$	$0.233 \pm 0.007 \text{ (stat.)} \pm 0.007 \text{ (syst.)}$

Nuclear excitation and  $\rho_0$  photo production are not completely independent

# $J/\psi$ PHOTOPRODUCTION IN AU+AU UPC EVENTS AT 200 GEV



- when  $Q^2 \sim 0$ ,  $p_T$  of  $J/\psi$  is directly related to momentum transfer ( $t \sim p_T^2$ )