



Improving results on transverse double spin asymmetries in the CNR region at STAR

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20th INTERNATIONAL SYMPOSIUM ON SPIN PHYSICS

JINR, Dubna, Russia

September 17 - 22, 2012



Matrix elements

$$\begin{aligned} \phi_1(s, t) &= \langle ++ | M | ++ \rangle \text{ spin non-flip} \\ \phi_2(s, t) &= \langle ++ | M | -- \rangle \text{ double spin flip} \\ \phi_3(s, t) &= \langle +- | M | +- \rangle \text{ spin non-flip} \\ \phi_4(s, t) &= \langle +- | M | -+ \rangle \text{ double spin flip} \\ \phi_5(s, t) &= \langle ++ | M | +- \rangle \text{ single spin flip} \\ \phi_i(s, t) &= \phi_i^{EM}(s, t) + \phi_i^{HAD}(s, t) \end{aligned}$$

$$A_{NN}(s, t) \frac{d\sigma}{dt} = \frac{4\pi}{s^2} \left\{ 2|\phi_5|^2 + \text{Re}(\phi_1^* \phi_2 - \phi_3^* \phi_4) \right\}$$

$$A_{SS}(s, t) \frac{d\sigma}{dt} = \frac{4\pi}{s^2} \text{Re} \left\{ \phi_1 \phi_2^* + \phi_3 \phi_4^* \right\}$$

Observables: cross sections and spin asymmetries

$$\sigma_{tot} = \frac{4\pi}{s} \text{Im} \{ \phi_1 + \phi_3 \}_{t=0} = \frac{4\pi}{s} \text{Im} \phi_+ |_{t=0}$$

$$\frac{d\sigma}{dt} = \frac{2\pi}{s^2} \left\{ |\phi_1|^2 + |\phi_2|^2 + |\phi_3|^2 + |\phi_4|^2 + 4|\phi_5|^2 \right\}$$

$$A_N(s, t) \frac{d\sigma}{dt} = \frac{-4\pi}{s^2} \text{Im} \left\{ \phi_5^* (\phi_1 + \phi_2 + \phi_3 - \phi_4) \right\}$$

$\phi_5^{HAD} \approx 0$ This experiment

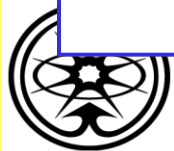
ϕ_5^{EM} small in CNI

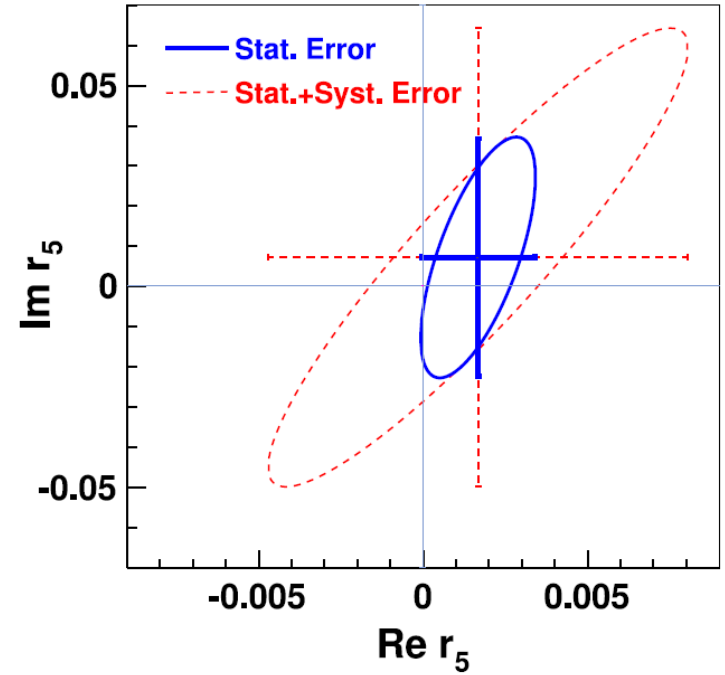
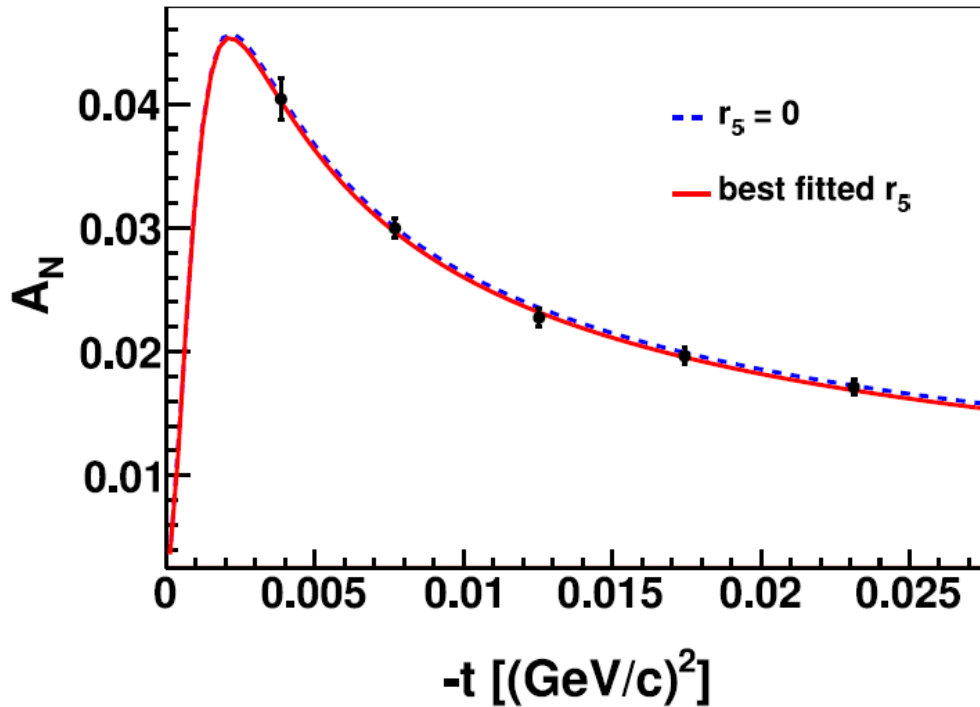
$$\frac{A_{NN} + A_{SS}}{2} \frac{d\sigma}{dt} = \frac{4\pi}{s^2} \text{Re} \left\{ \phi_1 \phi_2^* \right\}$$

$\phi_4^{HAD} \sim t \rightarrow 0$

$A_{NN} - A_{SS} \approx 0$

Probe for ϕ_2^{HAD}





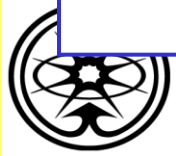
This experiment:

- Statistical error of 3% in each of 5 points
- Highest accuracy in extraction of r_5

arXiv:1206.1928 [nucl-ex]
also submitted for publication

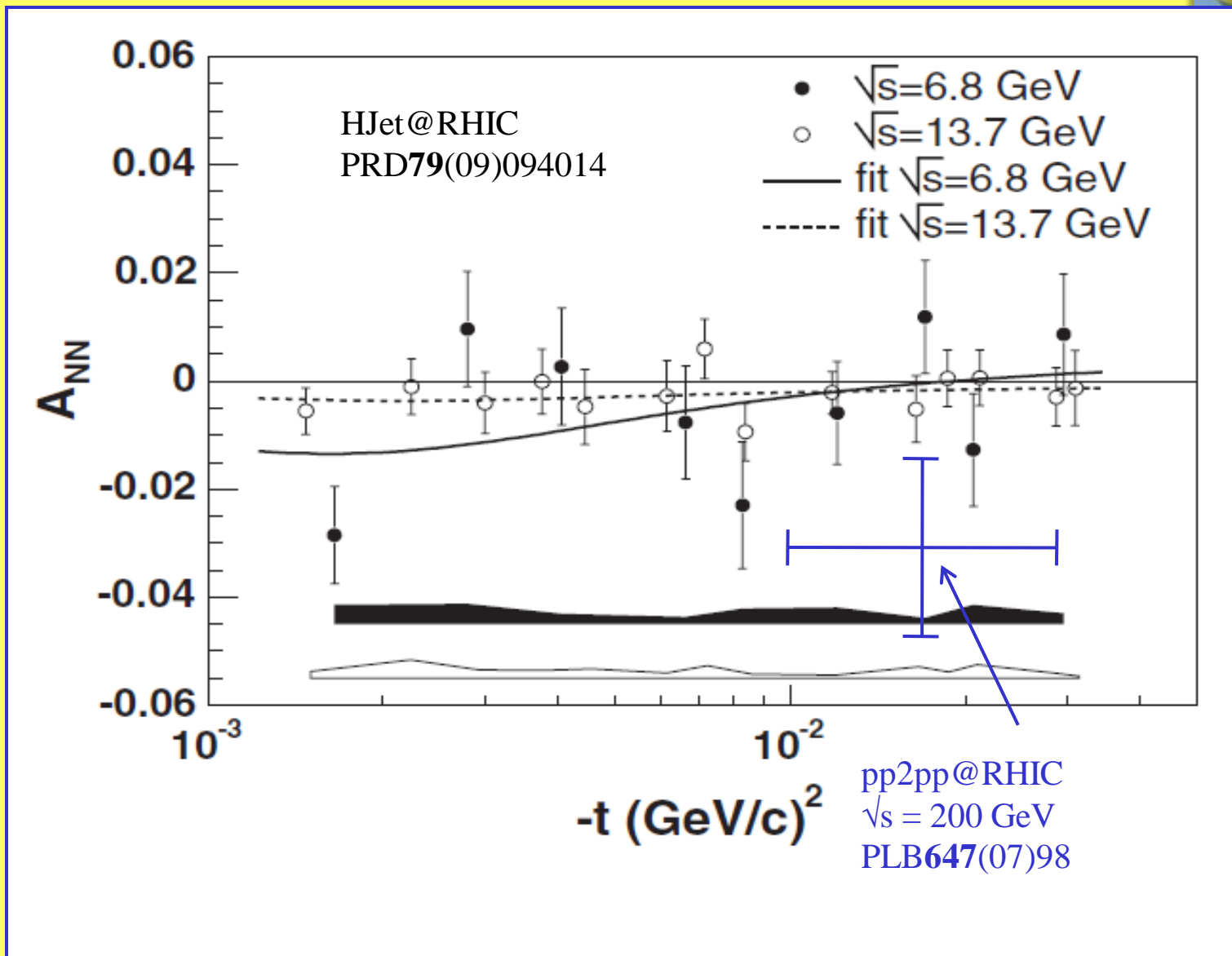
$$r_5 = \frac{2m_p}{\sqrt{-t}} \frac{\phi_5^{had}}{\text{Im}(\phi_1^{had} + \phi_3^{had})}$$

$$\begin{aligned} \text{Re } r_5 &= 0.0017 \pm 0.0063 \\ \text{Im } r_5 &= 0.007 \pm 0.057 \end{aligned}$$



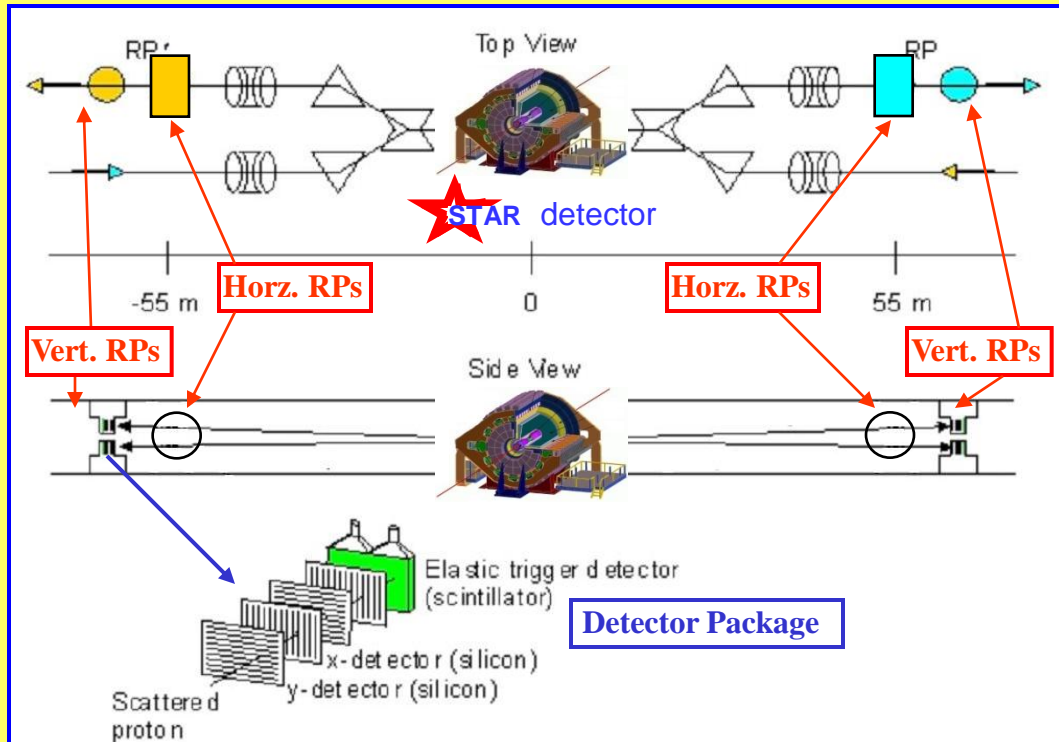


Previous measurements of A_{NN}



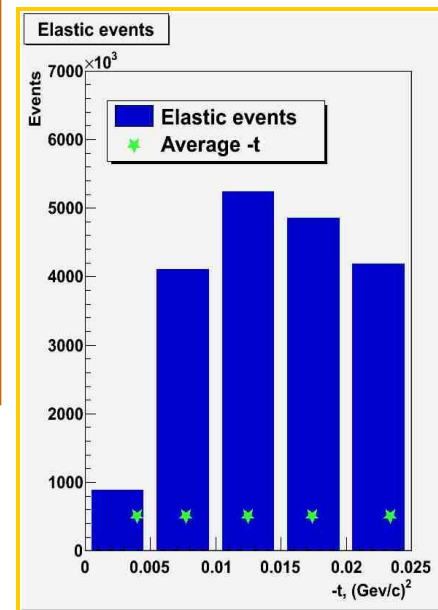
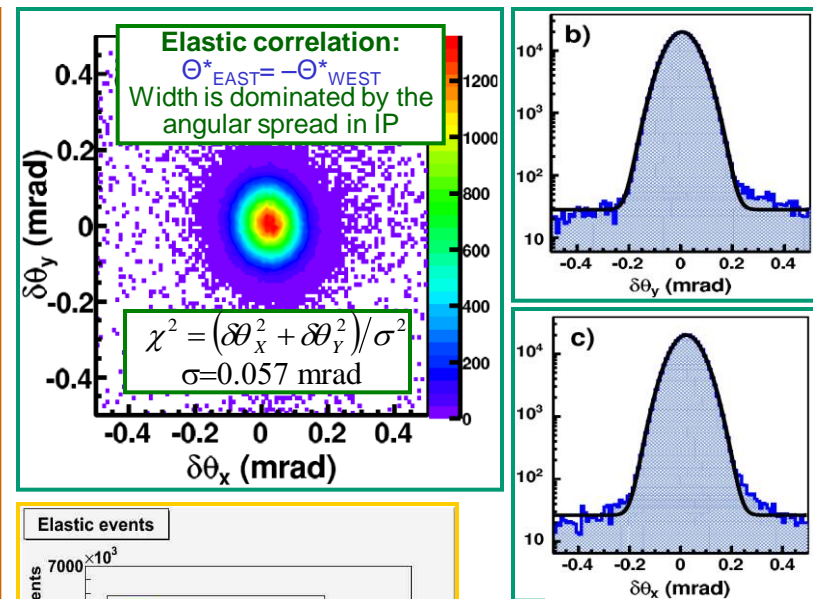
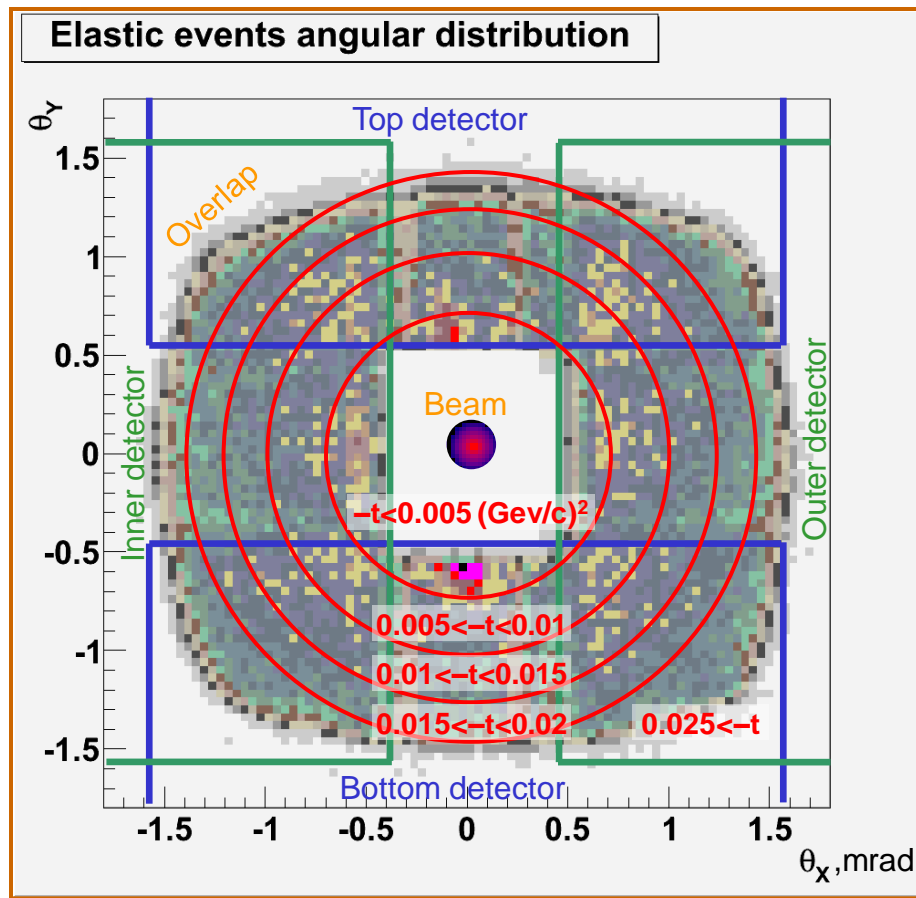


Detector layout and experimental conditions



- ✓ Roman Pots integrated with STAR detector – closest proximity to the beam.
- ✓ CNI region – $0.003 < -t < 0.03$.
- ✓ Ideal beam optics: $\beta^* = 21\text{m}$ and parallel to point focusing – terms other than L_{EFF} in the transport matrix were very small.
- ✓ High transverse polarization of both beams $\sim 60\%$.
- ✓ Excellent detector performance – nearly 100% efficiency and only 5 dead/noisy strips per ~ 14000 active strips.





✓ $2 \cdot 10^7$ elastic events with background $< 1\%$.

✓ 5 $(-t)$ ranges up to 0.03 (GeV/c)^2

- ✓ 2π acceptance in ϕ .
- ✓ Exactly the same sample of elastic events as for A_N studies

Cross-section azimuthal angular dependence for transversely polarized beams:

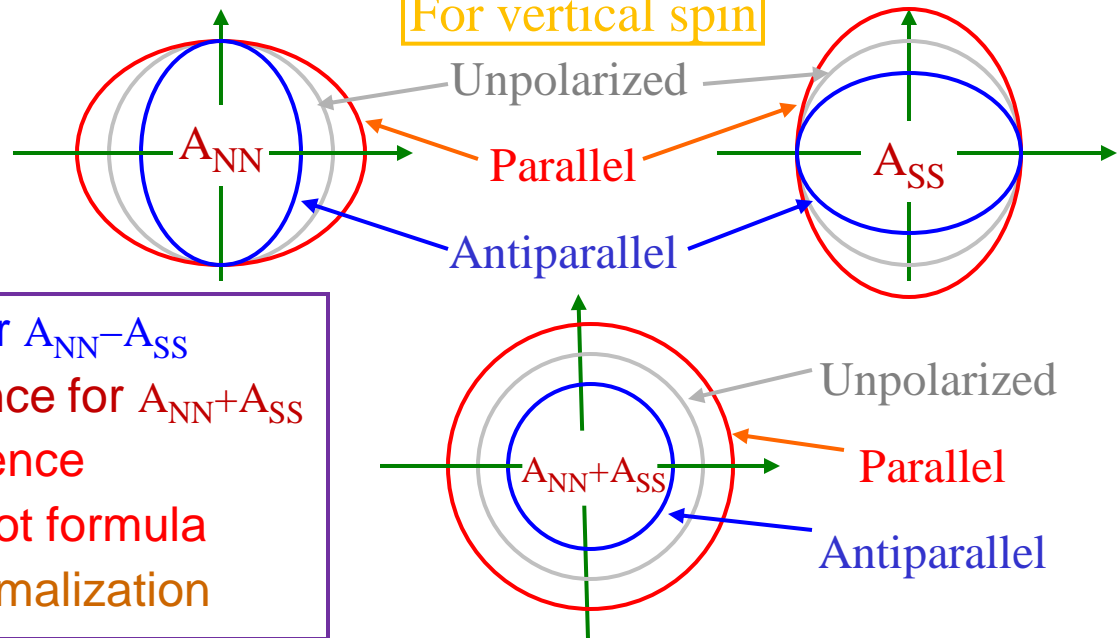
$$\sigma = \sigma_0 \left[1 + A_N (\vec{P}_B + \vec{P}_Y) \cdot \vec{n} + A_{NN} (\vec{P}_B \cdot \vec{n})(\vec{P}_Y \cdot \vec{n}) + A_{SS} (\vec{P}_B \cdot \vec{s})(\vec{P}_Y \cdot \vec{s}) \right]$$

\vec{n} - vector normal to the scattering plane $\vec{P}_B ; \vec{P}_Y$ - polarizations of the two beams

$\vec{s} = \frac{\vec{n} \times \vec{p}}{|\vec{n} \times \vec{p}|}$ - is the vector in the scattering plane, normal to the initial momentum

$$2\pi \frac{d^2\sigma}{dtd\varphi} = \frac{d\sigma}{dt} \cdot \left(1 + (P_B + P_Y)A_N \cos \varphi + P_B P_Y (A_{NN} \cos^2 \varphi + A_{SS} \sin^2 \varphi) \right)$$

For vertical spin



- ✓ $\cos 2\varphi$ dependence for $A_{NN} - A_{SS}$
- ✓ NO angular dependence for $A_{NN} + A_{SS}$
- Just cross-section difference
- ✓ Cannot use square root formula
- ✓ Must use external normalization



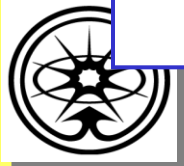
Normalization

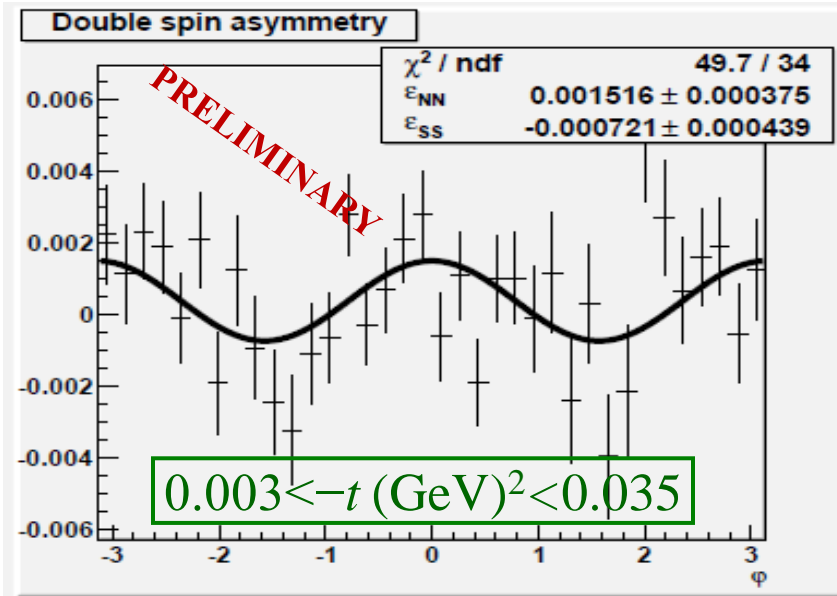


- Normalization is based on “inelastic” event counts assuming their negligible polarization dependence, yet this needs to be proved
- Normalized counts: $K^{+/-} = N^{+/-}/L^{+/-}$, $N^{+/-}$ -- elastic event counts for a certain spin combination, $L^{+/-}$ -- normalization factor
- Cannot use square root formula – have to rely on normalized counts $K^{+/-}$

$$\begin{aligned}\varepsilon_{NN}(\varphi) &= P_B P_Y (A_{NN} \cos^2 \varphi + A_{SS} \sin^2 \varphi) = \\ &= \frac{(K^{++}(\varphi) + K^{--}(\varphi)) - (K^{+-}(\varphi) + K^{-+}(\varphi))}{(K^{++}(\varphi) + K^{--}(\varphi)) + (K^{+-}(\varphi) + K^{-+}(\varphi))}\end{aligned}$$

- Study of several independent STAR subsystems, all having nearly or fully 2π acceptance for forward particles in both directions (East and West)
 - BBC – beam-beam counters
 - VPD – vertex position detector
 - ZDC – zero degree calorimeter
 - In addition WCM – RHIC wall current intensity monitor





- Based on BBC and VPD normalizing counts.
- The values are of the order 10^{-3} .
- Large normalization uncertainty 0.25%.
- Needed much better analysis of the normalization possibilities.

For further normalization studies use INDEPENDENT ratios instead of $L^{+/-}$:

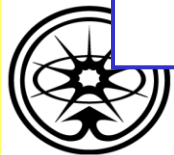
$N = N^{++} + N^{-} + N^{+-} + N^{+}$ – total counts of the normalizing subsystem

$R_2 = (N^{++} + N^{-})/N$ – fraction of parallel spin interactions (double spin)

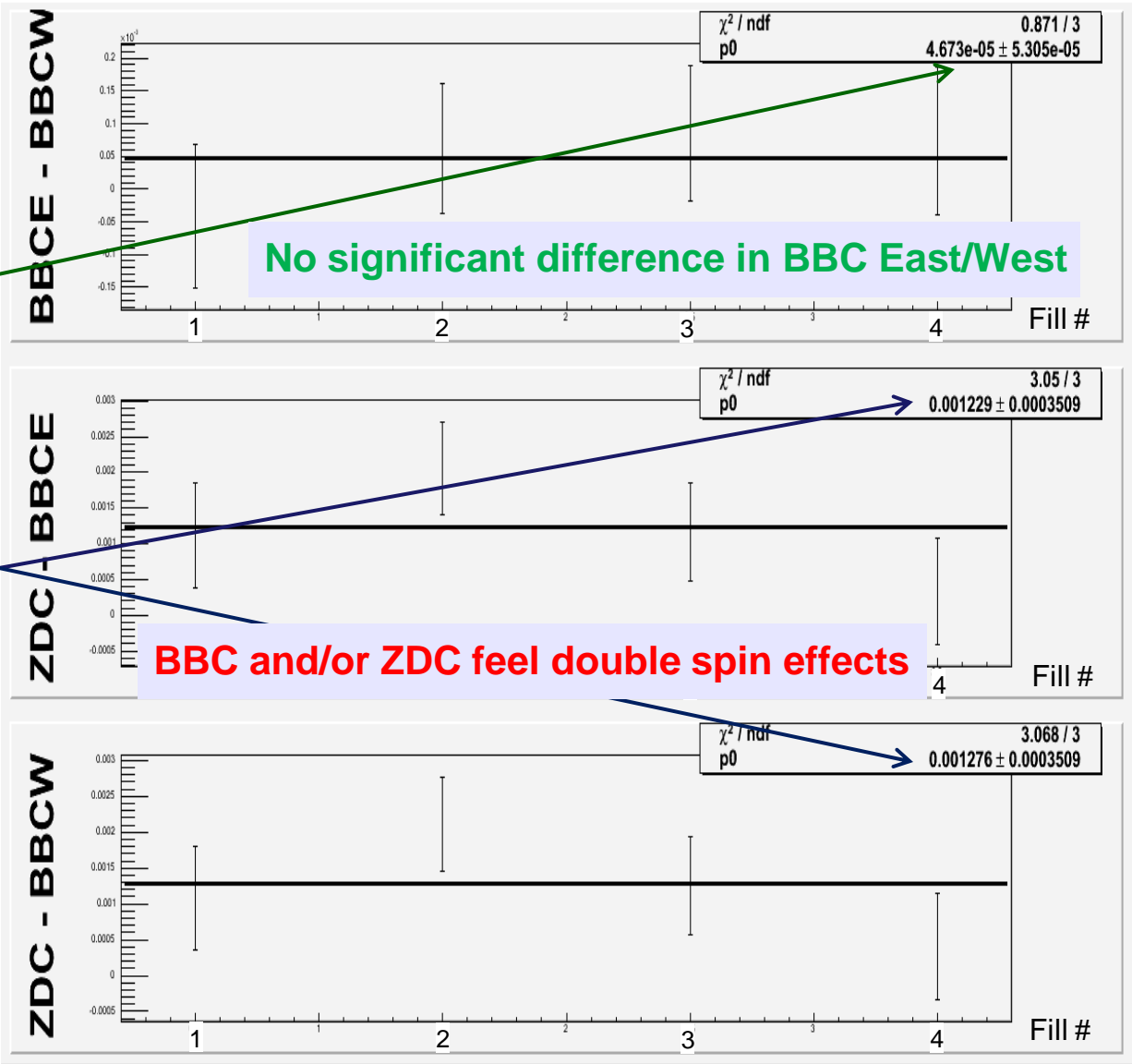
$R_B = (N^{++} + N^{+-})/N$ – fraction of interactions with spin UP in Blue beam

$R_Y = (N^{++} + N^{+})/N$ – fraction of interactions with spin UP in Yellow beam

R_2 is most important: $\delta \epsilon_{NN} \approx 2\delta R_2$, while R_B and R_Y contribute negligibly to the systematic uncertainty in ϵ_{NN}



- Of the four systems under study BBC and ZDC were chosen as most reliable
- BBC East is the same as BBC West at least to the level 10^{-5} .
- DS difference between BBC and ZDC $\sim 10^{-3}$ for 4 fills of our run with transverse polarization
- Checked for fake polarization pattern – DS averages out
- Also checked for 15 longitudinal runs: $2 \cdot 10^{-4}$.
- Systematic uncertainty $\sim 2 \cdot 10^{-3}$ in raw double spin asymmetry: too bad.



- Have access to single BBC tiles and coincidence of East and West BBC arms in any combinations.
- Choose several different physics processes, which give the same normalization.
- Double spin asymmetry for two different processes could be the same only if it is zero.
- Counts analyzed and corresponding physics:

Leading particle at small angle in East/West:

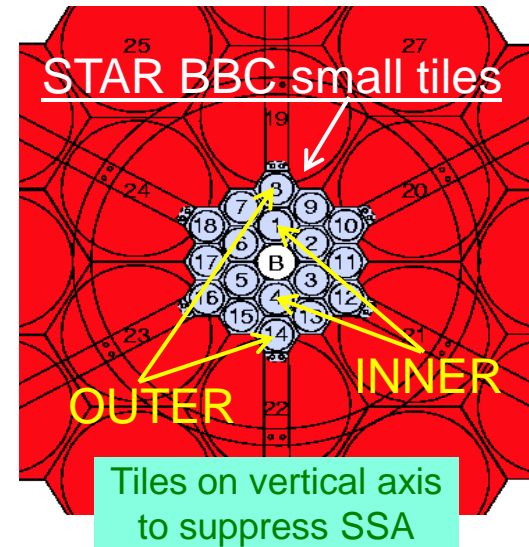
- **BBCIE** = (exactly 1 hit in INNER EAST tiles) & (at least 1 hit in WEST arm)
- **BBCIW** = (exactly 1 hit in INNER WEST tiles) & (at least 1 hit in EAST arm)

Leading particle at twice this angle in E/W:

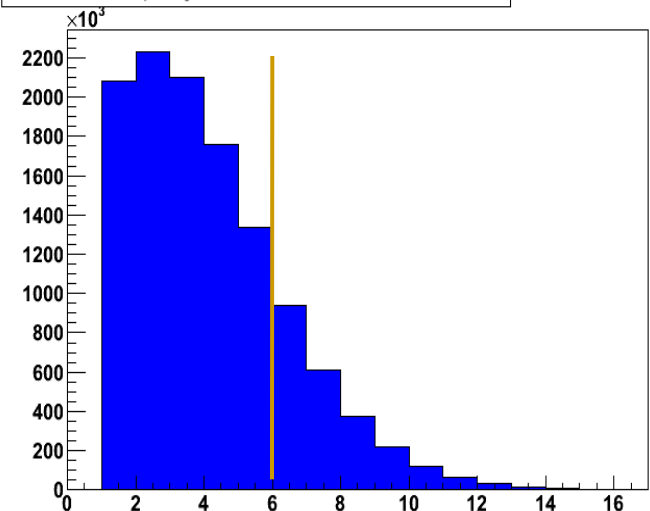
- **BBCOE/W** = (exactly 1 hit in OUTER E/W tiles) & (at least 1 hit in opposite arm)

Large forward multiplicity:

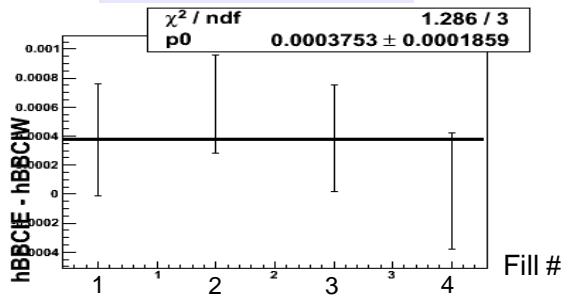
- **BBCXE/W** = (BBC E/W > 5 hits) & (at least 1 hit in opposite arm)



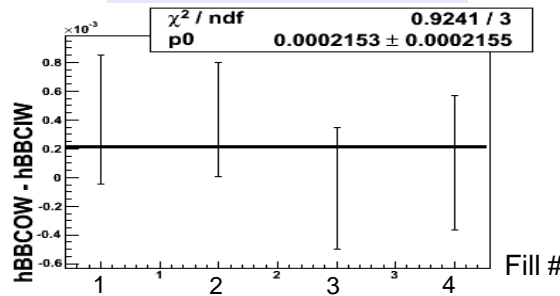
BBC east multiplicity with both BBC hit for run 10182005



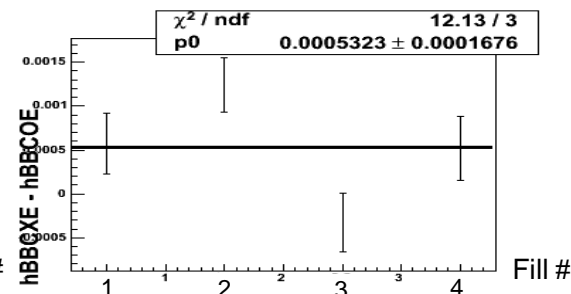
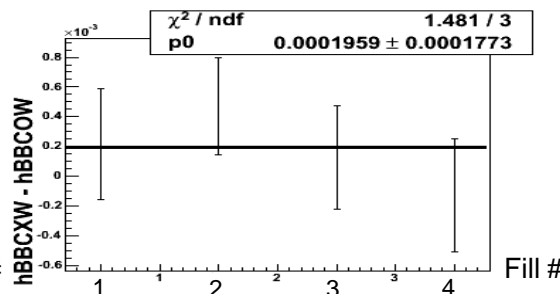
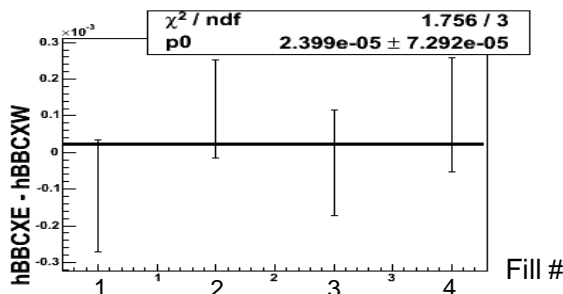
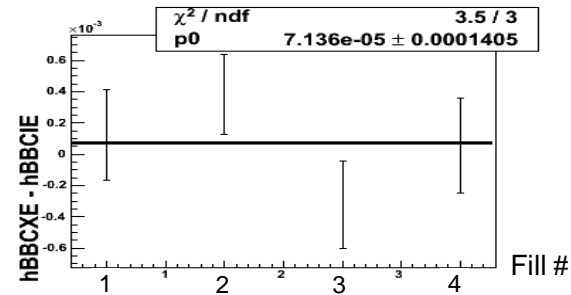
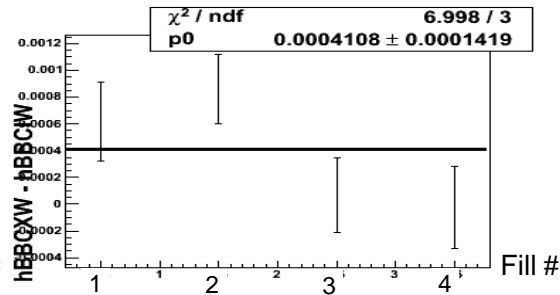
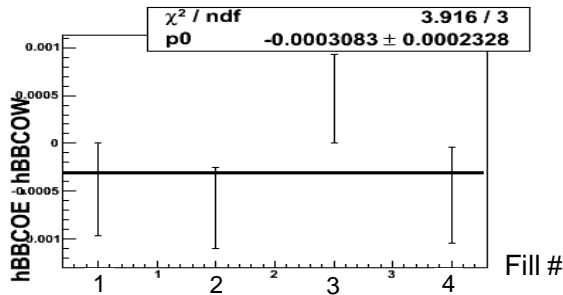
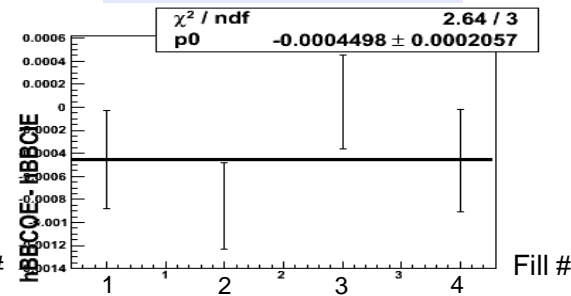
East / West



I / O / X West



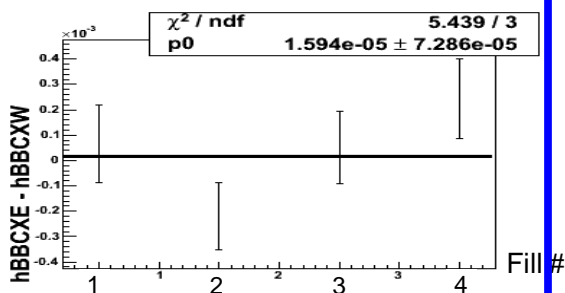
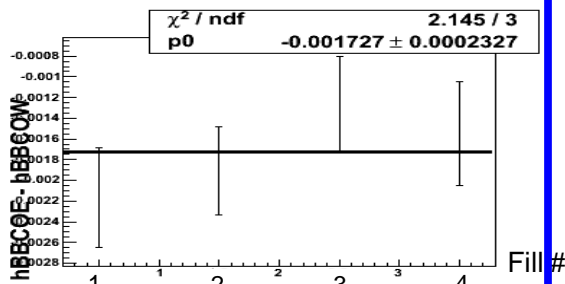
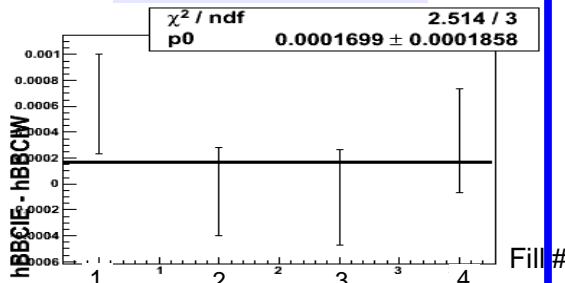
I / O / X East



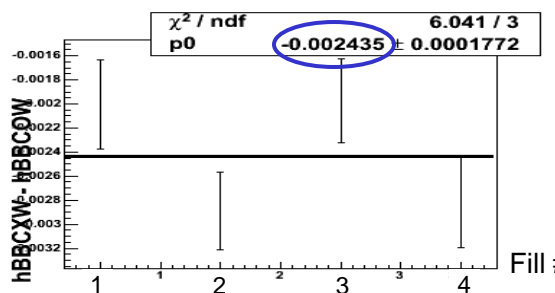
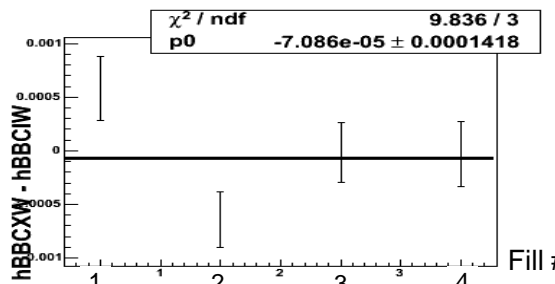
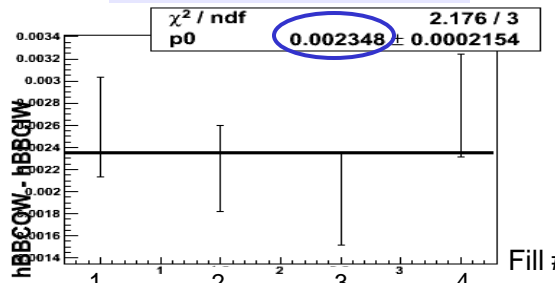
All shifts below $5 \cdot 10^{-4}$ \Rightarrow small or no DS effects in BBC
Choose BBC coincidence for normalization and use I/O/X for uncertainties



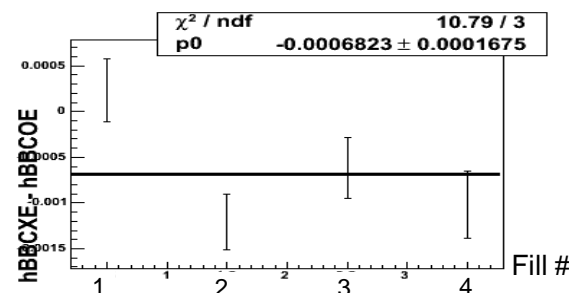
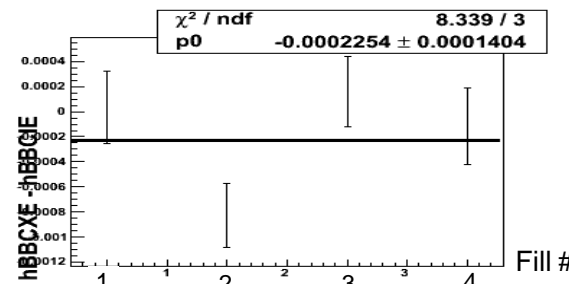
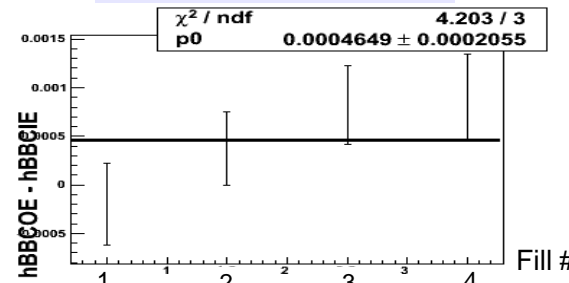
East / West



I / O / X West



I / O / X East



Maximum difference for West arm ($\sim 2 \cdot 10^{-3}$) – Blue beam fragmentation direction. Similar conclusion for Yellow single spin fraction ΔR_Y .

□ Very clean data set of ~20 million pp -elastic events is taken with Roman Pots integrated into the STAR detector for low t studies and both single and double spin asymmetries are being extracted with unprecedented accuracy in 5 t -ranges at $\sqrt{s}=200$ GeV.

□ Careful study of several STAR subsystems for extracting of normalization counts showed that BBC is nearly free of double spin effects and can be used for normalization.

□ Analysis of 3 processes in BBC with sensitivity to various physics allows to estimate the uncertainty in R_2 normalization ratio at the level of $2 \cdot 10^{-4}$.

□ Uncertainties of R_B and R_Y are 5 times larger but their contributions to the systematic error are strongly suppressed.

□ Need even more checks prior to release of the final result

□ Further plans include data taking at $\sqrt{s}=500$ GeV and measurements of A_{LL} .

