

First-order event plane correlated directed and triangular flow from fixed-target energies at RHIC-STAR

1

5

6

7

Sharang Rav Sharma (for the STAR Collaboration)

Indian Institute of Science Education and Research (IISER) Tirupati

January 26, 2024

Abstract

We report the measurement of first-order event plane correlated directed flow (v_1) 8 and triangular flow (v_3) for identified hadrons $(\pi^{\pm}, K^{\pm}, \text{ and } p)$, net-particle (net-K, 9 net-p), and light nuclei (d and t) in Au+Au collisions at $\sqrt{s_{\rm NN}} = 3.2, 3.5, \text{ and } 3.9 \text{ GeV}$ 10 in fixed-target mode from the second phase of beam energy scan (BES-II) program at 11 RHIC-STAR. The v_1 slopes at mid-rapidity for identified hadrons and net-particles 12 except π^+ are found to be positive, implying the effect of dominant repulsive baryonic 13 interactions. The slope of v_1 for net-kaon undergoes a sign change from negative to 14 positive at a lower collision energy compared to net-proton. An approximate atomic 15 mass number scaling is observed in the measured v_1 slopes of light nuclei at mid-16 rapidity, which favours the nucleon coalescence mechanism for the production of light 17 nuclei. The v_3 slope for all particles decreases in magnitude with increasing collision 18 energy, suggesting a notable integrated impact of the mean-field, baryon stopping, and 19 collision geometry at lower collision energies. 20

21 1 Introduction

The primary objective of ultra-relativistic heavy-ion collisions at the Relativistic Heavy 22 Ion Collider (RHIC) and the Large Hadron Collider (LHC) is to create and characterize a 23 novel state of matter with partonic degrees of freedom, known as the Quark-Gluon Plasma 24 (QGP). This state of strongly interacting matter is hypothesized to have been present 25 during the initial microseconds following the Big Bang, and gaining an understanding of 26 its properties holds the potential to offer insights into the evolution of the universe [1]. 27 The lattice Quantum Chromodynamics (QCD) predicts a crossover region between the 28 hadron gas and QGP at higher temperature (T) and low baryon chemical potential (μ_B) 29 [3]. At lower temperatures and higher μ_B , QCD-based models suggest a first-order phase 30 transition concluding at a conjectured QCD critical point [4]. Numerous experimental 31 observables measured at RHIC and LHC have presented compelling evidence of QGP 32 formation for matter near $\mu_B = 0$. However, experimental confirmation of the existence of 33 a critical point and a first-order phase transition at higher μ_B is still pending. 34

Numerous signatures of QGP formation and associated characteristics of the medium have been proposed. This paper will briefly delve into one of the suggested signatures, namely, anisotropic flow. The patterns of azimuthal anisotropy in particle production are commonly referred to as flow. The azimuthal anisotropy in particle production stands out as one of the most distinct experimental signature of collective flow in heavy-ion collisions. It can be obtained by studying the Fourier expansion of the azimuthal angle (ϕ) distribution of produced particles with respect to the event plane angle (Ψ_n).

⁴² The particle azimuthal angle distribution is written in the form of a Fourier series [4],

$$E\frac{\mathrm{d}^3 N}{\mathrm{d}p^3} = \frac{\mathrm{d}^2 N}{2\pi p_T \mathrm{d}p_T \mathrm{d}y} \left\{ 1 + \sum_{n\geq 1} 2v_n \cos\left[n\left(\phi - \Psi_n\right)\right] \right\},\tag{1}$$

where p_T , y, ϕ , and Ψ_n are particle transverse momentum, rapidity, azimuthal angle of the particle and the n^{th} order event plane angle, respectively. The various (order n) coefficients in this expansion are defined as:

$$v_n = \langle \cos[n(\phi - \Psi_n)] \rangle \tag{2}$$

The angular brackets in the definition denote an average over many particles and events
[4]. The sine terms in the distribution become zero due to the reflection symmetry concerning the reaction plane.

The flow anisotropy parameters (v_n) offer an insight into collective hydrodynamic expansion and transport properties of the produced medium at higher collision energies, while they are sensitive to the compressibility of the nuclear matter and nuclear EOS at lower collision energies. The first three Fourier expansion coefficients v_1 (directed flow), v_2 (elliptic flow) and v_3 (triangular flow) are sensitive probes for studying the properties of the

- ⁵⁴ matter created in high-energy nuclear collisions.
- At higher energies (nucleon-nucleon center-of-mass energy $\sqrt{s_{NN}} \gtrsim 27$ GeV), where the
- transit time of colliding nuclei $2R/\gamma\beta$ (where R is the radius of the nucleus, γ is the Lorentz
- factor, and β is the velocity of the nuclei) is smaller than the typical production time of
- ⁵⁸ particles [5], flow harmonics are predominantly influenced by the collective expansion of
- ⁵⁹ the initial partonic density distribution [6]. Conversely, at lower energies, the shadowing
- 60 effect caused by passing spectator nucleons becomes significant. For $\sqrt{s_{NN}} \lesssim 4$ GeV, nu-
- clear mean-field effects contribute to the observed azimuthal anisotropies [7]. Numerous
- ⁶² studies indicate that flow coefficients are notably sensitive to the incompressibility of nu-
- clear matter (κ) in the high baryon density region [8]. Comparing experimental data with
- results from theoretical transport models can provide constraints on κ , offering valuable
- ⁶⁵ insights into nuclear EOS.
- ⁶⁶ The directed flow (v_1) , sensitive to early collision dynamics, is proposed as a signature

⁶⁷ of first-order phase transition based on a hydrodynamic calculations. These calculations, ⁶⁸ whose EOS incorporates a first-order phase transition from hadronic matter to QGP,

- ⁶⁸ whose EOS incorporates a first-order phase transition from hadronic matter to QGP, ⁶⁹ predict a non-monotonic variation of the slope of the directed flow of baryons (and net-
- ⁷⁰ baryons) around midrapidity as a function of beam energy [9].
- The tradional v_3 , third order flow coefficient typically results from fluctuations in shape of the initial condition and is not correlated to the reaction plane. In contrast to this, initial observations were made by HADES, followed by the STAR collaboration, in Au+Au collisions at $\sqrt{s_{\text{NN}}} = 2.4$ GeV and 3 GeV, respectively. A noticeable triangular flow, correlated with the first order event plane (Ψ_1) was observed [6].

The v_3 is also observed to be sensitive to the EOS and can serve as a new tool to explore the time dependence of the pressure during the heavy-ion collision [10]. The evolution of v_3 is influenced by two crucial factors: the first involves the appropriate geometry determined by stopping, the passing time of spectators, and the expansion of the fireball; the second entails a potential within the responsive medium that propels the collective motion of particles.

⁸² 2 STAR Fixed-Target Program

⁸³ The fixed target (FXT) setup was implemented at Solenoidal Tracker at RHIC (STAR) ⁸⁴ to explore the region of high μ_B on the QCD phase diagram. This data was collected ⁸⁵ during the second phase of the Beam Energy Scan program (BES-II) (2019-2020) after ⁸⁶ incorporating various detector upgrades.

87 2.1 Experimental Setup

The STAR FXT comprises a 0.25 mm thick gold foil (equivalent to a 1% nuclear interaction 88 probability) mounted on a half collar with two aluminum support rods. Positioned at 89 the west edge of the TPC, the target is longitudinally 200 cm away from the nominal 90 interaction point at the center of the TPC. Placed at the bottom of the beam pipe, 91 the top edge of the gold foil is situated 2 cm below the center of the beam pipe. This 92 configuration is crucial to prevent unintended collisions between the beam and the target 93 during collider mode operation. In fixed-target mode, the accelerator technicians lowered 94 the beam by 1.8 cm until the trigger rate reached 2 kHz, which is the limit of the Data 95 Acquisition (DAQ) system. 96

During fixed-target mode operation, the accelerator utilizes only one cycling beam. In 97 this setup, the beam is filled with only 12 bunches, with each bunch containing 7×10^9 98 ions. This limitation on the number of bunches serves to separate out-of-time pileup by a 99 sufficiently large distance and also restricts the DAQ rate. The rationale behind limiting 100 the number of bunches is to avoid instances where two collisions occur too close together 101 temporally. In such cases, the vertices may appear too close together longitudinally in the 102 TPC and might be reconstructed as a single vertex with a multiplicity equal to the sum of 103 the two independent collision multiplicities. To prevent these out-of-time pileup vertices, 104 a reduction in the number of bunches ensures spatial separation. 105

106 2.2 Fixed-Target Conventions

In contrast to collider mode collisions, in FXT collisions in the STAR coordinate system, the target is situated at the edge of TPC, and midrapidity is not zero. To convert the measured rapidity (y) in the coordinate system to the rapidity in the center of mass frame (y_{cms}) , it is necessary to boost the measured rapidity by the beam rapidity. The beam rapidity (y_b) for a given center of mass energy is calculated by the equation,

$$y_b = \cosh^{-1} \left[\frac{\sqrt{s_{\rm NN}}}{2m_p} \right],\tag{3}$$

where the $\sqrt{s_{\rm NN}}$ is center of mass energy (e.g. 3.2 GeV), and the m_p is proton mass (0.938 GeV). In STAR convention, the beam-going direction is the positive direction (the target is located in the negative rapidity direction for 3.2 GeV at $y_{\rm b} = -1.127$). To match the STAR conventions, when calculating rapidity in the center of the mass frame and shifting by midrapidity, we also need to flip the sign of rapidity.

$$y_{cms} = -(y_{lab} - y_b).$$
 (4)

3 Dataset and Event Selection Cuts

In this paper, we present the results of first order event plane (Ψ_1) correlated v_1 and v_3 for identified hadrons (π^{\pm} , K^{\pm} , and p), net-particle (net-K, net-p), and light nuclei (dand t) in Au+Au collisions at $\sqrt{s_{NN}} = 3.2, 3.5$, and 3.9 GeV using the FXT data from the STAR experiment. In FXT mode, we apply a vertex cut along the z-direction (v_z) within [198, 202] cm. For the x and y directions, we set the V_r ($\sqrt{V_x^2 + V_y^2}$) less than 2 cm centered around (0, -2).

125 4 Analysis Details

126 4.1 Track Quality Cuts

To ensure the quality of primary tracks, tracks with the transverse momentum $p_T < 0.2$ GeV/c are excluded. Additionally, we mandate the utilization of a minimum of 15 fit points and 52% of the total possible fit points in the track fitting process. The selection criterion involves choosing dE/dx hit points \geq 10. Furthermore, the distance of closest approach (DCA) is set to < 3 cm.

132 4.2 Particle Identification

The identification of charged particles in STAR is done by the combination of Time Projection Chamber (TPC) and Time of Flight (TOF) detectors. For low-momentum particles, TPC is used, whereas for particles with intermediate or high momenta ($p_T > 1 \text{ GeV/c}$), the TOF is used. TPC uses the ionization energy loss (dE/dx) of the charged particles passing through it for particle identification. Using dE/dx information, the z variable is defined:

$$z = ln(\frac{\langle dE/dx \rangle}{\langle dE/dx \rangle_X^B}),\tag{5}$$

where $\langle dE/dX \rangle_X^B$ is the expected energy loss based on the Bichsel function and X is the particle type [11]. The raw yield from the TOF are obtained using the variable mass square (m^2) , given by

$$m^2 = p^2 \left(\frac{c^2 T^2}{L^2} - 1\right),\tag{6}$$

where, p, T, L, and c are the momentum, time of travel by the particle, path length and speed of light, respectively. The left panel of Fig 1 shows the average dE/dx of measured charged particles plotted as a function of "rigidity" (i.e., momentum/charge) of the particles. The curves represent the Bichsel expectation values. The right panel of Fig 1 shows the inverse of particle velocity in unit of the speed of light $1/\beta$, as a function of rigidity. The expected values of $1/\beta$ for charged particles are shown as the curves.



Figure 1: $\langle dE/dx \rangle$ from TPC (left panel) and $1/\beta$ from TOF (right panel) for charged particles in Au+Au collisions at $\sqrt{s_{\rm NN}} = 3.2$ GeV.

In this analysis, for the identification of pion, kaon, and proton, we require TPC $n\sigma$ (z/R; R: TPC resolution) and TOF m^2 cuts which are listed in the Table 1. In addition to m^2 cut a momentum dependent z cut is implemented for light nuclei identification.

pion	$ n\sigma_{\pi} < 3$ and $-0.1 < m^2 < 0.15 \ ((GeV/c^2)^2)$
kaon	$ n\sigma_K < 3$ and $0.16 < m^2 < 0.36 \ ((GeV/c^2)^2)$
proton	$ n\sigma_p < 2$ and $-0.6 < m^2 < 1.2 \; ((GeV/c^2)^2)$
deuteron	momentum dependent z cut and $3.15 < m^2 < 3.88 \ ((GeV/c^2)^2)$
triton	momentum dependent z cut and 7.01 $< m^2 < 8.75 ((GeV/c^2)^2)$

Table 1: Particle identification cuts

152 4.3 Event Plane Reconstruction

151

The event plane angle can be estimated from the particle azimuthal distribution on an event-by-event basis. In our calculations, we have used the first-order event plane angle Ψ_1 , which is measured using the Event Plane Detector (EPD). EPD is designed to measure the pattern of forward-going charged particles emitted in a high-energy collision between heavy nuclei. In order to calculate the first-order event plane angle, firstly, we construct the Q vector from particle's azimuthal angle.

$$\vec{Q} = (Q_x, Q_y) = \left(\sum_i w_i \cos(\phi_i), \sum_i w_i \sin(\phi_i)\right),\tag{7}$$

¹⁵⁹ The first-order event plane angle ψ_1 is defined as:

$$\psi_1 = \tan^{-1}(Q_y/Q_x),$$
 (8)

where sum extends over all detected hits i, and ϕ_i is the azimuthal angle in the laboratory frame, and w_i is the weight for the i^{th} hits, here we use the nMip as the weight, which is the calibrated ADC value. In order to mitigate acceptance correlations arising from the imperfect detector, it is essential to render the event plane angle distribution isotropic or flat. Consequently, a procedure for flattening the event plane angle distribution becomes necessary. In this analysis, we have implemented re-centering and shift corrections to extract a flat event plane angle distribution [4].

In the re-centering correction, the Q-vector averaged over multiple events is subtracted 167 from the Q-vector of each individual event. Subsequently, the event plane angle is cal-168 culated. However, performing only the re-centering correction is insufficient. After the 169 implementation of re-centering, a shift correction is additionally applied to ensure that 170 the event plane angle distribution becomes flat. In shift correction, one fits the non-flat 171 distribution of ψ_n averaged over many events with a Fourier expansion and calculates the 172 shifts for each event ψ_n necessary to force a flat distribution on average. The re-centering 173 and shift correction are all performed run-by-run and centrality-by-centrality in this anal-174 vsis. 175

176 4.4 Event Plane Resolution

The finite number of detected particles in detectors produces a limited resolution in the measured event plane angle. So, the observed flow coefficients must be corrected up to what they would be relative to the real reaction plane. This is done by dividing these coefficients by the event plane resolution, estimated from the correlation of the planes of independent subevents.

$$v_n = \frac{v_n^{obs}}{R_n} = \frac{v_n^{obs}}{\langle \cos[n(\psi_n - \Psi_R)] \rangle},\tag{9}$$

where the R_n is resolution, v_n is the n^{th} harmonic azimuthal anisotropy parameter, and ψ_n is the n^{th} harmonic order event plane, Ψ_r is reaction plane angle. The angle brackets denote an average over all particles in all events [4].

In fixed target mode, the final state particle's acceptance is not symmetric around midrapidity. Therefore, the commonly used 2-sub event method, employed in the collider BES-I analysis, cannot be used to calculate the resolution. This method necessitates each subevent to have similar multiplicity and resolution. Consequently, in this analysis, we opt for the 3-sub event method to calculate the resolution. Figure 2 shows the calculated firstorder event plane resolution R_{11} and the third-order event plane resolution R_{13} estimated from the first-order event plane for v_3 calculation, as functions of collision centrality.



Figure 2: Collision centrality dependence of R_{11} (circles) R_{13} (squares) in Au+Au collisions at $\sqrt{s_{\rm NN}} = 3.2$ (left panel), 3.5 (middle panel), and 3.9 GeV(right panel).

¹⁹² 5 Systematic Uncertainties

The systematic uncertainties associated with the measured flow harmonics stem from 193 the charged track selection method, particle identification, and event plane resolution. 194 These uncertainties are evaluated point-by-point on v_1 and v_3 as a function of y for each 195 identified hadron and light nuclei. The systematic uncertainties arising from track selection 196 are assessed by varying the selection requirements. Those linked to particle misidentifi-197 cation are determined by varying the z and m^2 cuts. A common systematic uncertainty 198 arising from event plane resolution is assessed by employing combinations of different η 199 sub-events. In the subsequent figures, the shaded boxes represent the total systematic 200 uncertainty for each data point. 201

²⁰² 6 Results and Discussion

203 6.1 Directed Flow (v_1)

The rapidity (y), centrality and collision energy dependence of v_1 for identified hadrons, net-particle, and light nuclei are measured at $\sqrt{s_{\rm NN}} = 3.2, 3.5$, and 3.9 GeV.

Figure 3 illustrates the the centrality dependence of π^+ for $\sqrt{s_{NN}} = 3.2$ GeV. The v_1 changes sign from negative to positive, moving from most central to peripheral collisions, implying the effect of dominant repulsive baryonic interactions and spectator shadowing. The energy dependence of proton v_1 involves an interplay between the directed flow of protons associated with baryon stopping and particle-antiparticle pair production at midrapidity. A means to distinguish between the two mechanisms would thus be to look at the net particle v_1 . The net particle represents the excess yield of a particle species over



Figure 3: v_1 as a function of y for pion in 0-10% (left panel), 10-40 % (middle panel), and 40-60% (right panel) centrality bin in Au+Au collisions at $\sqrt{s_{\rm NN}} = 3.2$ GeV. The line represents third order polynomial fit to distribution.

²¹³ its antiparticle. The net particle's v_1 is defined as

$$v_{1,net} = \frac{v_{1,p} - rv_{1,\bar{p}}}{1 - r},\tag{10}$$

where $v_{1,p}$, $v_{1,\bar{p}}$ corresponds to v_1 of particle and anti-particle, and r represents the ratio of anti-particles to particles [9].

In relativistic heavy-ion collisions, the production of light nuclei can occur through two 216 mechanisms. The first mechanism involves the direct production of nucleus-antinucleus 217 pairs in elementary nucleon-nucleon (NN) or parton-parton interactions. Due to their 218 small binding energies, the directly produced nuclei or antinuclei are likely to undergo 219 dissociation in the medium before escaping. The second and presumably dominant mech-220 anism for the production of nuclei and antinuclei is through the final state coalescence of 221 produced nucleons and antinucleons or participant nucleons [13]. In this process, nucleons 222 and antinucleons combine to form light nuclear and antinuclear clusters during the final 223 stages of kinetic freeze-out. The probability of formation is proportional to the product 224 of the phase space densities of its constituent nucleons [14]. Therefore, the production of 225 light nuclei yields information about the size of the emitting system and its space-time 226 evolution. Due to the longer passing time of the colliding ions in the few GeV regime, the 227 interference between the expanding central fireball and the spectator remnants becomes 228 more significant than at higher energies. 229

Figure 4 shows the y dependence of identified hadrons (left panel), net-particles (middle panel), and light nuclei (right panel) for 10-40% centrality. The magnitude of v_1 increases with increasing rapidity for all particles, and a mass ordering is also observed in the magnitude of v_1 .

The p_T -integrated $v_1(y)$ slope at mid-rapidity, $dv_1/dy|_{y=0}$, is obtained by fitting the data $v_1(y)$ with a third-order polynomial. Figure 5 shows the collision energy dependence of



Figure 4: v_1 as a function of y in 10-40% centrality for identified hadrons (left panel), net particles (middle panel) and light nuclei (right panel) in Au+Au collisions at $\sqrt{s_{\rm NN}} = 3.2$ GeV. The line represents 3rd order polynomial fit to distribution.

 $dv_1/dy|_{y=0}$ for identified particles (left panel), net-particle (middle panel), and light nuclei (right panel) in mid-central (10 - 40%) collisions. The extracted slope parameters, $dv_1/dy|_{y=0}$, are scaled by A for light nuclei to compare with protons. The magnitude of the slope decreases with increasing collision energy for all particles, including net-particles and light nuclei.

At low energies, the transit time (τ) is comparable to the formation time of particles. 241 Consequently, the spectators are not sufficiently distant from the collision volume, and the 242 medium does not have the freedom to expand freely. This results in interactions between 243 baryon-dominated spectator particles and the produced particles. Among the produced 244 particles, pions, being one of the lightest, are particularly affected. The flow for π^+ is ob-245 structed by the spectator particles, leading to a negative value for its v_1 slope. In contrast, 246 π^- , influenced by Coulomb interactions from the baryons (protons), acquires a positive v_1 247 slope value. 248

The slope of v_1 for net-kaon undergoes a sign change from negative to positive at a lower collision energy range ($\sqrt{s_{\rm NN}} = 3.9 - 7.7 \text{ GeV}$) compared to net-proton ($\sqrt{s_{\rm NN}} = 11.5 - 19.6 \text{ GeV}$).

The light nuclei v_1 slope exhibits an approximate mass number (A) scaling, consistent with the nucleon coalescence mechanism for the production of light nuclei at low collision energies.

255 6.2 Triangular Flow (v_3)

The y and collision energy dependence of v_3 for identified hadrons and light nuclei are measured at $\sqrt{s_{\rm NN}} = 3.2$, 3.5, and 3.9 GeV. The left panel of Fig. 6 shows the rapidity dependence of v_3 for identified hadrons. The magnitude of v_3 increases with increasing



Figure 5: Collision energy dependence of v_1 slope $dv_1/dy|_{y=0}$ for identified hadrons (left panel), net particles (middle panel) and light nuclei (right panel) in Au+Au collisions at RHIC for 10 - 40% centrality. The published data are shown in open markers [12].

rapidity. The distribution is fitted with a polynomial of order three to extract the slopeparameter.

The middle panel of Fig. 6 shows the slope of v_3 , $dv_3/dy|_{y=0}$, for identified hadrons as a function of collision energy. The magnitude of $dv_3/dy|_{y=0}$ decreases with increasing collision energy. It may indicates that the combined effect of the mean-field, baryon stopping, and collision geometry is considerably significant at the low collision energies [15].

The right panel of Fig. 6 shows the extracted slope parameters, $dv_3/dy|_{y=0}$, scaled by mass number (A) for light nuclei. The magnitude of the slope decreases with increasing collision energy. The light nuclei v_3 slope also exhibits an approximate mass number (A) scaling within systematic uncertainties, consistent with the nucleon coalescence mechanism for the light nuclei production.

$_{270}$ 7 Conclusion

In summary, the rapidity, centrality, and collision energy dependence of directed flow 271 (v_1) of identified hadrons, net particle, and light nuclei in Au+Au collisions at $\sqrt{s_{NN}} = 3.2$, 272 3.5, and 3.9 GeV is reported. The magnitude of v_1 increases with increasing rapidity for all 273 particles. The extracted v_1 slope of all the particles decreases in magnitude with increasing 274 collision energy. A positive v_1 slope at mid-rapidity for identified hadrons and net particles, 275 excluding π^+ , suggests prevalent repulsive baryonic interactions and spectator shadowing. 276 As collision energy decreases, a non-monotonic trend is observed in the slope of both net-277 kaon and net-proton. The v_1 slope for net-kaon experiences a transition from negative to 278 positive at a collision energy lower than that observed for net-proton. The light nuclei v_1 279 slope exhibits an approximate mass number scaling consistent with the nucleon coalescence 280



Figure 6: v_3 as a function of y in 10–40% centrality bin for identified hadrons (left panel) in Au+Au collisions at $\sqrt{s_{\rm NN}} = 3.2$ GeV. Collision energy of $dv_3/dy|_{y=0}$ for identified particles (middle panel) and light nuclei (right panel) in Au+Au collisions at RHIC for 10 – 40% centrality. The published data are shown in open markers [10].

mechanism for the production of light nuclei. The magnitude of slope of v_3 decreases with increasing collision energy, indicating a substantial collective impact of the meanfield, baryon stopping, and collision geometry at lower collision energies. Similar to the v_1 slope of light nuclei, the v_3 slope of light nuclei also displays an approximate scaling with mass number (A) within the systematic uncertainties. This trend supports the nucleon coalescence mechanism as a favorable explanation for the production of light nuclei.

287 **References**

- ²⁸⁸ [1] H. Kastrup, P. Zerwas, eds., QCD 20 yrs later, World Scientific, Singapore (1993).
- [2] A. Aprahamian et. al. DOE/NSF Nuclear Science Advisory Panel (NSAC) Report,
 (2015).
- ²⁹¹ [3] F.R. Brown, et. al., Phys. Rev. Lett. 65, 2491 (1990).
- ²⁹² [4] A. M. Poskanzer and S. A. Voloshin, Phys. Rev. C 58, 1671 (1998).
- ²⁹³ [5] A. Bialas, M. Gyulassy, Nucl. Phys. B 291, 793 (1987).
- ²⁹⁴ [6] B. I. Abelev et al. (STAR Collaboration), Phys. Rev. Lett. 99, 112301 (2007).
- ²⁹⁵ [7] H. Liu et al. (E895 Collaboration), Phys. Rev. Lett. 84, 5488 (2000).
- ²⁹⁶ [8] P. Danielewicz, R. Lacey, and W. G. Lynch, Science 298, 1592 (2002).
- ²⁹⁷ [9] H. Sorge, Phys. Rev. Lett. 78, 2309 (1997).

- ²⁹⁸ [10] J. Adamczewski-Musch et al. (HADES), Phys. Rev. Lett. 125, 262301 (2020).
- ²⁹⁹ [11] H. Bichsel, Nucl. Instrum. Meth. A 562, 154 (2006).
- ³⁰⁰ [12] L. Adamczyk et al. (STAR Collaboration), Phys. Rev. Lett. 120, 062301 (2018).
- ³⁰¹ [13] H.H. Gutbrod et al., Phys. Rev. Lett. 37, 667 (1976).
- ³⁰² [14] L. P. Csernai and J. I. Kapusta, Phys. Rep. 131, 223 (1986); A. Z. Mekjian, Phys.
- ³⁰³ Rev. C 17, 1051 (1978).
- ³⁰⁴ [15] STAR Collaboration arXiv: 2309.12610 [nucl-ex] (2023).