MEASUREMENT OF PSEUDORAPIDITY DISTRIBUTIONS WITH THE STAR EPD AT BES-II ENERGIES

BALAZS KORODI for the STAR Collaboration

The Ohio State University

WPCF 2024 Toulouse

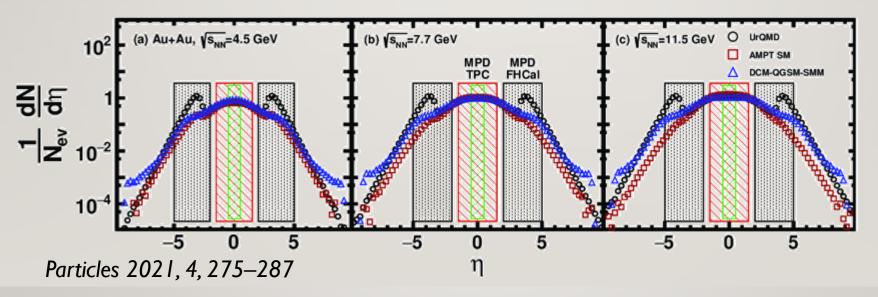






2/13 MOTIVATION

- $dN_{ch}/d\eta$ important for the tuning of models
 - Midrapidity: models give similar predictions to each other and to data
 - Forward rapidities: models disagree
- Few measurements at forward rapidities
 - None at all BES-II energies



C CTZ

Summary

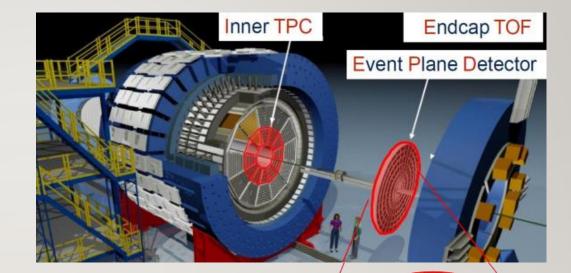
Introduction

3/13 THE STAR EVENT PLANE DETECTOR

- Part of the upgrade for BES-II
- Motivations:
 - Centrality determination
 - Event plane resolution
 - Triggering
- Characteristics:
 - Detects charged particles
 - Located at ±375 cm from the interaction point (East and West EPD)

deta

- Large pseudorapidity coverage: 2.14<|η|<5.09
- High η and φ segmentation:
 - 16 radial segments (rings)
 - 24 azimuthal segments (sectors) -
- Can be used to measure $dN_{ch}/d\eta$ for 2.14<| η |<5.09



Summa

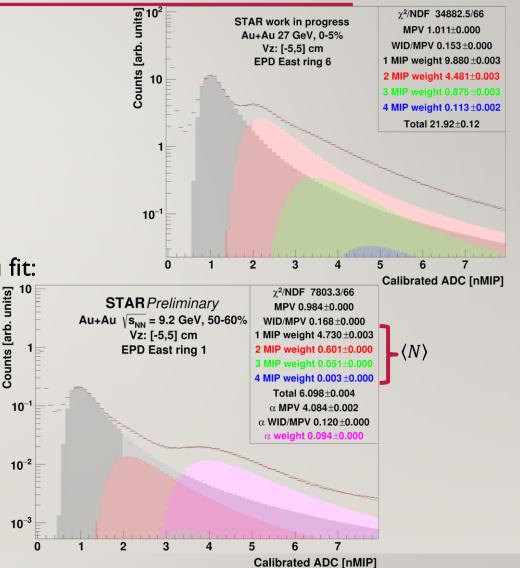
Introduction

4/13 LANDAU FITS – DETERMINING #MIPS IN EACH RING

- EPD measures energy deposited via ionization
- Mostly Minimum Ionizing Particles (MIPs)
- Deposited energy has Landau distribution for single MIP
- Multiple MIPs → convoluted Landau distributions
- Average #MIPs in each ring $(N(i_{Ring}))$ from convoluted Landau fit:
 - $N(i_{Ring}) = \sum_{n} n \cdot nMIPweight$
- Inner rings:

ntroductio

- α particles from projectile remnants need to be considered
- Poisson distribution of nMIP weights enforced
- Use same fit method in all cases for consistency

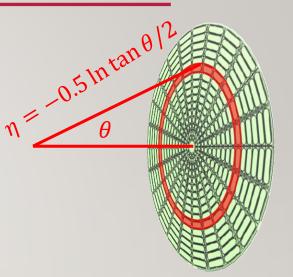


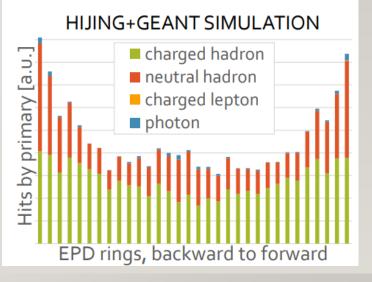
Summa

Introductio

5 /13 HOW (NOT) TO MEASURE $dN_{ch}/d\eta$ WITH THE EPD

- We could calculate $dN_{ch}/d\eta$ from raw EPD hit numbers, based on η corresponding to each ring
- This would not take into account scattering and decays:
 - Charged particles scatter in detector material, creating secondaries
 - Secondaries have large contribution to $dN_{ch}/d\eta$
 - Neutral particles contribute through decays (e.g., $\Lambda \rightarrow p + \pi$) and secondaries
 - Neutral particles also have a large contribution!





Summai

Introduction

6/13 MEASURING $dN_{ch}/d\eta$ WITH THE EPD

- From Landau fits: number of hits in each ring: $N(i_{Ring})$
- Given the underlying $dN/d\eta$, $N(i_{Ring})$ can be calculated as

$$N(i_{Ring}) = \int R(\eta, i_{Ring}) \frac{dN}{d\eta} d\eta$$

Summary

- Here R is the response matrix: no. of hits in given ring originating from primary particle at η
- Calculate R via simulations, then determine $dN/d\eta$ via unfolding
 - Bayesian iterative unfolding, G. D'Agostini, Nucl. Instr. Meth. A362 (1995) 487

7/13 UNFOLDING PROCEDURE

- . Create the response matrix
 - HIJING+GEANT simulation at same energy as data
 - For each primary track create list of EPD hits originating from that primary
 - If no EPD hit for a primary: ResponseMatrix->Miss(TrackEta)
 - For each EPD hit of a primary: ResponseMatrix->Fill(EPDRingNumber, TrackEta)
 - Do not take particles created inside the EPD into account
- 2. Perform Bayesian iterative unfolding
 - Implemented in RooUnfold
- 3. Apply multiple counting correction
 - Correct for multiple hits originating from the same primary
- 4. Apply charged fraction correction

Introductio

- Correct for hits originating from neutral primaries
- 3 possible methods: correct then unfold, unfold then correct, RooUnfold "Fakes" method

$R(\eta, i_{Ring})$ **STAR** *Preliminary* HIJING+GEANT Au+Au VS_{NN} = 14.6 GeV Г 10 East EPD West EPD

Summa

10

10³

102

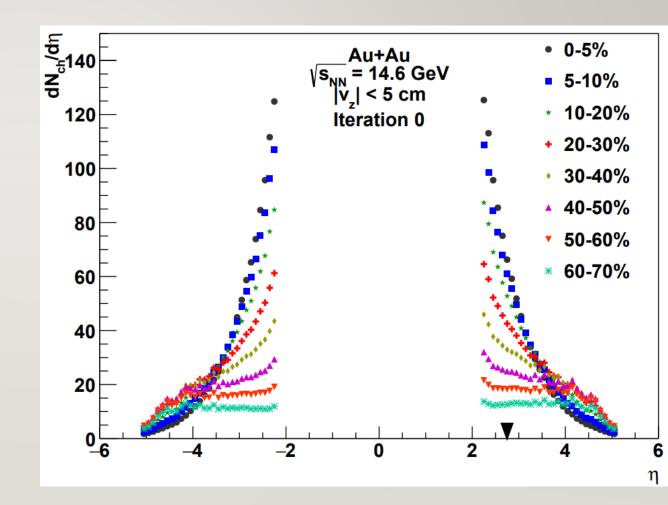
10

8/13 ITERATIVE APPROACH FOR INPUT $dN/d\eta$

- Unfolded $dN/d\eta$ depends on HIJING $dN/d\eta$
- Scale $dN/d\eta$ in simulation to unfolded $dN/d\eta$
- Unfold experimental data again
- Iterate until input = unfolded
- Issue:

Introductio

- Unfolding does not work at midrapidity (not in EPD range)
- Influences results in EPD range especially through many iterations
- Significant unwanted effect after the first iteration
- Only do 1 iteration for now



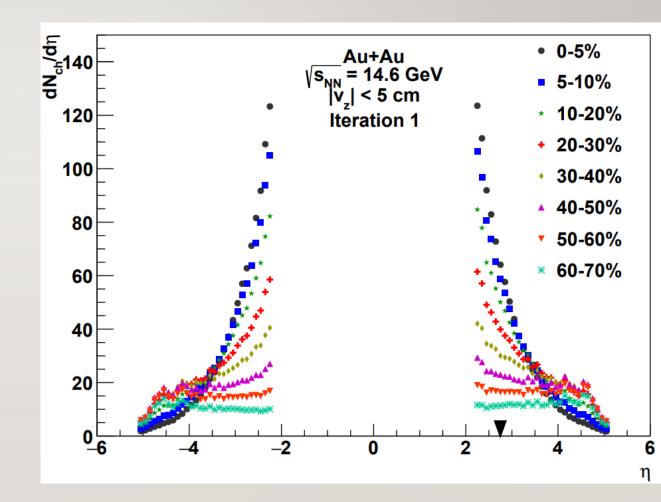
Summai

9/13 ITERATIVE APPROACH FOR INPUT $dN/d\eta$

- Unfolded $dN/d\eta$ depends on HIJING $dN/d\eta$
- Scale $dN/d\eta$ in simulation to unfolded $dN/d\eta$
- Unfold experimental data again
- Iterate until input = unfolded
- Issue:

Introductio

- Unfolding does not work at midrapidity (not in EPD range)
- Influences results in EPD range especially through many iterations
- Significant unwanted effect after the first iteration
- Only do 1 iteration for now



Summai

Introductio

10/13 SYSTEMATIC UNCERTAINTIES

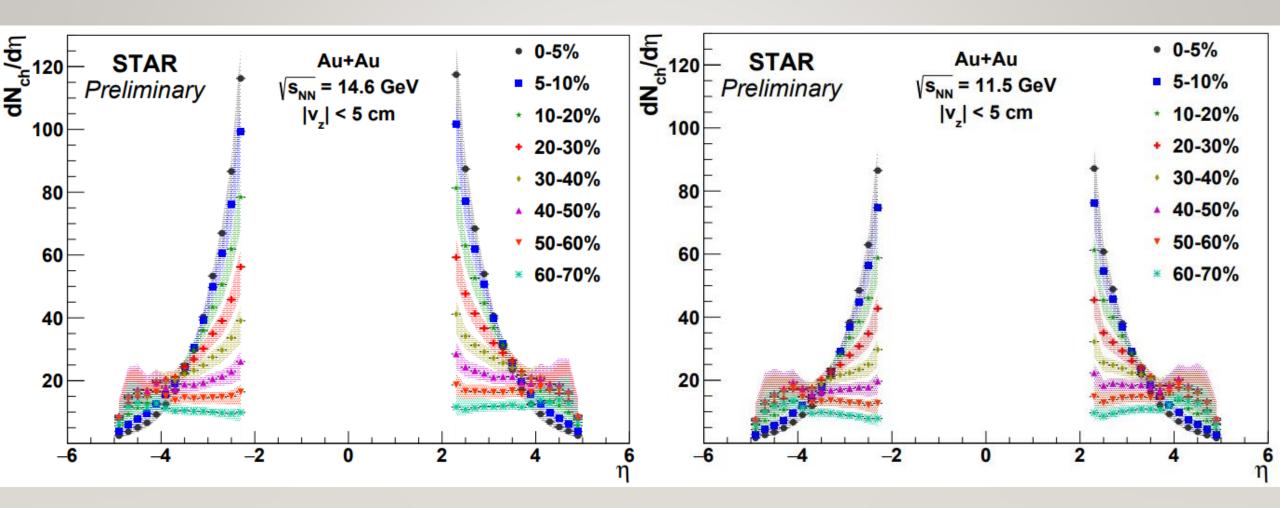
- Systematic checks in the unfolding:
 - Scale 27 GeV MC to 200 GeV and 7GeV MC:
 - Charged/neutral ratio $\rightarrow 4\%$
 - Baryon/meson ratio $\rightarrow 4\%$
 - Input $dN/d\eta \rightarrow 6\%$
 - Momentum distribution $\rightarrow 3\%$
 - Unfolding method (3 different methods for the charged fraction correction) \rightarrow 7%
- Centrality selection ($\pm 5\%$ change) $\rightarrow 3\%$
- z-vertex resolution ($\pm 5 \text{ cm shift}$) \rightarrow 1%
- z-vertex choice (± 40 cm from geometric center) $\rightarrow 6\%$
- Landau fit (fit without α peak, fit without enforcing Poisson weights) \rightarrow 5%

Calculated using 27 GeV data

Summa

Introduction

II /13 RESULTS AT 14.6 AND 11.5 GEV



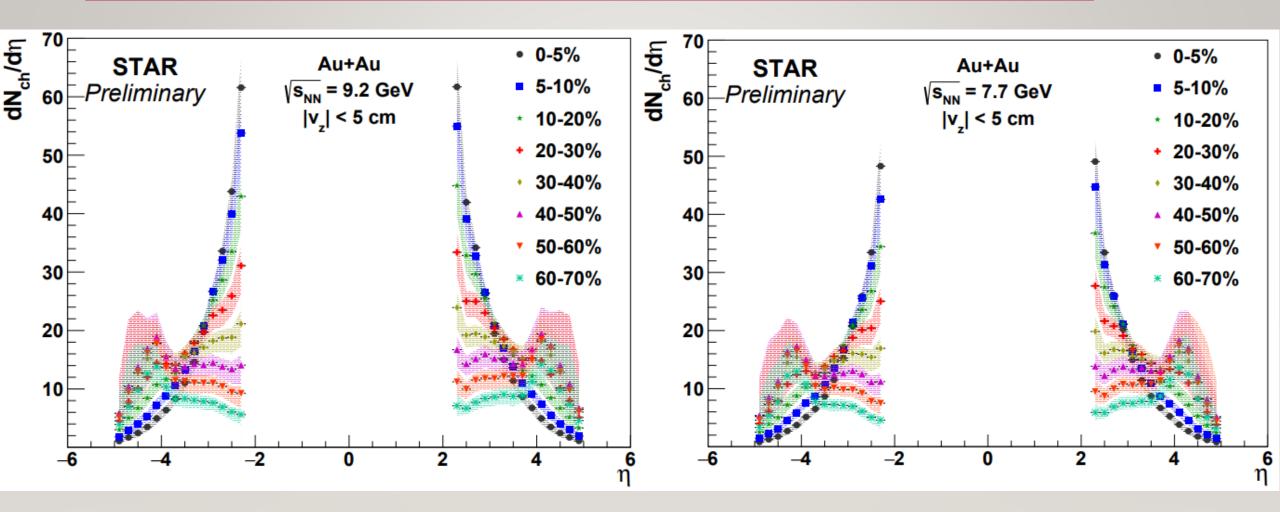
Bump appears at large $\eta \rightarrow$ caused by spectators

detai

Results

Introduction

12/13 RESULTS AT 9.2 AND 7.7 GEV



Bump becomes relatively larger at smaller energies

ceta

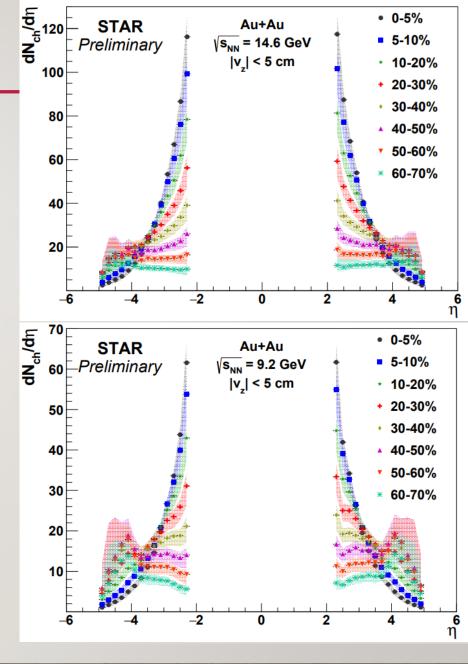
Results

13/13 SUMMARY

Introductio

- Measurement of $dN_{ch}/d\eta$ with the EPD
 - Beam Energy Scan II energies
 - Roughly expected η , centrality and $\sqrt{s_{NN}}$ dependence
 - Spectator contribution apparent
 - Detailed systematic studies
- Final steps for publication:
 - Try simulation with spectators, e.g., UrQMD
 - Refine iterative procedure
 - Calculate systematics separately for each energy

ieta



THANK YOU FOR YOUR ATTENTION!

BACKUP SLIDES

I6 DATA SETS

- All BES-II energies: Au+Au@ 27, 19.6, 14.6, 11.5, 9.2, 7.7 GeV
- MinimumBias, 8 centrality classes: 0-70%
- |zvtx| < 5 cm

Introduction

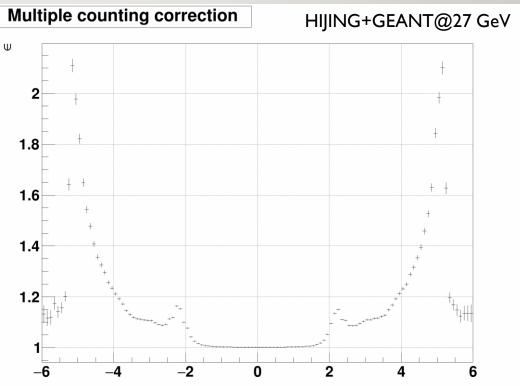
- Only data from EPD
- All available data used
 - Negligible statistical uncertainties

Analysis details

Introductio

17 MULTIPLE COUNTING CORRECTION

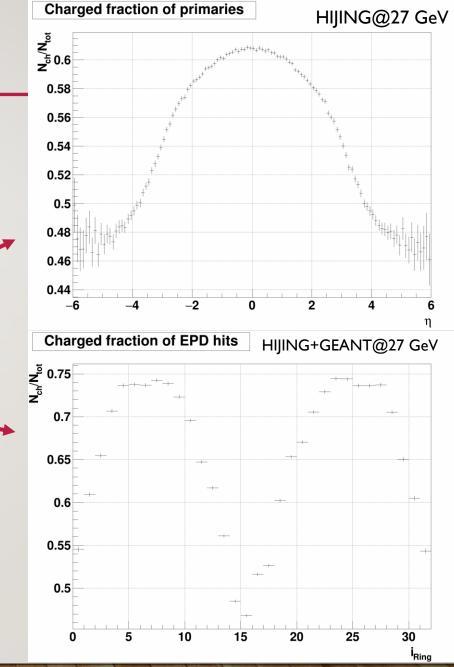
- Need to correct for multiple counting (multiple hits from one primary track)
- Check "inverse efficiency": how many hits on average from primary particles at given η Multiple counting correction HIJING+GEANT@2
- Largest at around $|\eta| \approx 5$
- Edge of EPD, support structures



Introductio

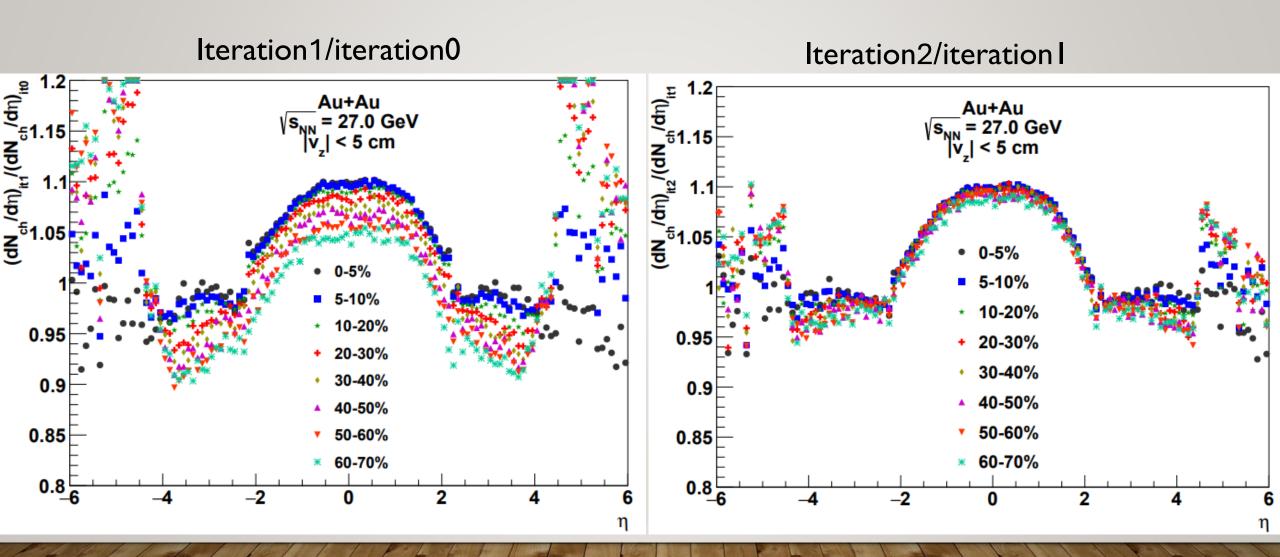
18 CHARGED FRACTION CORRECTION

- From simulations: charged particle fraction
 - For primary tracks and for EPD hits (based on primary cause)
- Applied 3 different methods:
 - I. Unfolding $dN/d\eta$; correcting via $N_{ch}(\eta)/N_{tot}(\eta)$
 - 2. Correcting via $N_{ch}(i_{ring})/N_{tot}(i_{ring})$, unfolding "corrected" EPD distribution
 - 3. Use RooUnfold's "Fakes" (neutrals ⇔ "fake" hits)
- Closure test works for all: MC input recovered when unfolding simulated EPD data
- Difference of methods: incorporated in systematics



Summa

19 CHANGE BETWEEN ITERATIONS



ceta

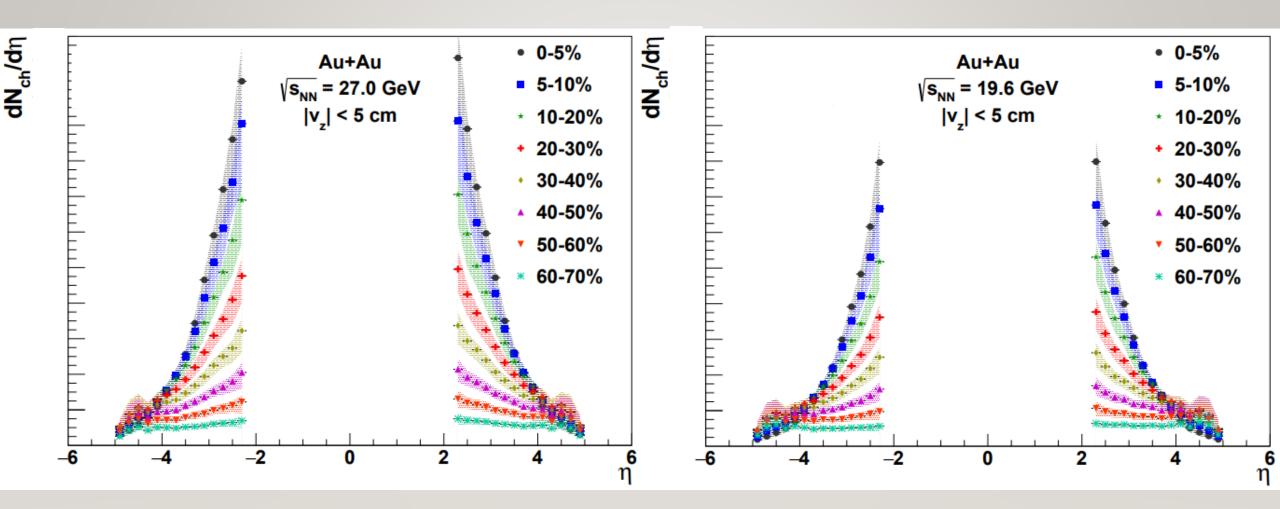
Results

Summary

Introduction

Introduction

20/15 NEW RESULTS AT 27 AND 19.6 GEV



Results

Summary

STAR already has Preliminary results at these energies therefore these new results cannot be officially shown

Cetai