

Exploiting jet topological differences in *pp* and Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV at RHIC



Raghav Kunnawalkam Elayavalli (raghav.ke@yale.edu),Yale/BNL, for the STAR Collaboration

Jets are multi-scale objects, and their characteristic evolution in each scale translates to a varied bounty of jet topologies. In this poster, we present novel measurements of the jet shower in *pp* and Au+Au collisions at STAR and discuss their connection to parton evolution. Formation times measurements in *pp* collisions point to onset of non-perturbative behavior. Measurements of subjet opening angles and formation times are presented for specially selected di-jet events in Au+Au collisions leading towards a first ever space-time study of jet quenching phenomena.

*

 $\frac{p_{\rm T,2}}{p_{\rm T,1} + p_{\rm T,2}}$

 $\theta = \Delta R \ (1,2); E = E_1 + E_2$

 $z \cdot (1-z) \cdot \overline{\theta^2 \cdot E}$

Jet formation time in pp collisions

Supported in part by the

U.S. DEPARTMENT OF

Office of Science

- SoftDrop selected splits correspond to mostly early splittings as compared to formation time defined using the two leading charged particles which are necessarily hadronized remants of the parton shower.
- Resolved splittings (where the leading and subleading charged particles are resolved in the C/A cluster tree) has similar shape as charged particle formation time at large values highlighting enhanced impact of non-perturbative QCD corrections at later formation times in the jet clustering tree.
- Subjet formation time has similar shape to the resolved splits with a large dynamic range (0.5 fm/c – 10 fm/c)



Subjet opening angle and formation time in Au+Au collisions

Au+Au collision dataset from in 2014 provides large statistics to scan across various opening angles and formation times.
Recoil jet's subjet opening angle and formation time for



R = 0.1 and R = 0.05 subjets with constituent p_T > 2.0 GeV/c
Ongoing analysis to study jet quenching observables for jets of different topologies.





Various opening angle and formation time selections with high statistics

Conclusions

- Comparison of the different splits highlights the transition from pQCD to npQCD.
- Resolved splits show similar shape as the charged particle split at large values occurring in the predominantly non-perturbative region.
- Scan across emission phase-space leads to first ever space-time tomography of the QGP – selections can be varied with different opening angles and formation times.



The STAR Collaboration https://drupal.star.bnl.gov/STAR/presentations