9th Conference on Chirality, Vorticity and Magnetic Fields in Quantum Matter ICTP-SAIFR, São Paulo, Brazil

Measurements of azimuthal correlations with spectator and participant planes to search for the chiral magnetic effect in STAR

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Why CME?

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PHYSICAL REVIEW LETTERS

20 JULY 1998

Possibility of Spontaneous Parity Violation in Hot QCD

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CME: Chiral symmetry restoration, Local P/CP violation, matter-antimatter asymmetry...

Fundamentally important physics

Heavy ion collisions are a good place to look for it

Outline:

- The $\Delta \gamma$ observable
- Flow-induced background
- Nonflow contamination
- Results
- Summary

How to look for it?

Voloshin, PRC 2004 STAR, PRL 2009, PRC 2010



Significant $\Delta \gamma$ observed

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Flow-induced background is large

Voloshin 2004 FW 2009 Bzdak, Koch, Liao 2010 Pratt, Schlichting 2010

$$\begin{split} \gamma_{\alpha\beta} = \left\langle \cos(\varphi_{\alpha} + \varphi_{\beta} - 2\psi_{RP}) \right\rangle \\ \Delta\gamma = \gamma_{OS} - \gamma_{SS} \end{split} \qquad \begin{aligned} dN_z / d\varphi \propto 1 + 2v_1 \cos \varphi^{\pm} + 2a_z \cdot \sin \varphi^{\pm} + 2v_2 \cos 2\varphi^{\pm} + ... \\ \gamma_{\alpha\beta} = \left[\left\langle \cos(\varphi_{\alpha} - \psi_{RP}) \cos(\varphi_{\beta} - \psi_{RP}) \right\rangle - \left\langle \sin(\varphi_{\alpha} - \psi_{RP}) \sin(\varphi_{\beta} - \psi_{RP}) \right\rangle \right] \\ + \left[\frac{N_{cluster}}{N_a N_{\beta}} \left\langle \cos(\varphi_a + \varphi_{\beta} - 2\varphi_{cluster}) \cos(2\varphi_{cluster} - 2\varphi_{RP}) \right\rangle \right] \\ = \left[\left\langle v_{1,a} v_{1,\beta} \right\rangle - \left\langle a_a a_{\beta} \right\rangle \right] + \frac{N_{cluster}}{N_a N_{\beta}} \left\langle \cos(\varphi_a + \varphi_{\beta} - 2\varphi_{cluster}) \right\rangle v_{2,cluster} \\ \Delta\gamma = 2 \left\langle a_1^2 \right\rangle + \left[\frac{N_{\rho}}{N_a N_{\beta}} \left\langle \cos(\varphi_a + \varphi_{\beta} - 2\varphi_{\rho}) \right\rangle v_{2,\rho} \right] \\ \\ Flow-induced charge-dependent background: nonflow coupled with flow \\ \Delta\gamma_{Bkg} \propto v_2 / N \end{split}$$

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f^{obs}_{CME} **removing flow background**

H.-j. Xu, et al., CPC 42 (2018) 084103 S.A. Voloshin, PRC 98 (2018) 054911 STAR, PRL 128 (2022) 092301



Midcentral 20-50%: ~2-3σ significance

Flow-induced background is removed by the SP/PP method

Nonflow contamination



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V_{2} {2} nonflow

STAR, PRRes 6 (2024) L032005, PRC 110 (2024) 014905



 V_{2} {ZDC}(η) measurement -> V_{2} {ZDC}($\Delta \eta$) -> fit

- Flow decorrelation $1-2F_2\Delta\eta$, $F_2=1.15\% \pm 50\%$ (syst)
- Flow fluctuations effect: assumed constant over η

0.8 STAR + + 0.6 STAR FE SE C)

0.4

0.2

- Nonflow models by two Gaussians
- Fit flow+nonflow to $V_2(\Delta \eta)$



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60

centrality (%)

80

STAR Run11+14+16 Au+Au 200GeV

STAR Preliminary

40

20

RP-independent 3-particle correlations







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Average <f cm >



Summary

- Flow-induced background is well understood and under control by SP/PP comparison measurements
- Additional backgrounds from nonflow v2 contamination and RP-independent 3-particle correlations
 - Decomposition of flow and nonflow via a fitting procedure
 - The genuine 3-particle/2-particle correlation ratio $(\xi = C_{3p}/V_{2p})$ has weak model dependency and is robust against collective radial flow
 - Assume model ξ to correct for RP-independent 3-particle correlations
- f_{CME} extracted. Further scrutiny, e.g. MC closure.





- f_{CME}^{ODS} published previously by STAR PRL 128 (2022) 092301
- This analysis used the same data, with improved analysis cuts and systematic studies:
 - \circ p_T-dependent efficiency correction is applied
 - \circ $\Delta\eta$ cuts between POIs and between POI and particle *c* are applied
 - Systematic uncertainties are assessed with corresponding efficiency corrections
- Results are ~consistent given the p_T -dependent efficiency correction and $\Delta \eta$ cuts

Systematic uncertainty assessment

- For a quantity x, its default measurement is x₀ with statistical uncertainty e₀, and it has systematic variations x_i ± e_i.
- The differences are $d_i = x_i x_0$.
- The systematic uncertainty from each variation is s_i. If we use Barlow's check

if
$$d_i \ge \sqrt{|e_i^2 - e_0^2|}$$
, then $s_i = \sqrt{d_i^2 - |e_i^2 - e_0^2|}$
else $s_i = 0$

- ► For n_j variations from the same sources (e.g., multiple cuts on one quantity, set Q_j) s_{i∈Q_j}, RMS is used by default.
- Combining all those variations in quadrature

		Systematic variations in data analysis								
		V_z [max]	DCA (cm)		NI	nits	no eff.			
		$\geq 0 < 0$	< 0.8 <	(2 < 3)	≥ 15	≥ 25	[max]			
	-	Systematic variations in v_2 nonflow fit								
		v_2^2 ZDC fit fixed fixed $\Delta \eta$ decorr. [max]								
	۔ ۔ ۱	a Evloa	$olv5$ $\mu_2 =$	= 0 0.1	6-1.6	ZDC	FMS			
		Systematic variations in ξ								
		HI IING HI IING Pythia \downarrow HI IING BW/B [max]								
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		default no quell. $DVVD-p_T$ $DVVD-r_T$								
										1
f _{CME}								-	default	
		STAR Preliminary							Vz<0	
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	0.5								DCA<2.0	
									DCA<3.0	
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Hijing closure test checks out

- Apply v_2 -modulated weight. Input $v_2 = 0.05 p_T$ for $p_T < 2$ GeV/c, saturate at $v_2 = 0.1$. No centrality dependence. ٠
- PP fluctuates randomly about RP (=0) event by event, with Gaussian sampling of width $\pi/6$.
- The default HIJING (without flow input) is taken as nonflow. ٠



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