

Search for the chiral magnetic effect in U+U & Isobar collisions

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(for the STAR Collaboration)

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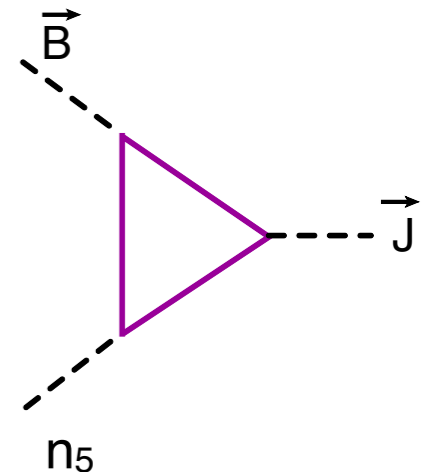
CME workshop, RHIC & AGS Annual Users' Meeting, June 7-10, 2016

Brookhaven National Laboratory, Upton, NY, USA



Outline

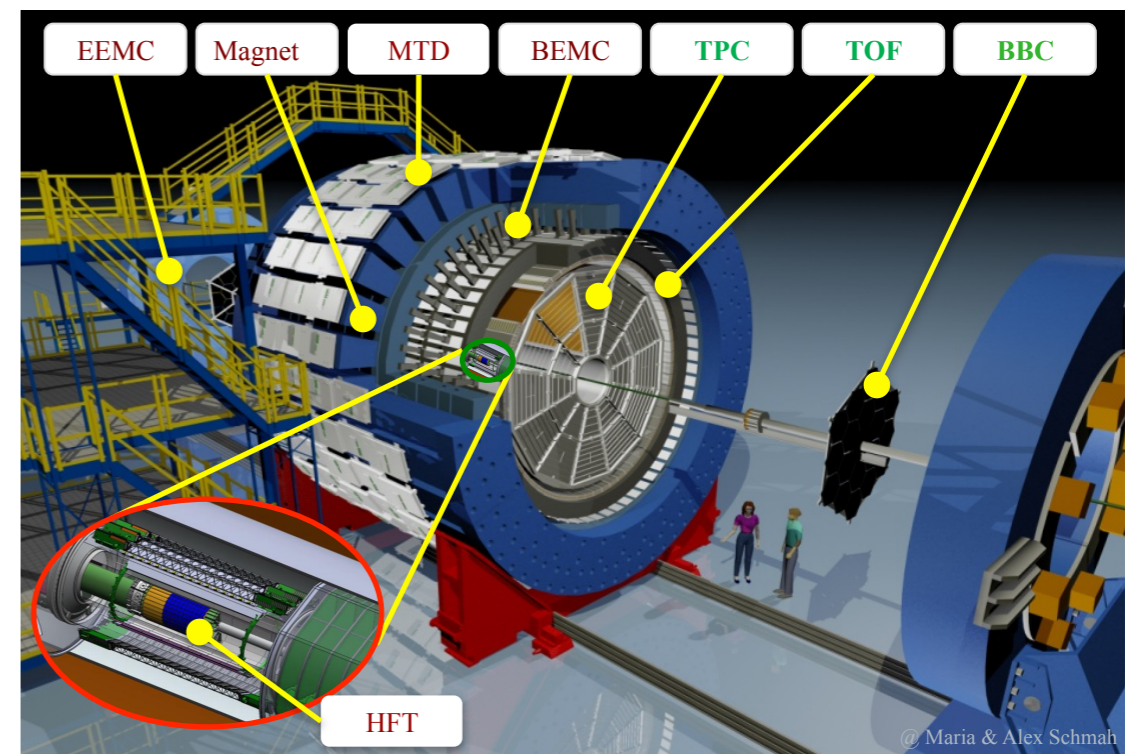
- Results for U+U & Au+Au collisions from STAR
- What else can we try with the existing data
- Outlook for collisions of Isobars @ RHIC



Flow is the dominant source of background for signals of CME

How can U+U collisions be used to disentangle the two effects ?

What else can we try : Isobars



What have we learned from the U+U collisions at RHIC ?

- Limitations of two-component model in MC-Glauber :

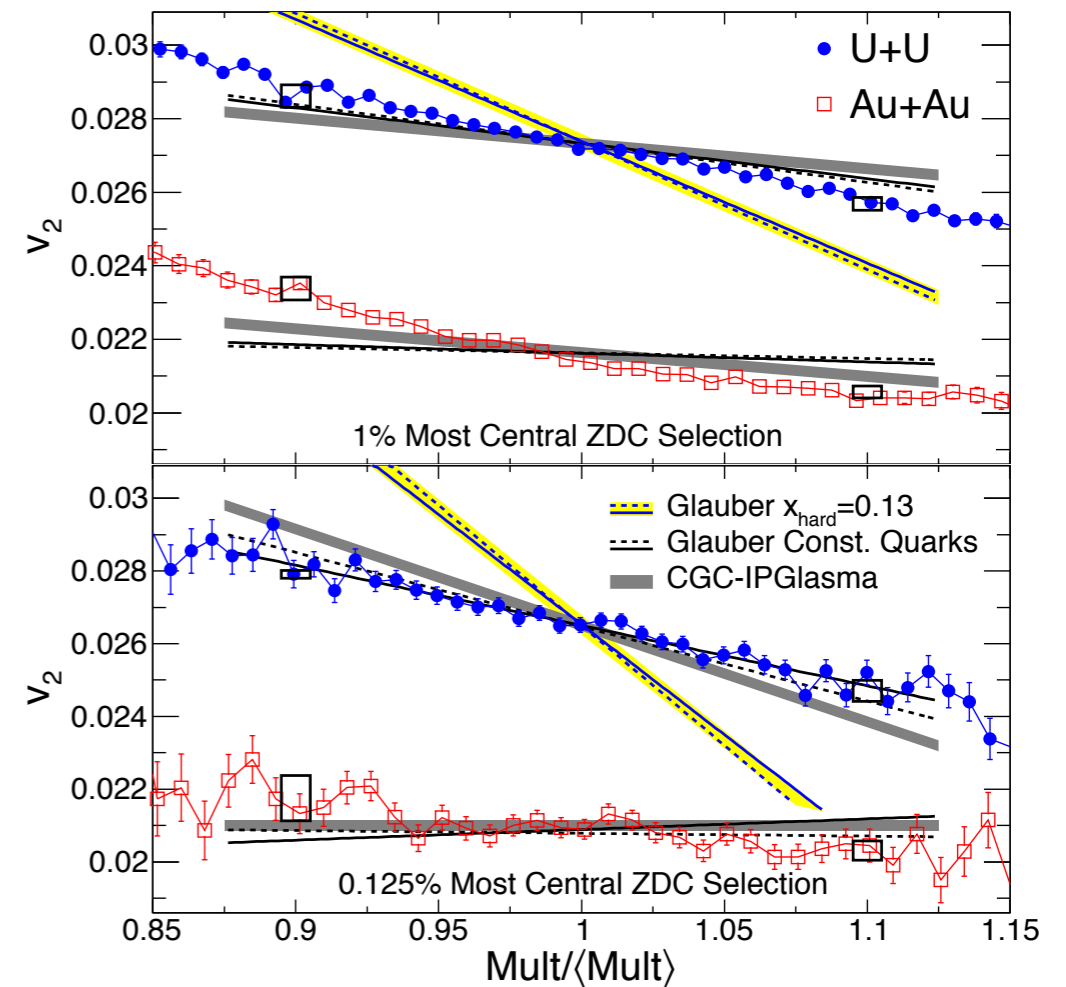
Modifications : Quark-Glauber (nucl-th/0302071, 1509.06727), TRENTO (1412.4708), Shadowed Glauber (1510.01311)

- Evidence of color coherence & CGC like initial state :

CGC \rightarrow Weak dependence of multiplicity on shape (Schenke, PT, Venugopalan 1403.2232)

- Dominance of fluctuations, small control in triggering shape : 35% variation in $dN/d\eta$ \rightarrow 12% variation in v_2 in ($<1\%$ ZDC)

L. Adamczyk et al. (STAR Collaboration)
Phys. Rev. Lett. 115, 222301 (2015)

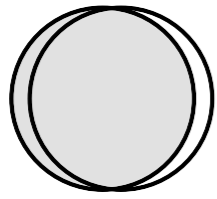


U+U data contradicts strong binary-collision dependence of multiplicity

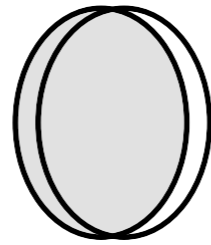
Next Step: Can we use U+U collisions to learn about CME ?

Qualitative picture

Correlation between B-field & eccentricity

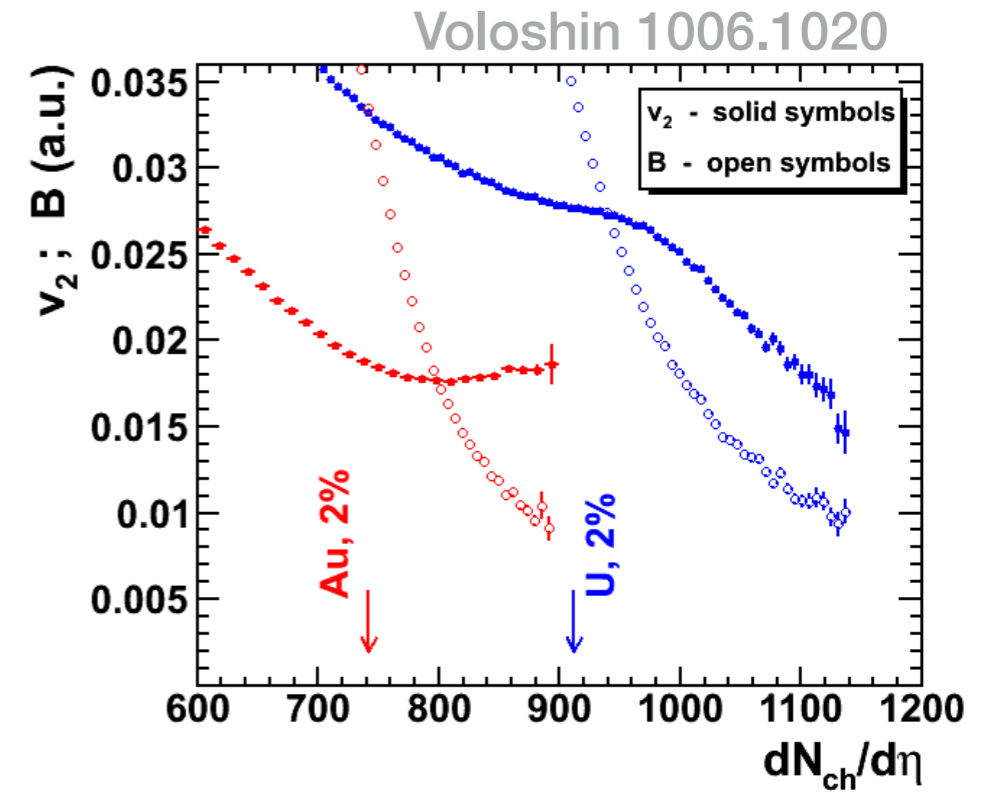
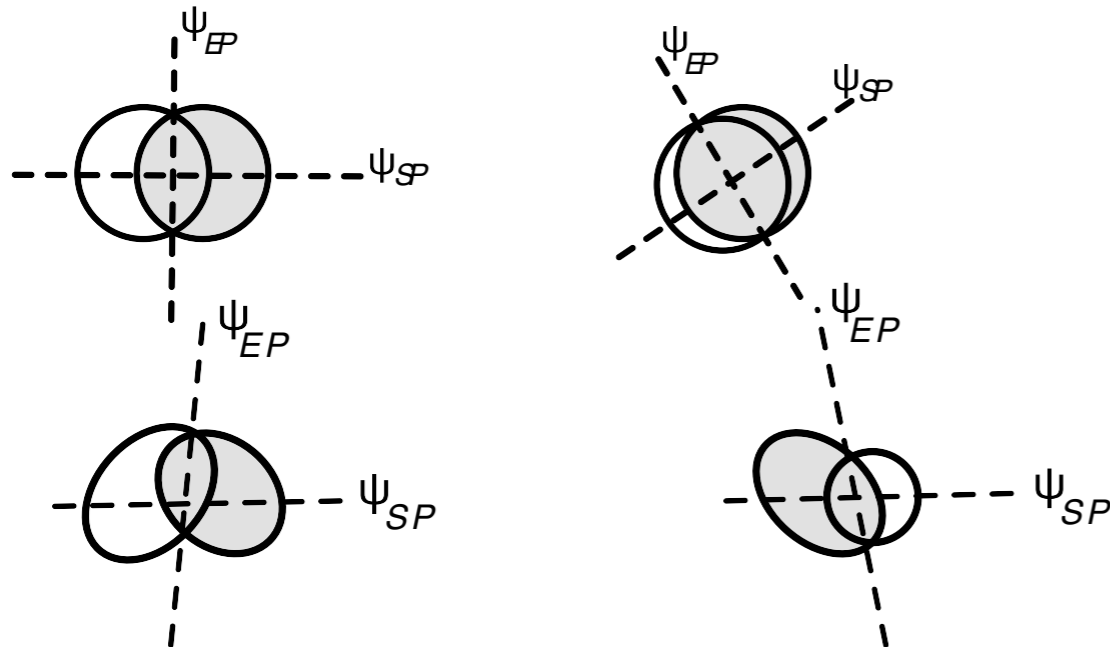


Au+Au (ultracentral)
 $\epsilon \sim 0, B \sim 0$



U+U (ultracentral)
 $\epsilon \neq 0, B \sim 0$

Search for non-zero v_2 & zero CME



Reaction plane & B-field direction is strongly correlated in Au+Au \rightarrow Not true for U+U

Can U+U collisions disentangle flow & signals of CME ?

Observables for CME

- General (3-particle) correlator :

$$C_{m,n,m+n} = \langle \cos((m\phi_1 + n\phi_2 - (m+n)\phi_3)) \rangle$$

- Lowest order (3-particle) **charge sensitive correlator** :

$$C_{112} = \langle \cos((\phi_1^\pm + \phi_2^\mp - 2\phi_3)) \rangle$$

- The CME correlator :

$$\gamma^{a,b} \sim \frac{\langle \cos(\phi_1^a + \phi_2^b - 2\phi_3) \rangle}{v_2\{2\}} \sim \langle \cos(\phi^a + \phi^b - 2\Psi_{RP}) \rangle$$

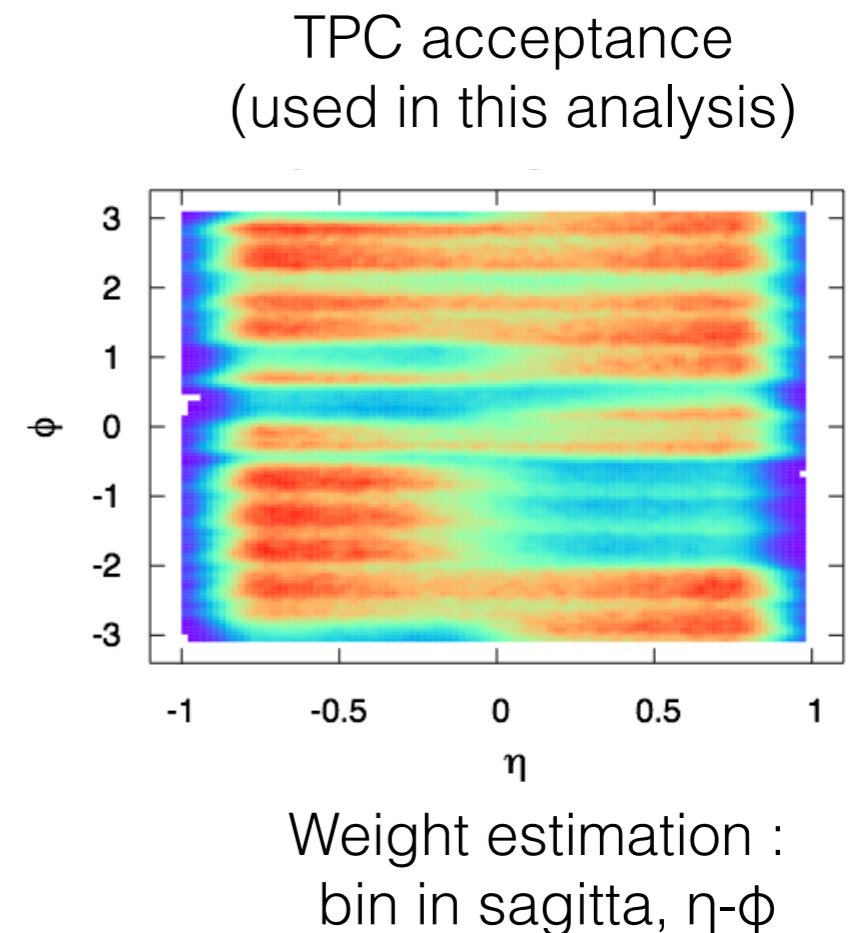
(3P-cumulant method)

(event-plane method)

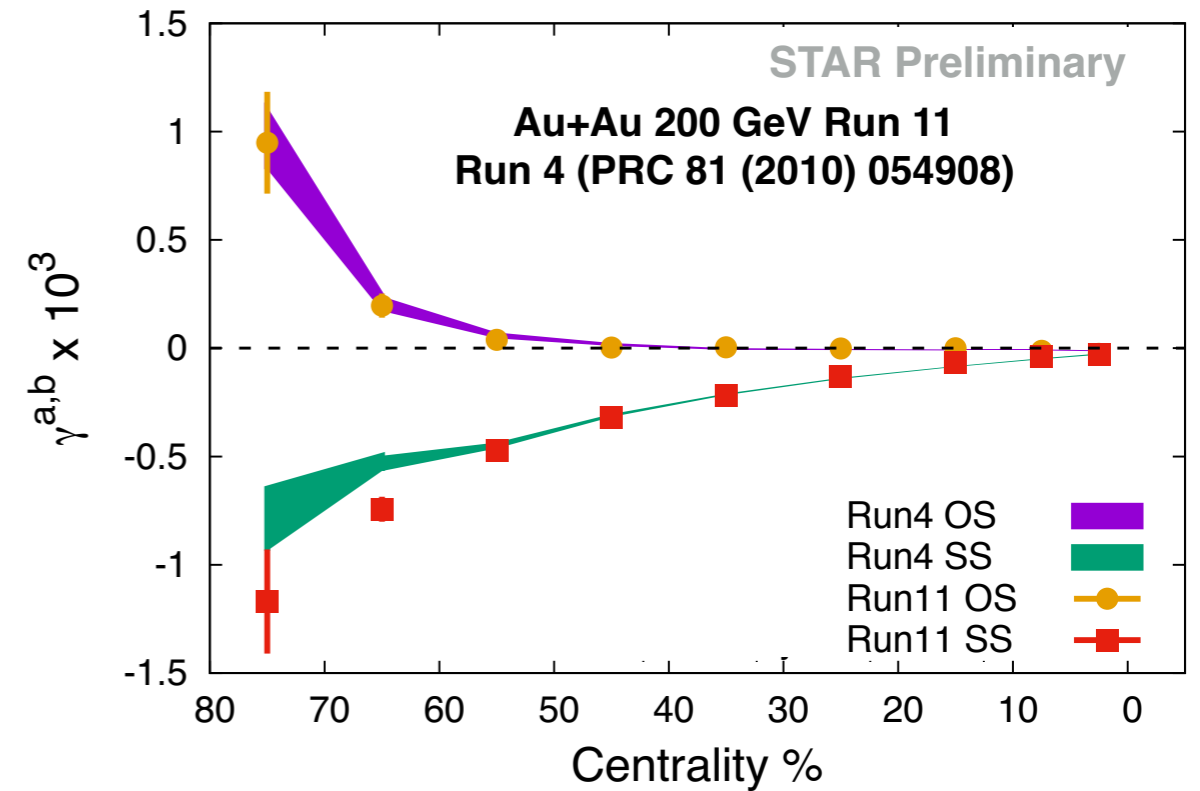
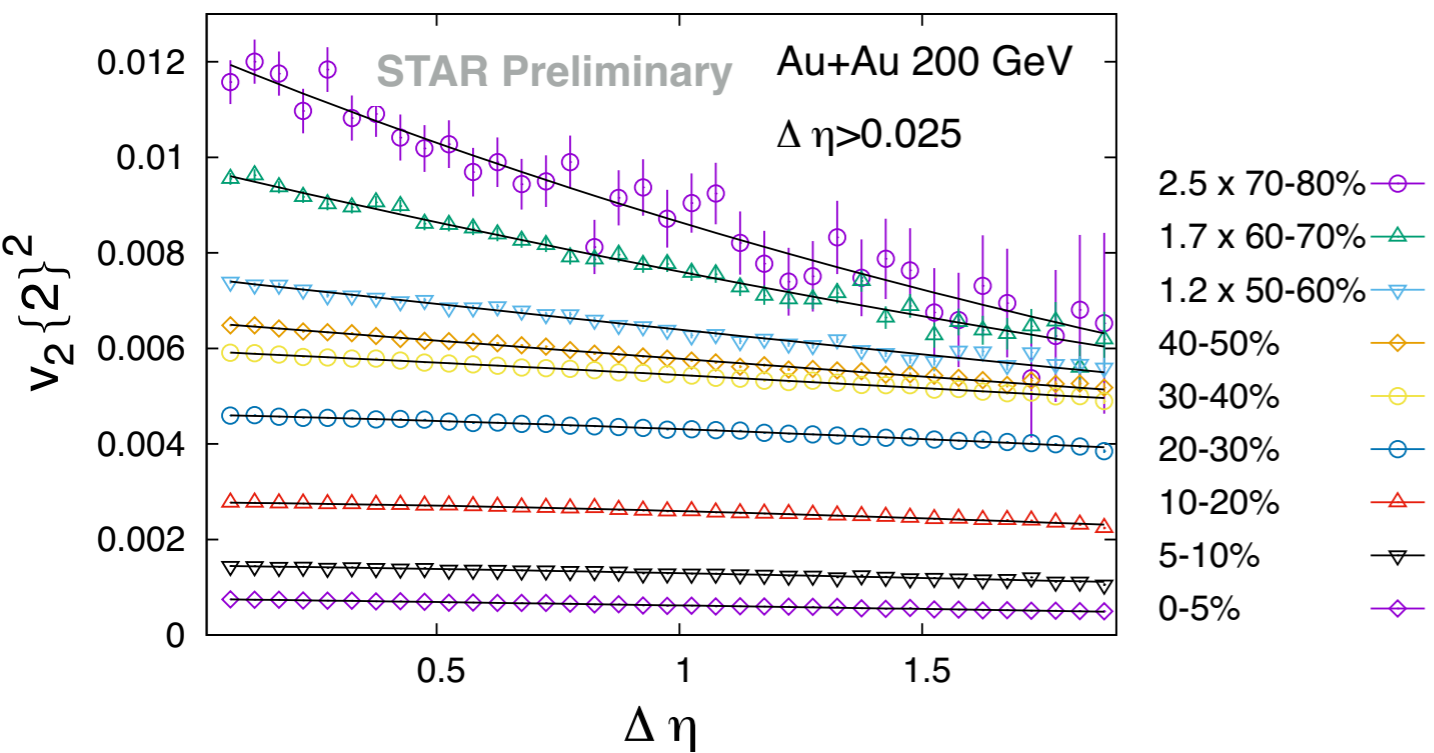
$$v_2\{2\}^2 = \langle \cos(2(\phi_1 - \phi_2)) \rangle$$

Details of the data set

- U+U 193 GeV : Year 2012 (Min-bias/ultra-central)
- Au+Au 200 GeV : Year 2004, 2007 (Min-bias), 2011 (ultra-central)
- Centrality selection :
 - TPC uncorrected multiplicity $|\eta| < 0.5$
 - ZDC East & West ADC
- Common QA cuts :
 - $|V_r| < 2$, $|V_z| < 20$, $|V_z - v_{pd}V_z| < 2$ cm
- Acceptance cuts: $|\eta| < 1$, $0.2 \text{ GeV}/c < |p_T|$



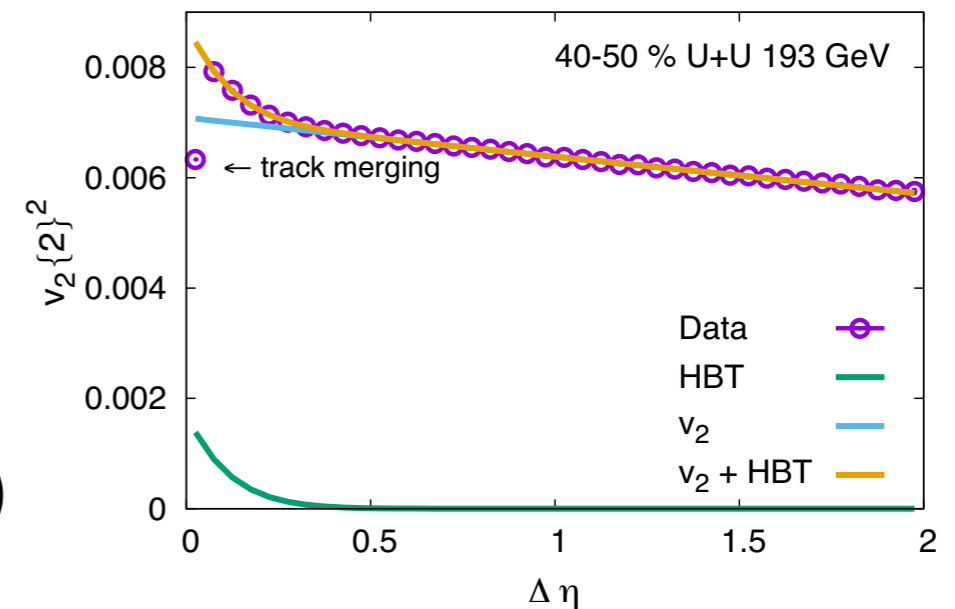
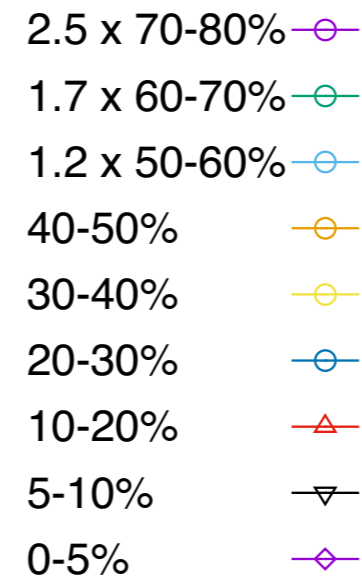
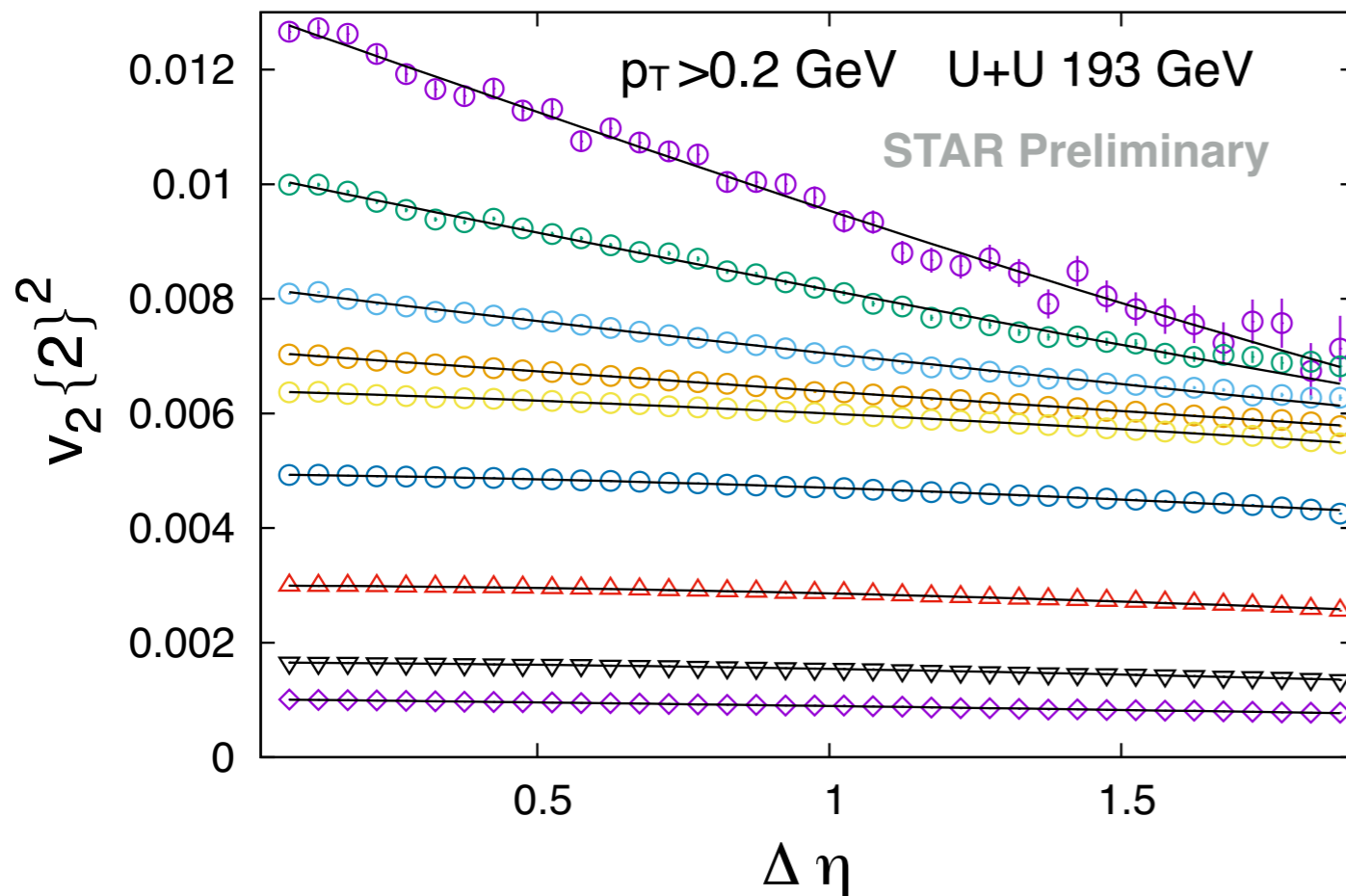
Measurement of $v_2\{2\}$ & γ^{ab} in Au+Au collisions



Au+Au results \longrightarrow baseline for measurements in U+U

Measurement of $v_2\{2\}$ in U+U collisions

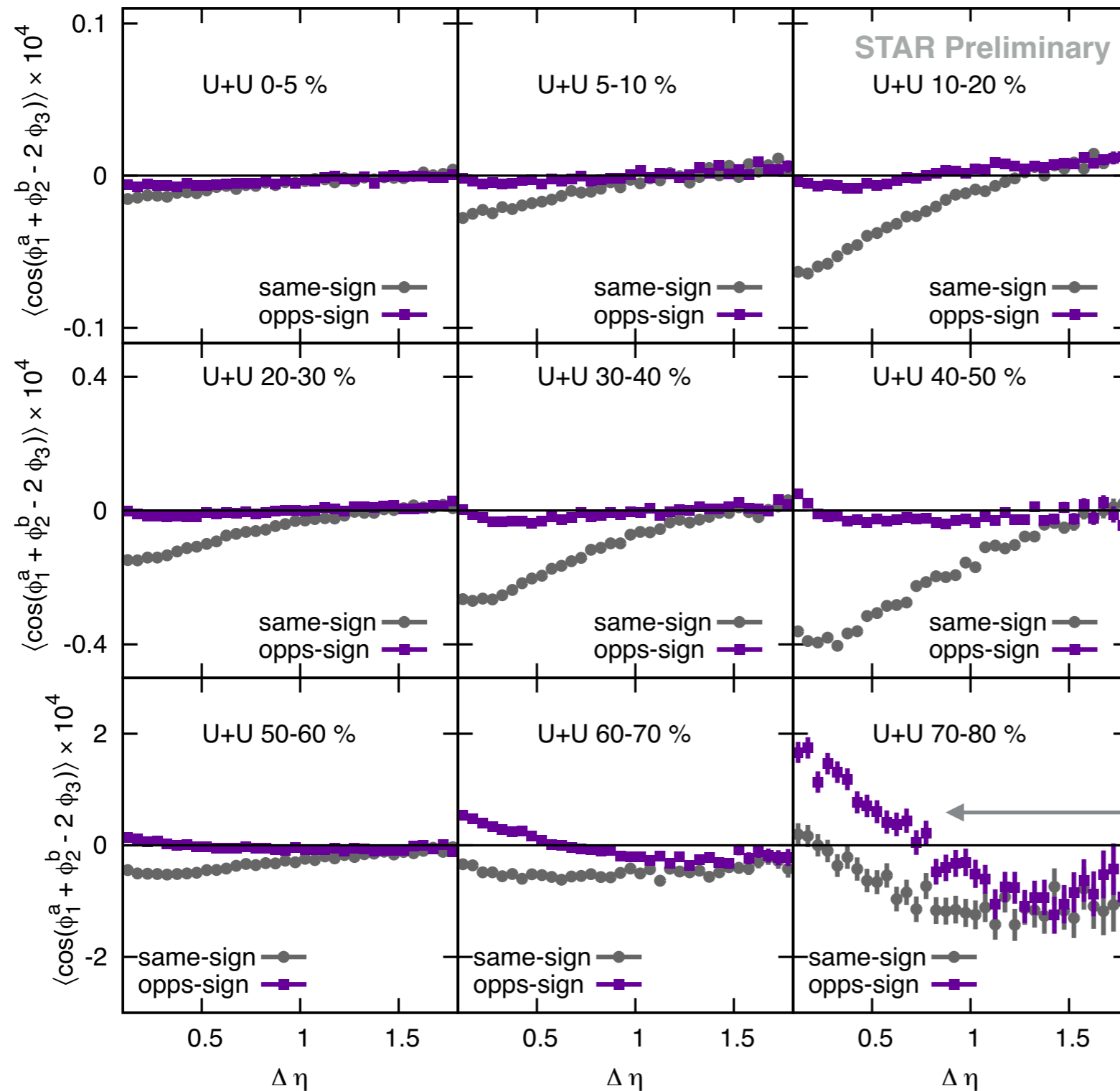
$$\gamma^{a,b} = \frac{\langle \cos(2(\phi_1^a + \phi_2^b - 2\phi_3)) \rangle}{v_2\{2\}}$$



Removing two major artifacts:

- Track merging (apply $\Delta\eta > 0.025$)
- Short range -correlations (Gaussian fit)

Differential measurement of the C_{112} correlator



$$C_{112} = \langle \cos(2(\phi_1^a + \phi_2^b - 2\phi_3)) \rangle$$

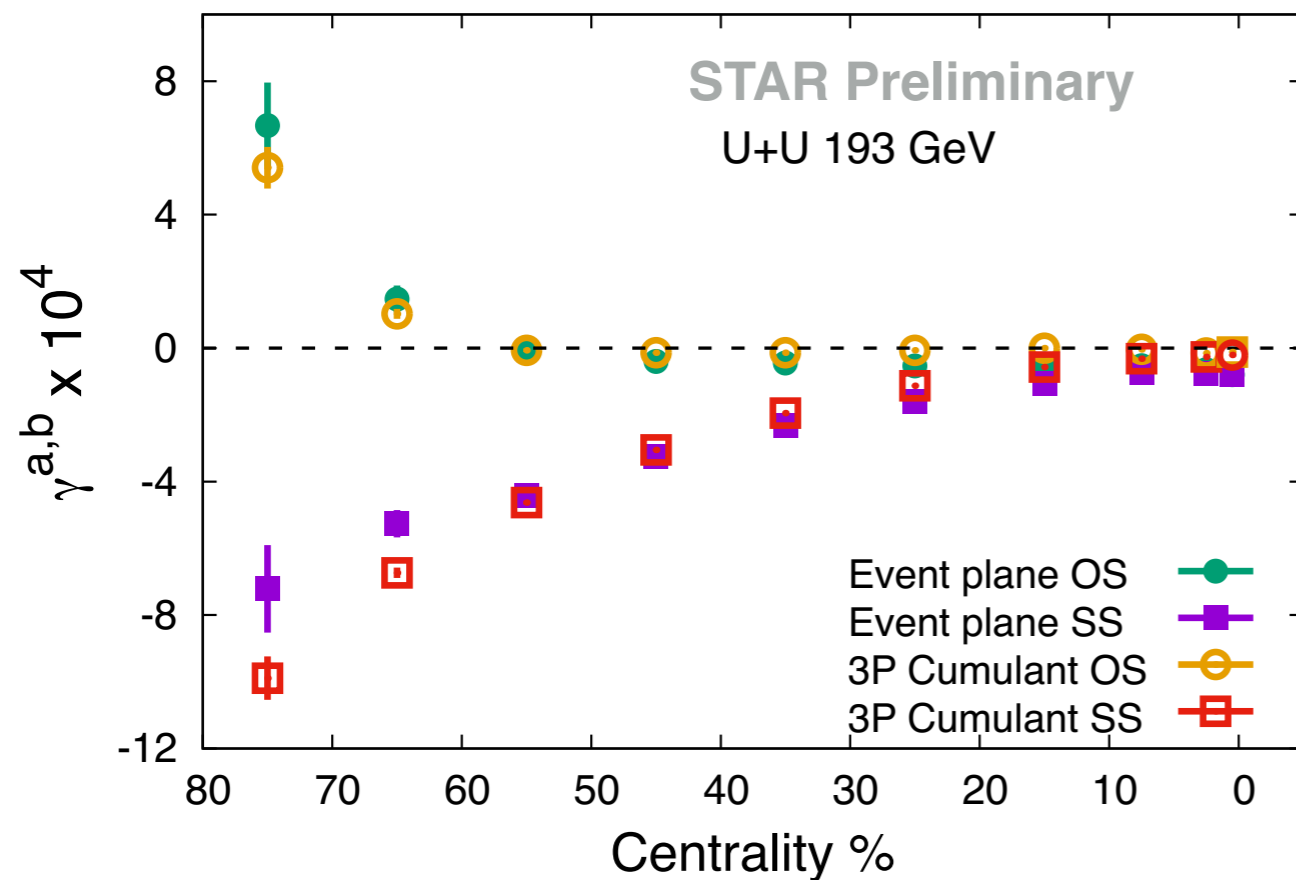
$$\Delta\eta = \Delta\eta_{1,2}$$

Need to remove two major artifacts :

- Track merging
 - apply $\Delta\eta > 0.025$
- Short-range - correlations
 - do : (OS - SS)

Results using Cumulant and Event-plane methods

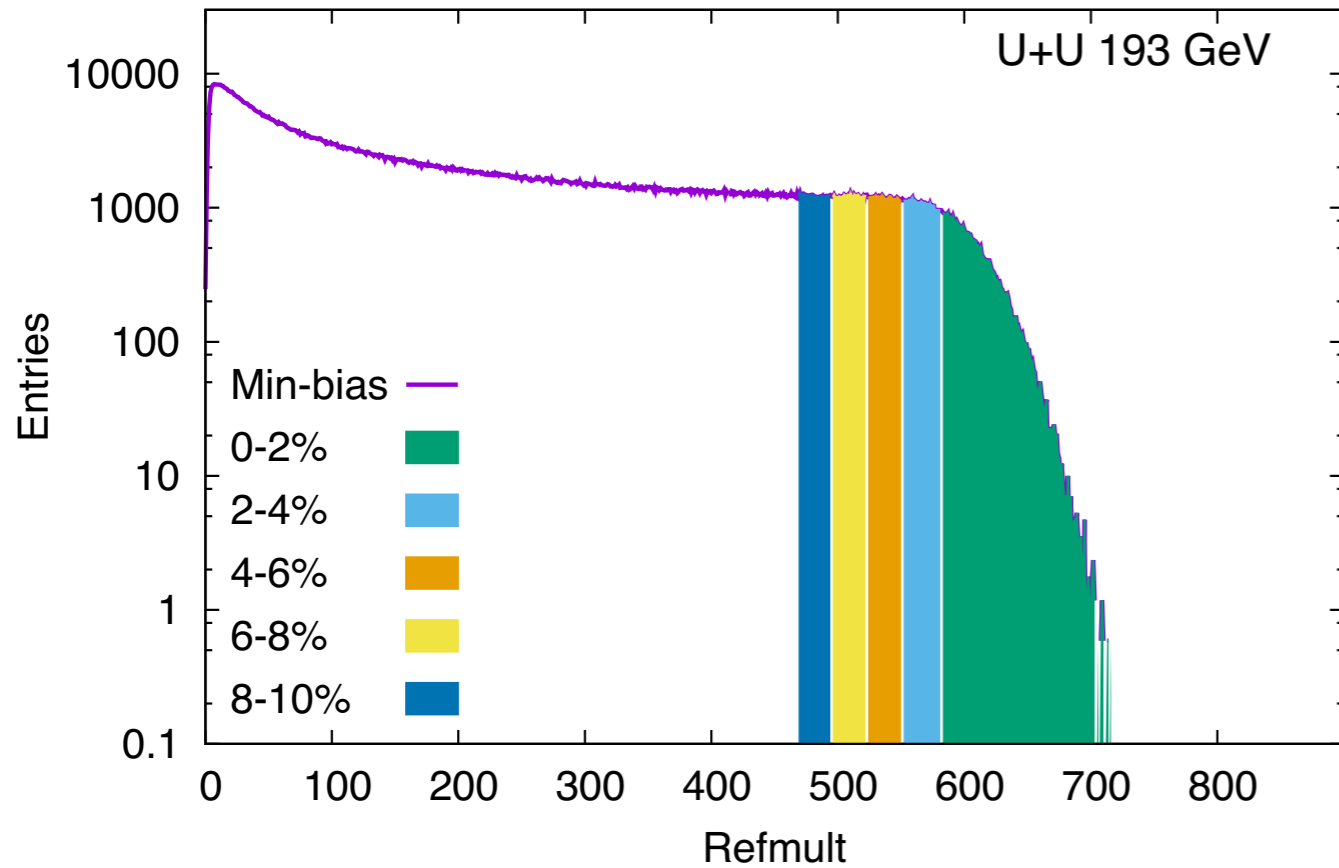
$$\gamma^{a,b} \sim \frac{\langle \cos(\phi_1^a + \phi_2^b - 2\phi_3) \rangle}{v_2\{2\}} \sim \langle \cos(\phi^a + \phi^b - 2\Psi_{RP}) \rangle$$



Centrality bins finer than 0-10% is needed to probe the shape of Uranium

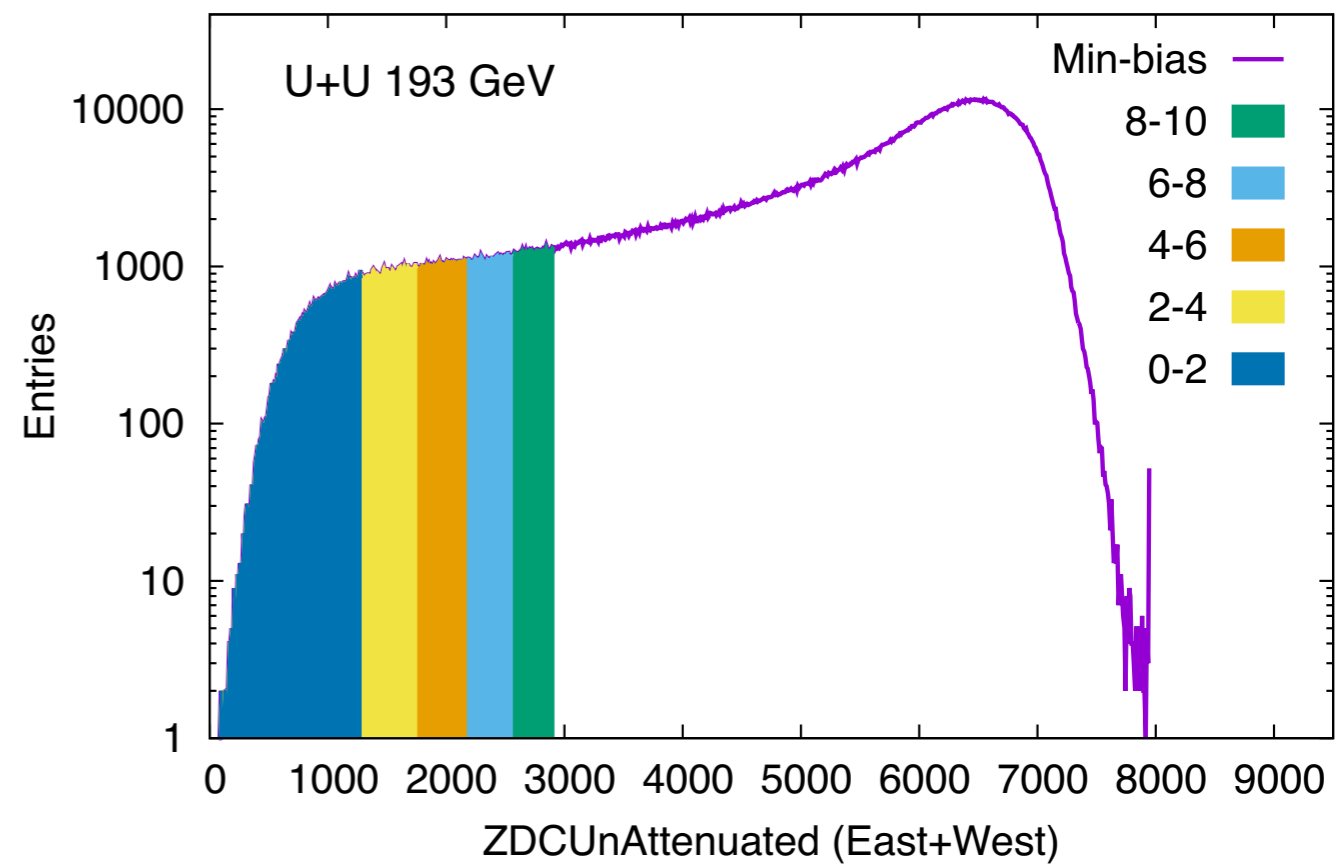
Centrality Selection in 0-10% events

-Method 1 (using RefmultCorr)



Binning on multiplicity

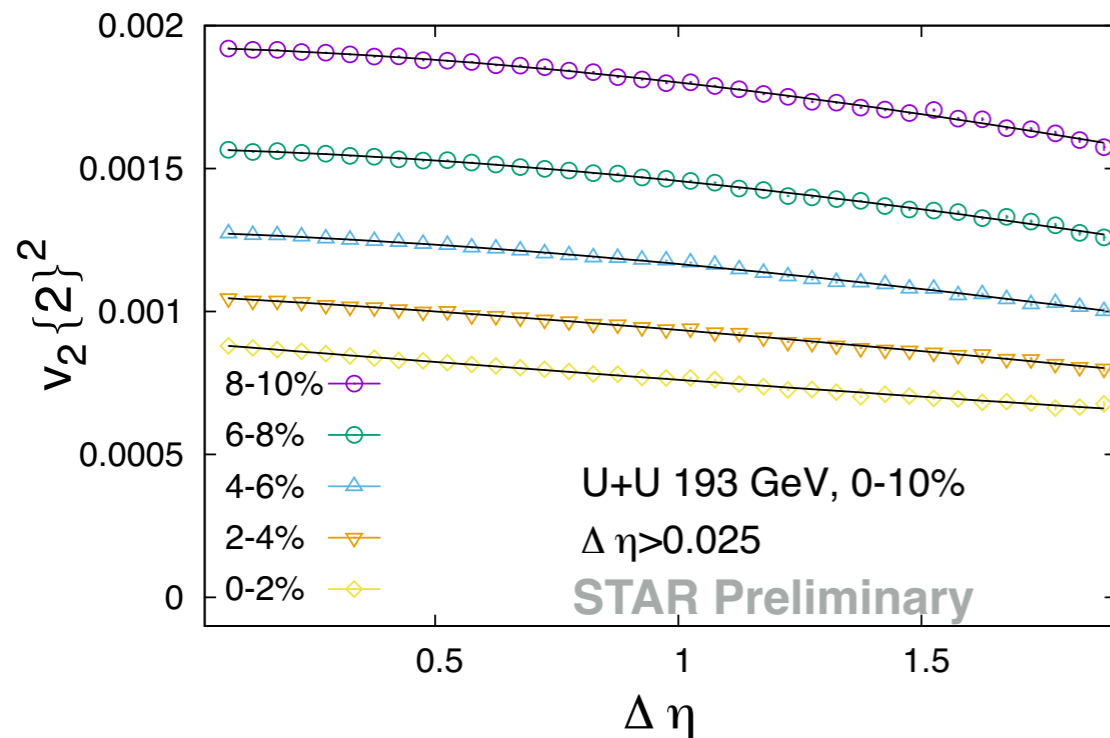
-Method 2 (using ZDCs)



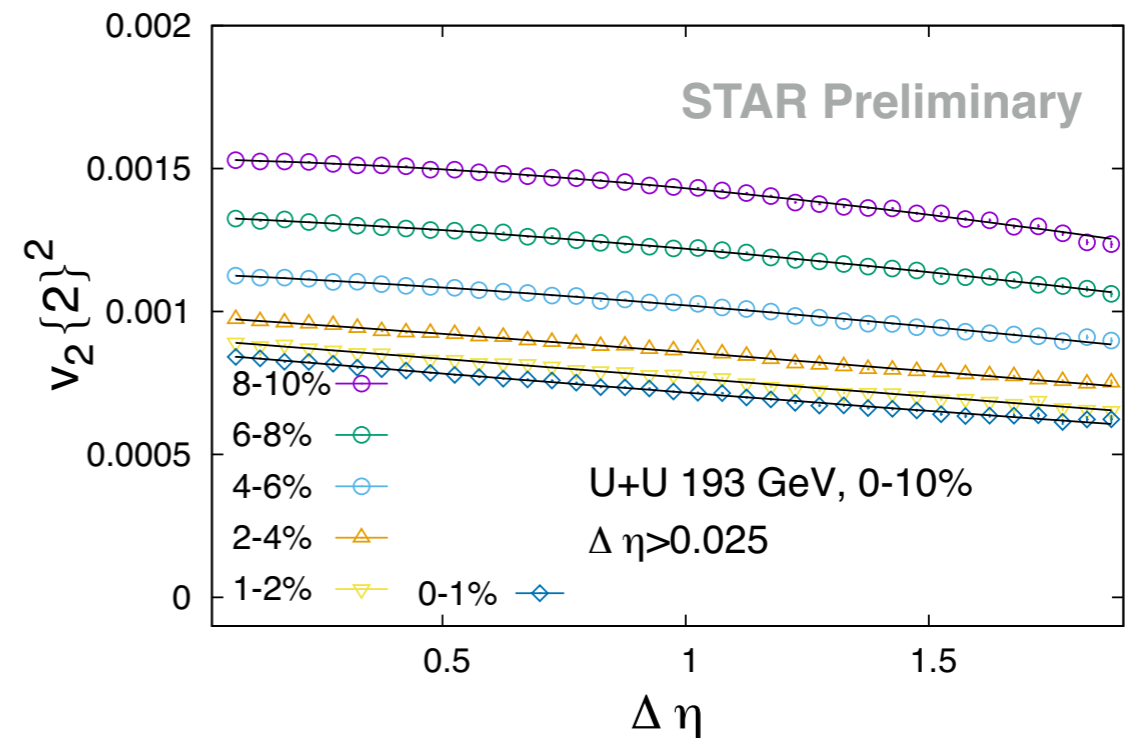
Binning on spectators

Estimation of $v_2\{2\}$ (varying multiplicity & spectators)

After removing track merging and HBT peak



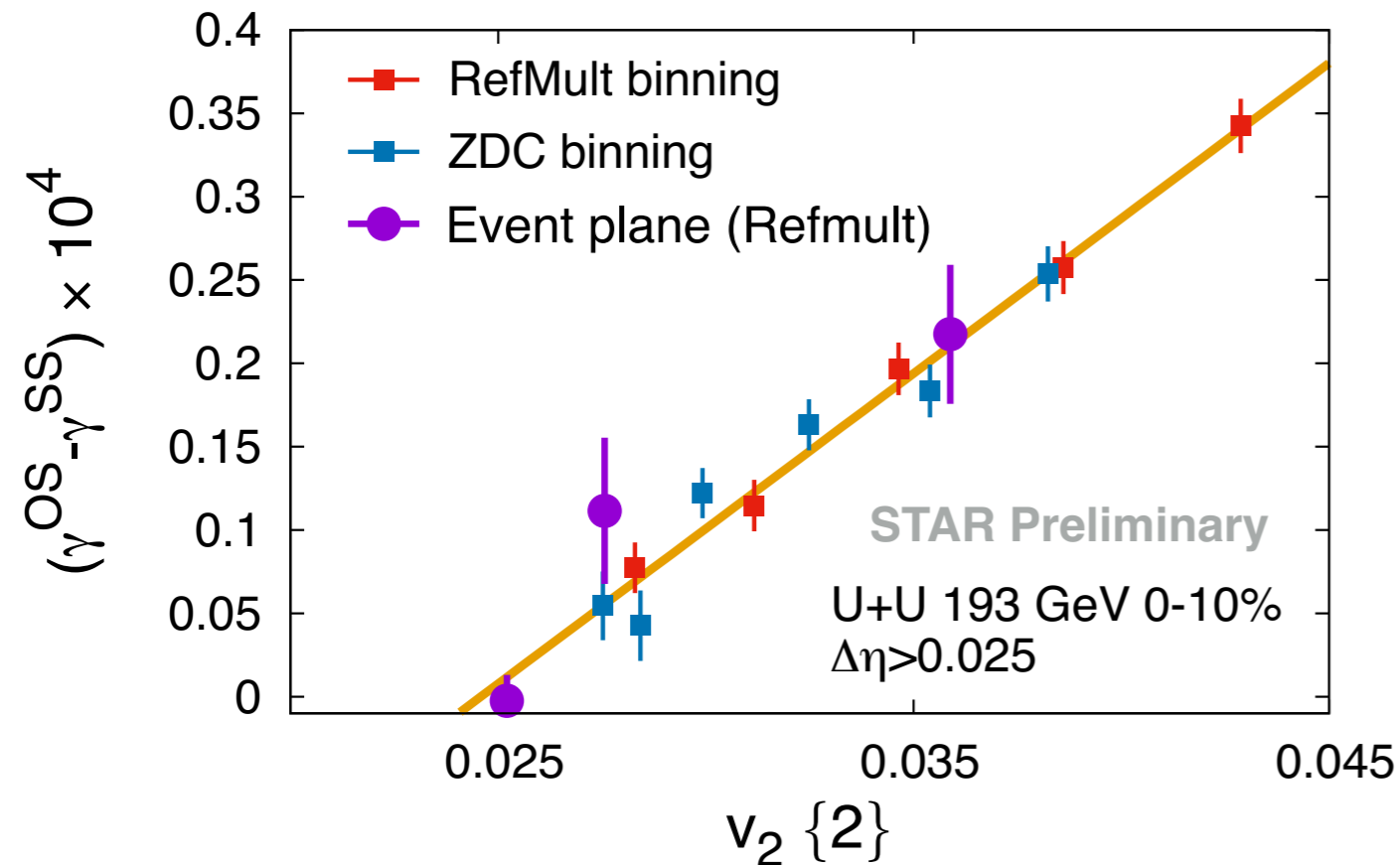
Refmult bins



ZDC bins

Stronger variation of v_2 with multiplicity compared to spectators

γ^{ab} - V_2 correlations (varying multiplicity & spectators)

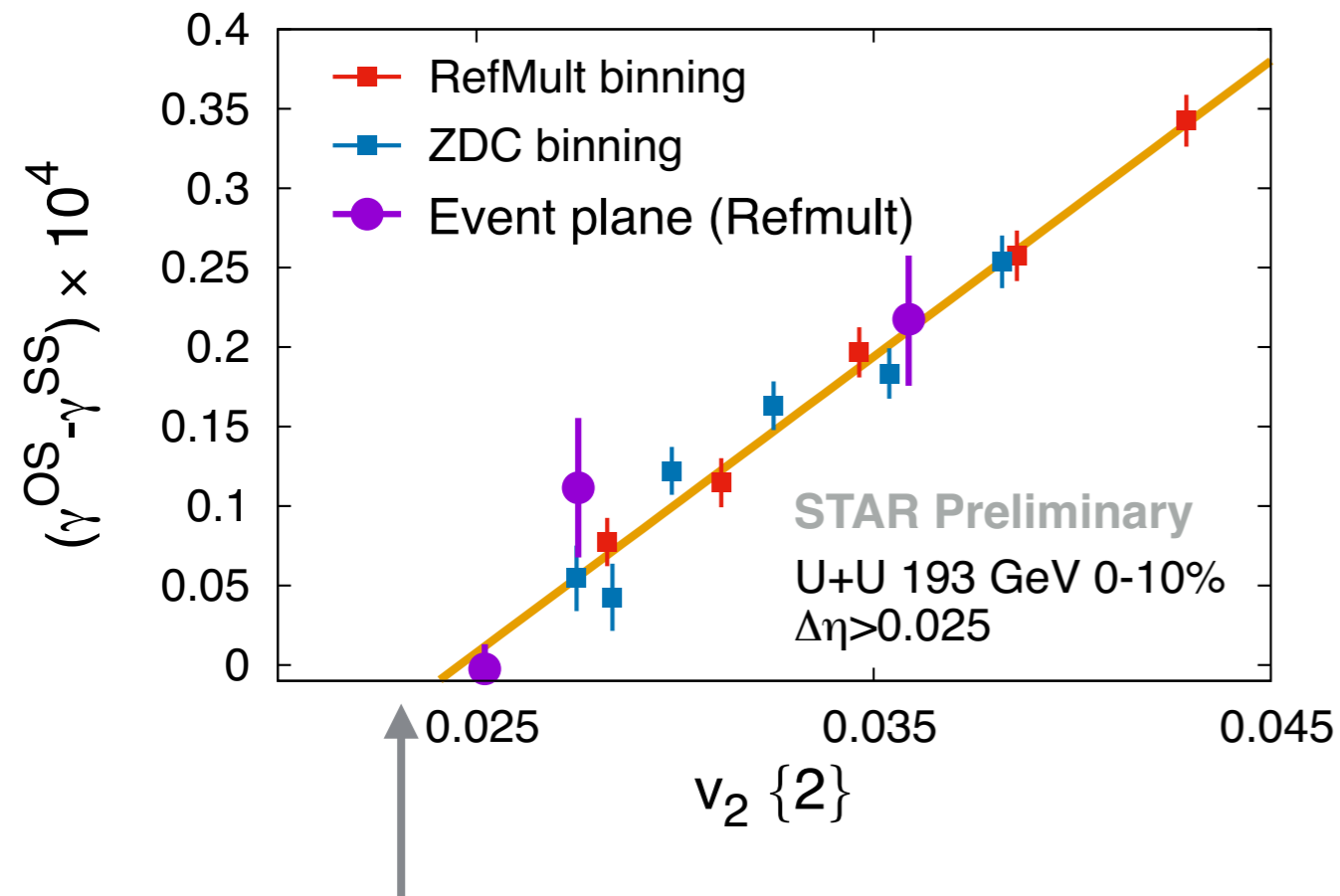


Observations in 0-10%:

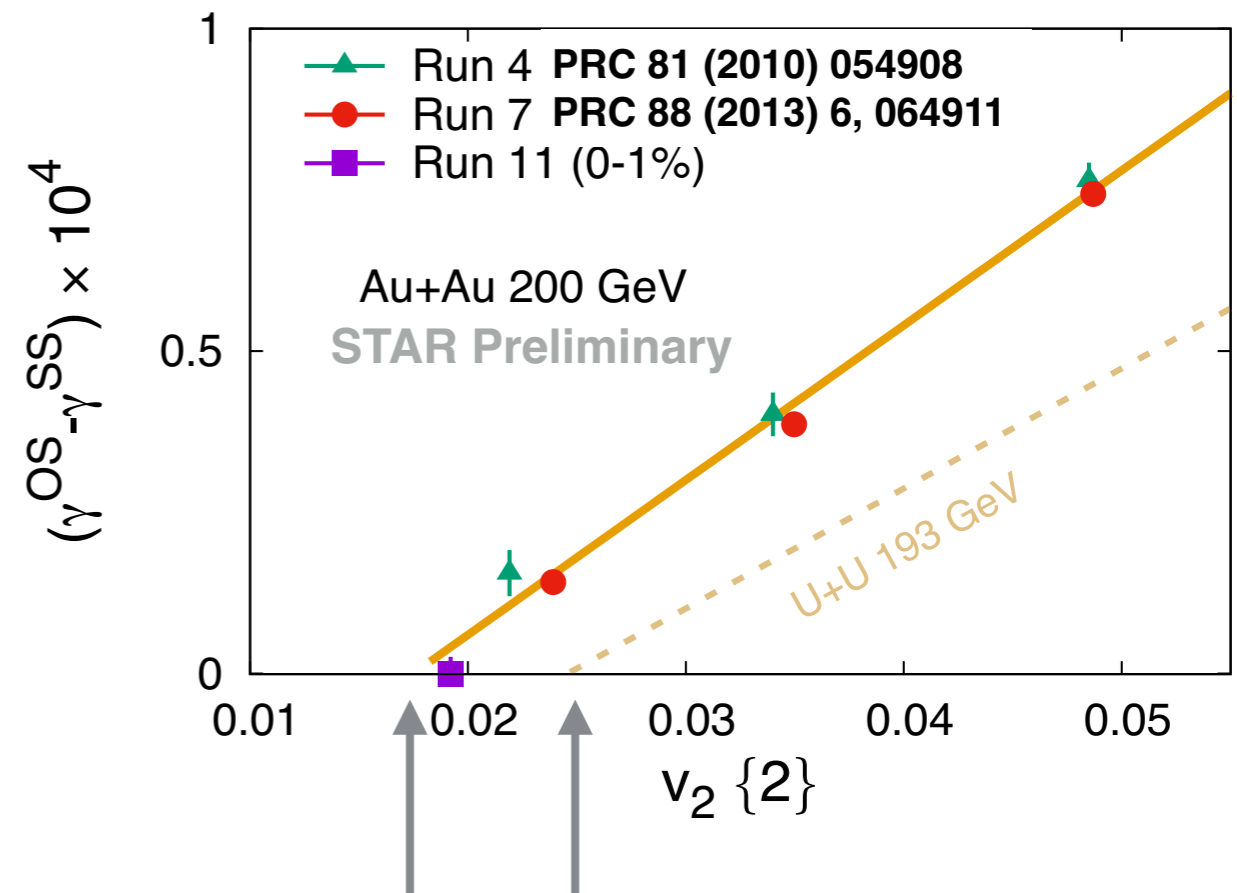
- Strong correlation : **nearly linear dependence** between γ^{ab} & v_2
- $\gamma^{ab} \sim 0$ for $v_2 \neq 0$

$\gamma^{ab}-V_2$ correlations (varying multiplicity & spectators)

U+U collisions



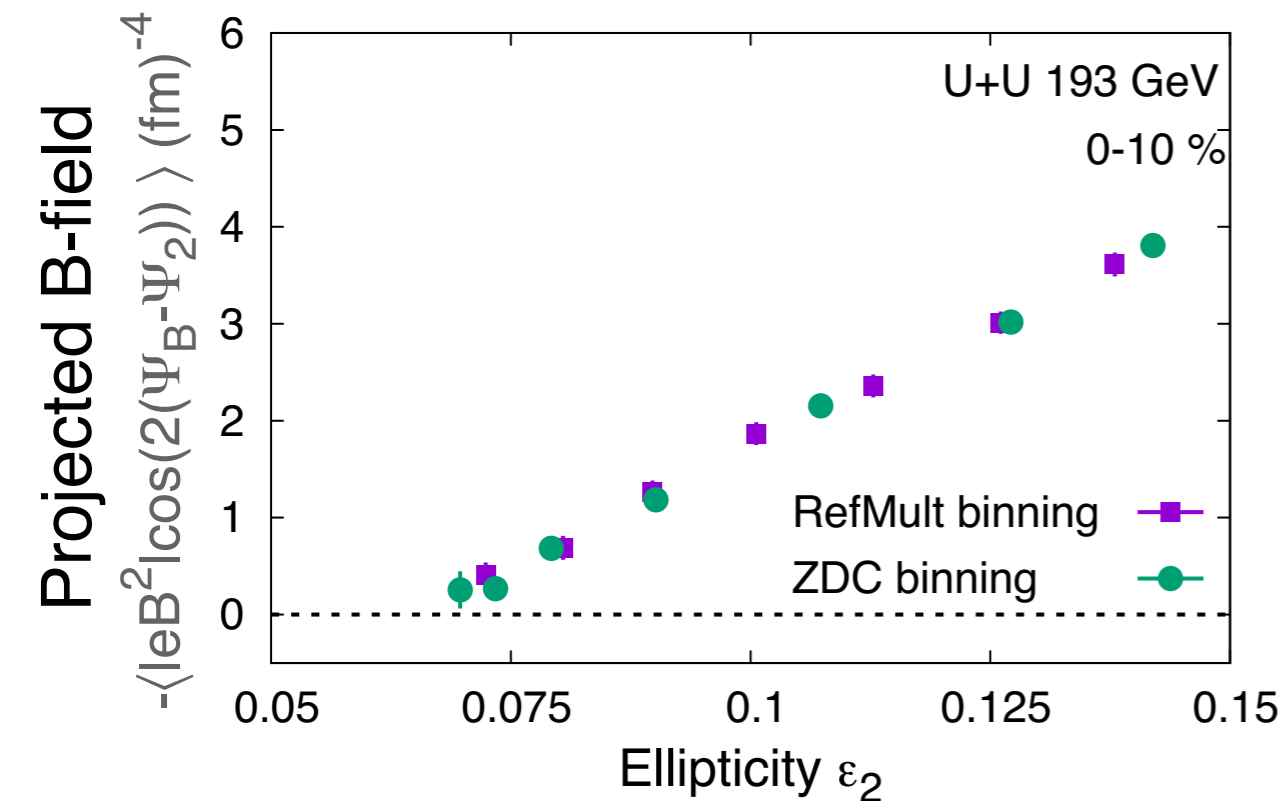
Au+Au collisions



- Observation-I : linear dependence ($\Delta\gamma-v_2$)
- Observation-II : $\Delta\gamma = (\gamma^{OS} - \gamma^{SS}) \sim 0$ for non-zero v_2

Can model calculations provide some insights ?

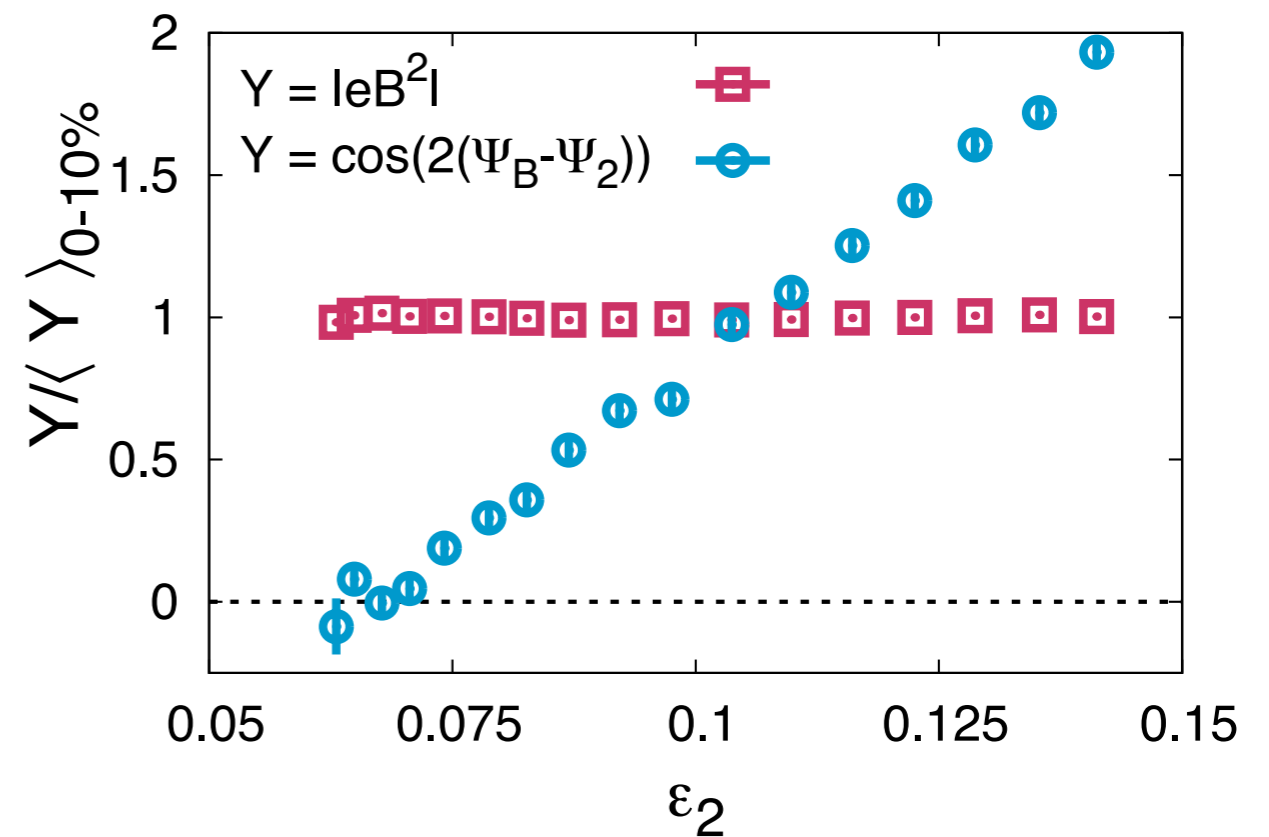
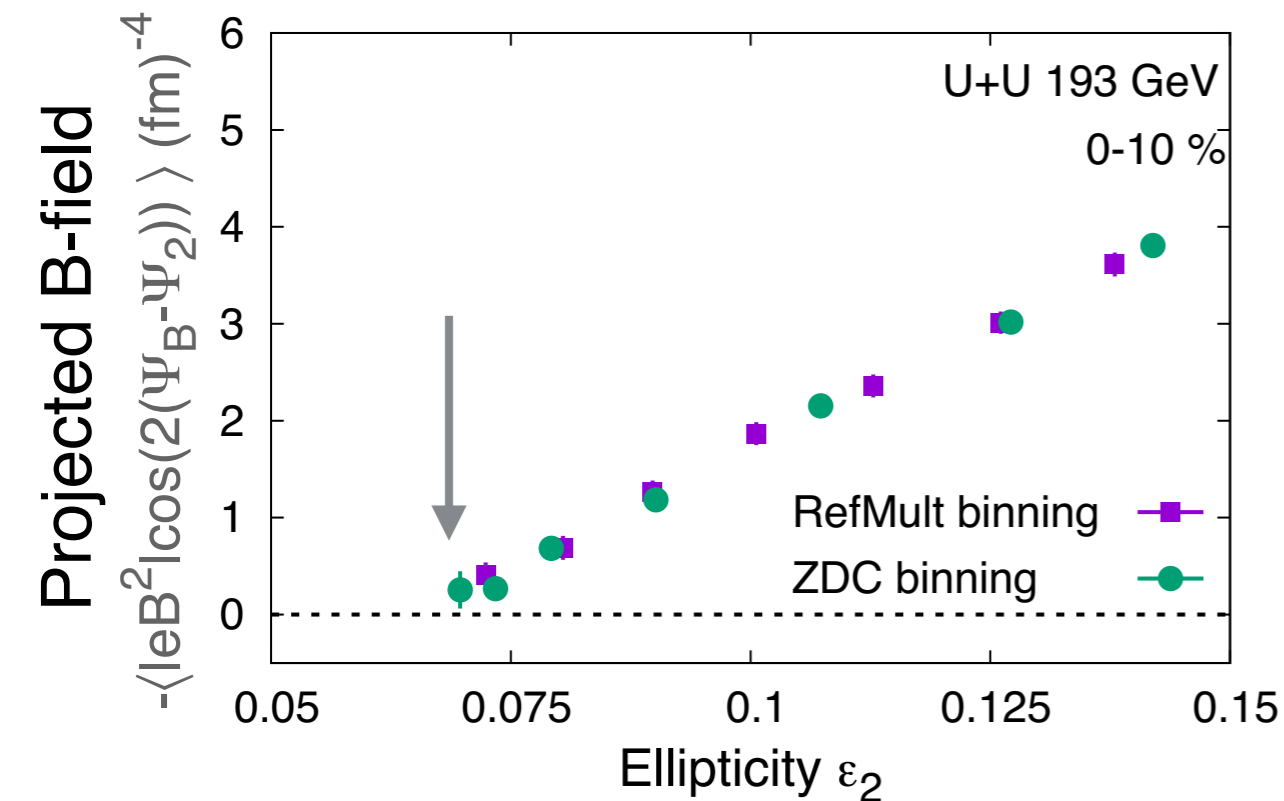
Proxy for γ^{ab} & v_2 \rightarrow MC-Glauber model simulation



Linear dependence ($\Delta\gamma-v_2$) + offset \rightarrow also seen in model ($B-\epsilon_2$)

Can model calculations provide some insights ?

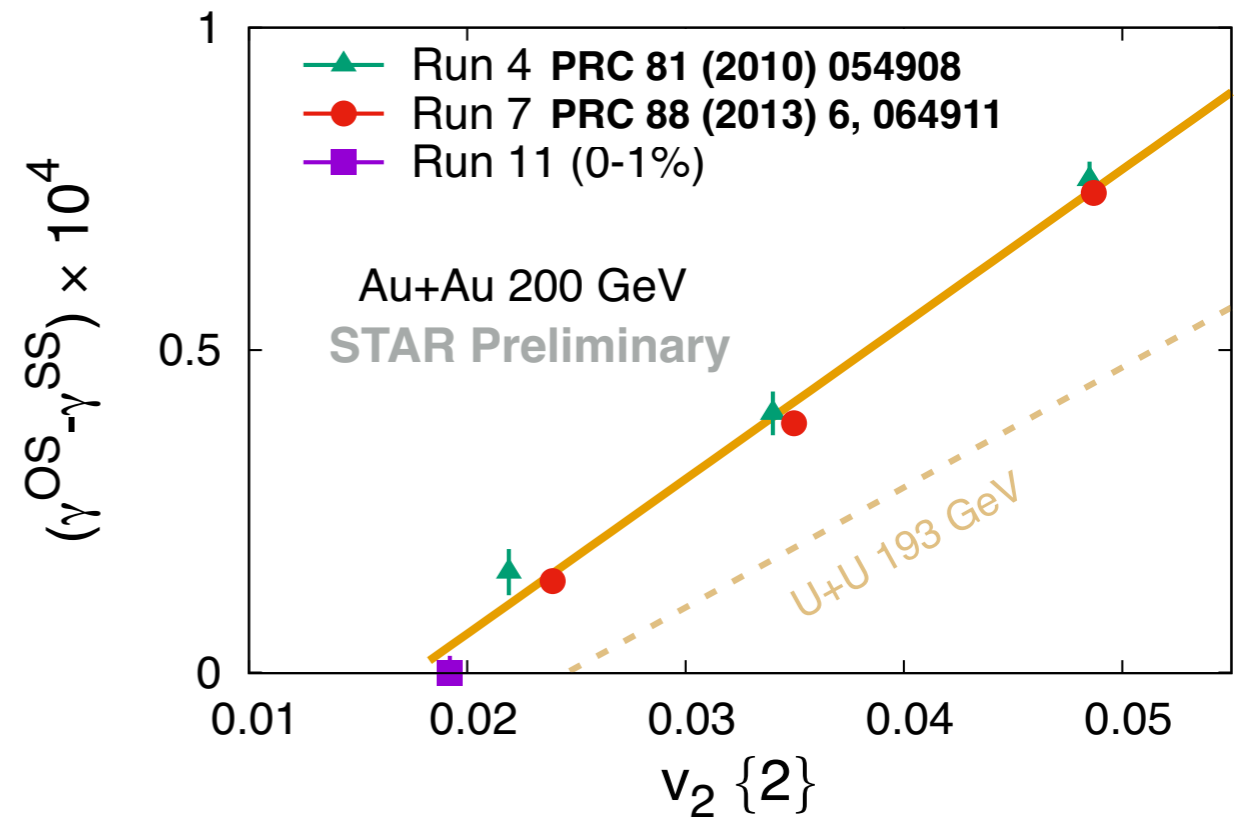
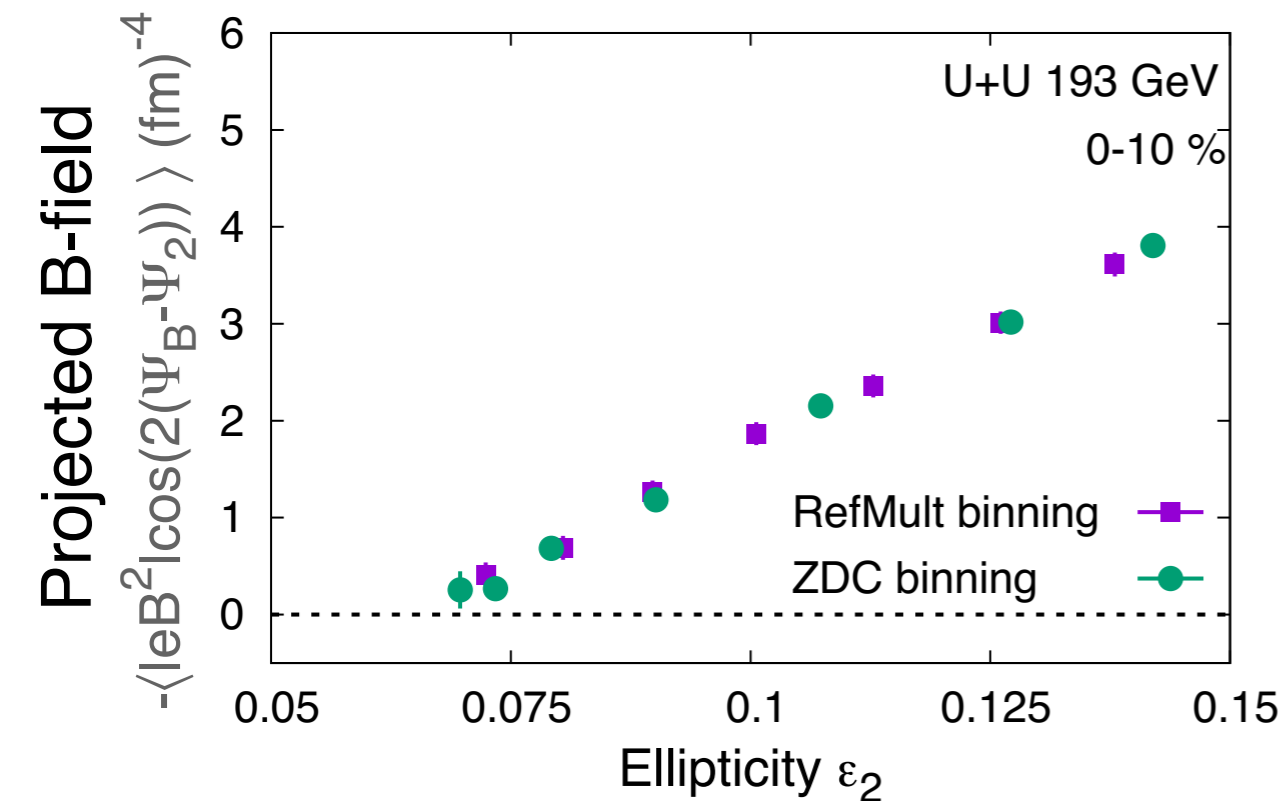
Proxy for γ^{ab} & v_2 \rightarrow MC-Glauber model simulation



Linear dependence ($\Delta\gamma$ - v_2) + offset \rightarrow also seen in model (\vec{B} - ϵ_2)
Central events \rightarrow direction of \vec{B} de-correlates with event-plane

Can model calculations provide some insights ?

Proxy for γ^{ab} & $v_2 \rightarrow$ MC-Glauber model simulation

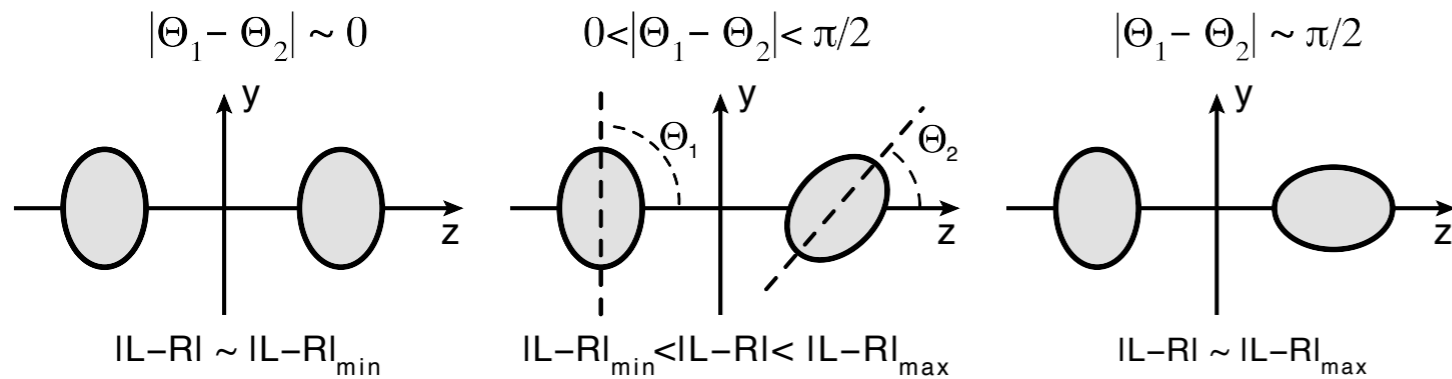


Background driven scenario : $\Delta\gamma \rightarrow 0$ when $v_2 \rightarrow 0$
 B-field driven scenario : $\Delta\gamma \rightarrow 0$ when $\cos(\Psi_B - \Psi_2) \rightarrow 0$ even if $v_2 \neq 0$

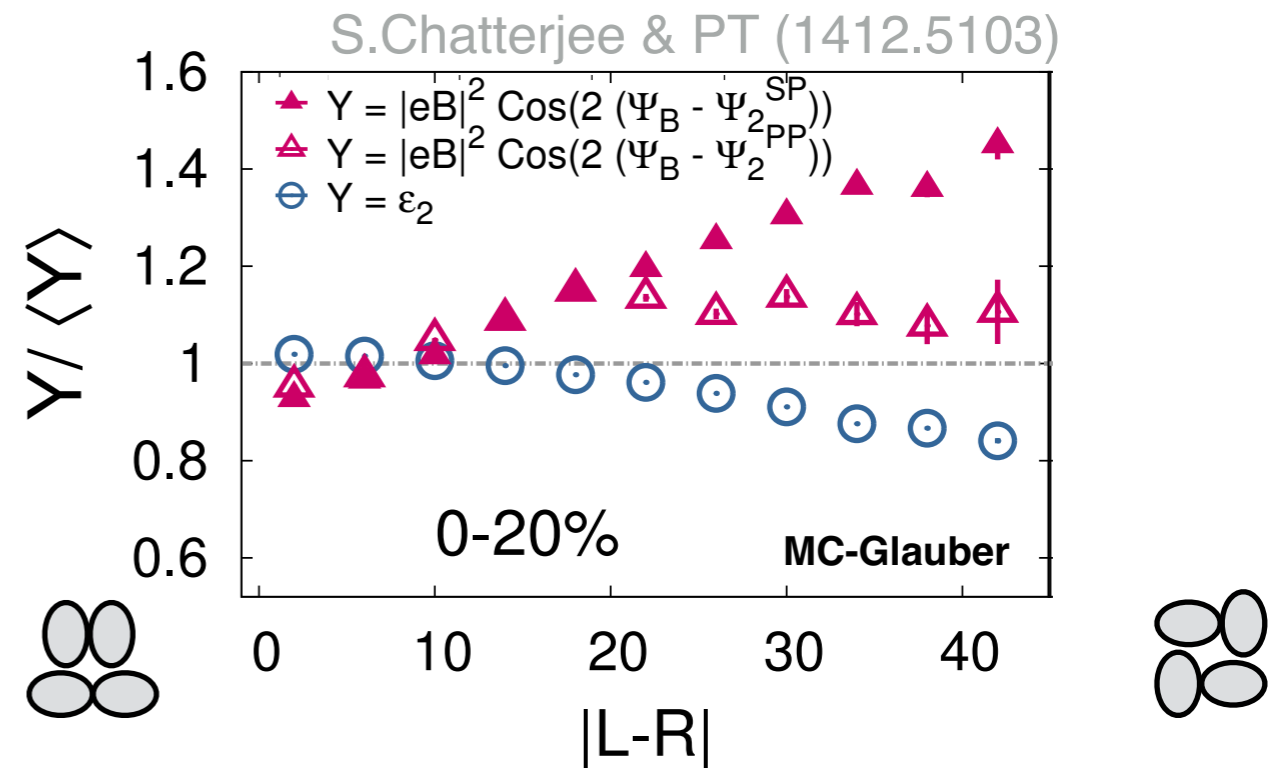
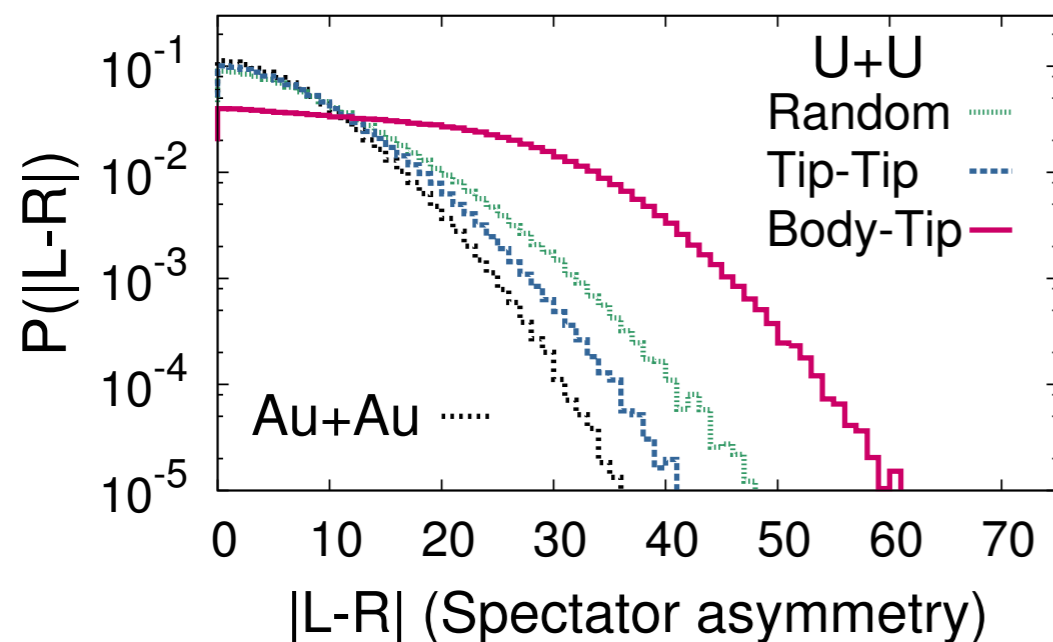
New challenge for 100% background driven models

What else can we try with the existing U+U data ?

A new tuning parameter to disentangle ε_2 and B-field



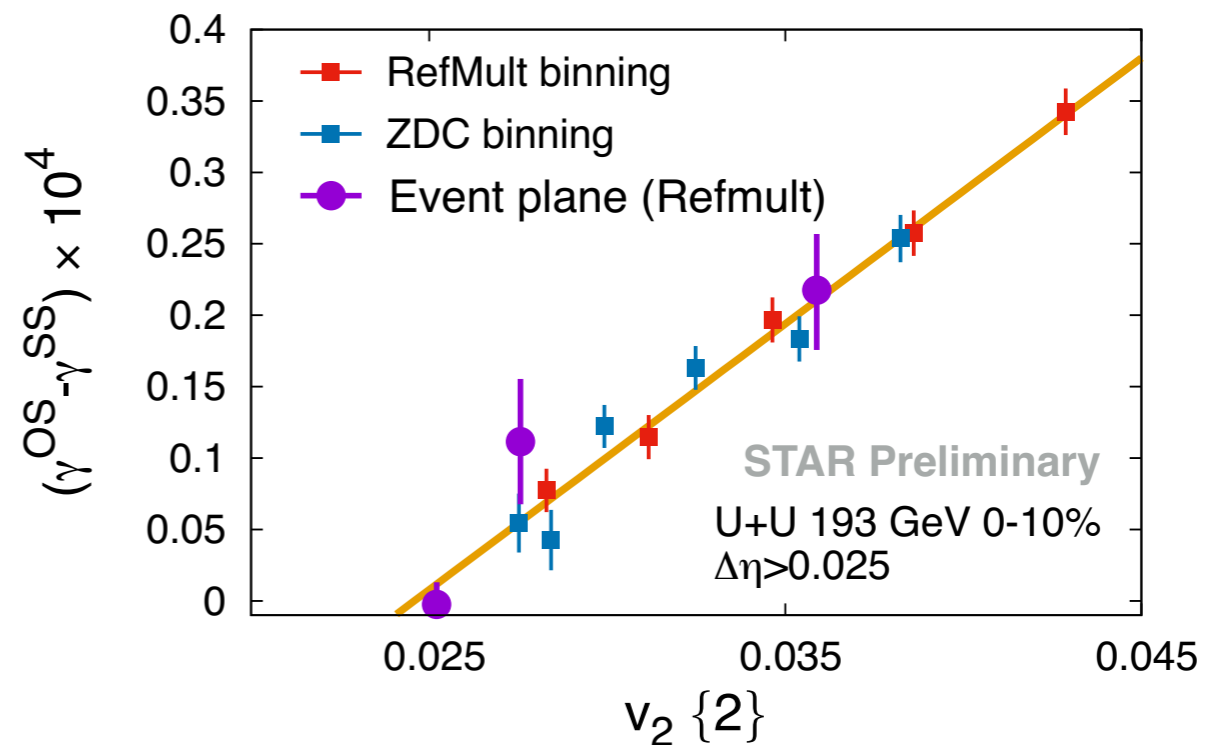
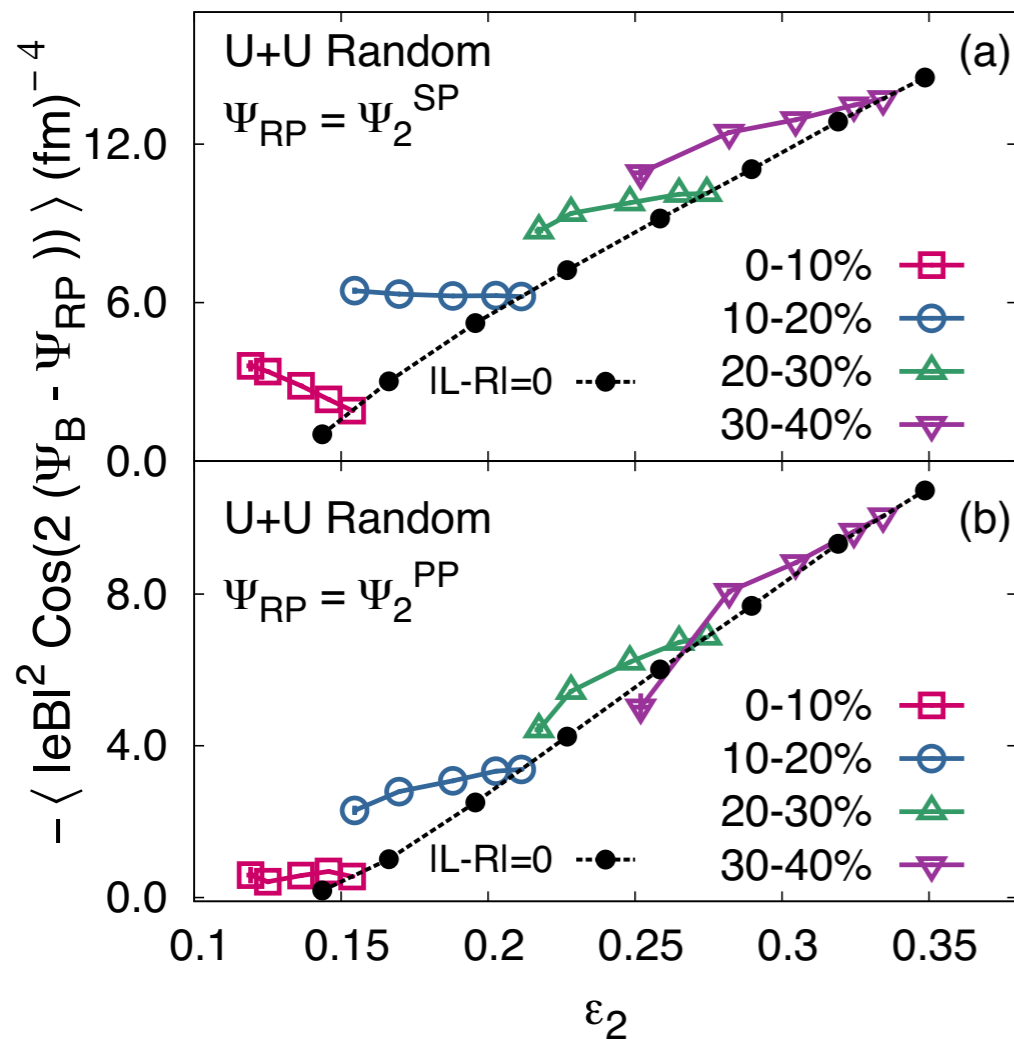
R: right going spectator neutrons
L: left going spectator neutrons



Binning in $IL-Rl$ it is possible to trigger body-tip events : $B \uparrow \varepsilon_2 \downarrow$

Spectator asymmetry in U+U to disentangle $\Delta\gamma$ & v_2

S.Chatterjee & PT (1412.5103)



Spectator asymmetry \rightarrow
 Triggers event with different B but
 same ε_2 & vice-versa

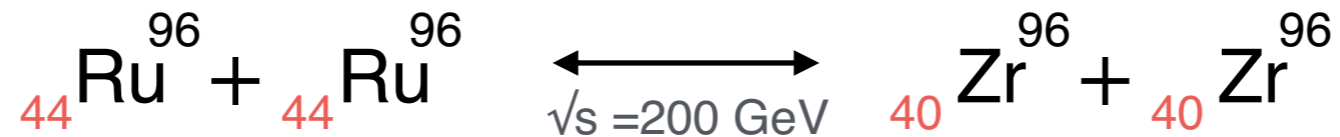


Look for similar trend between
 $\Delta\gamma$ & v_2 with ZDC asymmetry

Analysis under progress (challenges : ZDC response to neutrons)

Outlook for Isobar collisions at RHIC

Idea is to change B-field without changing background



Different B-field with same flow background is expected

e-A scattering experiment

	Zr	Ru
R0	5.05	5.13
a(d) [fm]	0.45	0.45
β_2	0.18	0.03
β_4	0.01	0.009

(Woods-Saxon parameters)

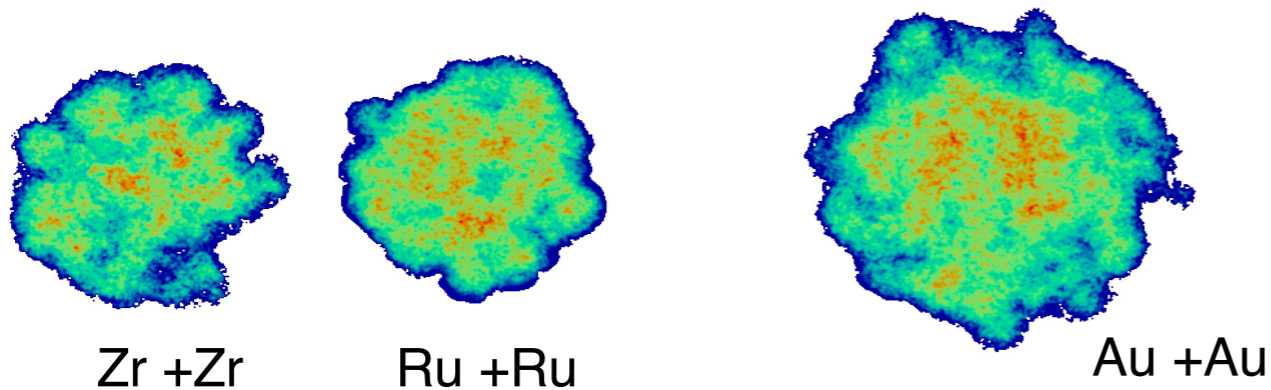
Q. Y. Shou *et al*
arXiv:1409.8375

comprehensive model deduction

	Zr	Ru
R0	5.07	5.14
a(d) [fm]	0.48	0.46
β_2	0.06	0.13
β_4	0	0

Ref: Gang Wang, QCD
Chirality workshop '2016

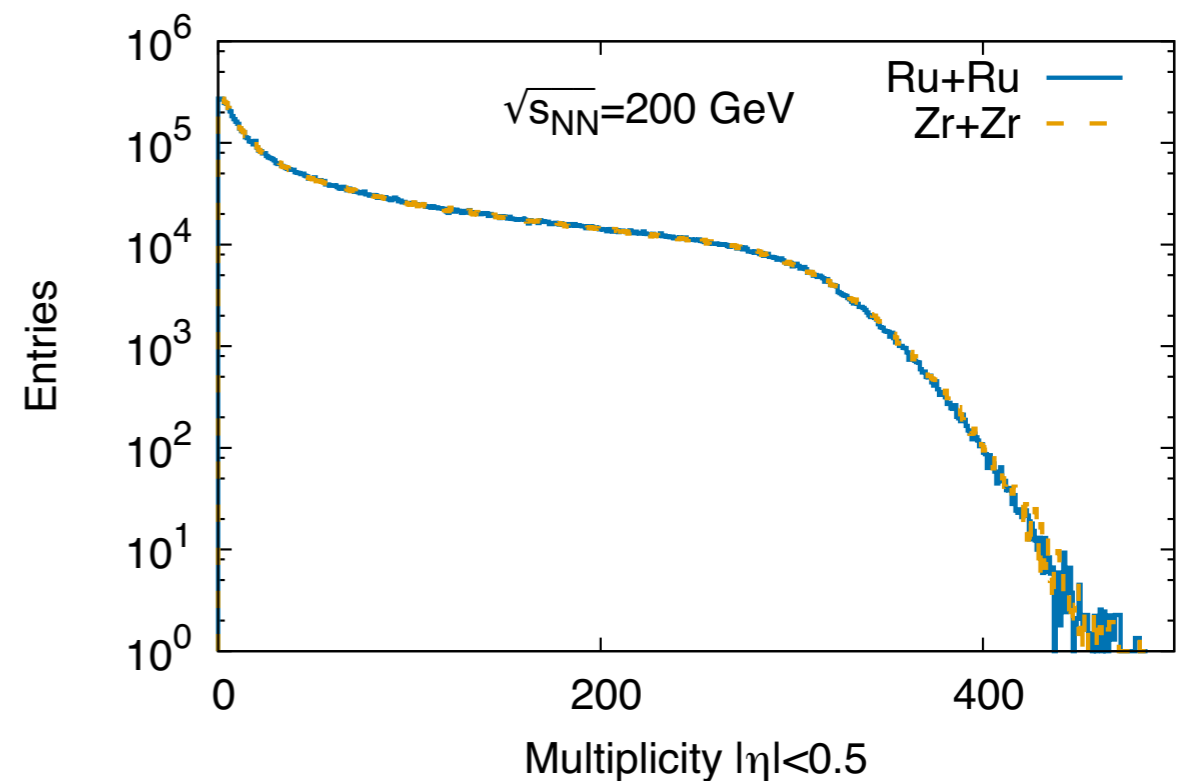
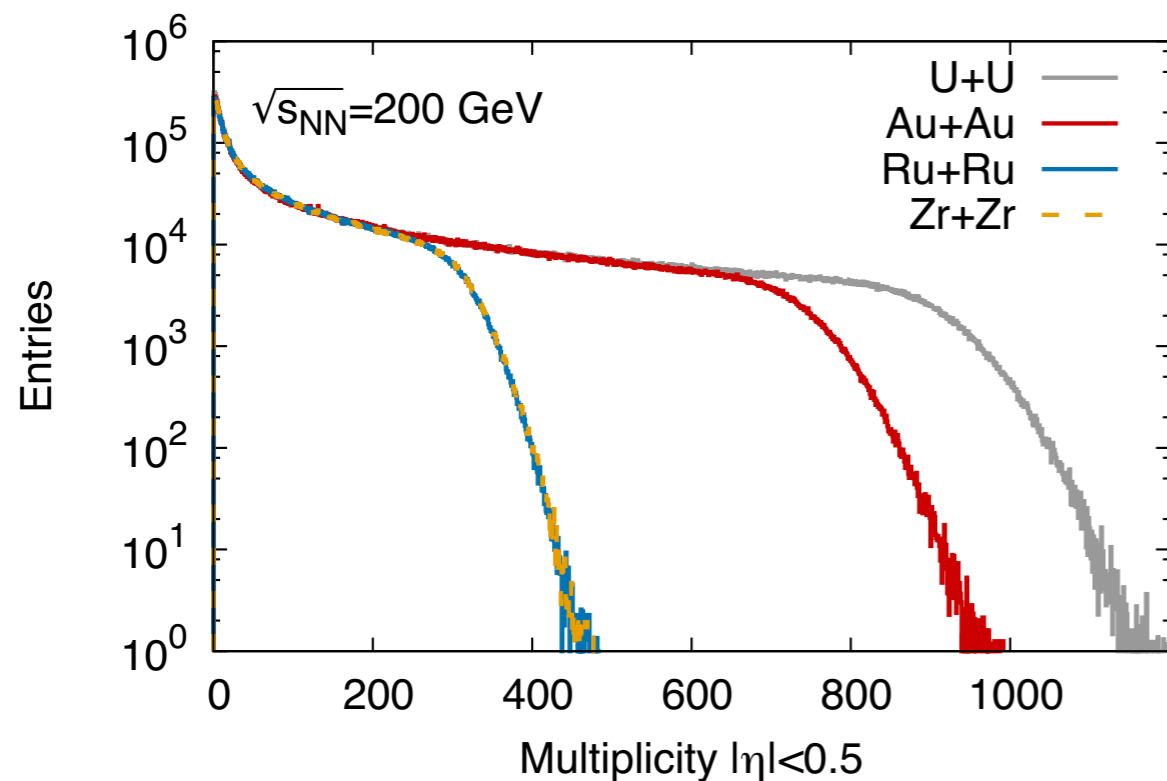
http://starmetings.physics.ucla.edu/sites/default/files/gang_wang.pdf



Single collision in IP-Glasma model (b=0)

Comparisons of multiplicities for centrality estimation

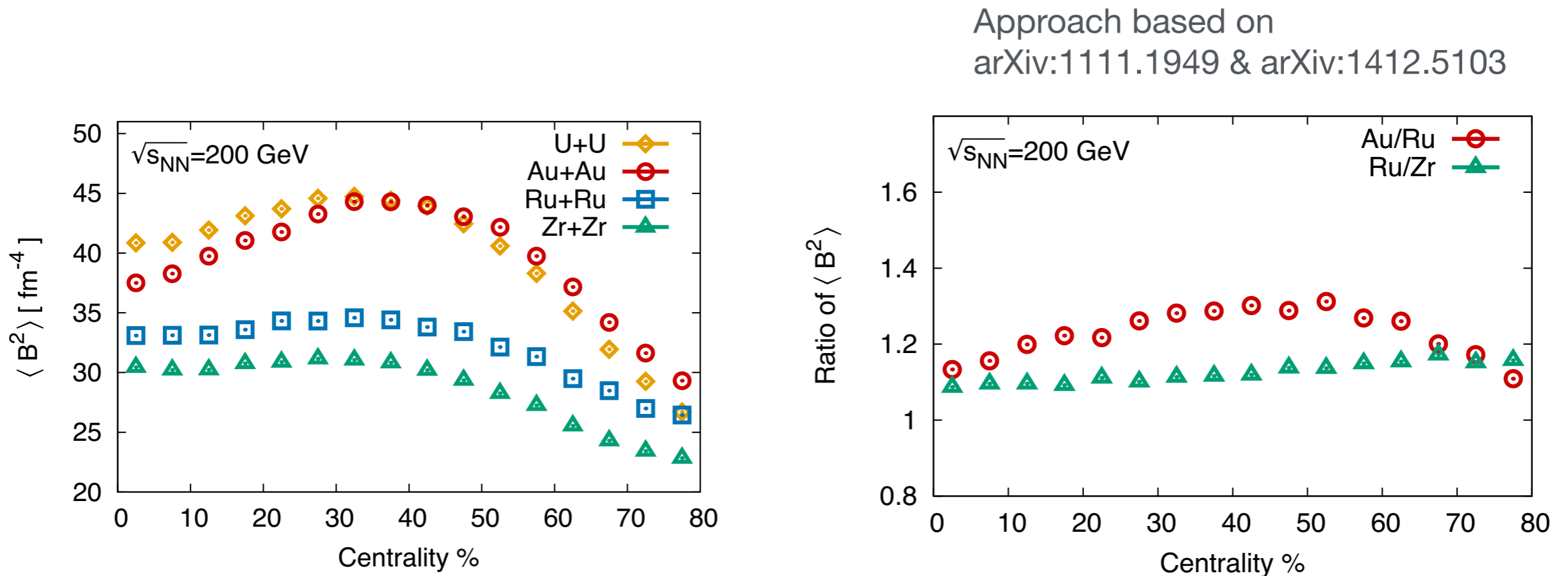
- Two component MC-Glauber corrected $dN/d\eta$
- Using parameters compatible to e-A scattering data



Comparison between different systems \rightarrow Zr+Zr & Ru+Ru similar

Comparisons of the magnetic fields

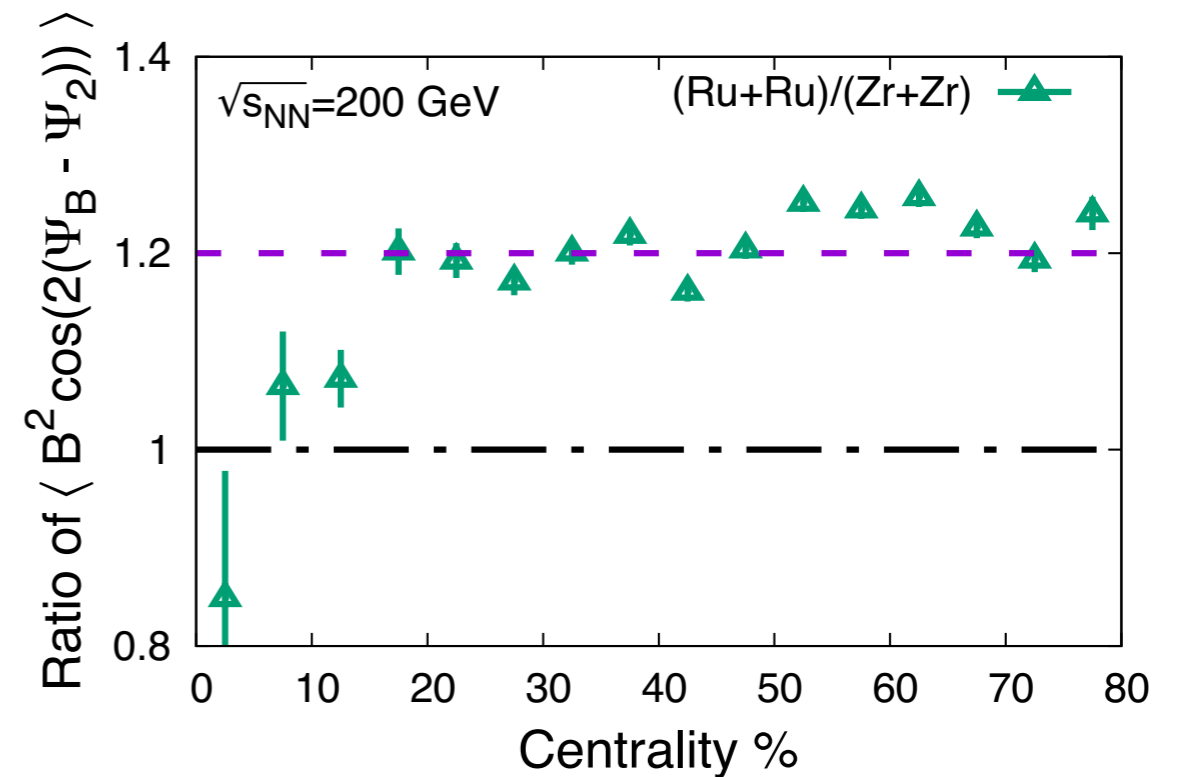
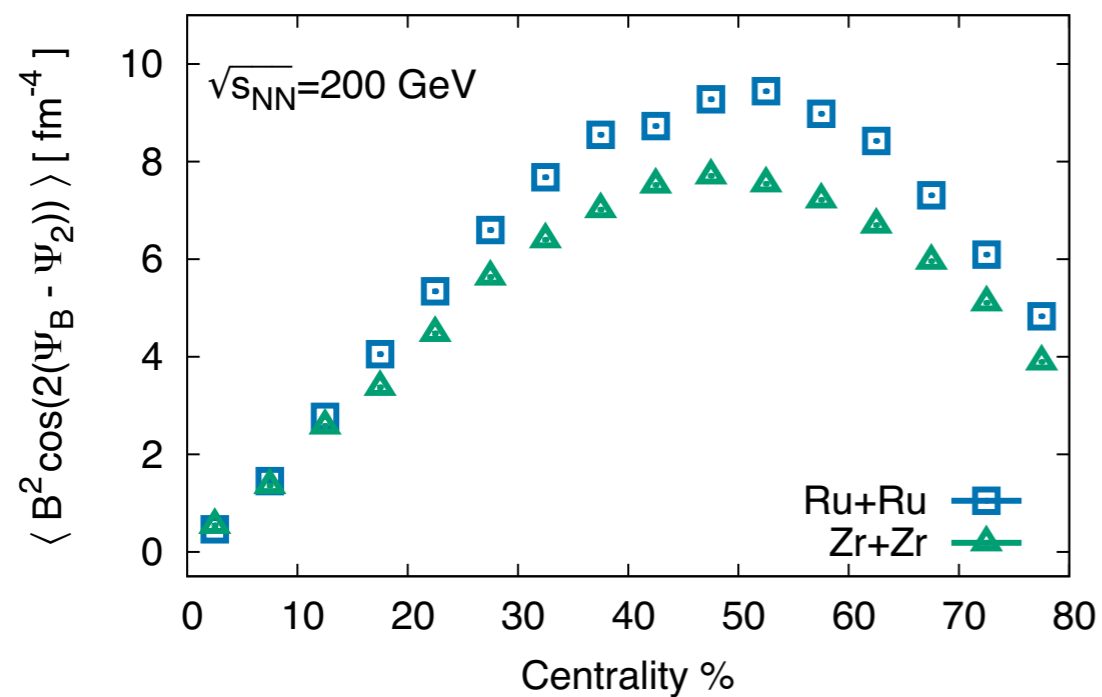
Estimation of B-field \rightarrow $t=0$, center of participant zone (in vacuum)



Estimation of B-field \rightarrow not effected by nuclear deformation

Comparisons of the projected magnetic fields

CME signal depends on both $|B|$ and direction Ψ_B



- Signal strength $\sim |B^2| \cos(2(\Psi_B - \Psi_2)) \rightarrow$ Projected field
- About 20% difference in Ru+Ru vs Zr+Zr

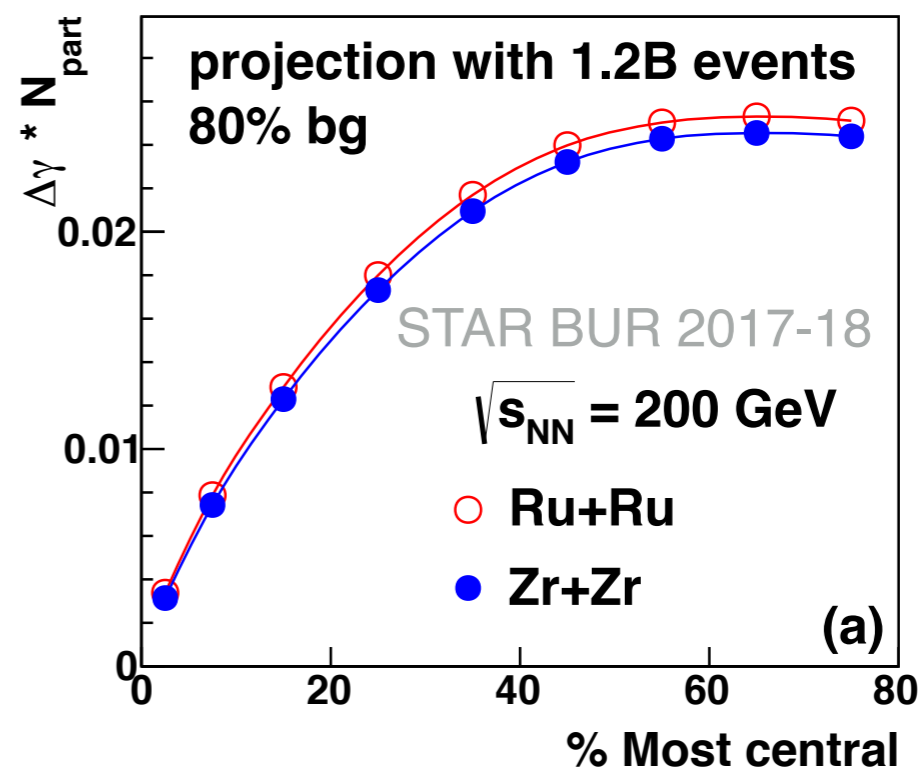
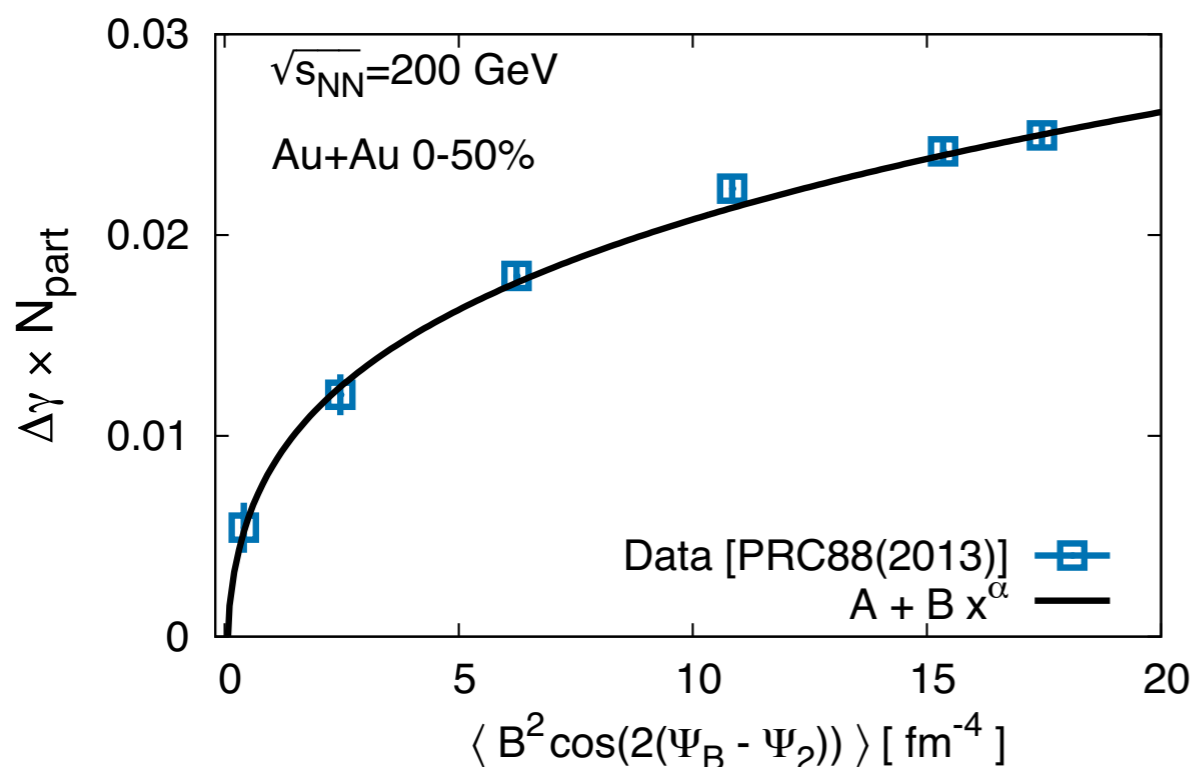
Projection for CME correlator in Isobar collisions

- Step-I \rightarrow parameterize Au+Au data in terms of B

$$Y = \Delta\gamma \times N_{\text{part}} = \alpha + \beta \langle B^2 \cos(2(\Psi_B - \Psi_2)) \rangle^\lambda$$

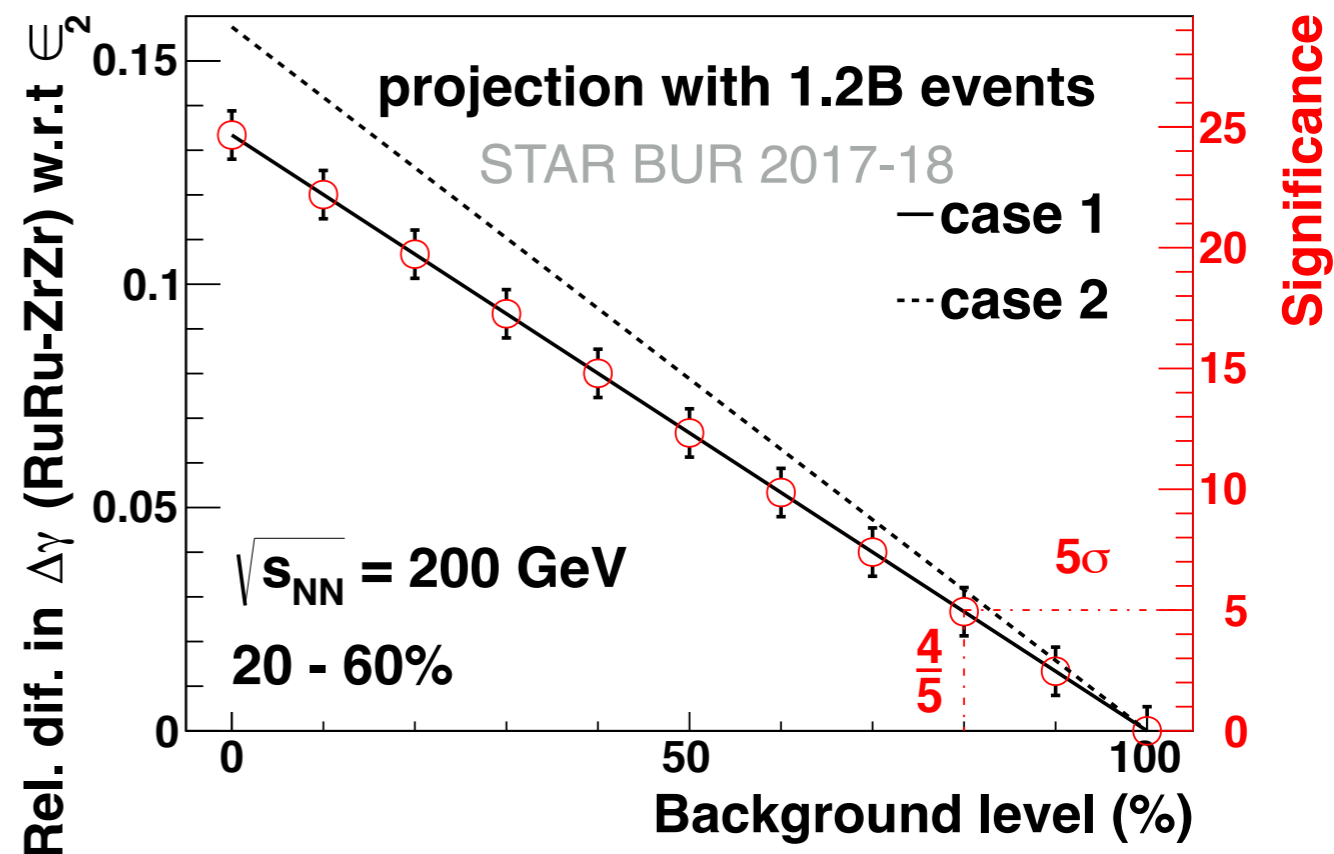
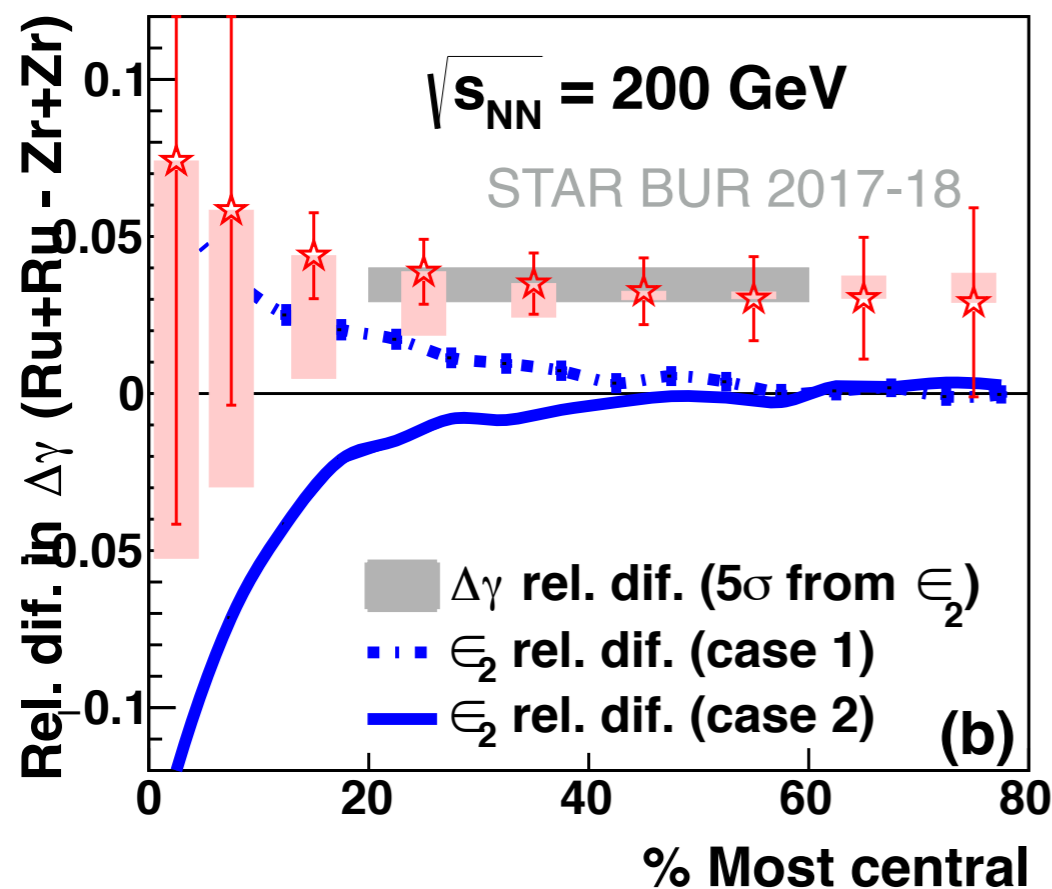
- Step-II \rightarrow make projections for Ru+Ru & Zr+Zr

Contribution from the background has to be incorporated



Significance of the expected signal in Isobar collisions

Ref: https://drupal.star.bnl.gov/STAR/system/files/STAR_BUR_Run1718_v19.pdf



Isobar ratio gives maximum 5σ (1.2 B) if CME is $\sim 20\%$ B-field driven

Equivalent running of RHIC : 3.5-weeks for each species (STAR proposal for BUR in Runs 18)

Summary / Outlook

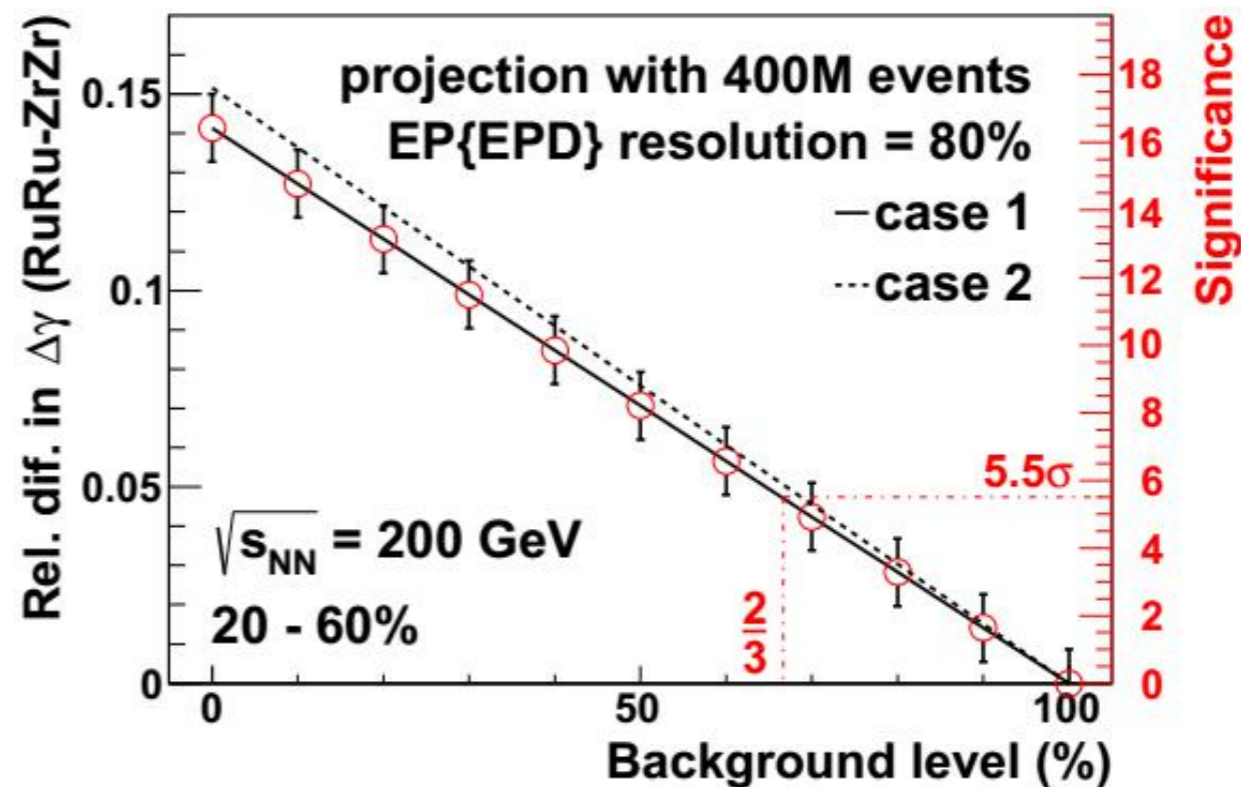
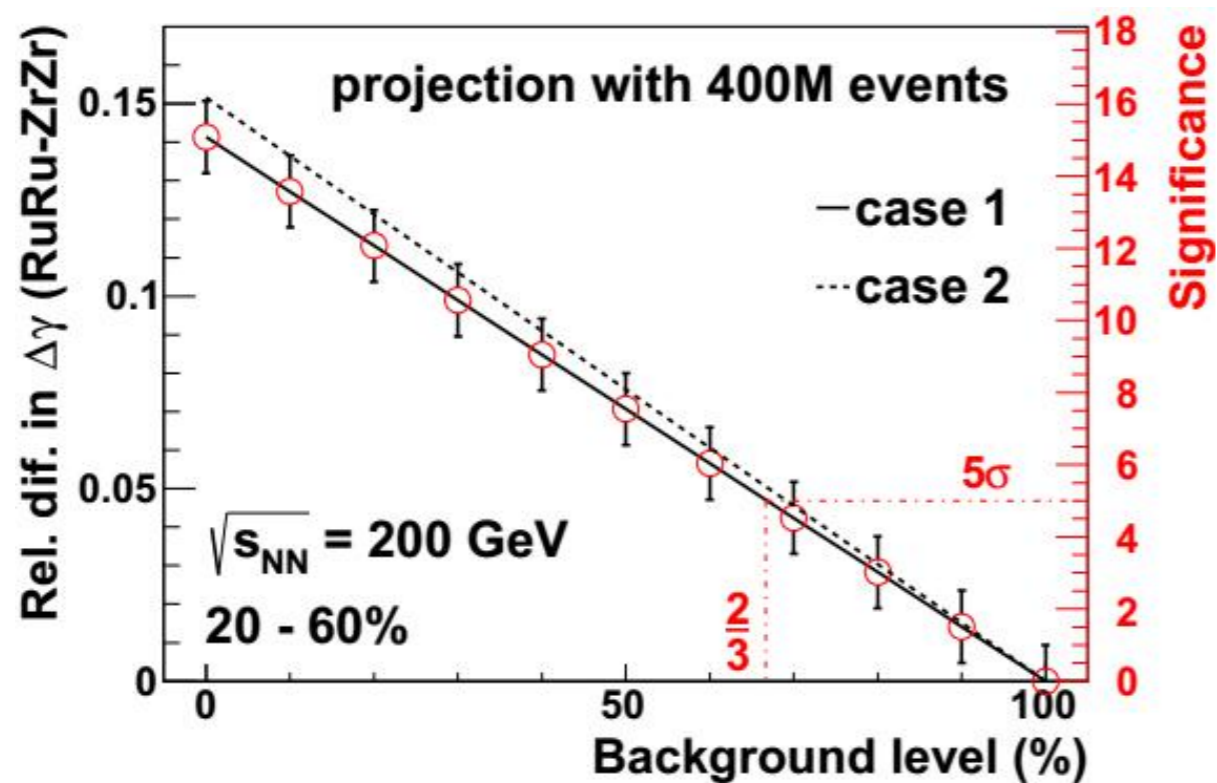
- Charge dependent azimuthal correlations have been studied in U+U & compared to Au+Au results.
- Strong correlation between γ^{ab} & v_2 observed in central events (<10%) with $\gamma^{ab} \sim 0$ for $v_2 > 0$ in both U+U and Au+Au collisions.
- New analyses under way to use U+U data to disentangle CME signal from backgrounds (specifically using spectator asymmetry).
- Collisions of Isobar looks very promising : (3.5 x 2) weeks of running with about (1.2 x 2) B events can provide about 5σ confidence of signal/bkg.

backup

Significance of the expected signal in Isobar collisions

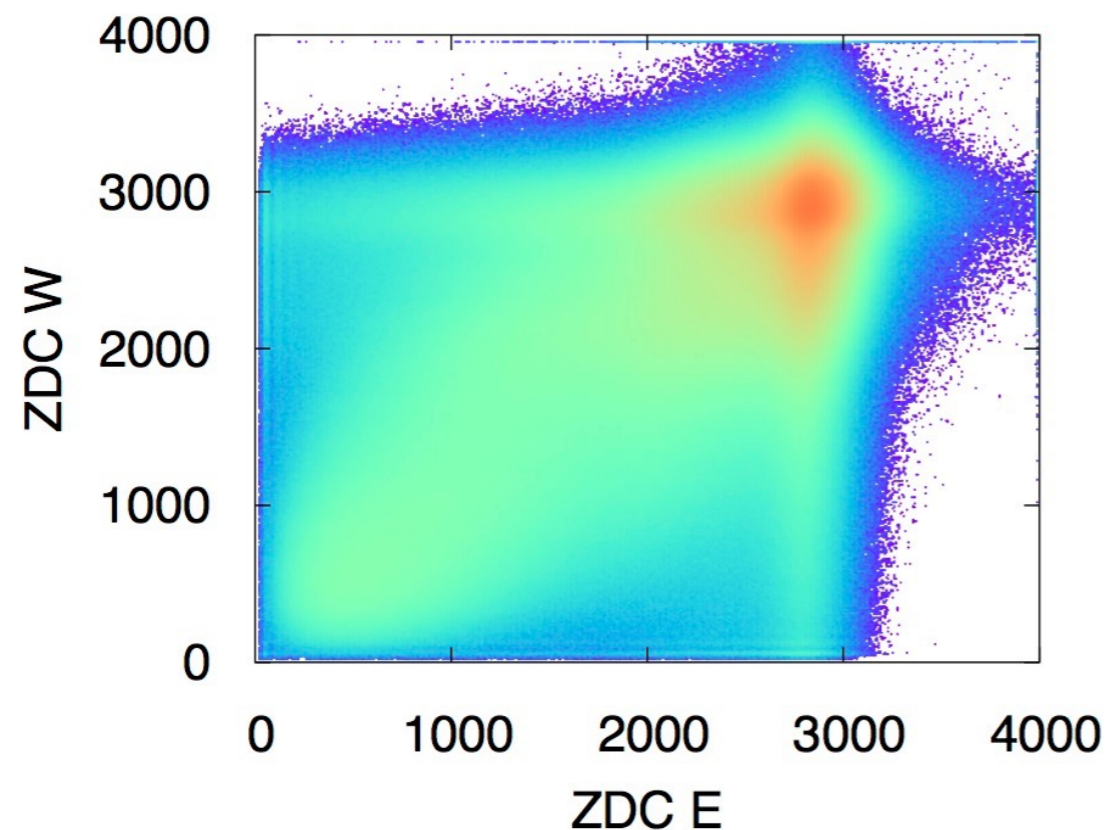
Projection for 400M events

Study done by Gang Wang



Spectator asymmetry in U+U collisions

Body-Tip events are experimentally triggered by asymmetry of ZDCs

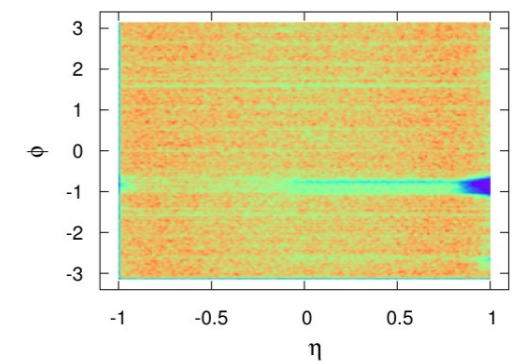
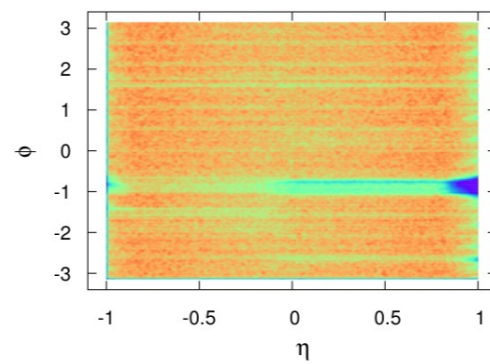
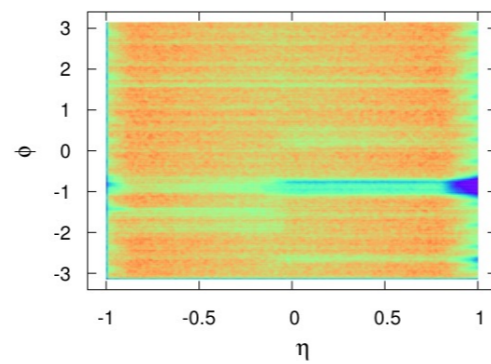
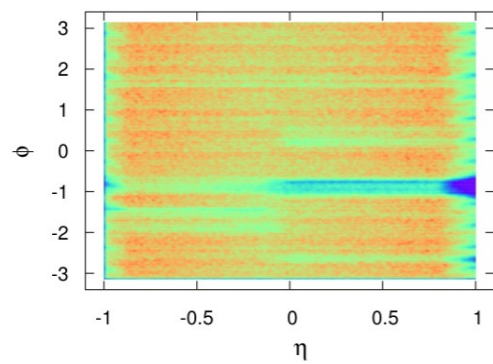
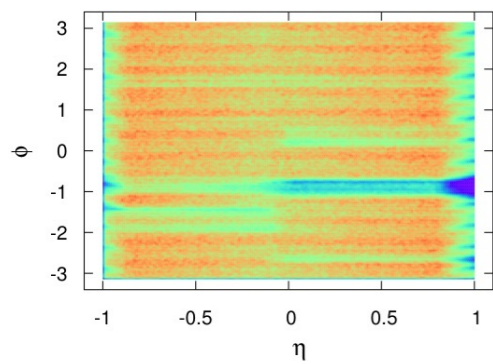
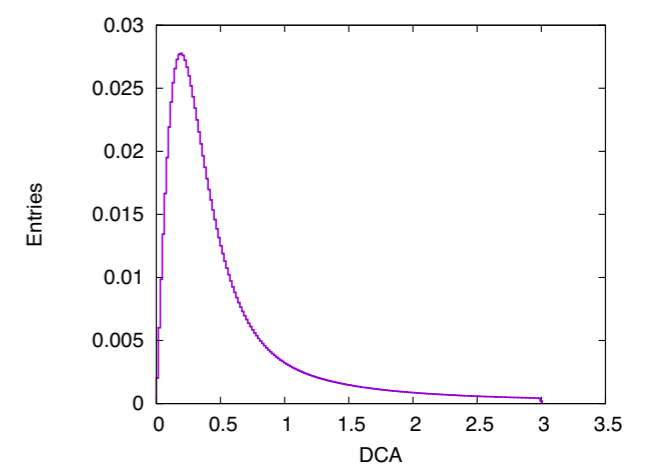
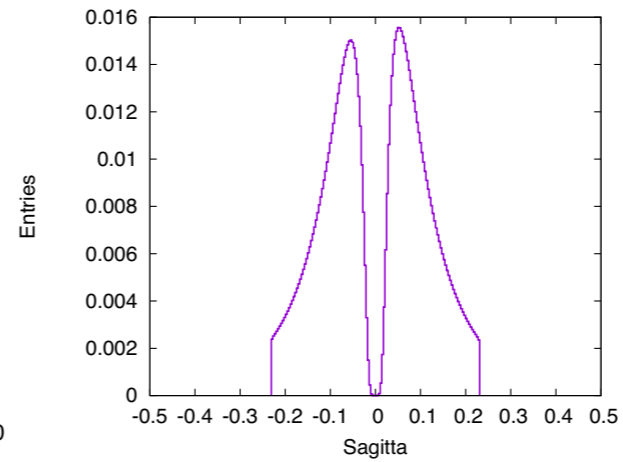
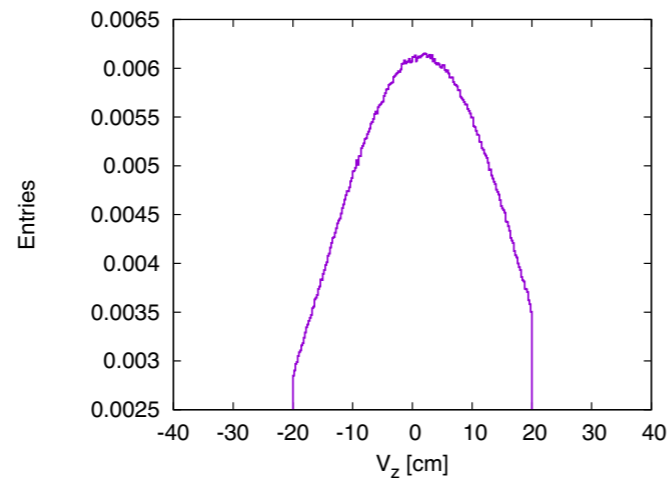
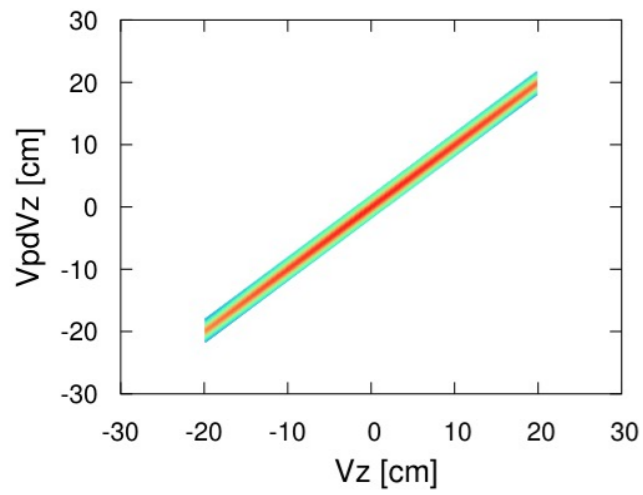
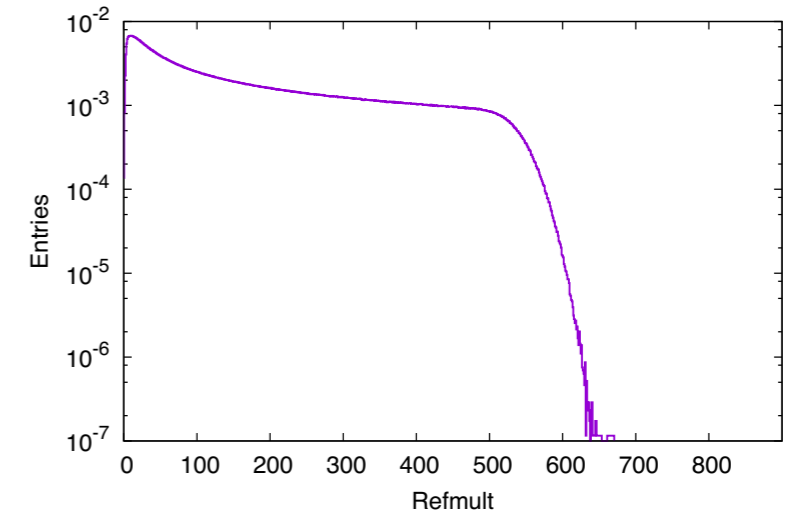
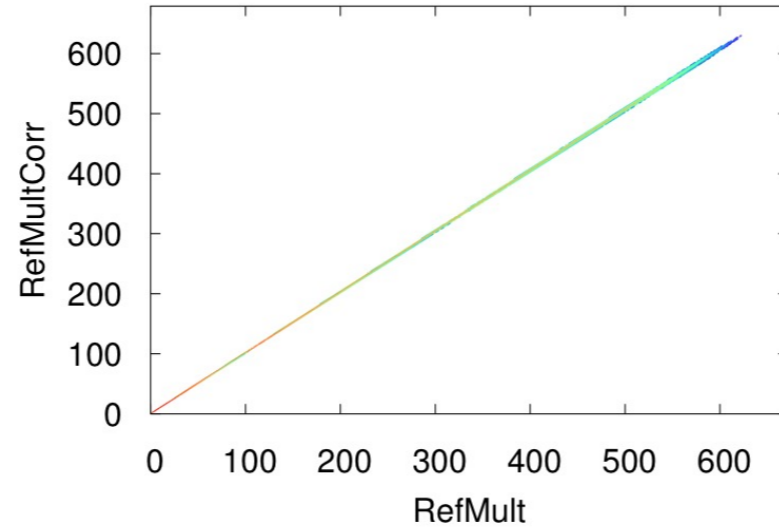
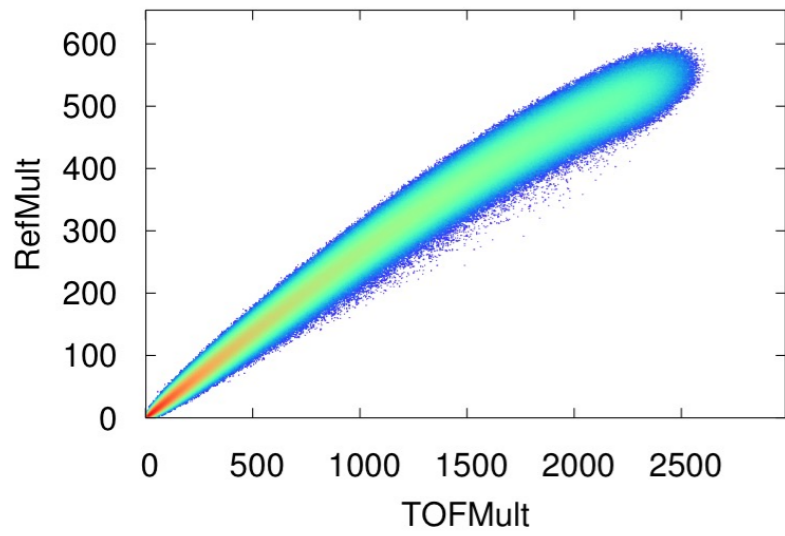


Experimental challenges :

- Response of ZDC to neutrons
- Clustering of nucleons that introduces artificial de-correlation

Analysis in this direction (separating signals of flow & CME) and systematic studies are under progress

A few QA plots



Weight estimation for cumulant calculations

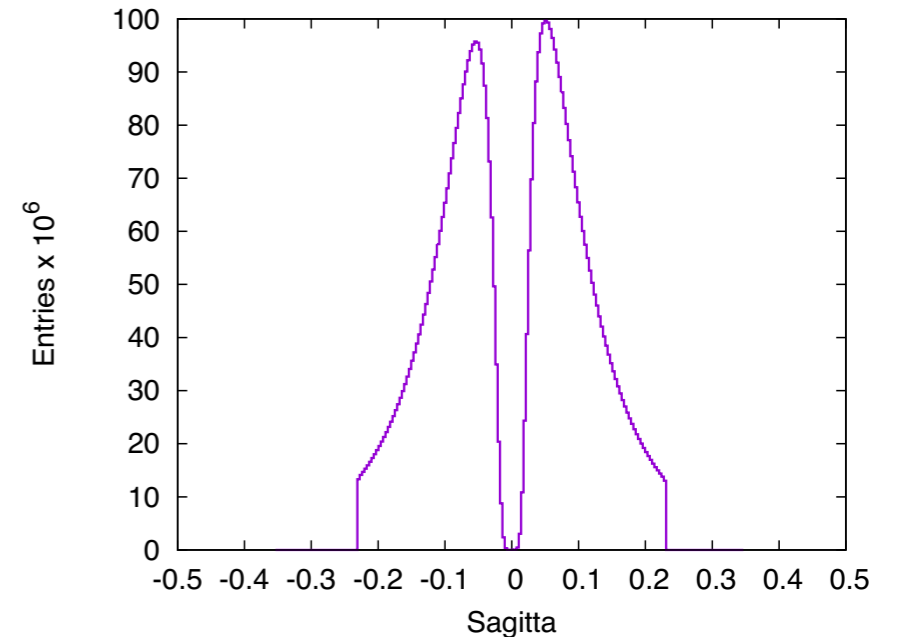
- Acceptance binning for weight calculation :

$$\text{Sagitta} = \text{charge} * ((20. * p_T / 3.) - \sqrt{((20. * p_T / 3.)^2 - 0.75^2)})$$

$$\text{Weight} = 1 / (\text{entries in } \eta\text{-}\phi) * 1 / \epsilon$$

The tracking efficiency :

$$\epsilon = \frac{C}{(1. + \exp(-(p_T + 0.1) / 0.15))}$$



B-field simulations : Dominance of fluctuations

$t=0, x=\langle x \rangle, y=\langle y \rangle, z=0, U+U$ collisions

