

Beam energy dependence of directed flow of deuteron in Au+Au collisions at RHIC

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Workshop on the QCD Phase Structure at High Baryon Density Region

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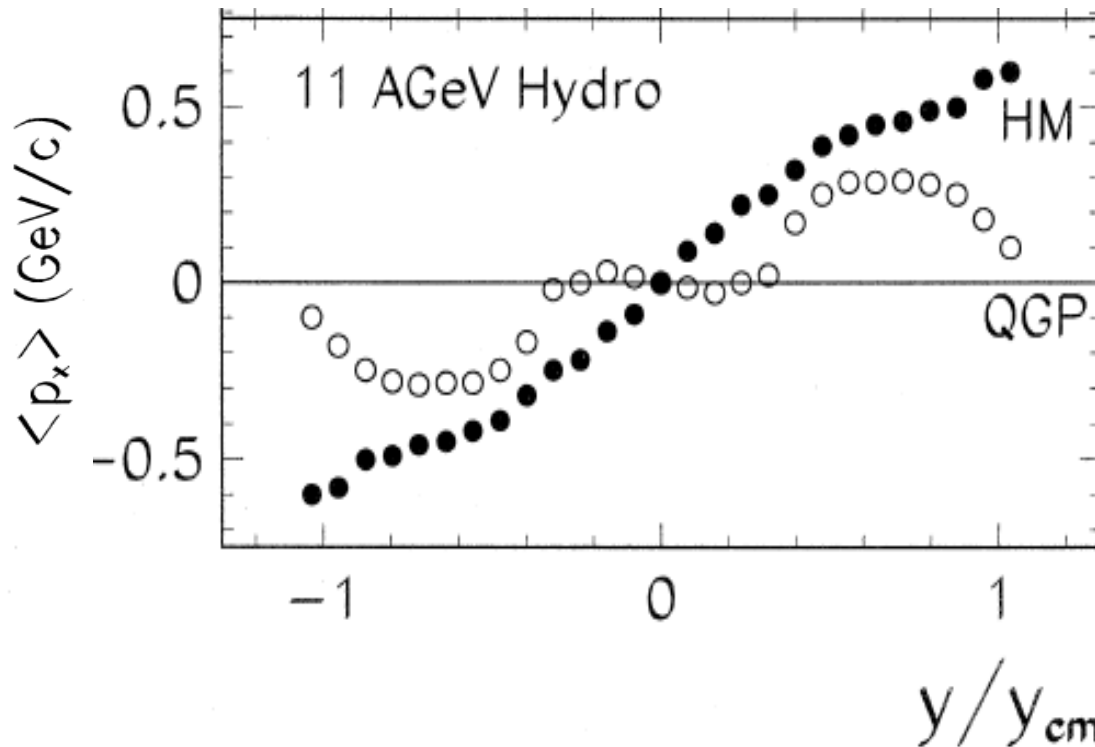
Outline

1. Introduction
2. Data analysis : STAR BES I
3. Results and discussion
4. Summary

Directed Flow

Directed Flow (v_1) : 1st harmonic in the Fourier expansion of particle azimuthal spectrum.

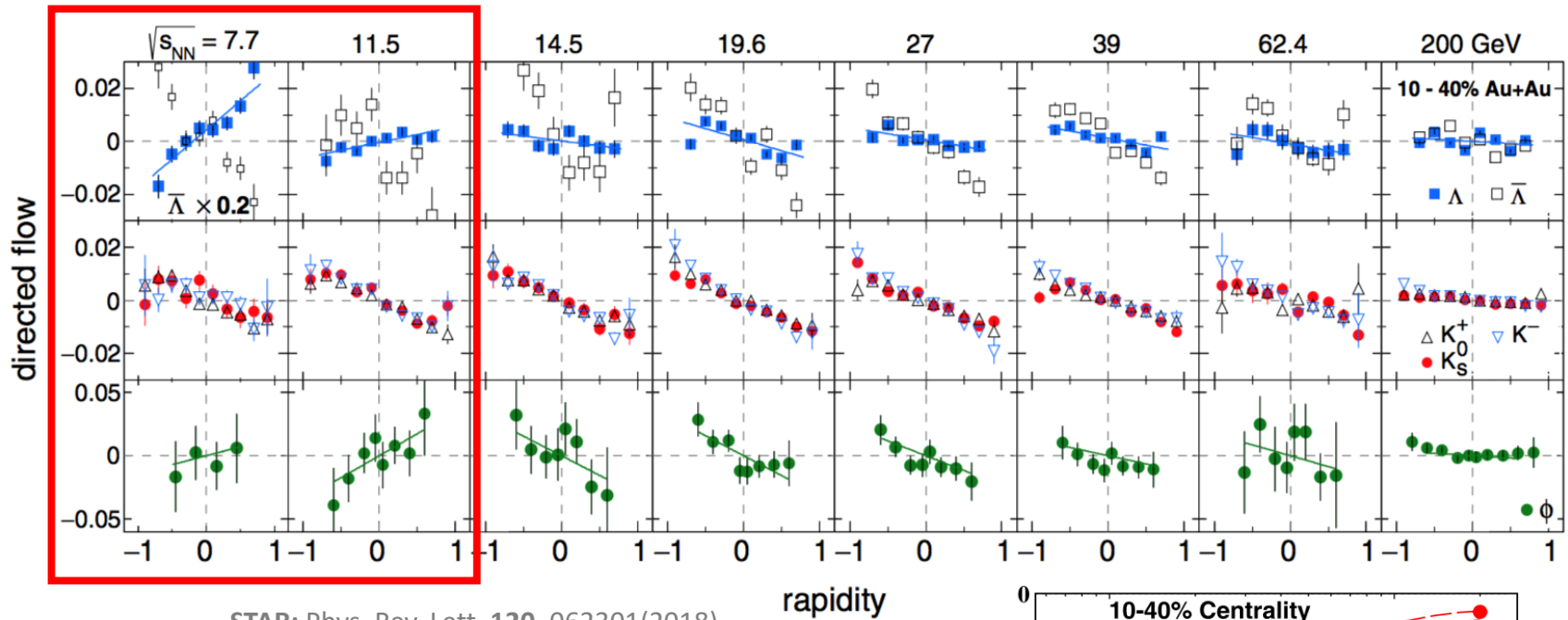
$$\frac{d^3N}{dp_T dy d\phi} = \frac{d^2N}{dp_T dy} (1 + 2v_1 \cos(\phi - \psi) + 2v_2 \cos 2(\phi - \psi) + \dots)$$



Phys. Lett. B 485, 454(1999)

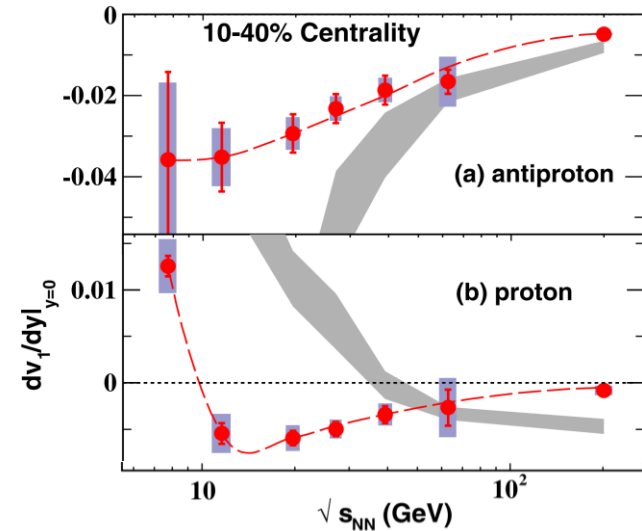
The **directed flow slope** at mid-rapidity is sensitive to softening of EOS?

Directed Flow v_1 in RHIC BES-I



Mesons and all produced baryons show negative slope except ϕ mesons when collisions energy < 14.5 GeV.

What about light nuclei??

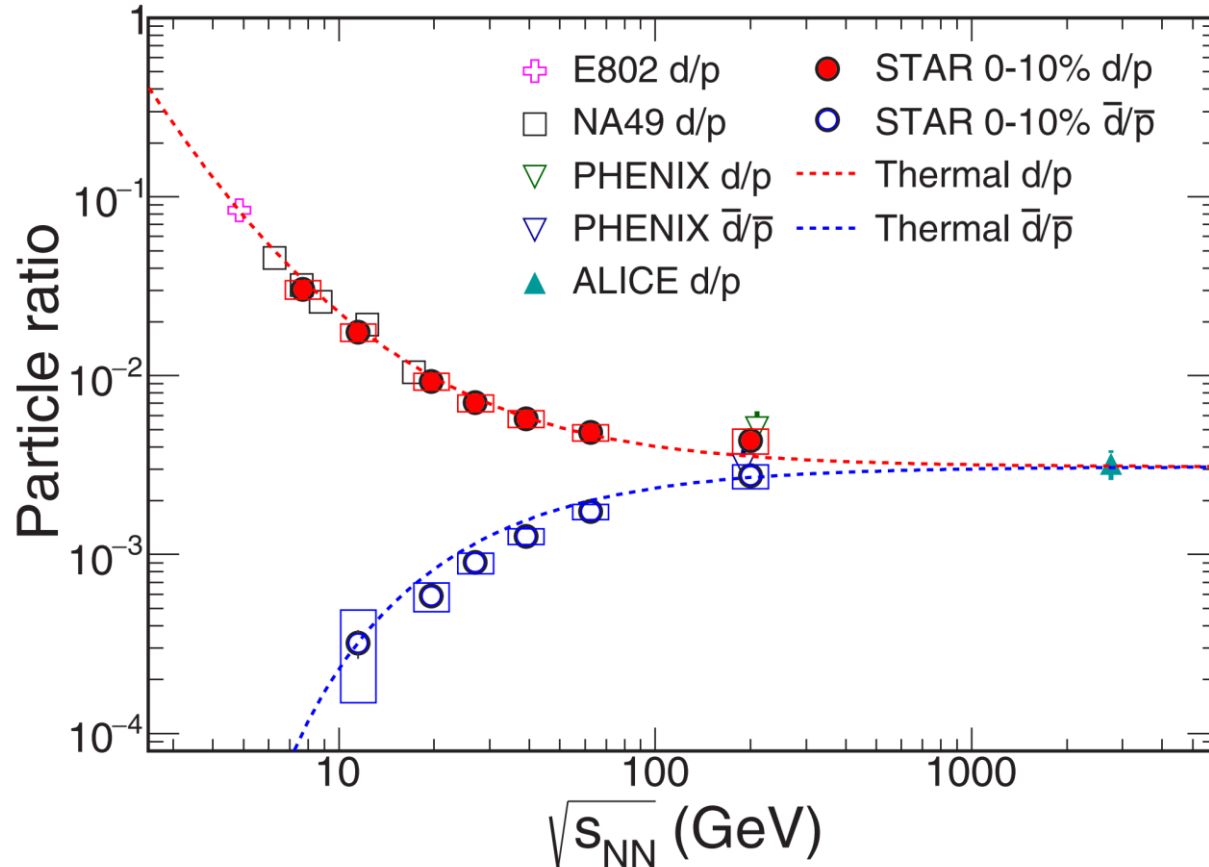


Deuteron Production

Thermal model

- Nuclei formed early | at hadronic freeze-out
- Deuteron yield can be described

STAR: Phys. Rev. C 99, 064905(2019)

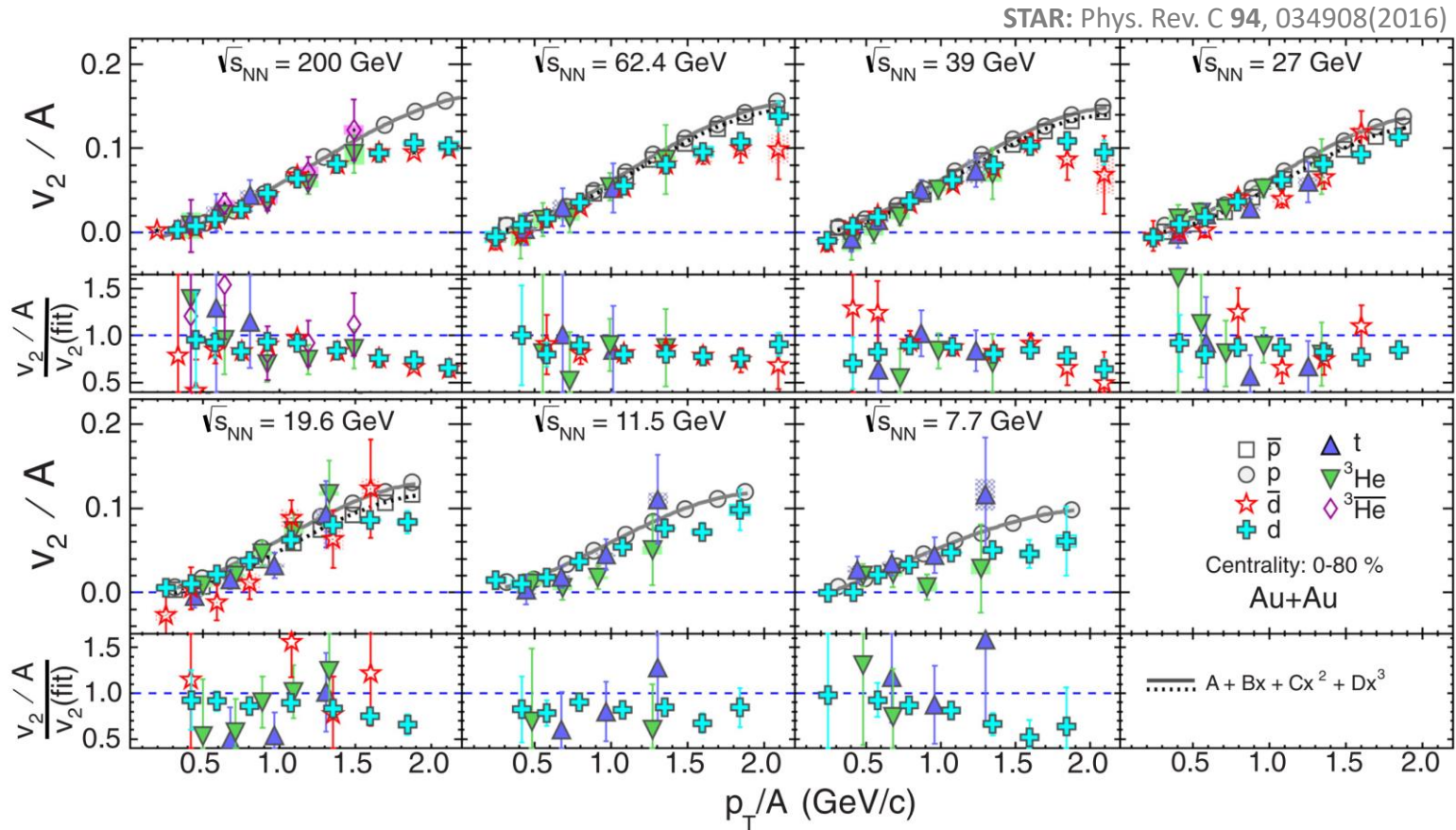


Deuteron binding energy is 2.2 MeV, how to keep bound in the fireball?

Light nuclei v_2

Coalescence Model

- Light nuclei formed at later stage of fireball evolution
- Through combination of protons and neutrons with close position and momentum

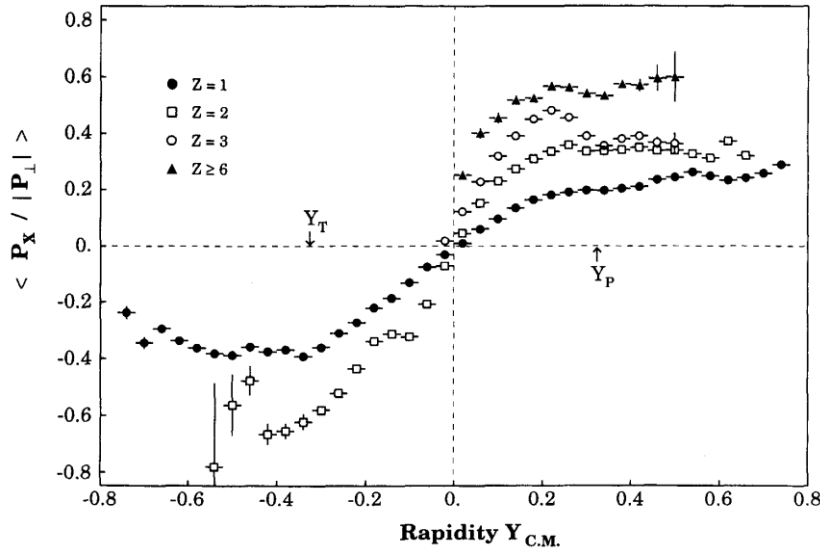


Mass scaling behavior of deuteron v_2 within $p_T < 3.0$ GeV/c nearly for all energies

Light Nuclei v_1 Measurements

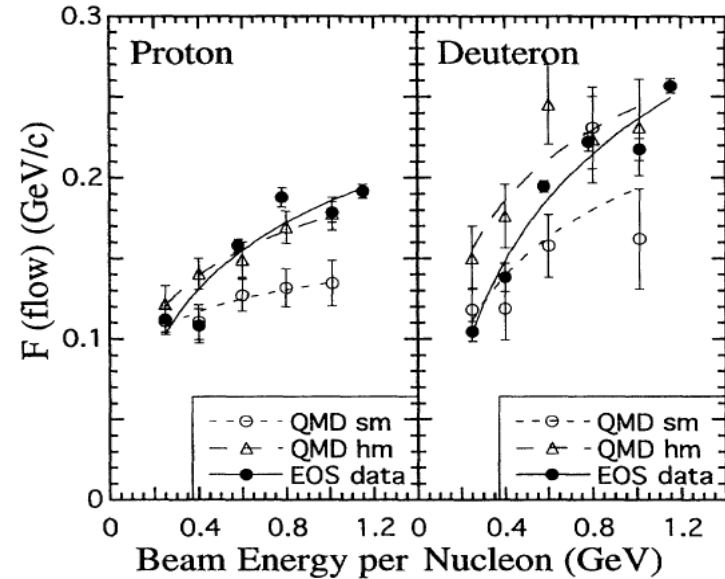
Phys. Rev. Lett. **59**, 2720(1987)

Au+Au 0.2 GeV/nucleon



Phys. Rev. Lett. **75**, 2100(1995)

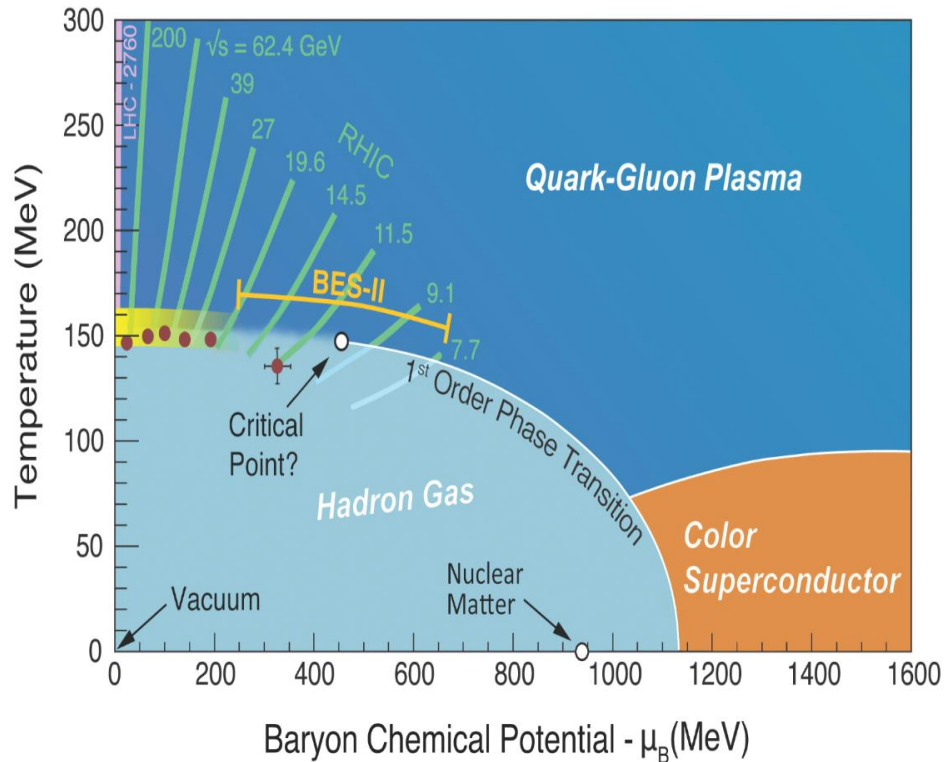
Au+Au 0.25A-1.15A GeV



- Stronger collective flow observed for heavier nuclei
- The proton and deuteron directed flow increase monotonically with rising beam energy
- The differences in fragment flow become larger with rising beam energy

How about BES program energies???

The Beam Energy Scan at RHIC/STAR



Map QCD phase diagram

- Search for **1st order phase transition**
- Search for **critical point**

Directed flow (v_1) is a key observable to search for the signature of 1st order phase transition.

Au+Au events usable for analysis

$\sqrt{s_{NN}}$ (GeV)	7.7	11.5	14.5	19.6	27	39
Events ($\times 10^6$)	4	12	10	36	70	130

Diagram of the STAR Detector

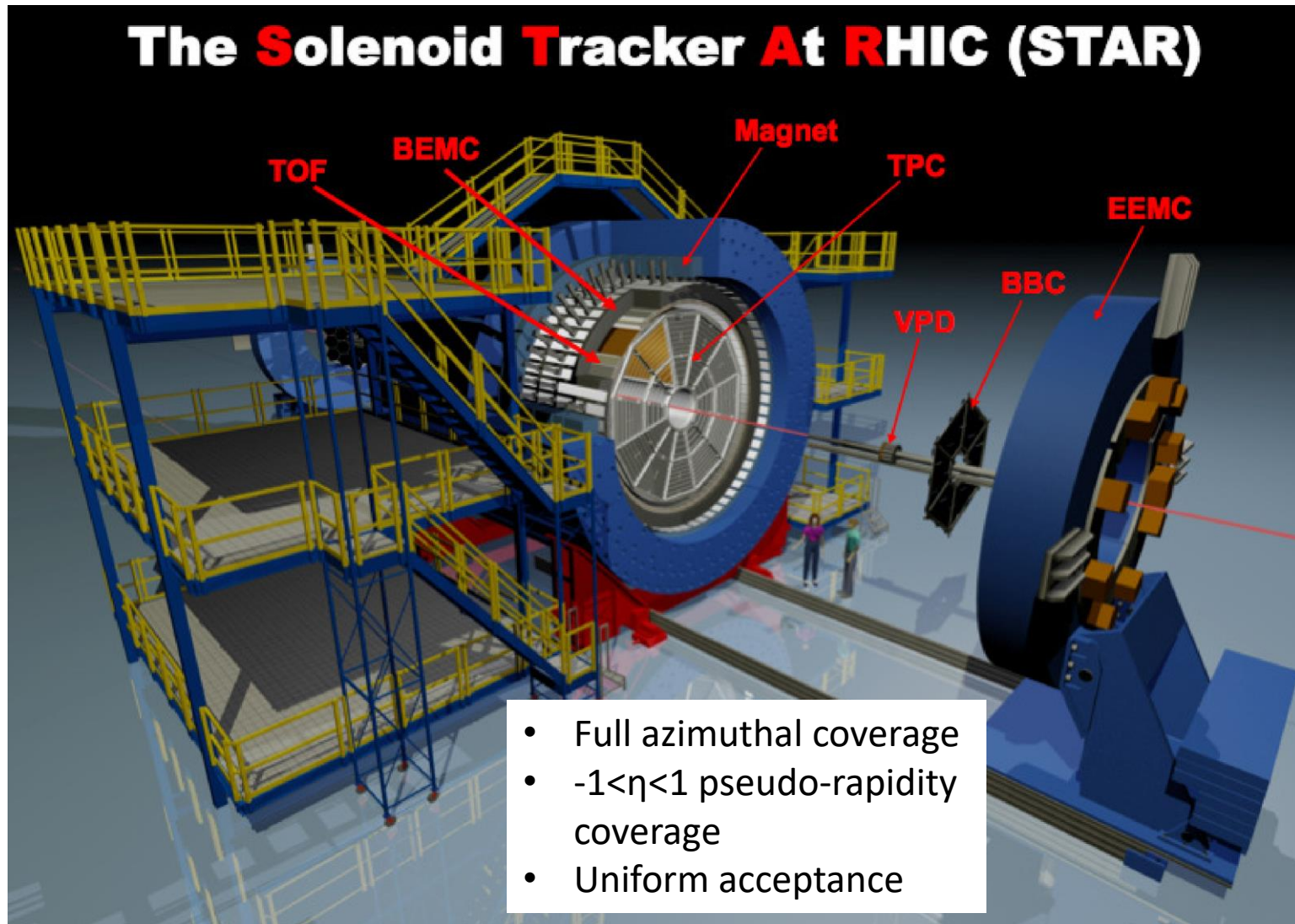


Diagram of the STAR Detector

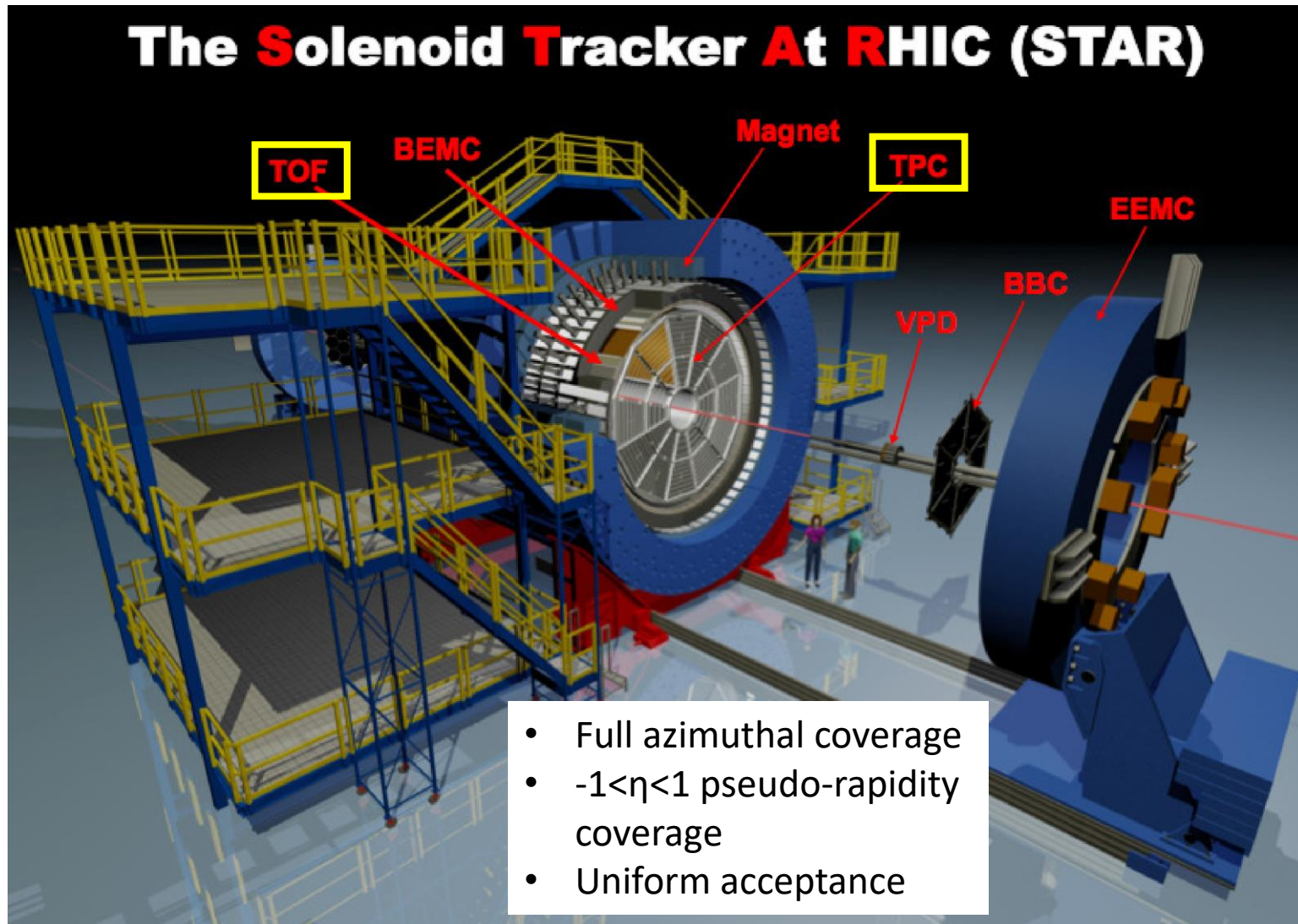
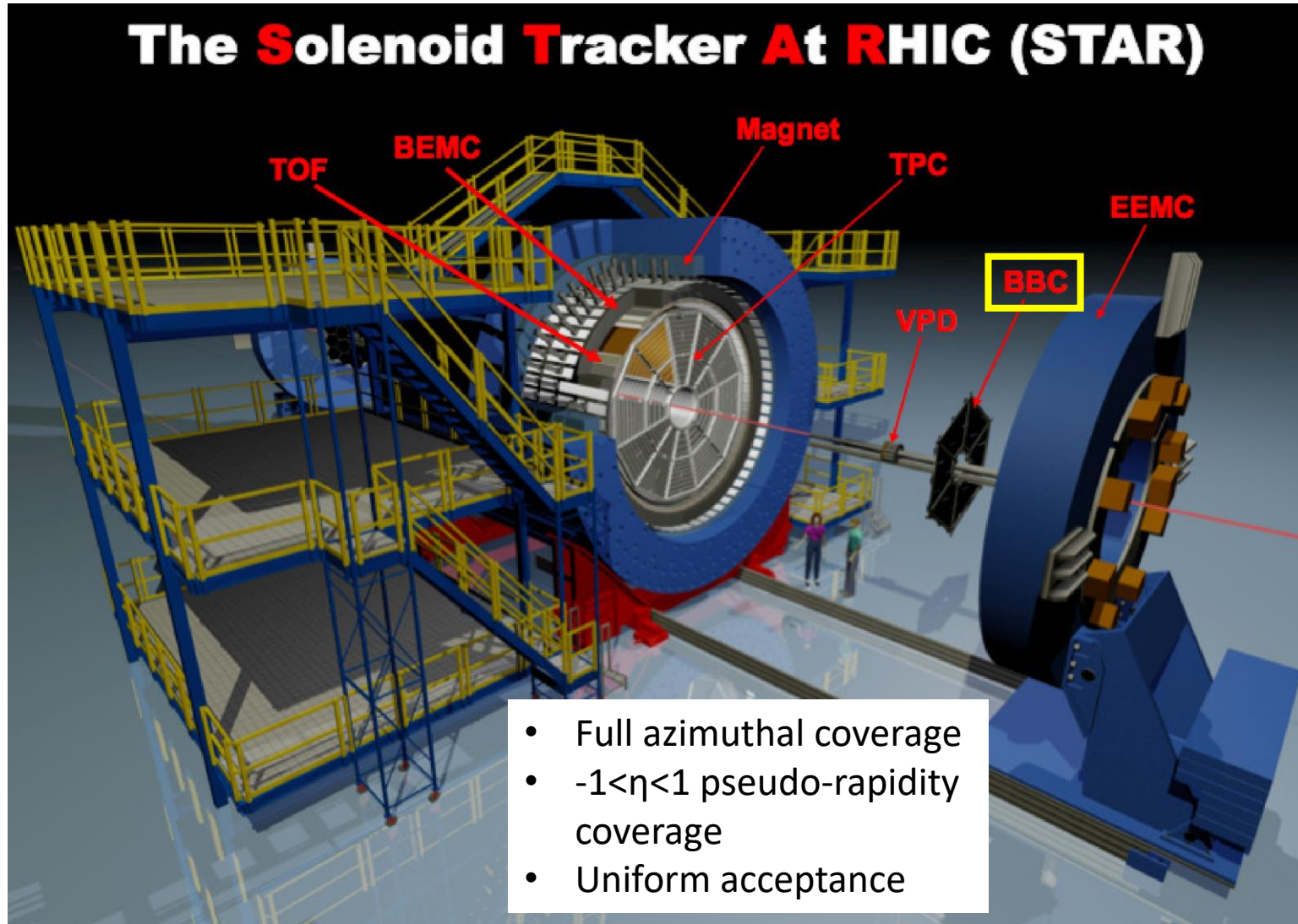
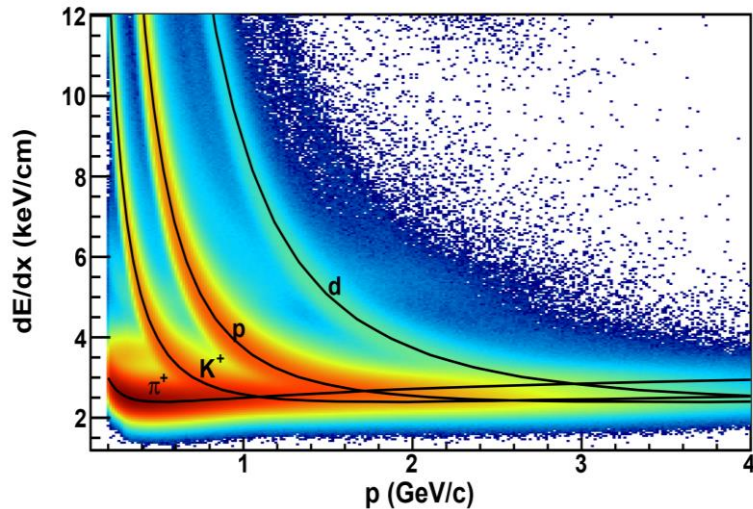


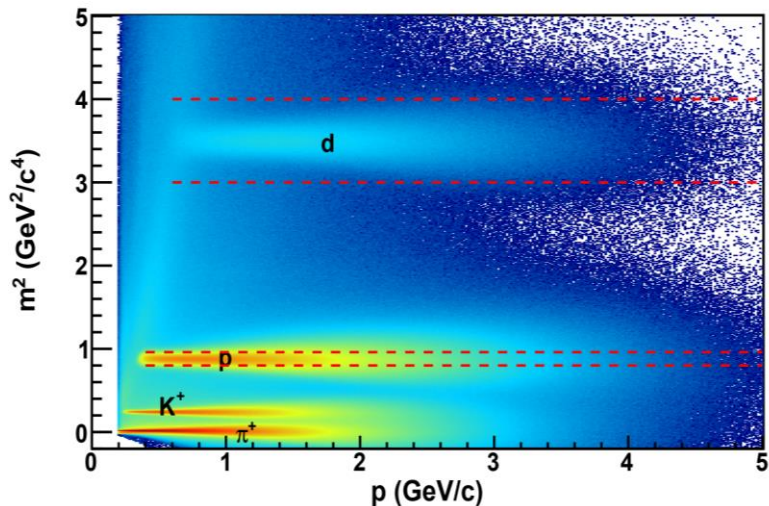
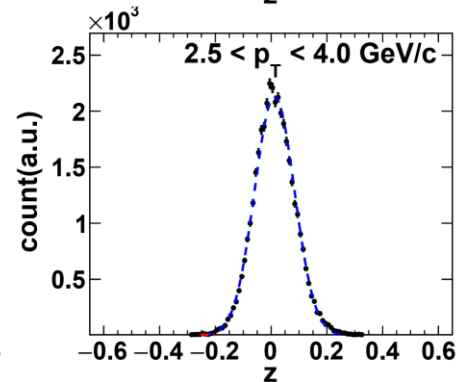
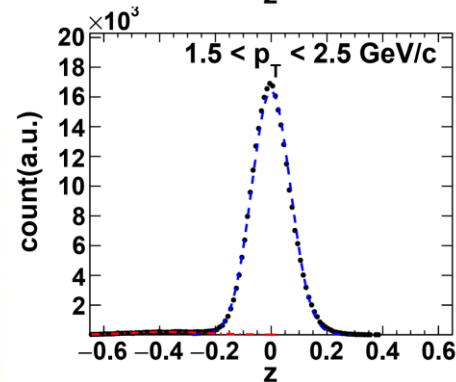
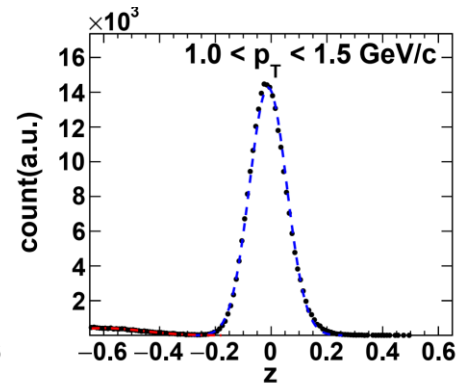
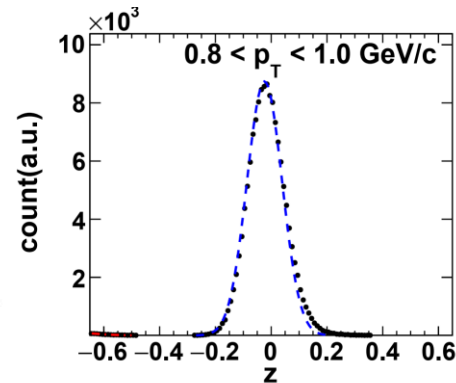
Diagram of the STAR Detector



Particle Identification

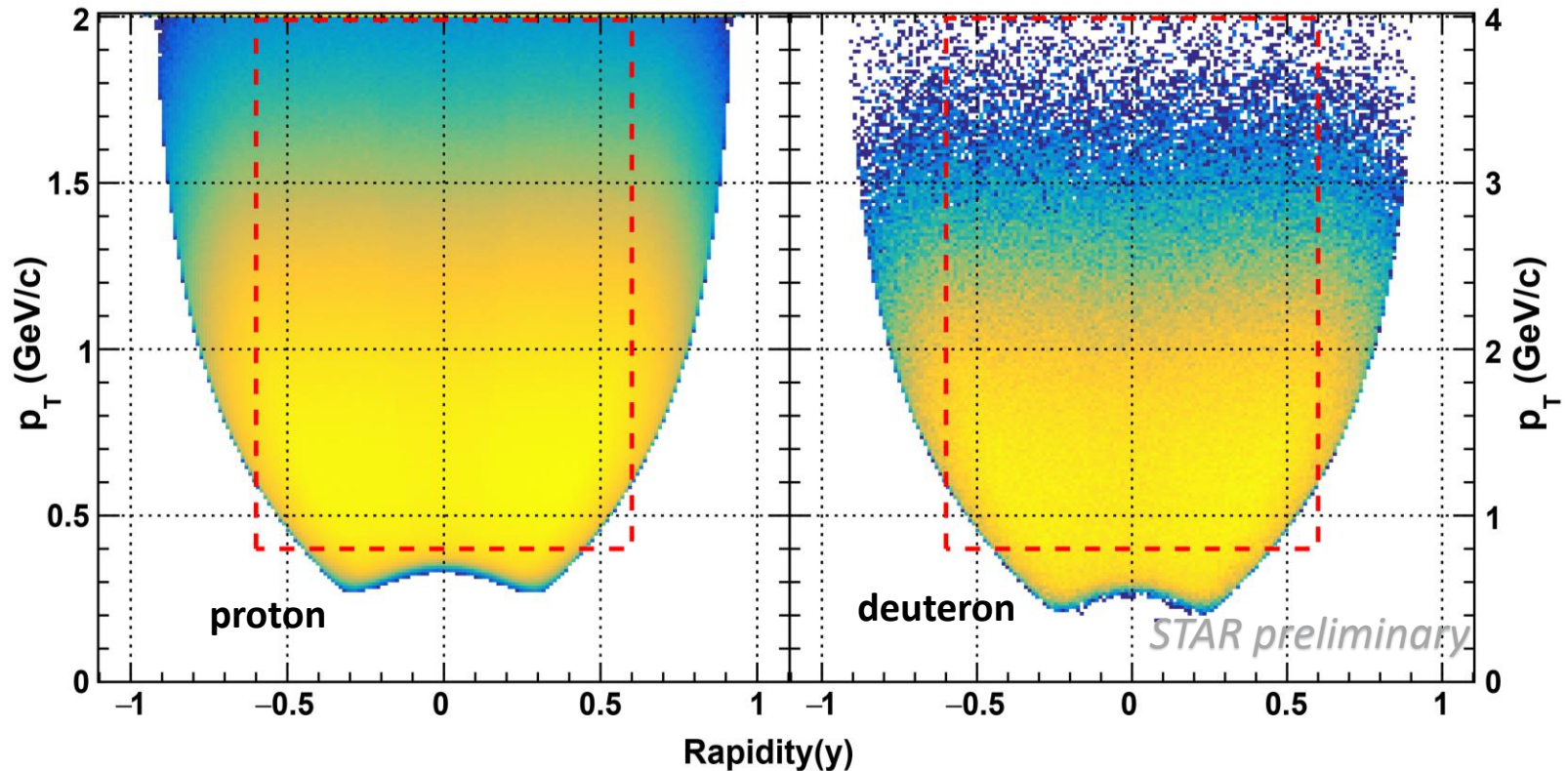


$$z = \ln \left(\frac{(dE/dx)_{\text{measured}}}{(dE/dx)_{\text{Bichsel}}} \right)$$



- Energy loss: $-0.2 < z < 0.2$
- m^2 : $3.0 < m^2 < 4.0$ (GeV/c 2) 2

Phase Space Distribution



For v_1 calculation

- Rapidity : $-0.6 < y < 0.6$
- Transverse momentum : $0.4 < p_T < 2.0$ GeV/c for proton;
 $0.8 < p_T < 4.0$ GeV/c for deuteron

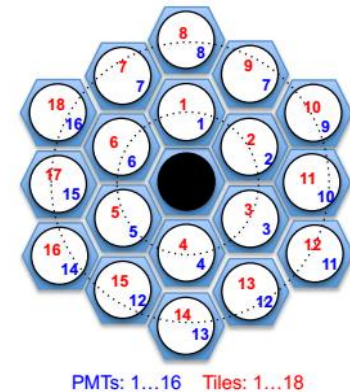
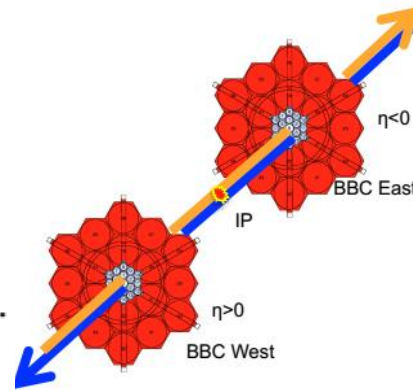
1st Order Event Plane Reconstruction

$$v_1 = \langle \cos(\phi - \Psi_{RP}) \rangle$$

$$Q_n \cos(n\Psi_n) = X_n = \sum_i w_i \cos(n\phi_i),$$

$$Q_n \sin(n\Psi_n) = Y_n = \sum_i w_i \sin(n\phi_i),$$

$$\Psi_n = \left(\tan^{-1} \frac{\sum_i w_i \sin(n\phi_i)}{\sum_i w_i \cos(n\phi_i)} \right) / n.$$

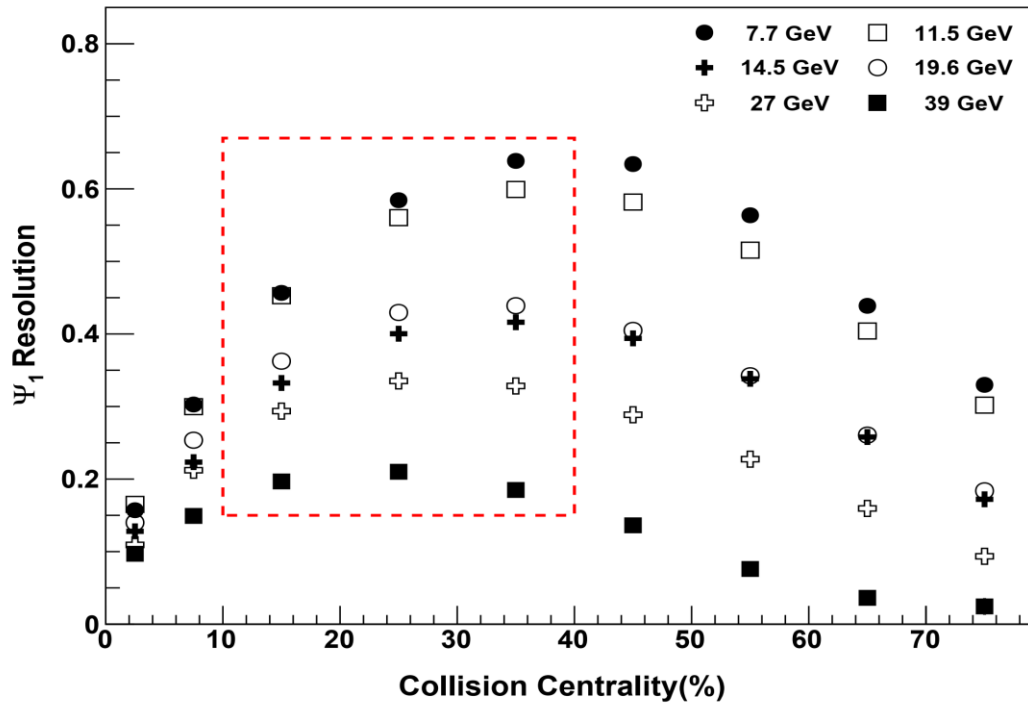


- 1st order event plane (ψ_1) estimated with east and west BBC detectors
 - BBC coverage $3.3 < |\eta| < 5.0$
 - large η gap between TPC and BBC reduces non-flow effects
- The raw ψ_1 distributions were flattened by shifting method

Event Plane Resolution

The estimated **event plane** with respect to the **real reaction plane** is calculated by the event plane resolution.

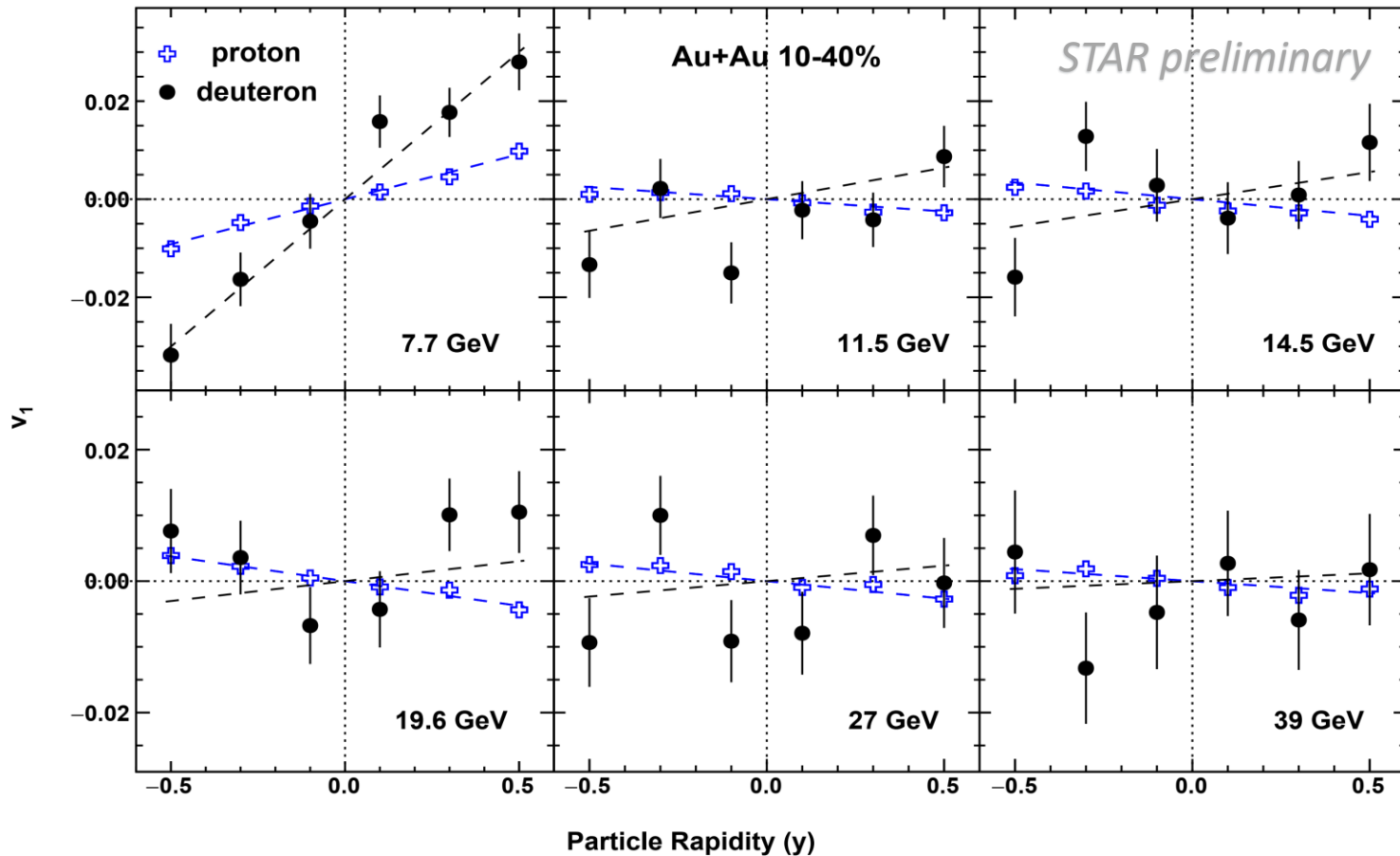
$$R_1 = \langle \cos(\psi_1 - \psi_{RP}) \rangle$$
$$\langle \cos(\psi_{\text{east}} - \psi_{\text{west}}) \rangle = \langle \cos(\psi_{\text{east}} - \psi_{RP}) \rangle \langle \cos(\psi_{RP} - \psi_{\text{west}}) \rangle$$



ψ_1 resolution improves at low collision energies because the stronger v_1 near the BBC rapidity coverage.

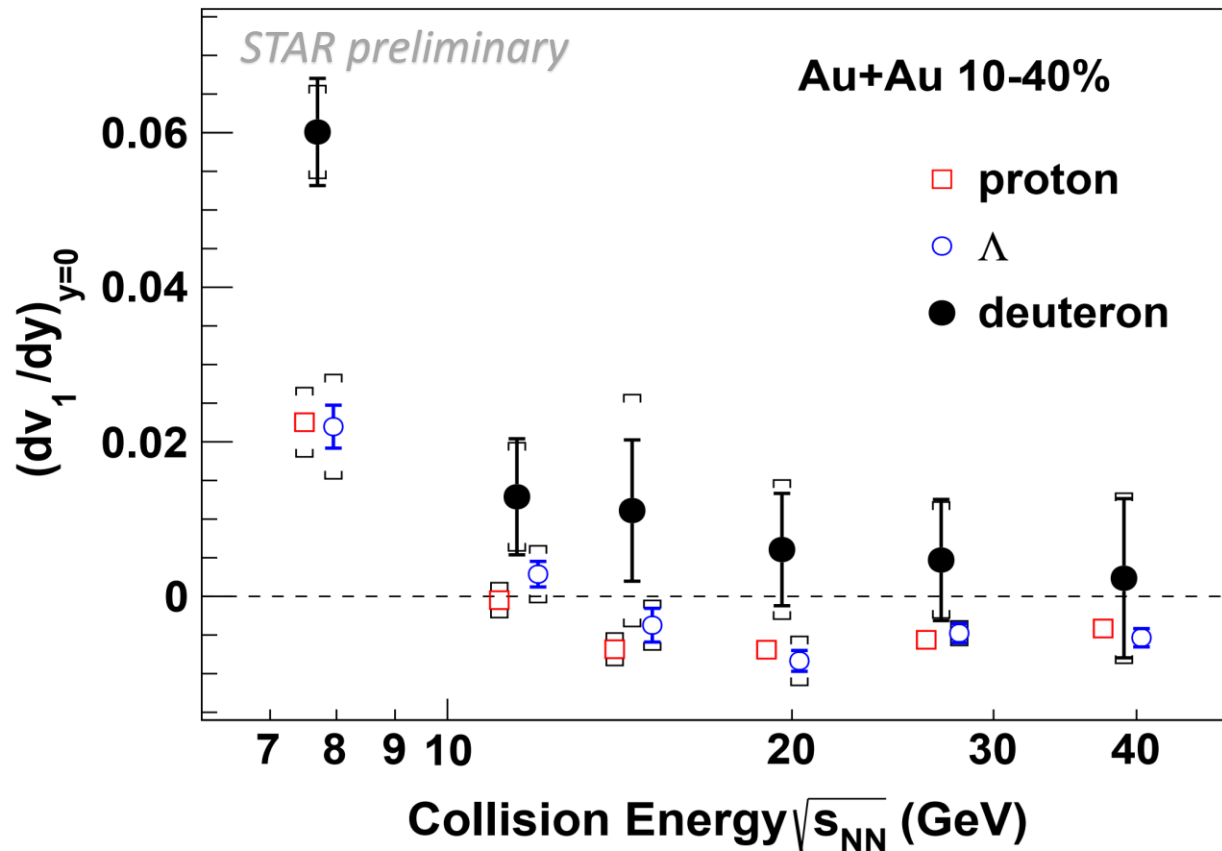
Rapidity Dependence of v_1

$$v_1 = \frac{v_1^{obs}}{R_1} = \frac{\langle \cos(\phi - \psi_1) \rangle}{R_1}$$



The slopes at the mid-rapidity ($|y| < 0.6$) were extracted with linear functions for mid-central collisions.

Energy Dependence of v_1 Slope



- The v_1 slopes at mid-rapidity of deuteron are positive for all energies
- Strong enhancement of deuteron v_1 slope observed at 7.7 GeV, while close to zero for 10 GeV

Coalescence Model

Deuteron v_1

constituent nucleons are close in space and have similar velocities.
At mid-rapidity:

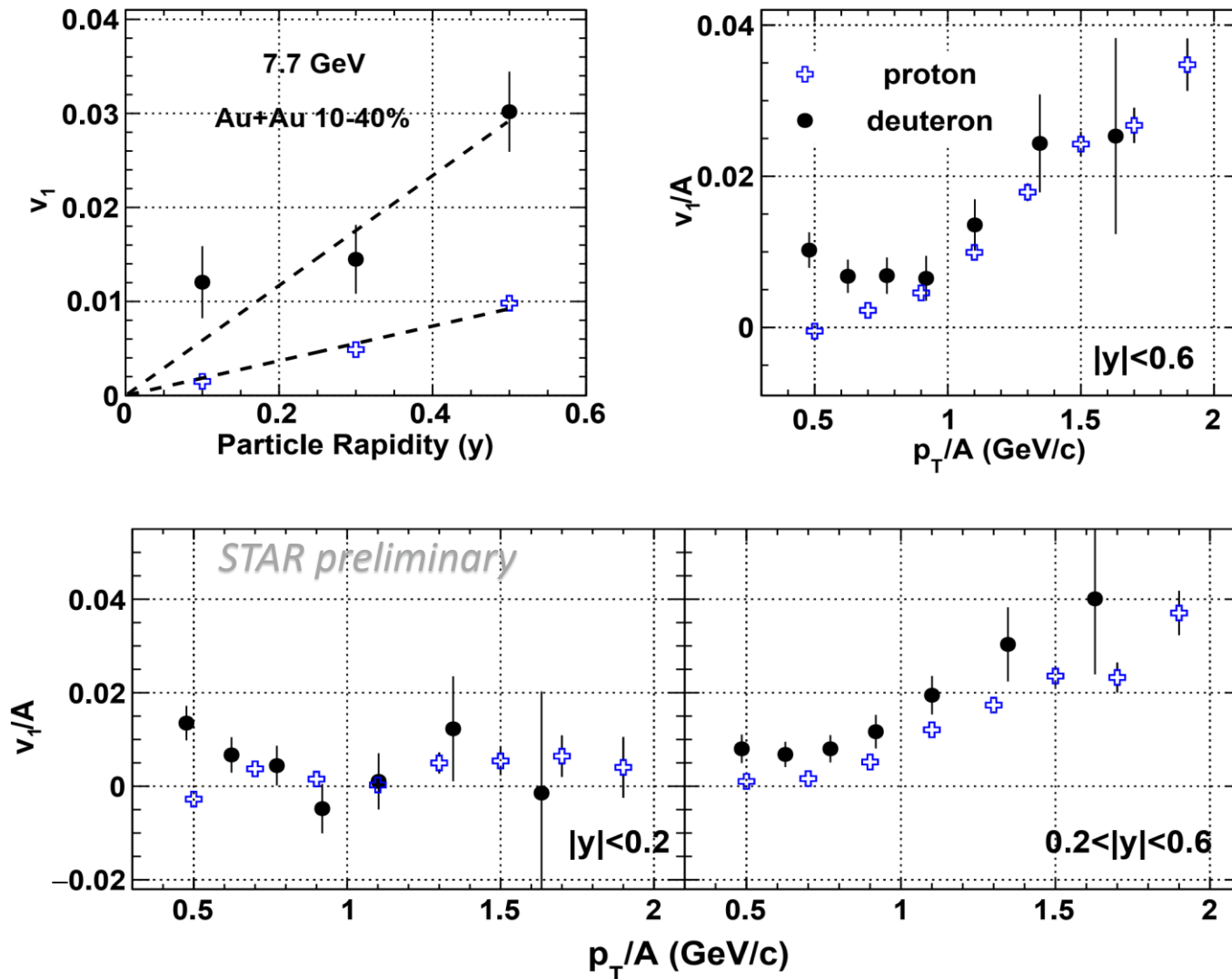
$$\vec{p}_T(d) \approx 2\vec{p}_T(p), \quad y(d) \approx y(p)$$

$$v_1^d(p_T, y) = \frac{2v_1^p\left(\frac{p_T}{2}, y\right)}{1 + \left(2v_1^p\left(\frac{p_T}{2}, y\right)\right)^2}$$

if $v_1 \ll 1$

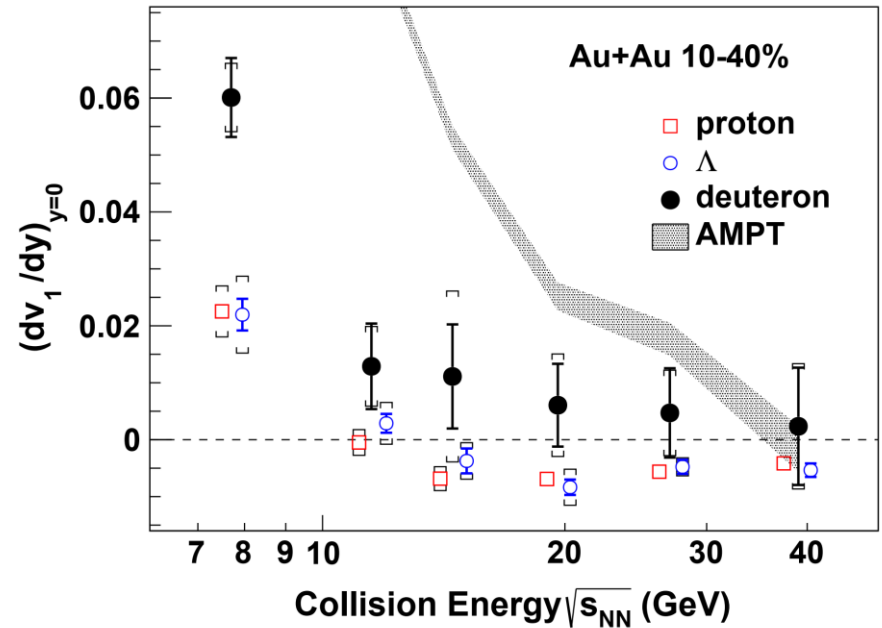
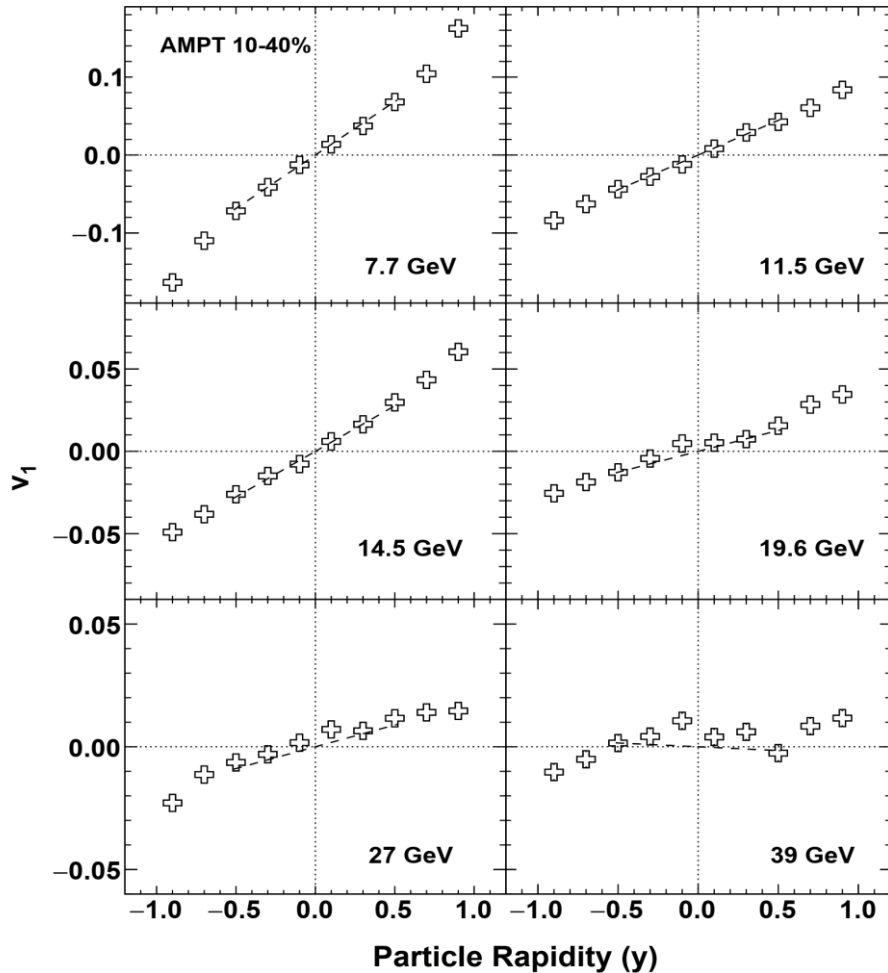
$$v_1^d(p_T, y) \approx 2v_1^p\left(\frac{p_T}{2}, y\right)$$

p_T Dependence of Deuteron v_1 at 7.7 GeV



Deuteron v_1 from AMPT Simulation

In AMPT, (anti-)deuterons are produced and dissolved via nuclear reaction in the hadronic transport stage of AMPT.



For AMPT simulation, values of dv_1/dy are also positive for all energies.

Summary

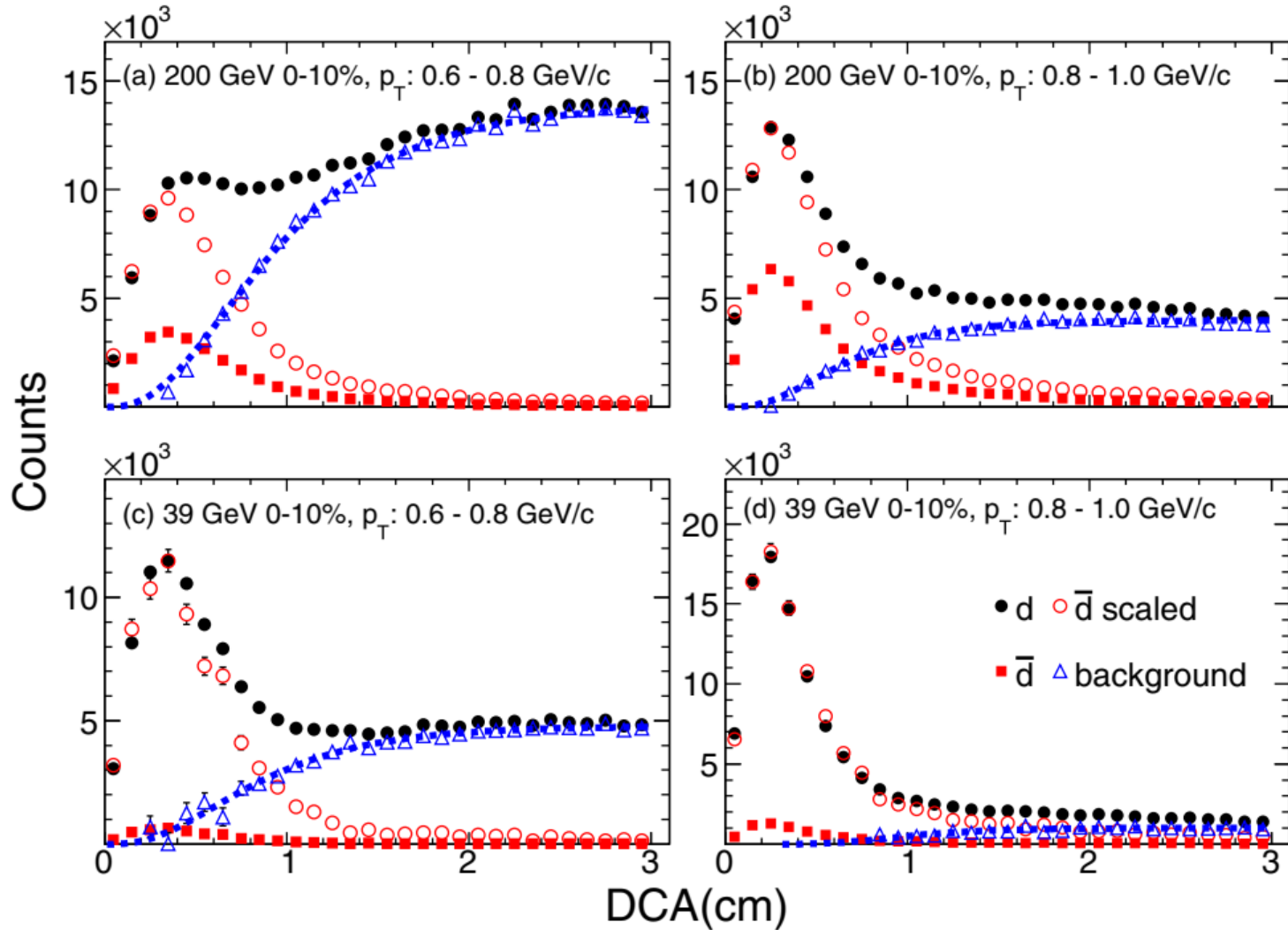
- The deuteron v_1 was measured in Au+Au collisions at $\sqrt{s_{NN}}=7.7$ -39 GeV. The slopes at midrapidity ($|y|<0.6$) were extracted
- The dv_1/dy of deuteron are positive for all energies. Strong enhancement observed at $\sqrt{s_{NN}} = 7.7$ GeV, while close to zero for $\sqrt{s_{NN}} > 10$ GeV
- AMPT simulation : the dv_1/dy are larger than the measurement for most energies
- **Outlook** : precise measurement of v_1 for light nuclei with BES II

Thank you for Your Attention!

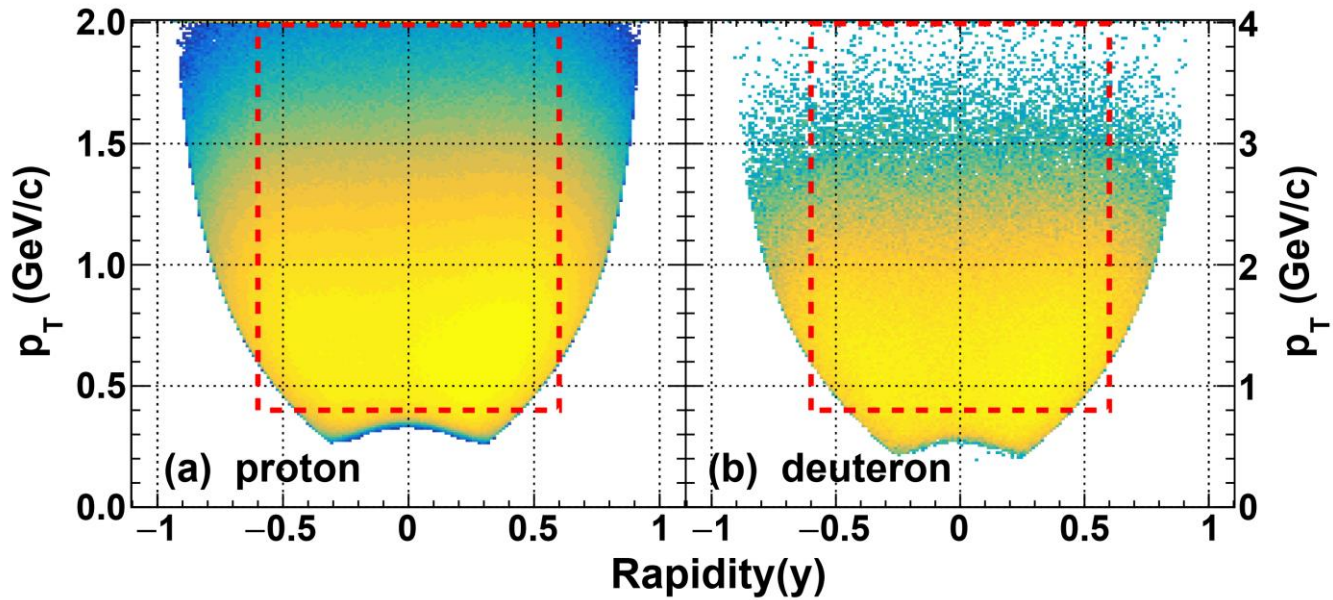
Back Up

DCA distribution

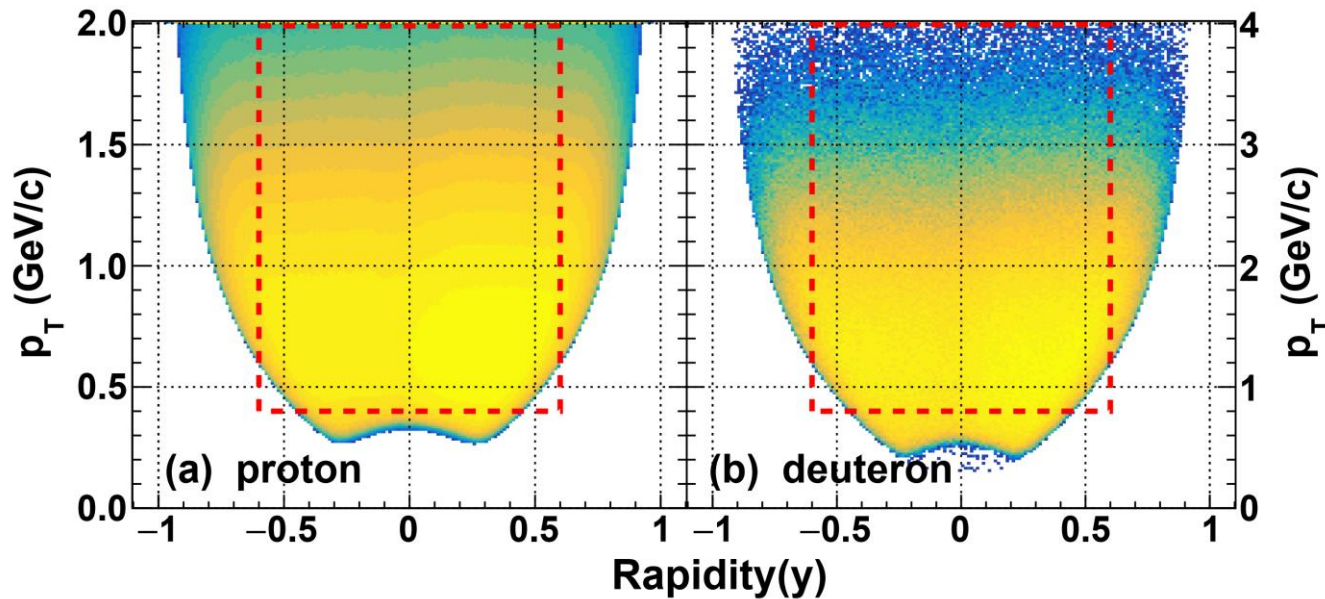
STAR: Phys. Rev. C **99**, 064905(2019)



Phase space distribution



7.7 GeV



39 GeV