

## Probing the Early Medium in Heavy Ion Collisions with the Energy and System-Size Dependence of Long-Range Multiplicity Correlations

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## Abstract

Long-range forward-backward multiplicity correlations have been measured in heavy ion collisions at RHIC. New results for short and long-range multiplicity correlations (forward-backward) are presented for several systems (Au+Au, Cu+Cu, and pp) at energies of √sNN = 62.4 and ~ 200 GeV. A strong, long-range correlation has been seen previously in central heavy ion collisions at  $\sqrt{\text{sNN}= 200 \text{ GeV}}$ , which vanishes in semi-peripheral events and pp collisions. The long-range component of the correlation has been attributed to a multiple partonic interactions, while the origin of the short-range correlation is due to independent sources. This scenario tends to support the view that there is a partonic core formed in heavy ion collisions surrounded by a hadronic corona.



- - N. Armesto, L. McLerran and C. Pajares , Nucl. Phys . A781, 201(2007)

## **SUMMARY**

Long range correlations have been observed in Au+Au and Cu+Cu system at collision energies of 62 and 200 GeV for most central events.

The LRC is more energy dependent than the system size. The correlation is same for like sign and unlike sign charges in central collisions indicating the possibility of cluster formation. Based on CGC one can argue that the long range rapidity correlations are due to the fluctuations of the number of gluons and can be created early time only.



In DPM long-range correlations are due to superposition of fluctuating number of strings. SRC  $D_{bf}^{2} = \langle N_{f} N_{b} \rangle - \langle N_{f} \rangle \langle N_{b} \rangle =$  $< n > (< N_{0f} N_{0b} > - < N_{0f} > < N_{0b} > + LRC$ 



In CGC LRC origin lies in the longitudinal color field created in Glasma similar to strings in DPM. Predicts increase of correlation with centrality.



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For gluons the LRC increases with energy and decreases with  $|y_1-y_2|$ . Baryons -> Smaller correlations.

The STAR Collaboration: http://drupal.star.bnl.gov/STAR/presentations