



Temperature Fluctuation and the Specific Heat in Au+Au Collisions at $\sqrt{s_{NN}} = 7.7 - 200$ GeV from AMPT model and STAR

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Abstract

In this poster, we will present the energy dependence of event-by-event temperature fluctuations and specific heat of the QCD matter created in Au+Au collisions at $\sqrt{s_{NN}} = 5, 7, 17.3, 39, 62.4,$ and 200 GeV from the AMPT model calculations and at $\sqrt{s_{NN}} = 7.7, 19.6,$ and 27 GeV from BES-I program of the STAR experiment.

Analysis method

In a system, for a range of p_T within a to b one can estimate $\langle p_T \rangle$ via [1]:

$$\langle p_T \rangle = 2T_{\text{eff}} + \frac{a^2 e^{-\frac{a}{T_{\text{eff}}}} - b^2 e^{-\frac{b}{T_{\text{eff}}}}}{(a+T_{\text{eff}})e^{-\frac{a}{T_{\text{eff}}}} - (b+T_{\text{eff}})e^{-\frac{b}{T_{\text{eff}}}}}$$

This equation links the value of $\langle p_T \rangle$ within a specified range (from a to b) of p_T to the effective temperature, T_{eff} .

From the temperature distribution [2] $f(T) \sim \exp\left[-\frac{C(T-\langle T \rangle)^2}{\langle T \rangle^2}\right]$, heat capacity [3]:

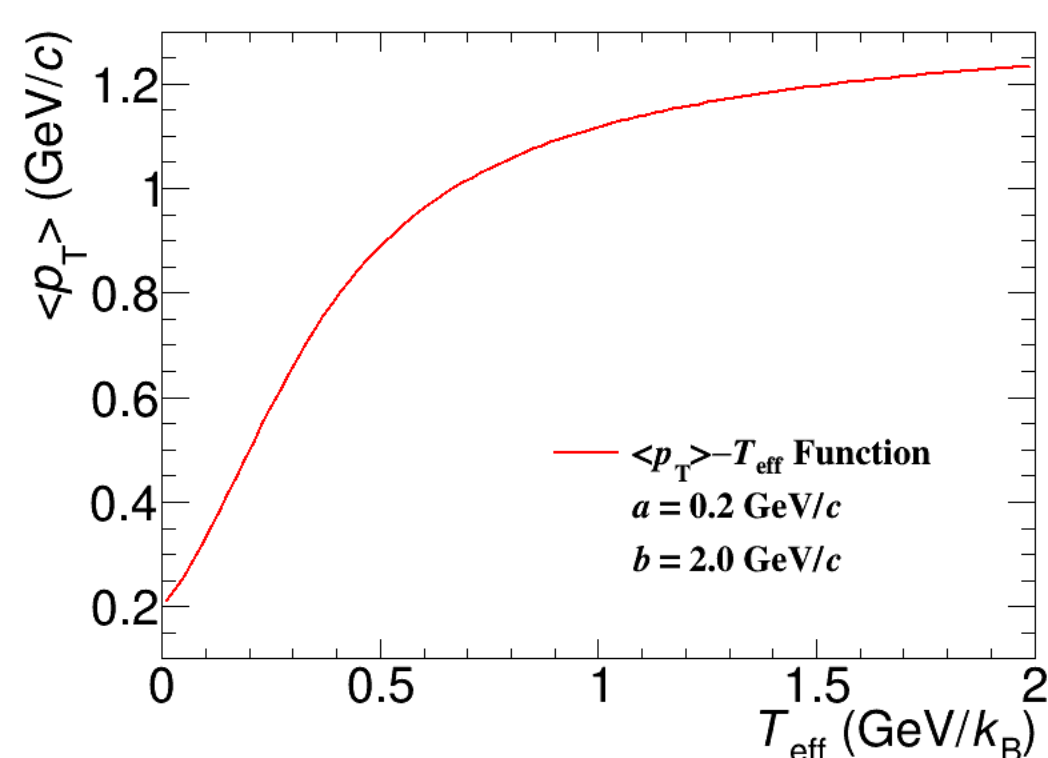
$$C = \frac{\langle T_{\text{kin}} \rangle^2}{(\Delta T_{\text{eff}}^{\text{dyn}})^2} = \frac{\langle T_{\text{kin}} \rangle^2}{(\Delta T_{\text{eff}}^{\text{stat}})^2}$$

mixed events are used to subtract statistical fluctuations.

The specific heat C_V is obtained as:

$$C_V = \frac{C}{N}$$

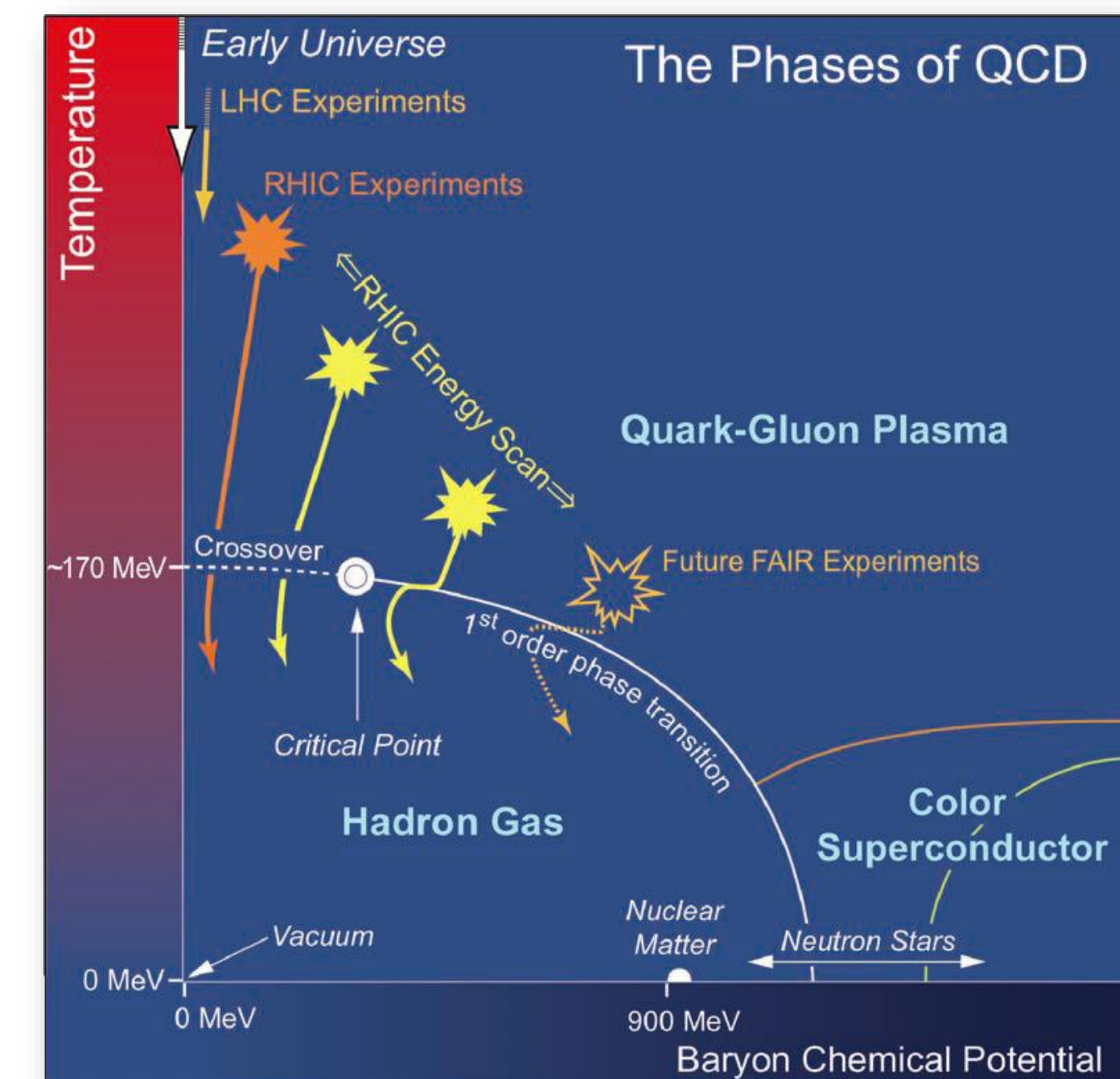
where N is the number of particles.



Motivation

The transition to QGP changes from a crossover to the 1st order phase transition results in the existence of a critical point (circle) in the QCD phase diagram. Lattice QCD estimates indicated that the critical point exists in the range $250 < \mu_B < 450$ MeV [8].

The specific heat, C_V , is a thermodynamic quantity that characterizes the equation of state of the system. For a system undergoing phase transition, C_V is expected to diverge at the critical point, which can be extracted from event-by-event temperature fluctuation. Thus the variation of thermal fluctuations with temperature can be effectively used to probe the QCD phase transition and QCD critical point.



Data Set

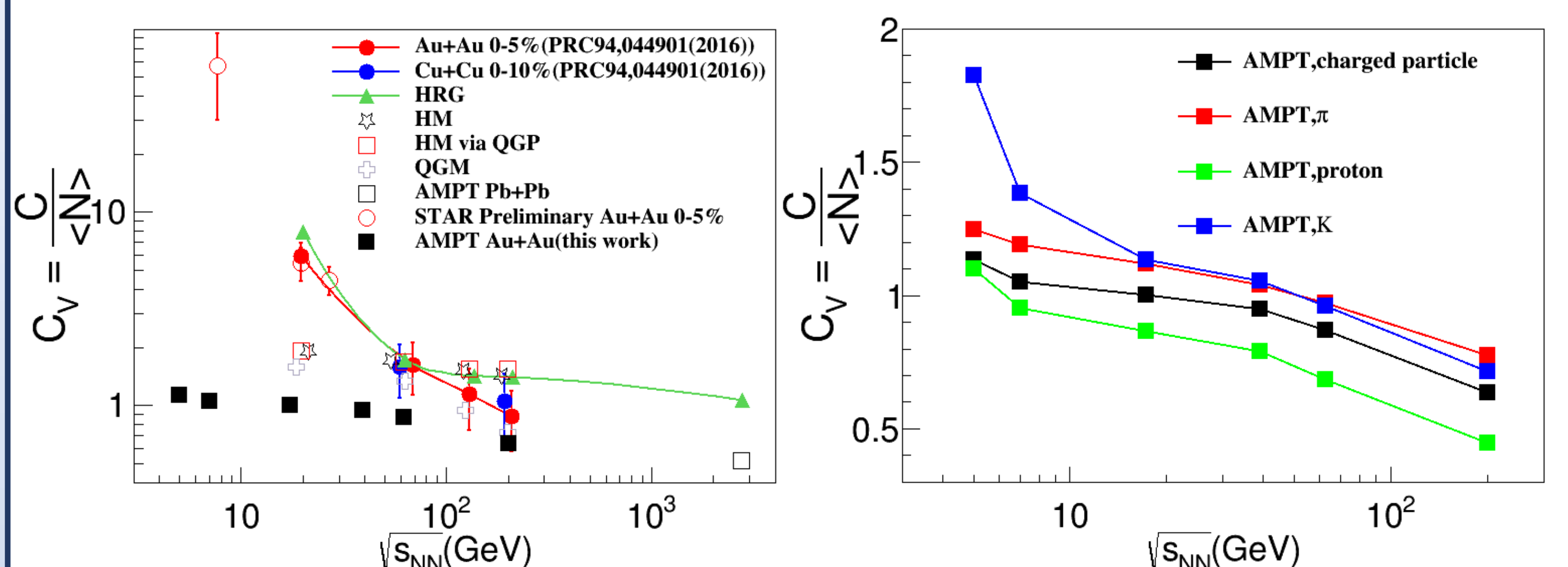
• AMPT model:

$\sqrt{s_{NN}}$ (GeV)	5	7	17.3	39	62.4	200
Centrality Selection	impact parameter $0 < B < 3$ fm (0-5%)					
Track Selection	$ \eta < 1$ $a = 0.15$ GeV/c $< p_T < 2$ GeV/c $= b$					

• STAR data:

$\sqrt{s_{NN}}$ (GeV)	7.7	19.6	27
Year	2010	2011	2011
Events (10^6)	3	15	32
Centrality Selection	0-5%		
Track Selection	$ \eta < 1$, $a = 0.15$ GeV/c $< p_T < 2$ GeV/c $= b$		

Results

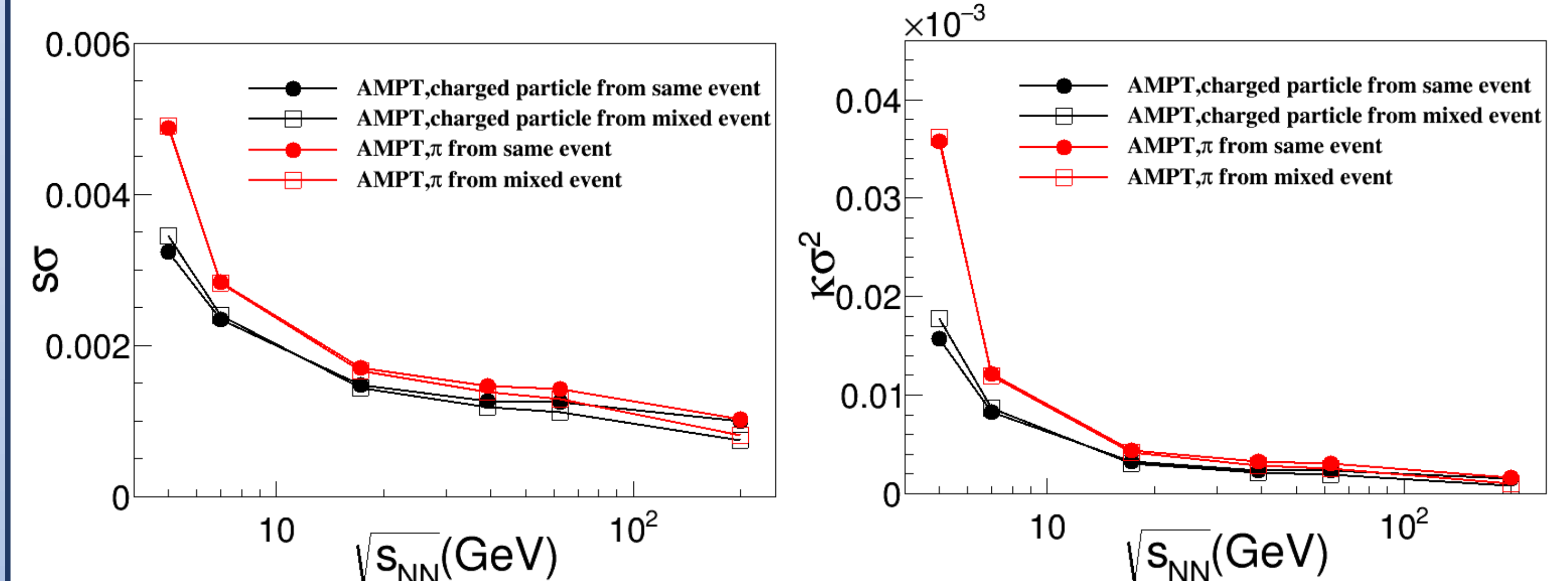
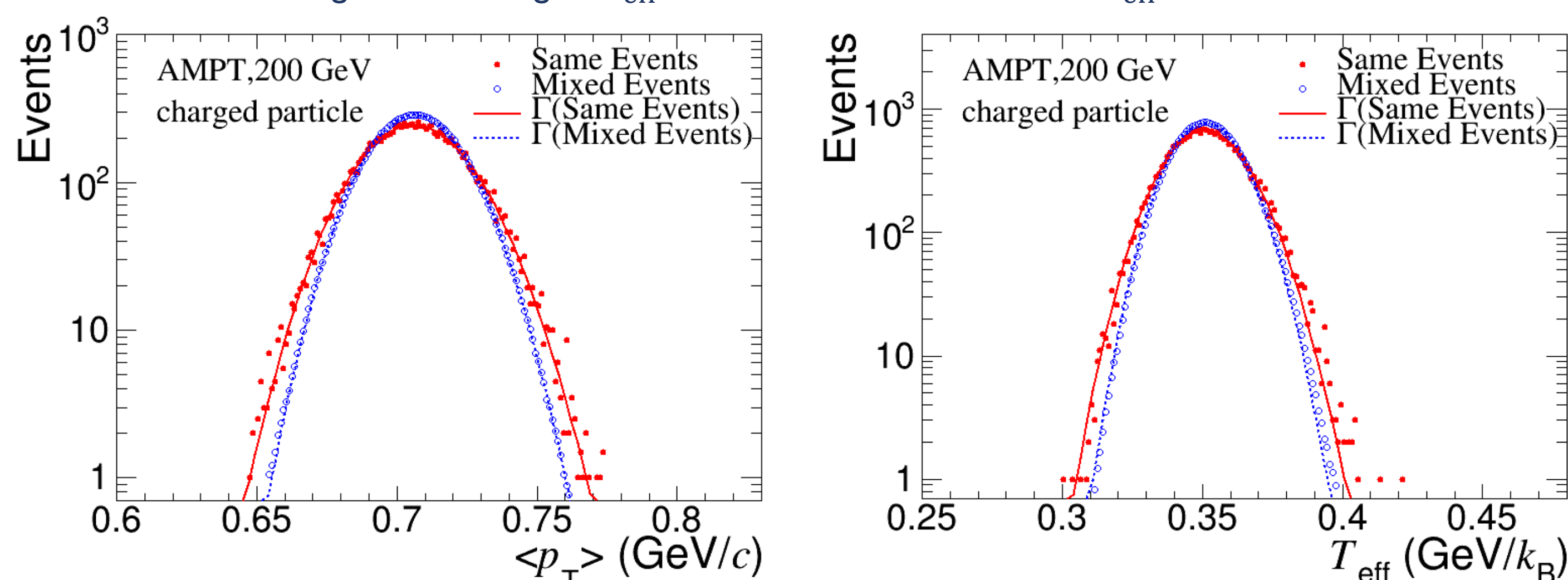


• The specific heat, C_V , as a function of collision energy at $0.2 \text{ GeV} < p_T < 2 \text{ GeV}$ in 0-5% centrality interval. Left panel: compares the charged particle C_V with published results. Right panel: shows the C_V from AMPT model.

$\langle p_T \rangle$ and T_{eff} distributions

Curves represent fitted gamma distributions [4]: $Ga(x) = \frac{x^{\alpha-1} e^{-x/\beta}}{\Gamma(\alpha)\beta^\alpha}$, $x > 0$ ($\alpha, \beta > 0$).

The mean (μ) and standard deviation (σ) of the distribution are related to the fit parameters (α and β) by $\mu = \alpha\beta$ and $\sigma^2 = \alpha\beta^2$. The T_{eff} distributions can be nicely described by the gamma distribution. The σ of gamma fitting of T_{eff} distributions are used as ΔT_{eff} .



• The 3rd and 4th order fluctuations of temperature as a function of from AMPT model.

Temperature high order fluctuations measures

High order cumulants describe the shape of distribution and quantify fluctuations. They are sensitive to the correlation length ξ . We use them to characterize the fluctuations of temperature. The cumulants of event-by-event temperature distributions can be expressed in terms of moments.

For a distribution of N (here N is T_{eff} , $\delta N = N - \langle N \rangle$), the cumulants can be expressed [5,7]:

$$\begin{aligned} C_1 &= \mu = \langle N \rangle, \\ C_2 &= \sigma^2 = \langle (\delta N)^2 \rangle, \\ C_3 &= S\sigma^3 = \langle (\delta N)^3 \rangle, \\ C_4 &= \kappa\sigma^4 = \langle (\delta N)^4 \rangle - 3\langle (\delta N)^2 \rangle^2, \end{aligned}$$

Directly connected to the susceptibility of the system: $S\sigma = \frac{C_3}{C_2}$, $\kappa\sigma^2 = \frac{C_4}{C_2}$.

For a gamma distribution $S\sigma = 2\beta$ and $\kappa\sigma^2 = 6\beta^2$.

Summary

Both C_V and higher order cumulants of the temperature fluctuations show monotonic distributions, which is expected that there is no phase transition critical point in the AMPT model. This provides a good reference for comparison with experimental data to search for the signal of critical point.

At low energies, a sharp drop of C_V from the STAR data is observed and it deviates from the AMPT results. Results will be further improved with more coming data at different energies.

References

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